

# Hyper-Kamiokande計画 現状と展望

田中秀和

ICRR, University of Tokyo

for Hyper-Kamiokande proto-collaboration

ニュートリノ研究会, 2018年2月24日



# Hyper-K proto-collaboration

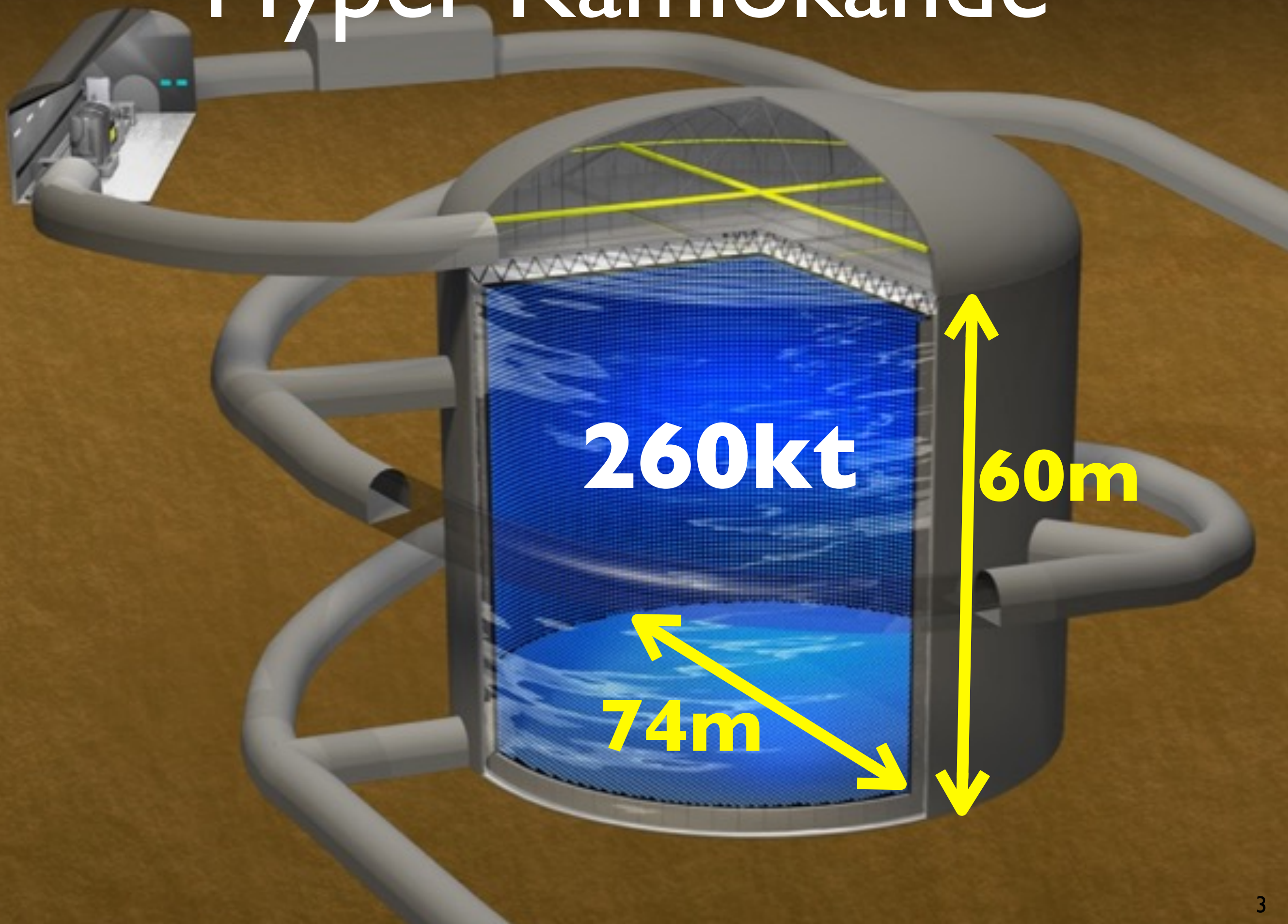


- ~300 collaborators
- 75 institutions from 15 countries
- ~70% of collaborators from overseas countries





# Hyper-Kamiokande





# Hyper-Kamiokande

- **Next generation water Cherenkov detector**

- Construct two detectors with staging
- **Realize the first detector as soon as possible**
  - Option of second detector in Korea
- The first detector (1 tank)
  - Filled with 260kton of ultra-pure water
    - 60m tall x 74 diameter water tank
- **Fiducial mass: 190kton**
  - **~10 x Super-K**
- **Photo-coverage: 40%** (Inner Detector)
  - 40,000 of **new 50cm $\phi$  PMTs**
    - **x2 higher photon sensitivity than SK PMT**
- All physics sensitivities shown in this talk assumes 1 tank (otherwise will be mentioned)

260kt

60m

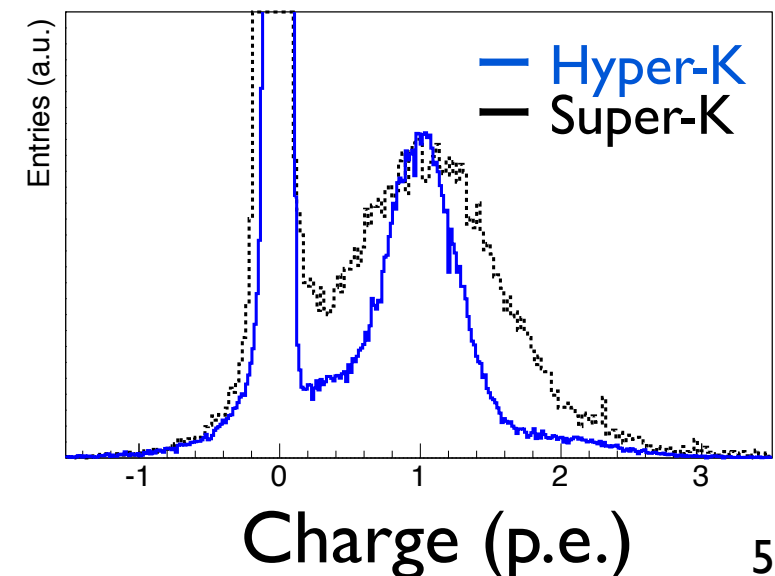
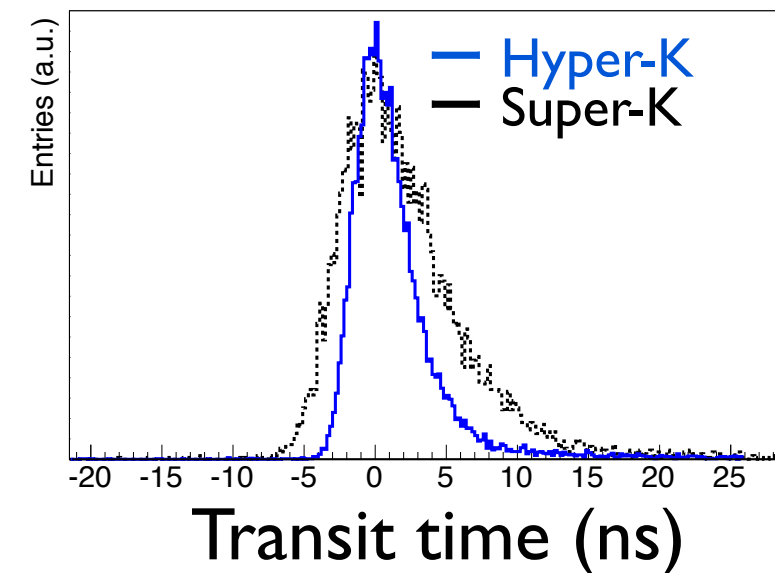
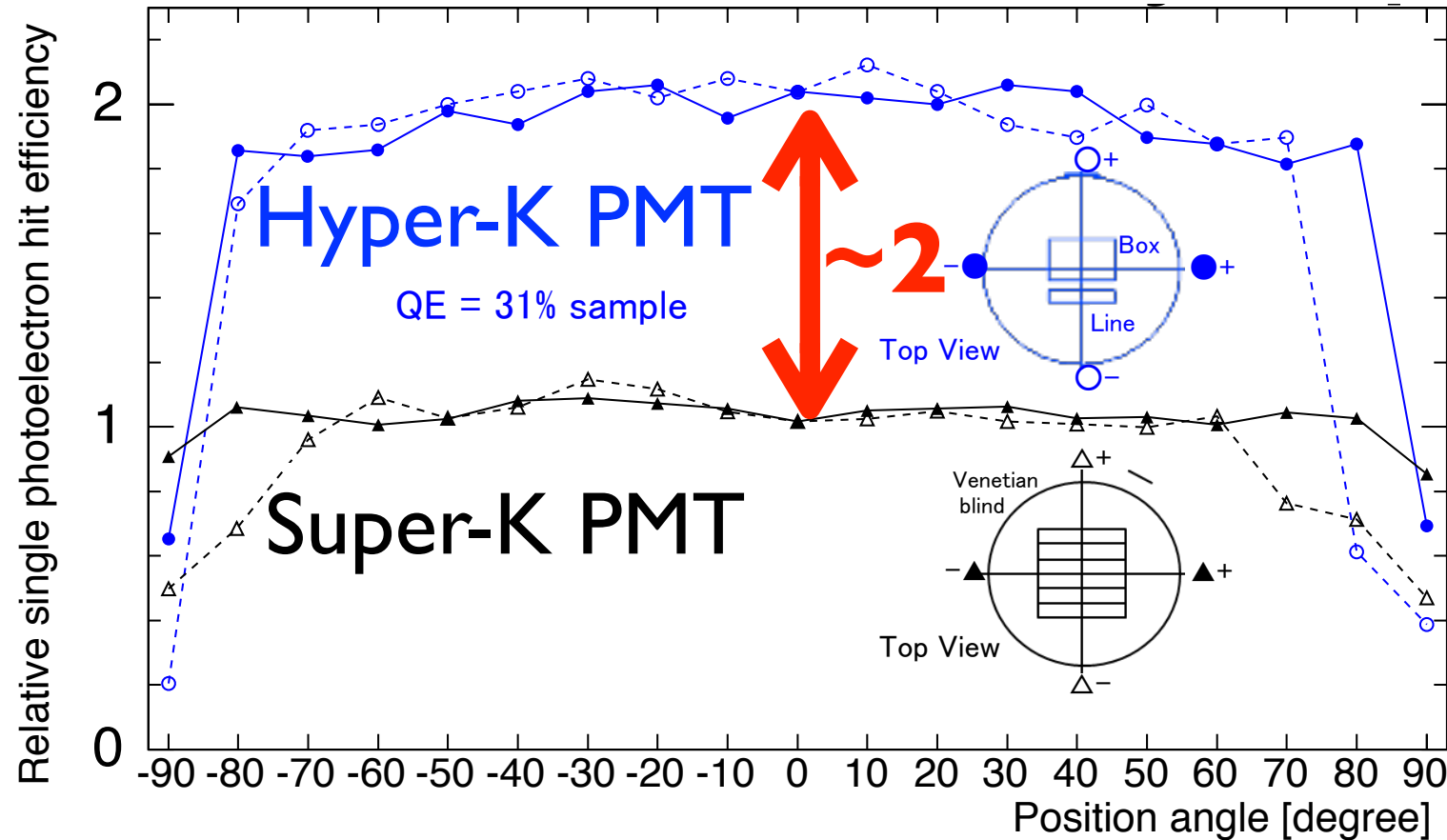
74m



# New 50cm $\phi$ PMT for Hyper-K

Box & line dynode PMT

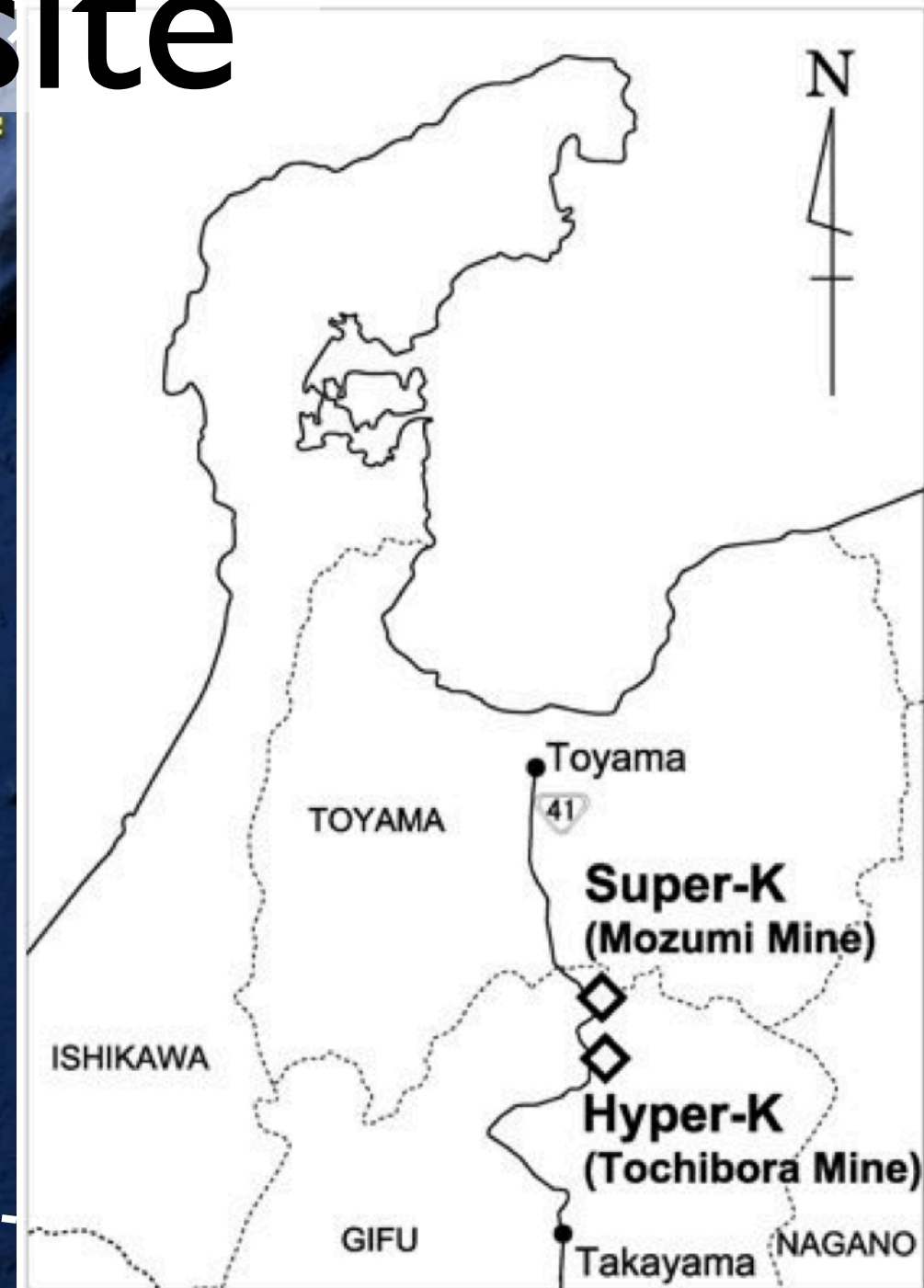
Photo-detection efficiency (l p.e.)



- Twice better photo-detection efficiency than SK PMTs
- Timing resolution (TTS): 1.1 ns
  - cf. SK PMT: 2.1 ns
- Higher pressure tolerance: >80m



# Detector site



- The candidate site locates under Mt. Nijugo-yama
- ~8km south from Super-K
- Identical baseline (295km) and off-axis angle (2.5deg) to T2K
- Overburden ~650m (~1755 m.w.e.)



# Hyper-K: multi-purpose detector

- **Comprehensive study of  $\nu$  oscillation**

- CPV: 76% of  $\delta$  space w/  $3\sigma$ ,  $<22^\circ$  precision
- MH determination for all  $\delta$  with J-PARC/Atm  $\nu$
- $\theta_{23}$  octant determination at  $|\theta_{23}-45^\circ|>2^\circ$
- $<1\%$  precision of  $\Delta m^2_{32}$
- Test standard  $\nu$  oscillation scenario w/ acc/atm  $\nu$

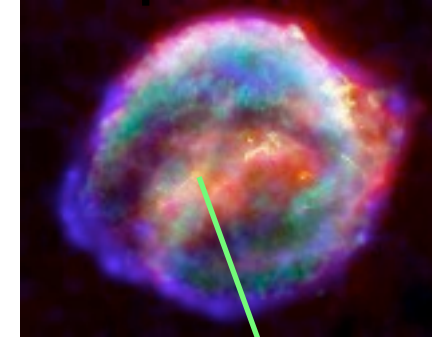
- **Proton decay  $3\sigma$  discovery potential**

- $1 \times 10^{35}$  years for  $p \rightarrow e^+ \pi^0$
- $3 \times 10^{34}$  years for  $p \rightarrow \nu K^+$

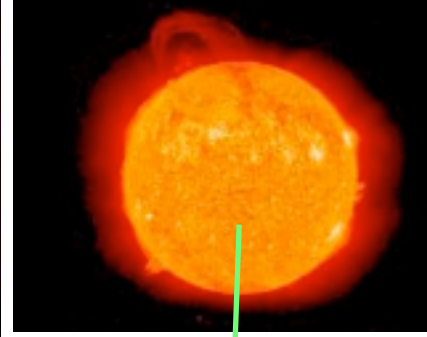
- **Astrophysical neutrino**

- Solar  $\nu$ : test standard matter effect (MSW) model
- Supernova  $\nu$ , supernova relic- $\nu$
- Dark matter neutrinos from Sun, Galaxy, Earth

Supernova



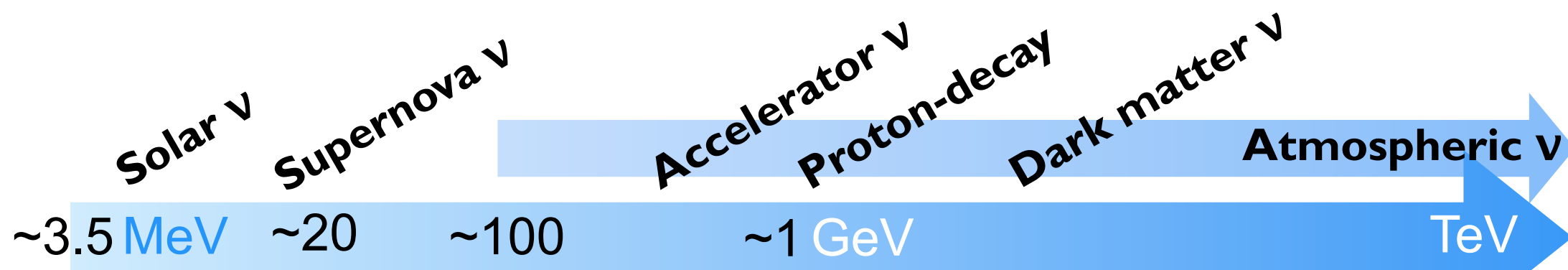
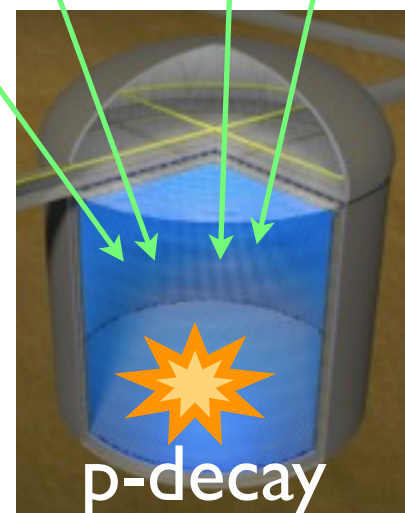
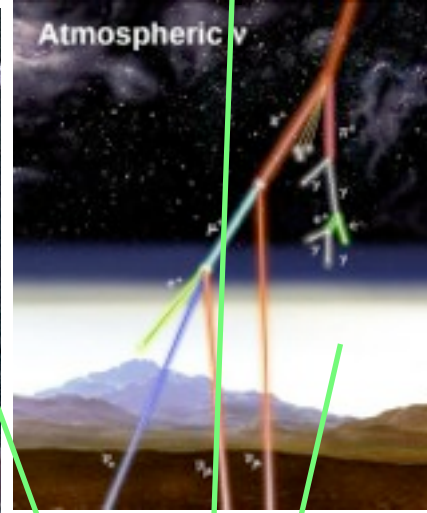
Sun



Accelerator



Atmospheric

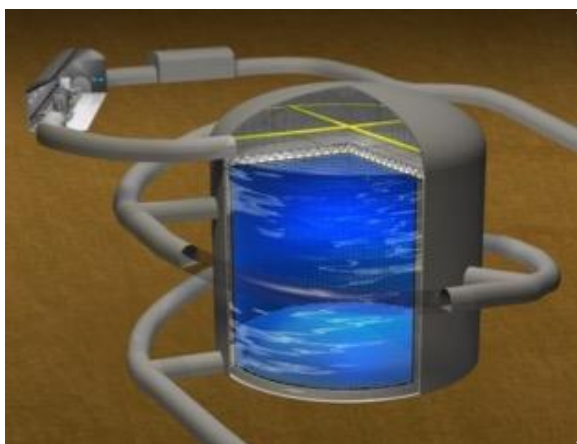
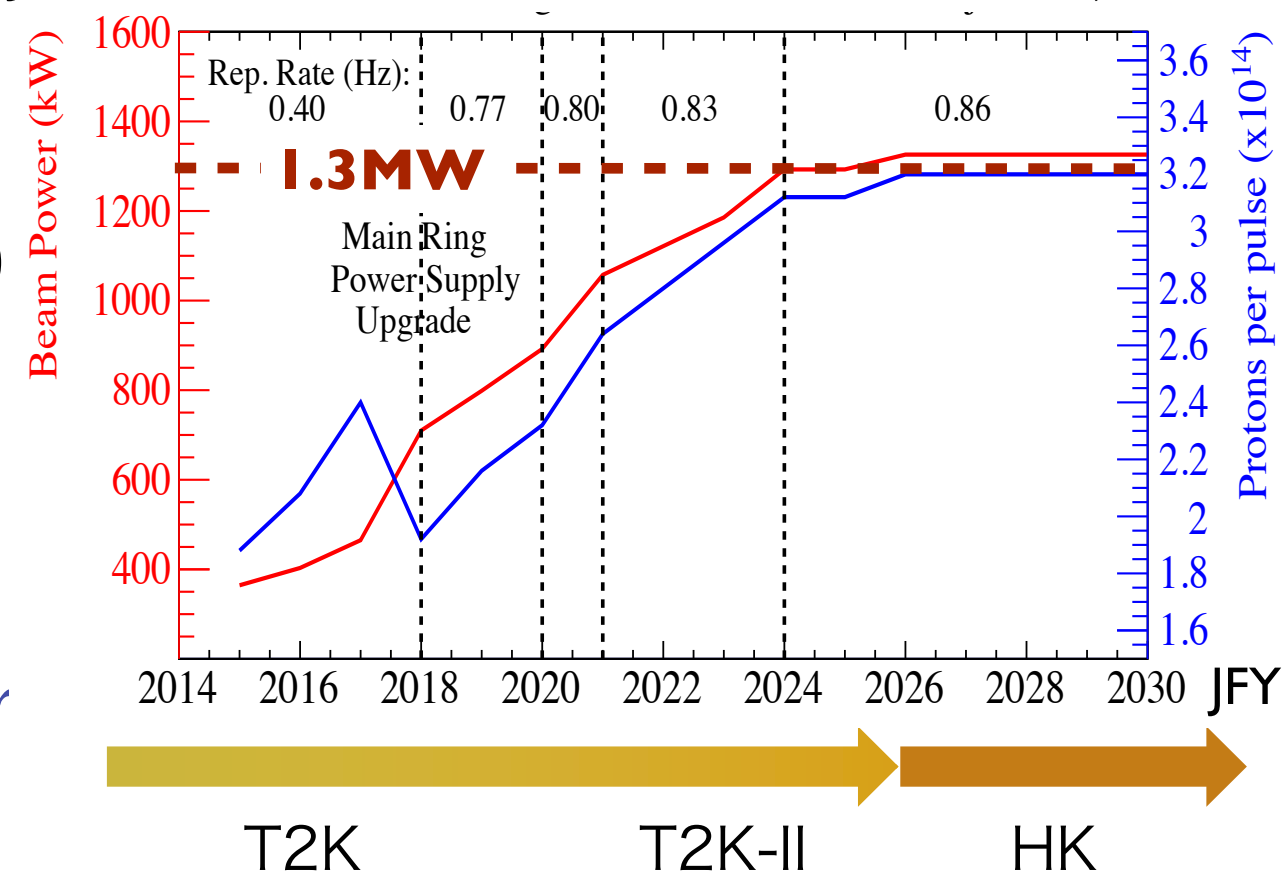




# Accelerator based neutrinos

- High quality & high intensity neutrino beam
- 2.5 deg. off-axis narrow band neutrino beam (identical to T2K)
- Beam power: 1.3MW (before Hyper-K starts)
  - KEK Project Implementation Plan: top priority on 'J-PARC upgrade for Hyper-K'

J-PARC MR Fast Extraction Power Projection



Hyper-K



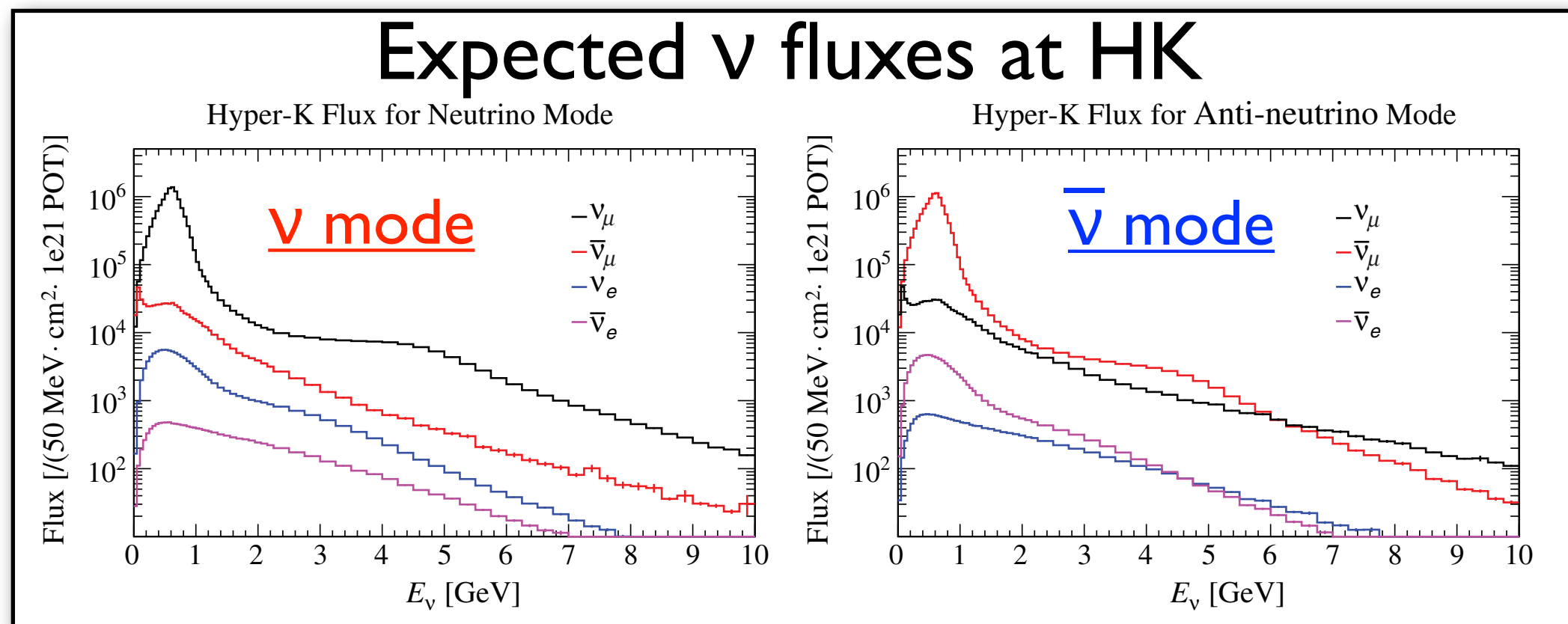
J-PARC  
Accelerator Complex





# J-PARC to Hyper-K

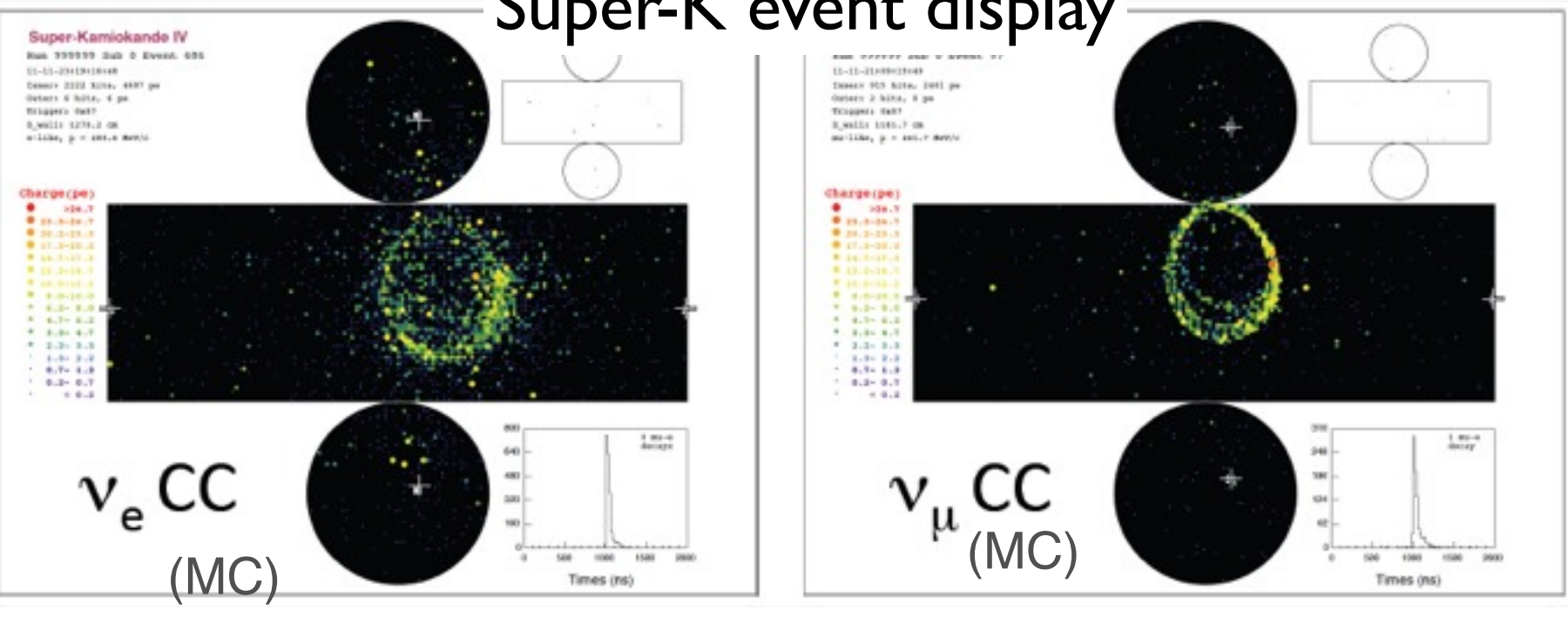
- Relatively short baseline (295km, same as T2K)
  - Less matter effect: good for CP measurement
  - Complementary to other experiments with longer baseline ( $>1000\text{km}$ )
- 2.5 deg. Off-Axis beam (same as T2K)
- Low energy ( $\sim 0.6\text{GeV}$ ) and narrow band beam
  - Peak around oscillation maximum
  - Good match for Water Cherenkov detector





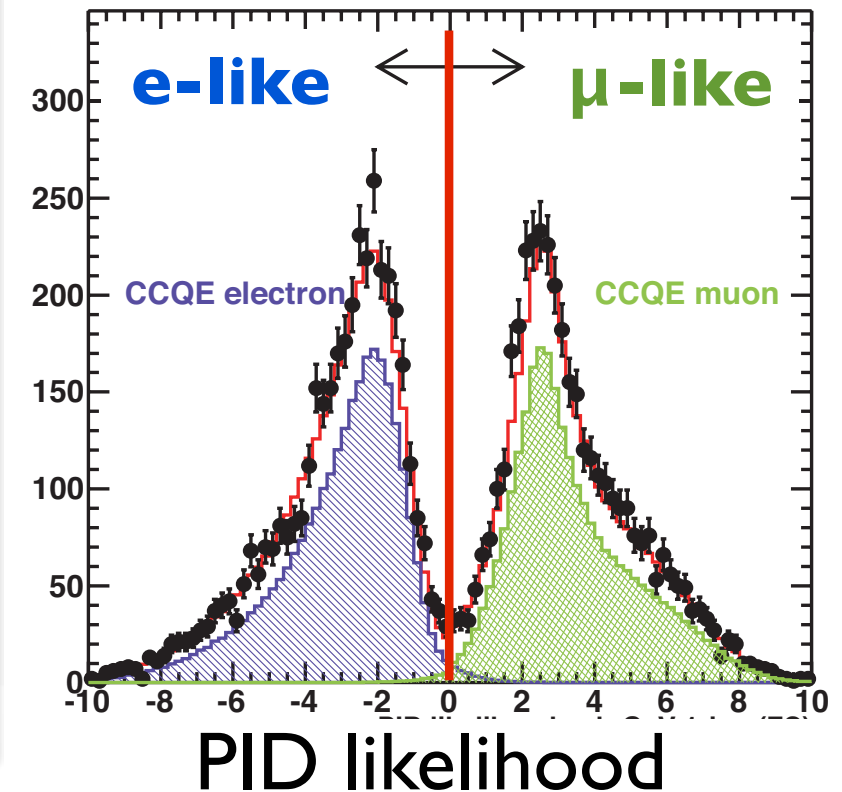
# Detector performance

## Super-K event display

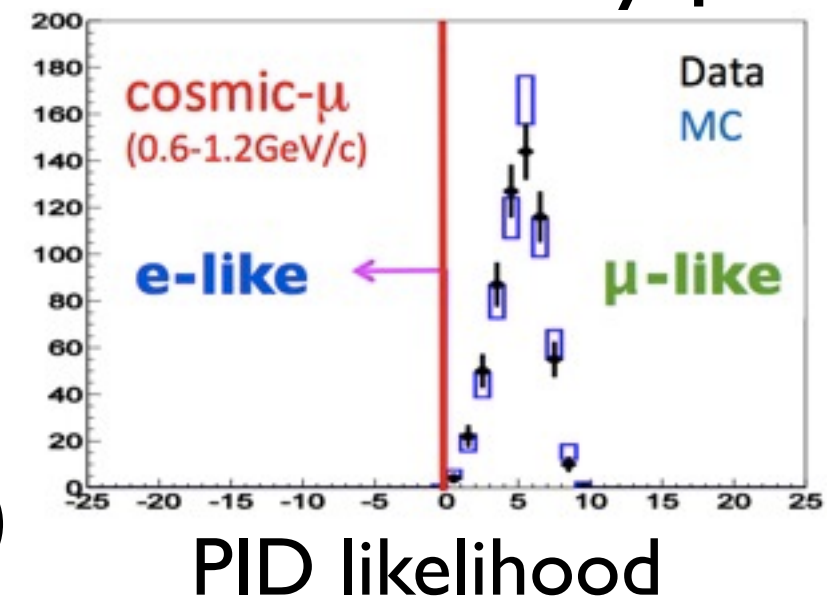


- **Large mass** ( $\sim 10 \times$  Super-K FV)
- Statistics is always critical
- **Excellent particle ID (e/ $\mu$ )**
  - Mis-identification  $< 1\%$
- Energy resolution e/ $\mu$   $\sim 3\%$
- Quasi-elastic is dominant (sub-GeV)
  - Clean one-ring event

## SK atmospheric $\nu$



## SK cosmic-ray $\mu$



mis-PID:

Data:  $0.00 \pm 0.16$  (stat.) %

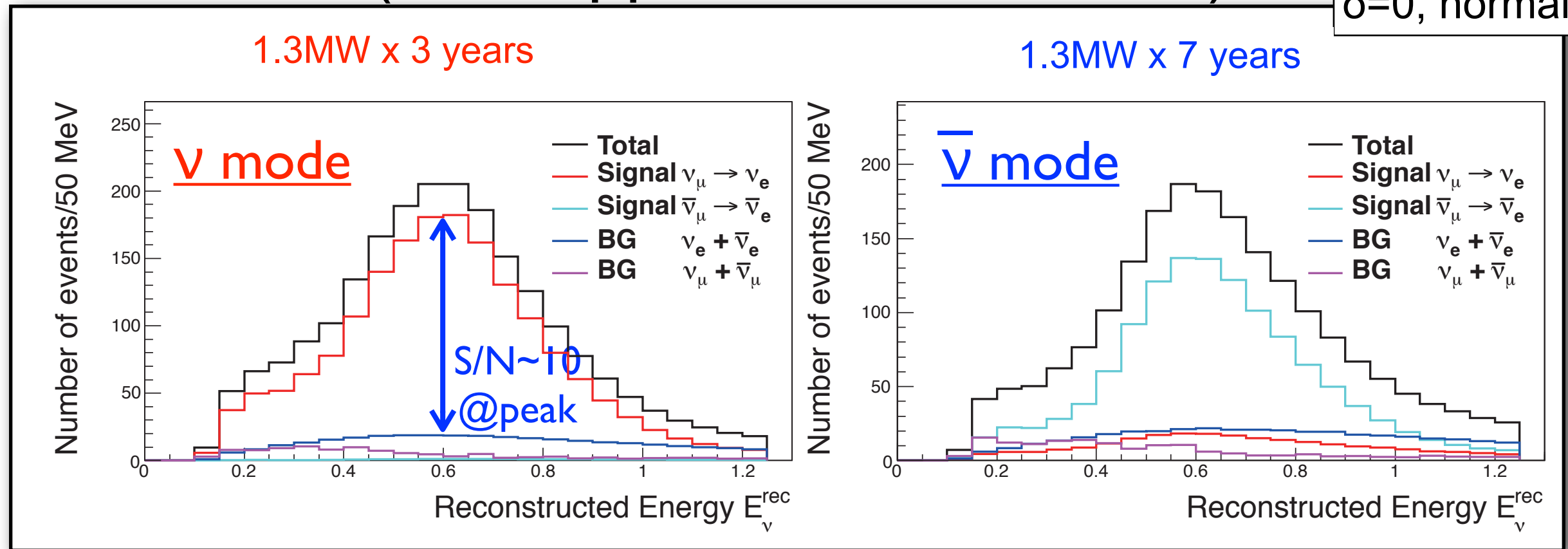
MC :  $0.10 \pm 0.10$  (stat.) %



# $\nu_e$ CC candidates

## Reconstructed $\nu$ energy distributions (after applied selection cuts)

$\sin^2(2\theta_{13})=0.1$ ,  
 $\delta=0$ , normal MH



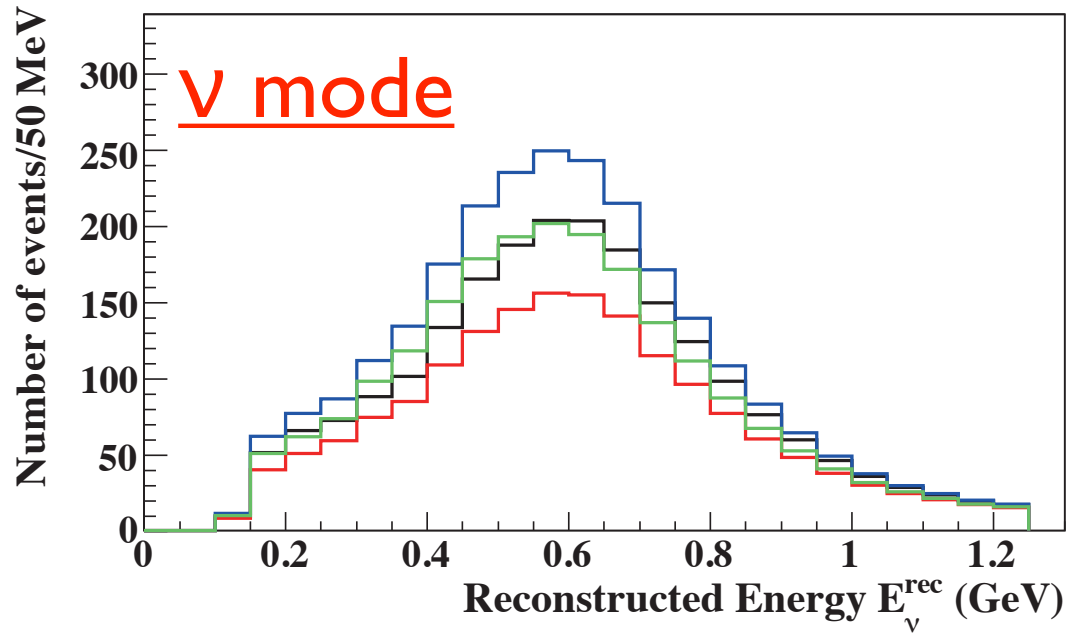
for $\delta_{\text{CP}} = 0$	Signal $\nu_\mu \rightarrow \nu_e$ CC	Wrong sign appearance	$\nu_\mu / \bar{\nu}_\mu$ CC	Beam $\nu_e / \bar{\nu}_e$ contamination	NC
$\nu$ beam	1,643	15	7	259	134
$\bar{\nu}$ beam	1,183	206	4	317	196

- $S/N \sim 10$  at peak ( $>99\%$   $\nu_\mu$  CC/NC bkg rejection)
- ex. NC  $\pi^0$  bkg rejection efficiency:  $>99.5\%$

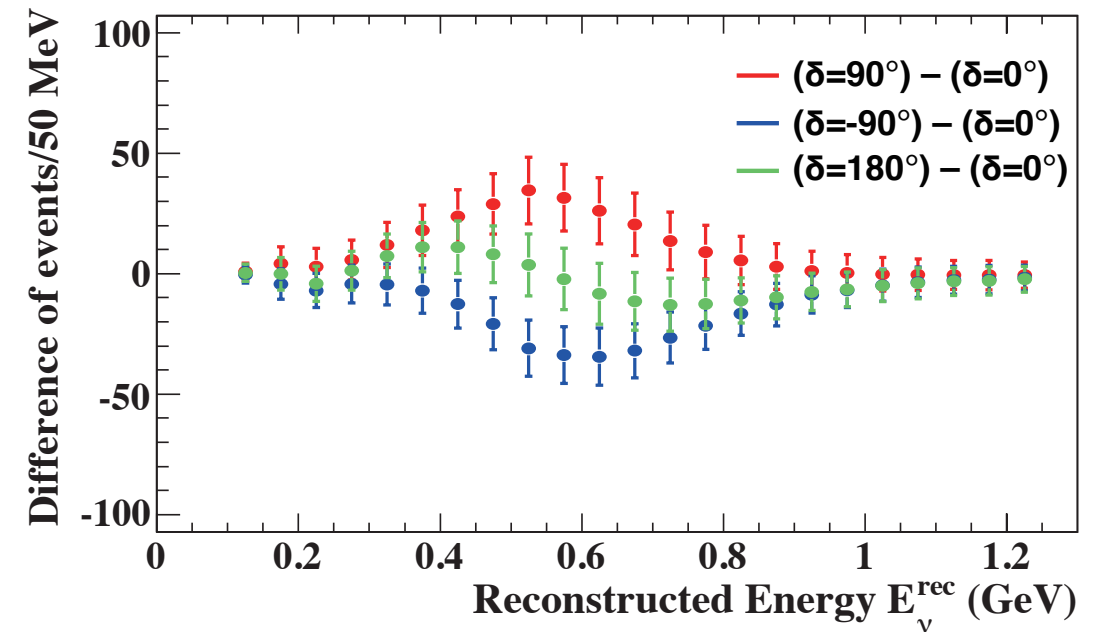
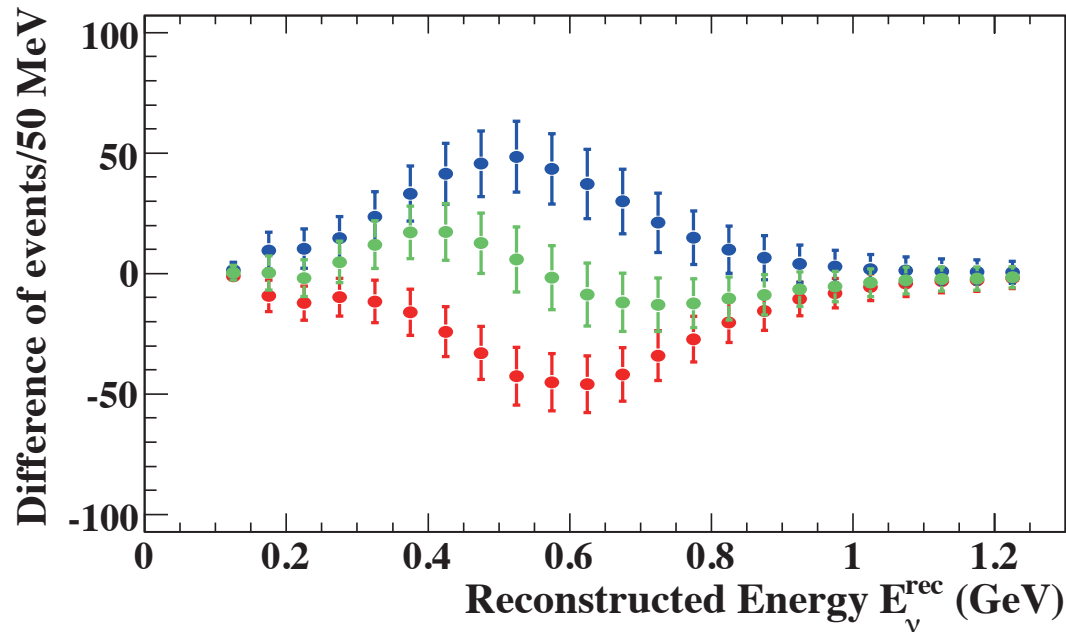
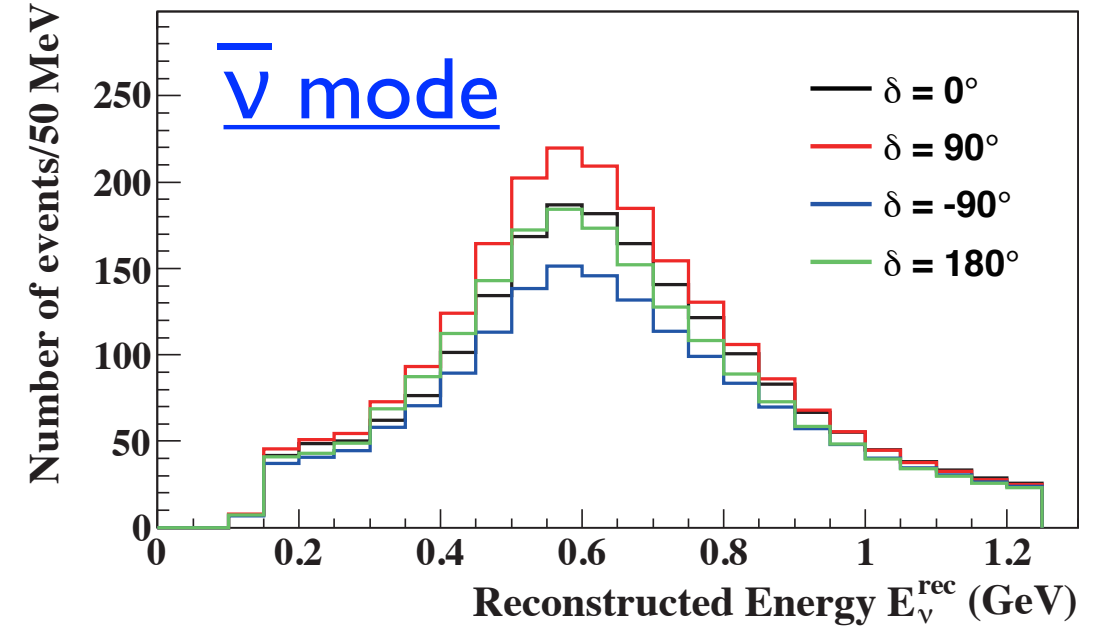


# Effect of $\delta_{CP} \neq 0$

Neutrino mode: appearance



Antineutrino mode: appearance

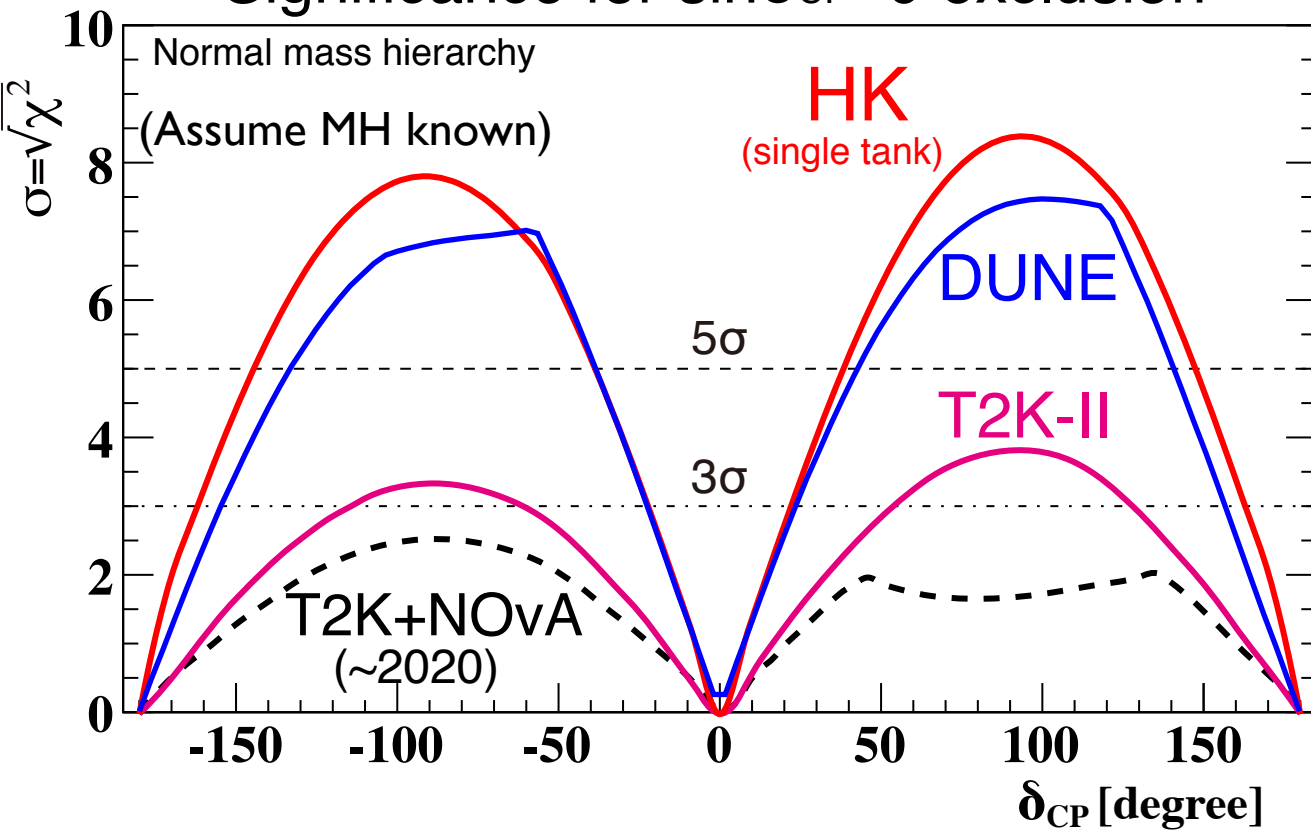


- Sensitive to all  $\delta$  values
- also sensitive to any non-standard CPV, if any

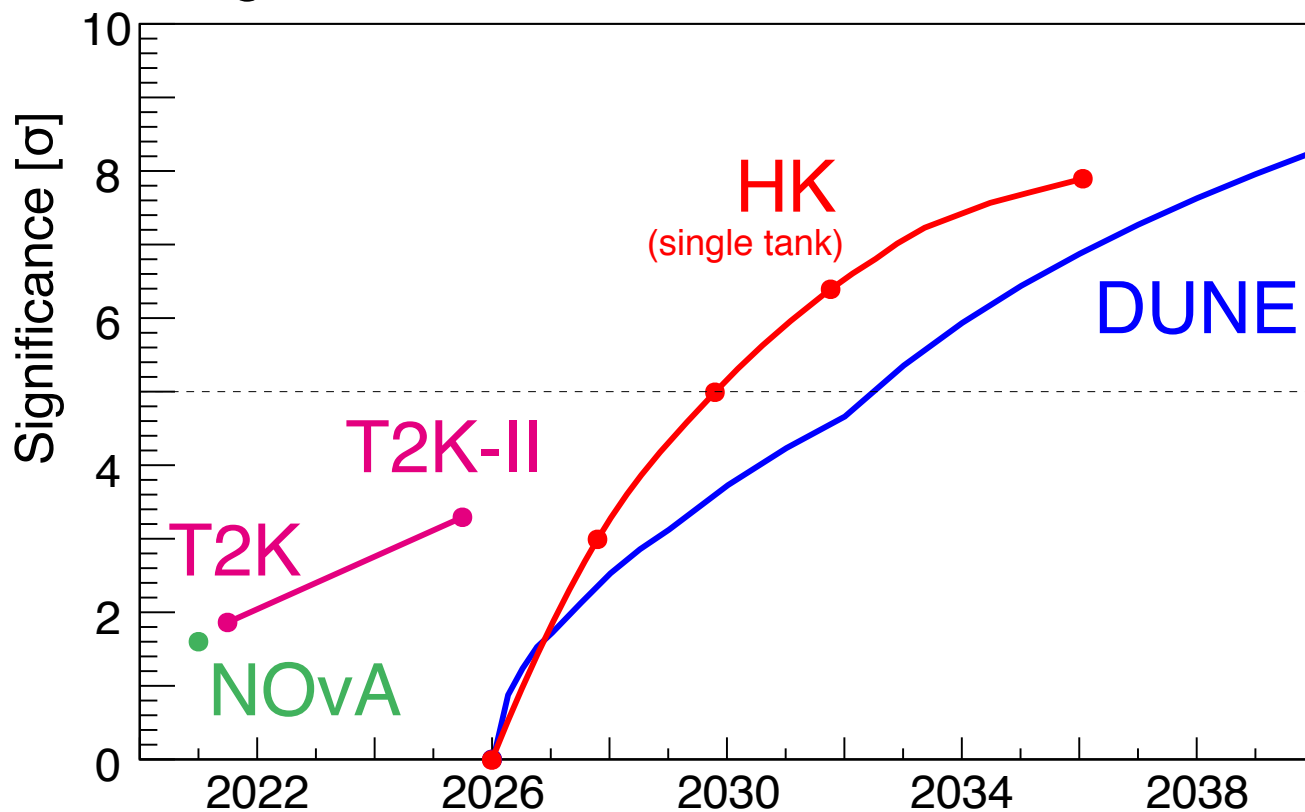


# Expected sensitivity for CPV

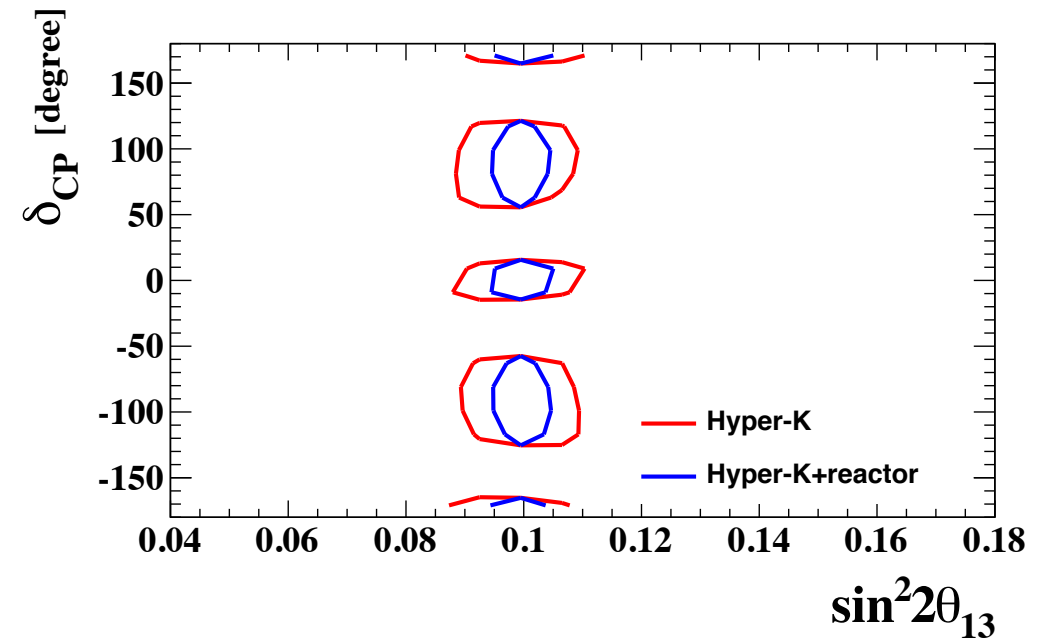
Significance for  $\sin\delta_{CP}=0$  exclusion



CPV significance for  $\delta_{CP}=-90^\circ$ , normal hierarchy



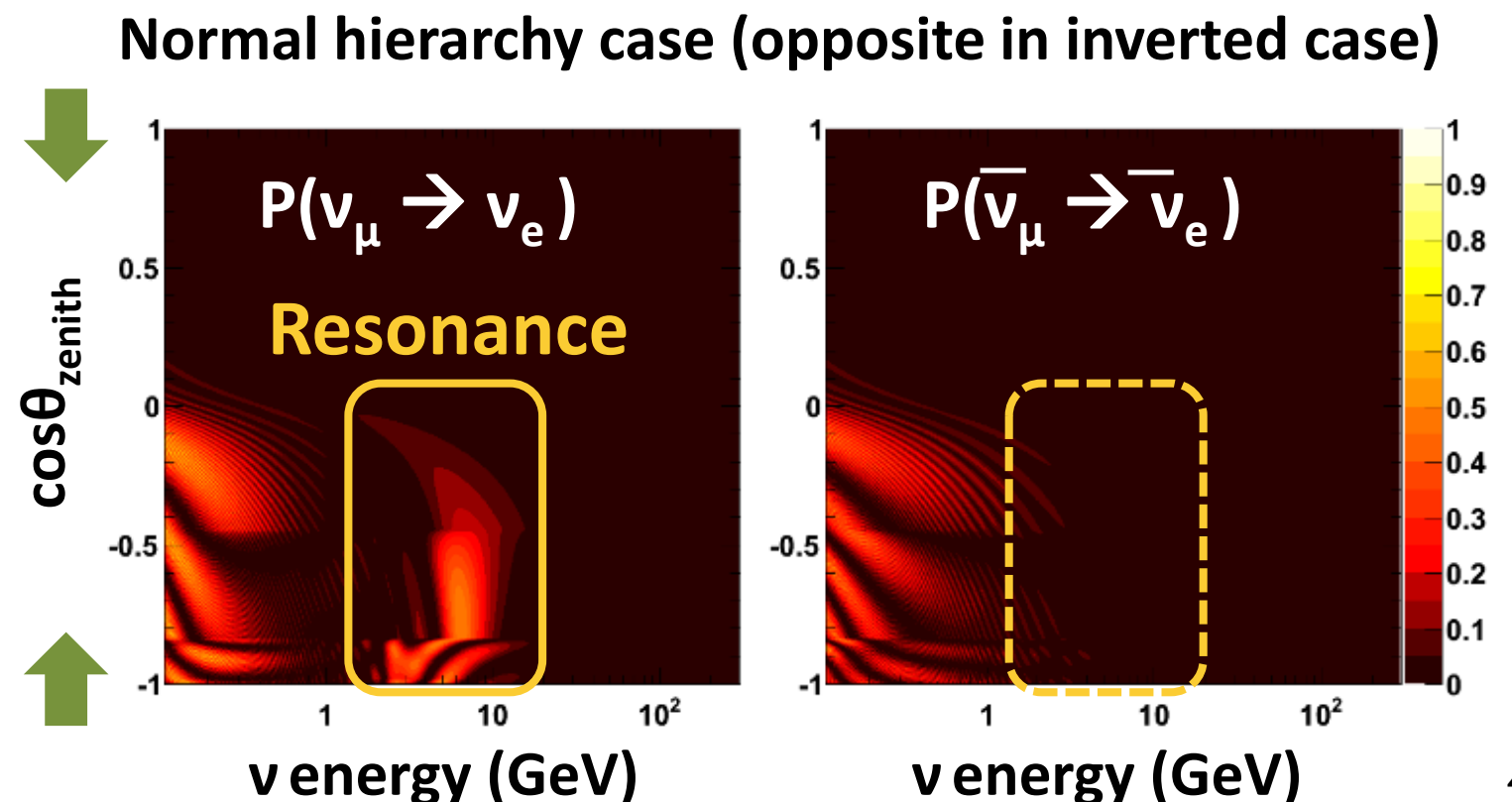
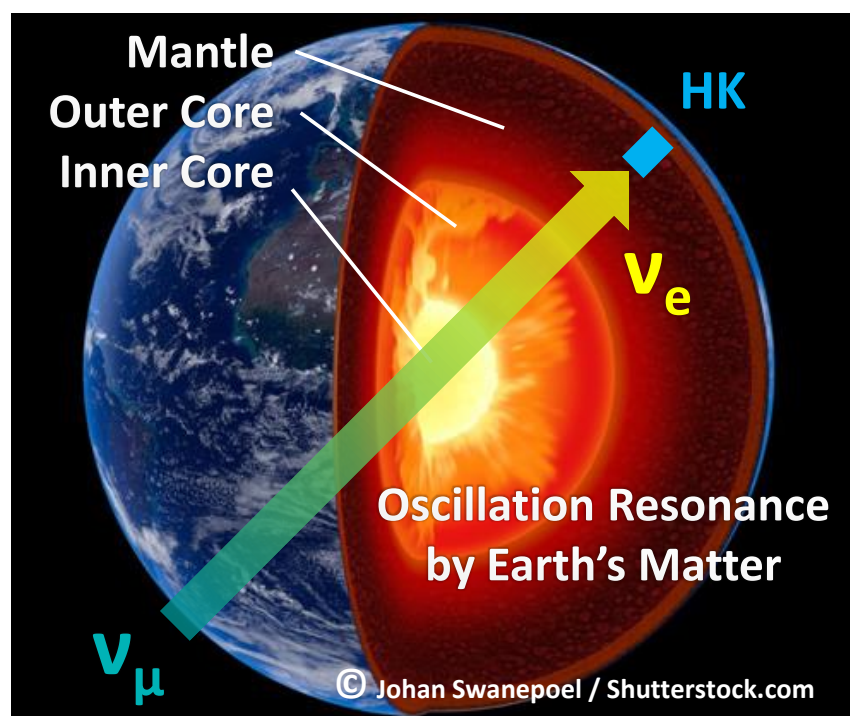
- $\sin\delta_{CP}=0$  exclusion:
    - $\sim 8\sigma$  significance if  $\delta_{CP}=\pm 90^\circ$
    - $\sim 6\sigma$  significance if  $\delta_{CP}=\pm 45^\circ$
  - Observe CPV for 76% (58%) of  $\delta_{CP}$  space w/  $3\sigma$  ( $5\sigma$ ) significance
  - $\delta_{CP}$  resolution:
    - $22^\circ$  at  $\delta_{CP}=\pm 90^\circ$
    - $7^\circ$  at  $\delta_{CP}=0^\circ / 180^\circ$
- Normal mass hierarchy





# Mass Hierarchy determination in Hyper-K

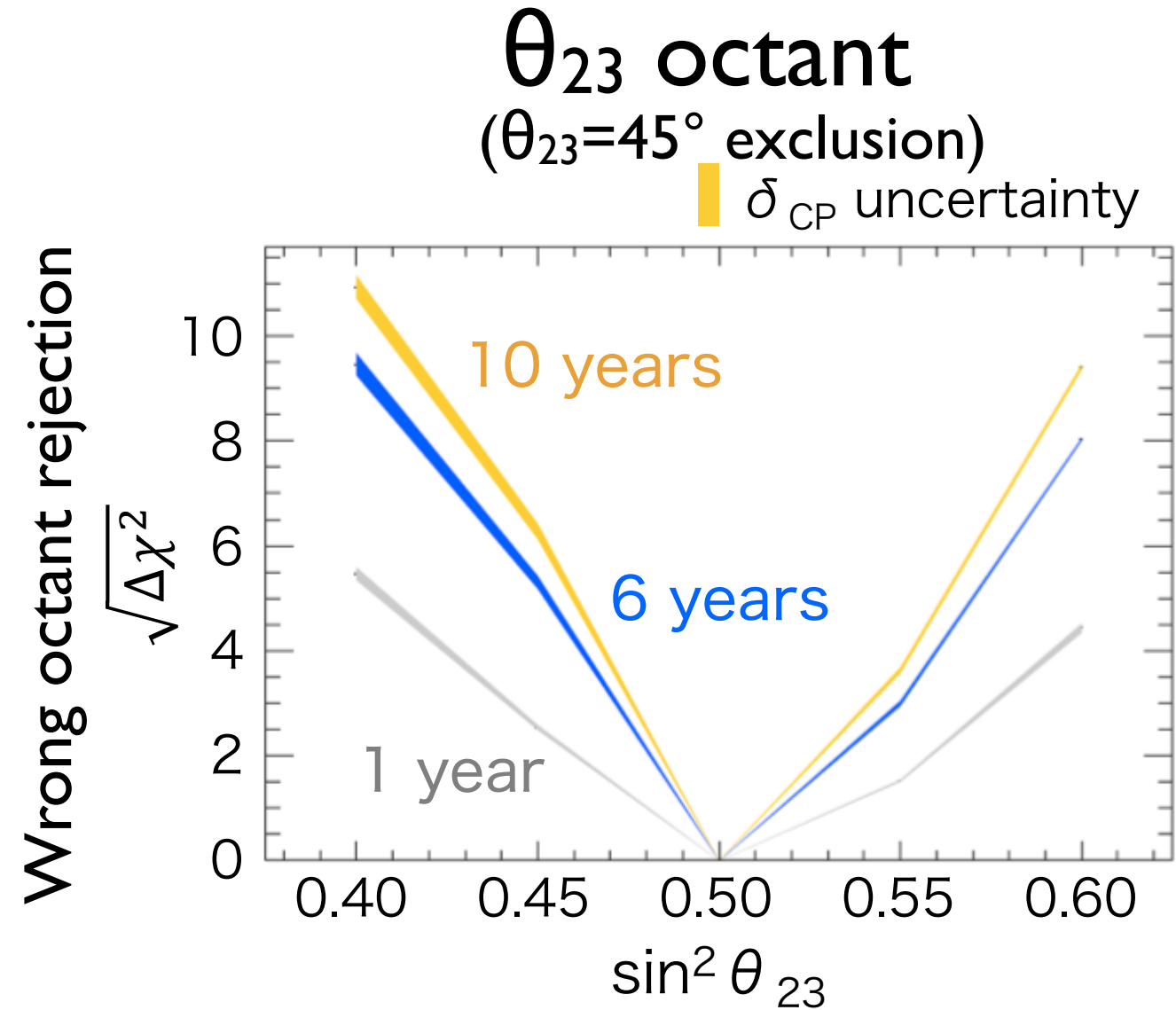
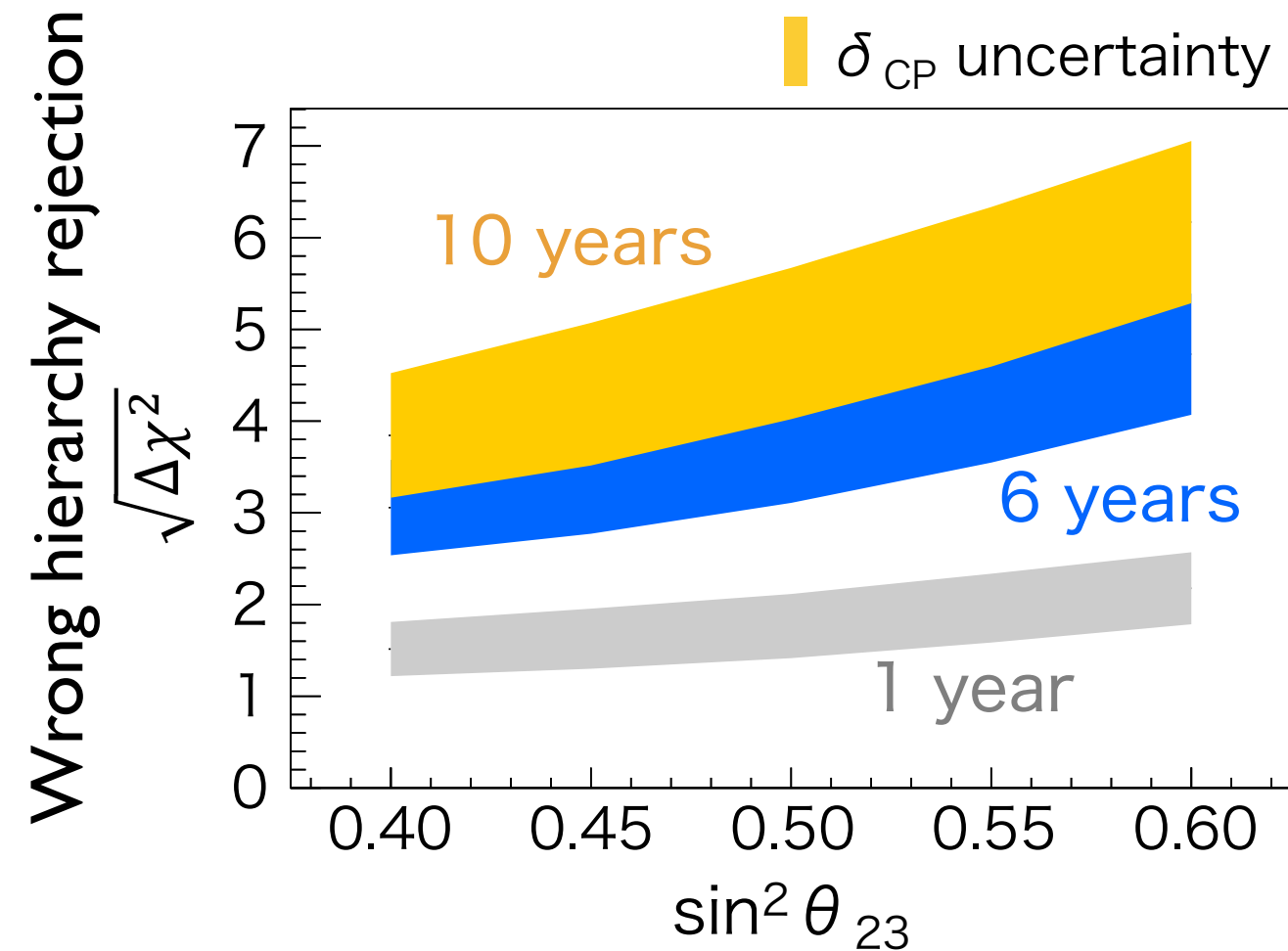
- Earth matter effect in upward-going multi-GeV  $\nu_e$  sensitive to mass hierarchy
  - Earth matter effect 'resonance' appears in  $\nu_e$  app. for NH, in  $\bar{\nu}_e$  app. in IH
- Combination of atmospheric  $\nu$  and beam  $\nu$  to determine mass hierarchy





# Mass Hierarchy sensitivity in Hyper-K

## Mass Hierarchy



- Mass hierarchy and  $\theta_{23}$  octant can be determined ( $\geq 3\sigma$ ) within several years for the nearly entire parameter space



# Systematic uncertainties

- Syst. errors based on SK/T2K experience and prospect
  - Error matrices of T2K adopted in HK sensitivity
  - Near detector: constrain cross section error w/ water target
  - Far detector: systematic error reduced by large stat. atm- $\nu$

## Assumed syst. uncertainties for sensitivity studies

		Flux & ND-constrained cross section	ND-independent cross section	Far detector	Total	T2K (2017)
$\nu$ mode	Appearance	3.0%	0.5%	0.7%	3.2%	6.1%
	Disappearance	3.3%	0.9%	1.0%	3.6%	4.4%
$\bar{\nu}$ mode	Appearance	3.2%	1.5%	1.5%	3.9%	6.5%
	Disappearance	3.3%	0.9%	1.1%	3.6%	3.8%

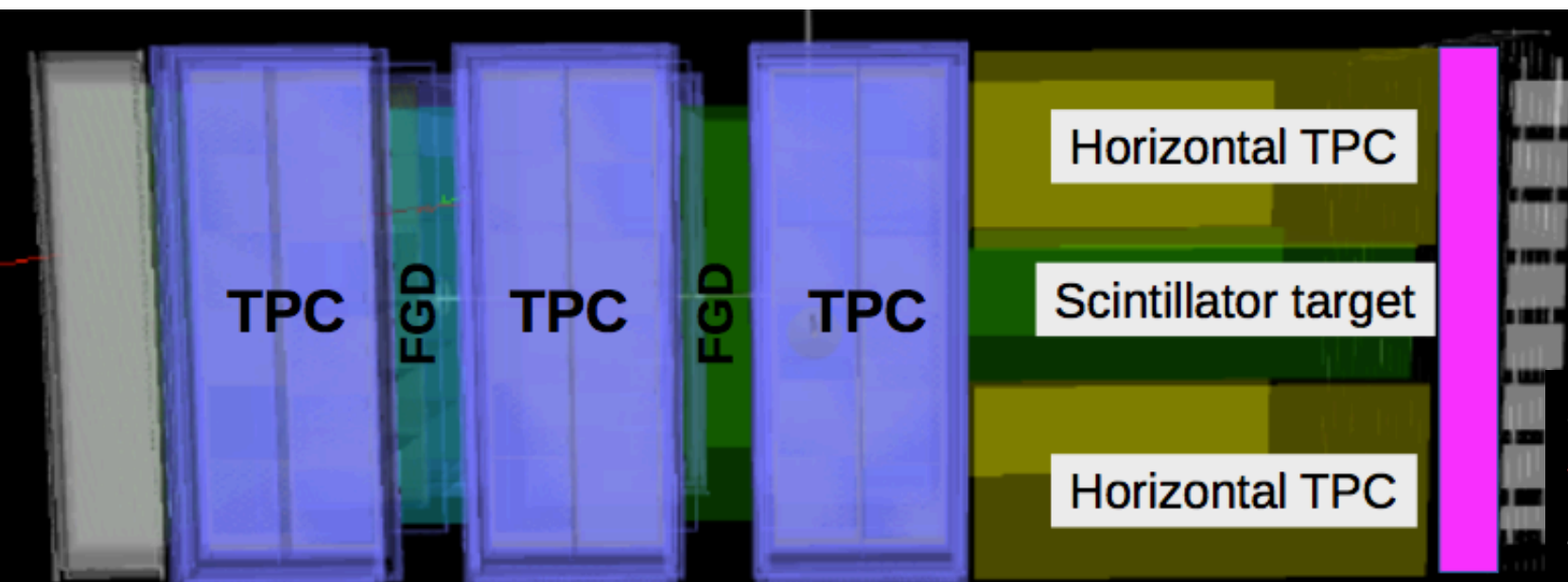
HK sensitivity studies assume uncertainties at 3~4% level

(cf. Uncertainty at the 6~7% level in T2K)

Dominant errors: electron (anti)neutrino cross section, near-to-far extrapolation of event rates, far detector modeling

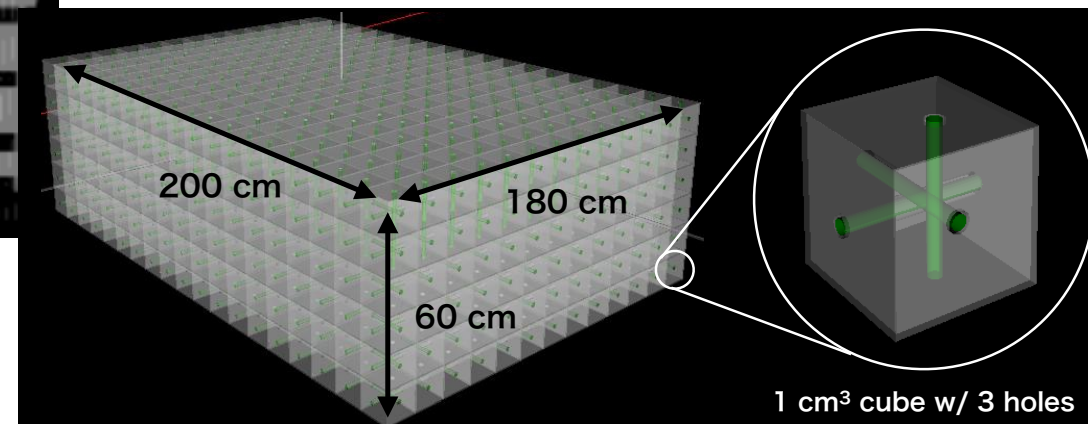
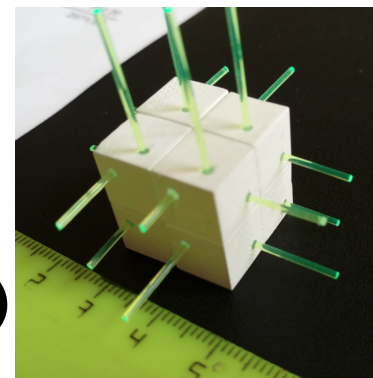
# Near Detector Upgrade

- Near and intermediate detectors for HK are being developed to control flux and cross-section systematic errors
- Will continue using the INGRID on-axis and magnetized ND280 off-axis detectors
- Work within T2K to upgrade ND280 detector
  - See Satashita-san's talk for details



→ High-granularity scintillator detector (SuperFGD) and 'high-angle' TPC

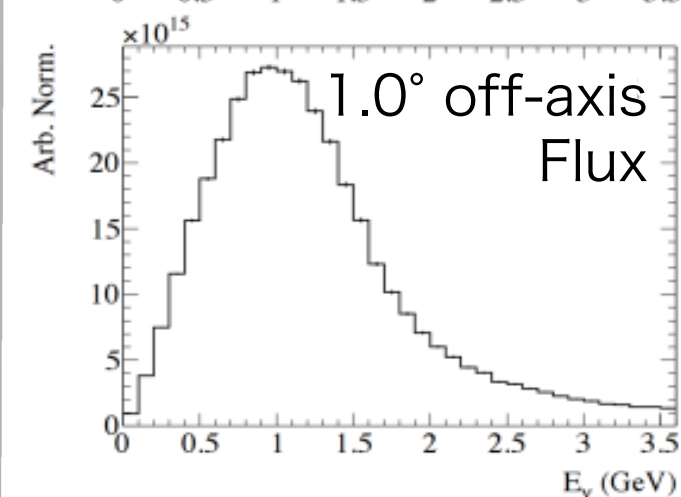
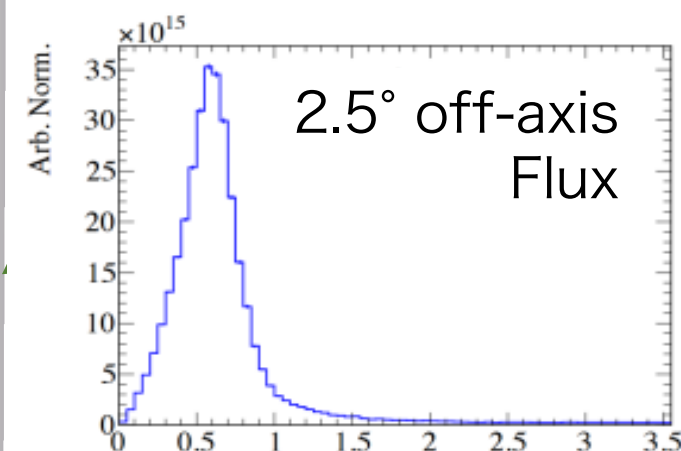
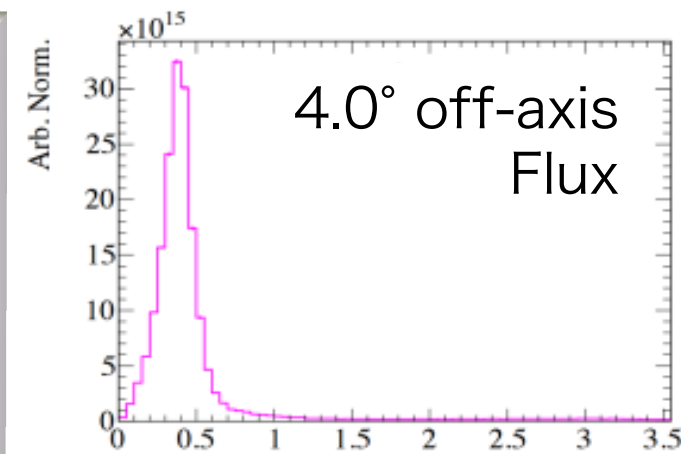
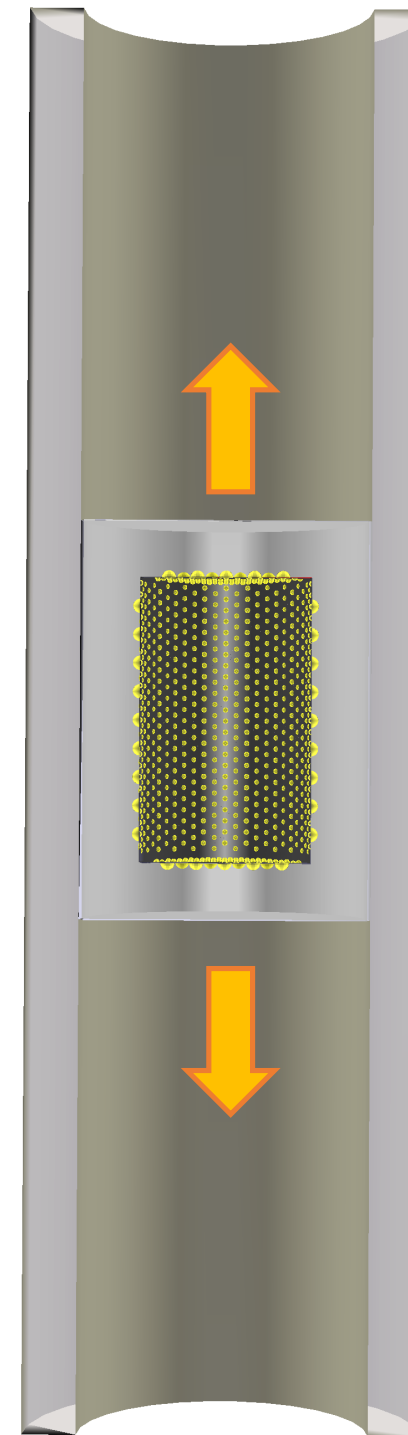
SuperFGD



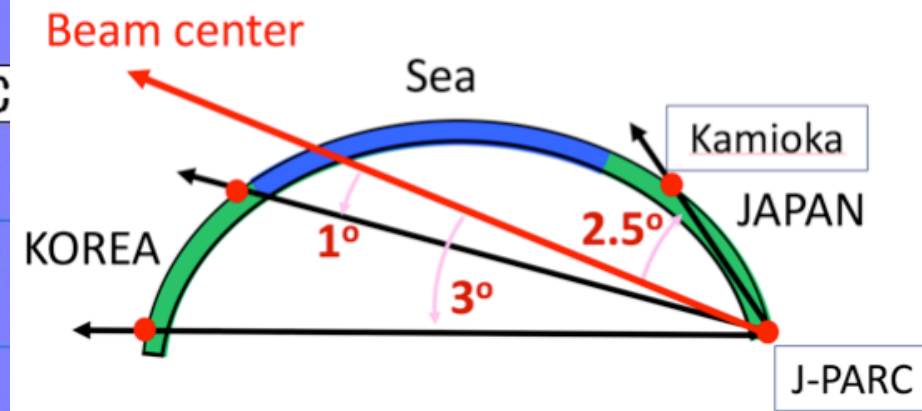
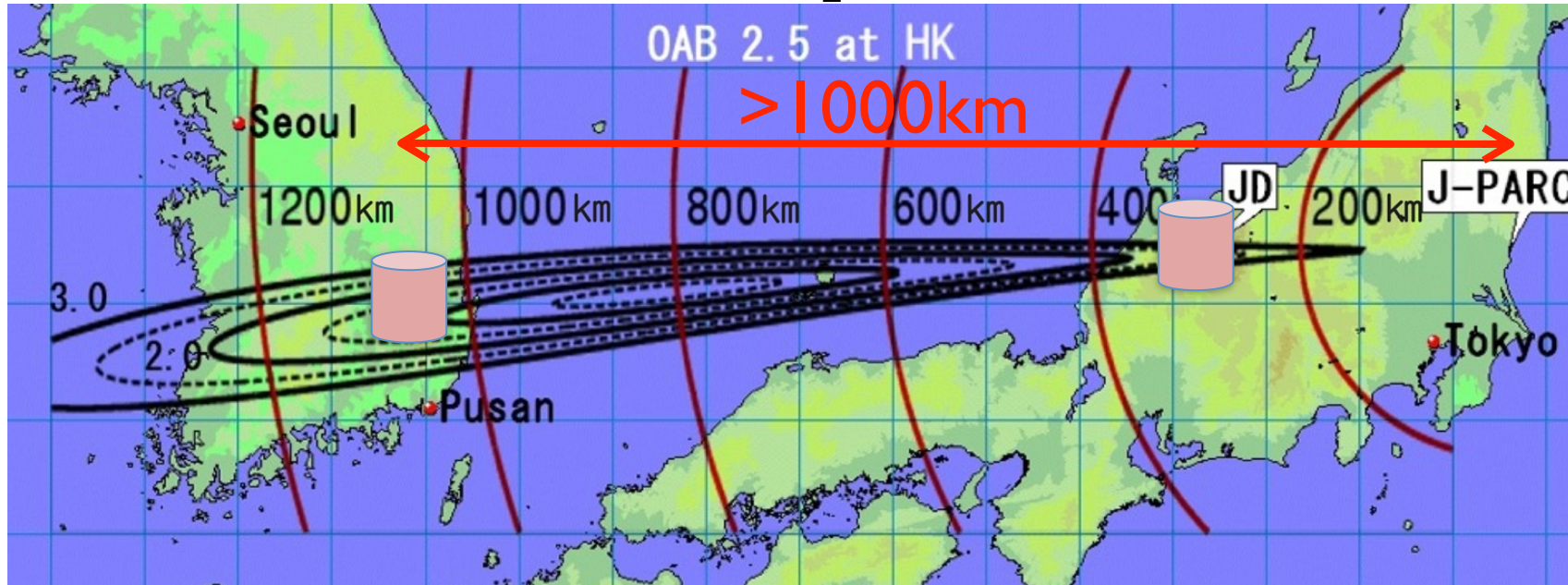


# Intermediate Water Č detector

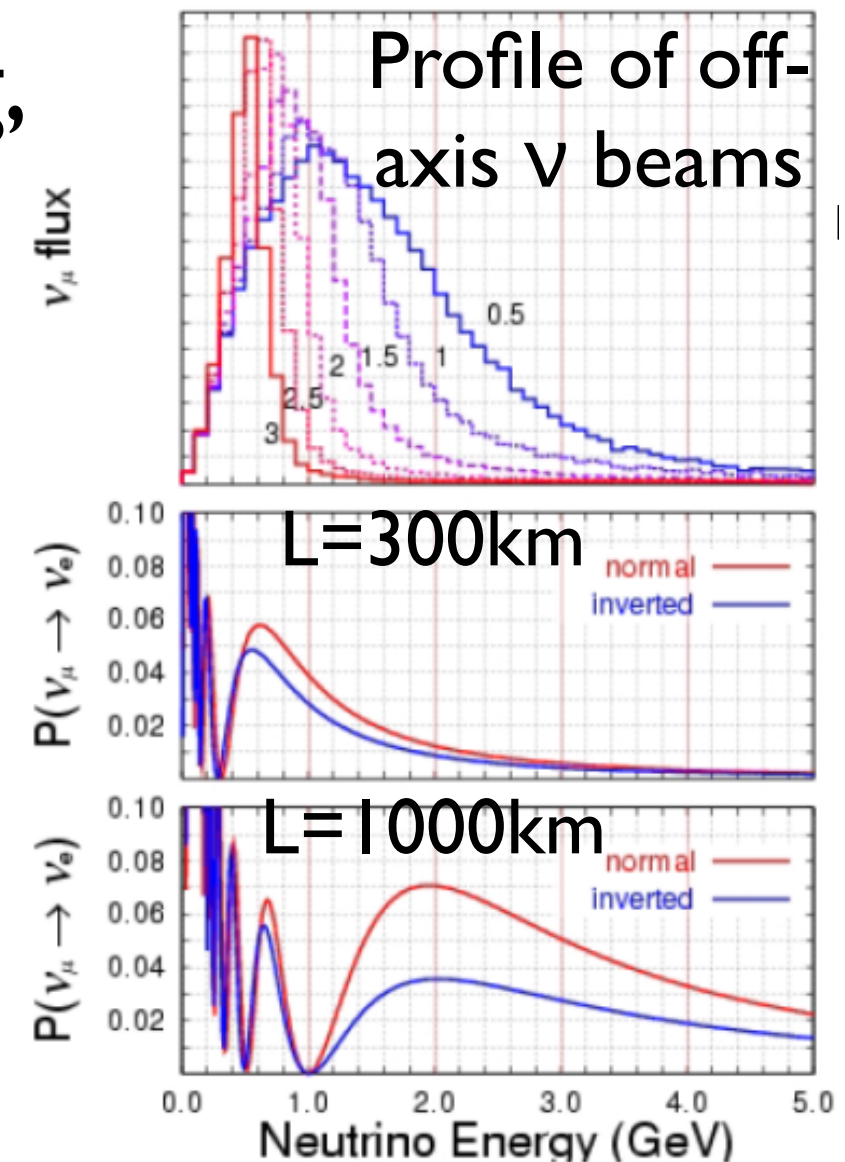
- Intermediate water Č detector proposed
  - Locate 1~2km downstream of J-PARC neutrino beam
  - J-PARC E61 (NuPRISM)
- Off-axis spanning ( $1^\circ \sim 4^\circ$ ) to probe neutrino energy vs. final state kinematics relationship
- Increase  $\nu_e$  purity at higher angles: larger contribution from Kaon
  - Electron neutrino cross-section
- Aim to  $\sigma(\nu_e)/\sigma(\nu_\mu)$  cross-section ratio error down to 2~3%



# Korean option for 2nd tank



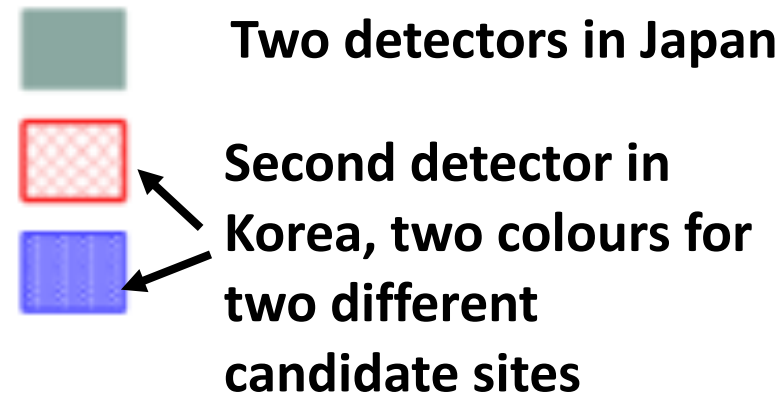
- Given the two tank design with staging, may benefit from building the second tank at a different baseline
- Option for second tank in Korea is being considered
- Advantage of 2nd detector in Korea:
  - CP effect at second oscillation maximum
  - Mass-hierarchy sensitivity to complement the measurement with atmospheric  $\nu$



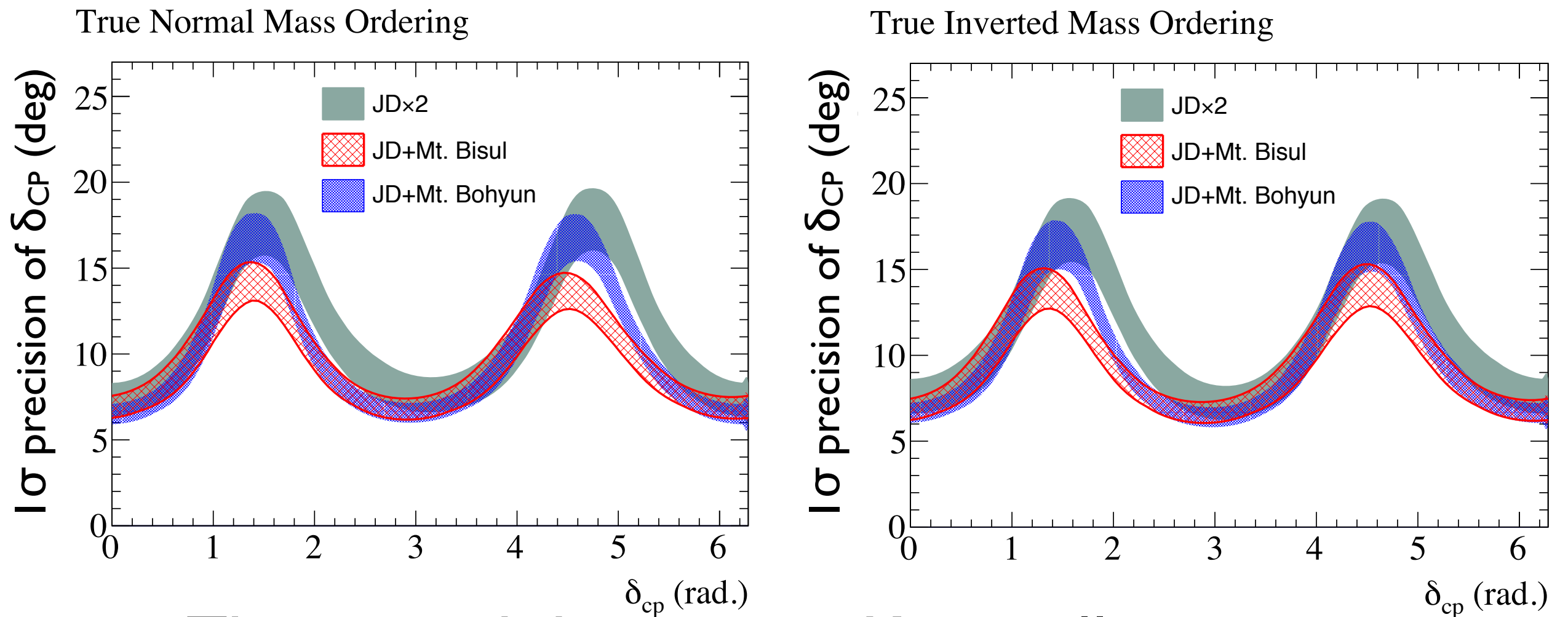


# $\delta_{CP}$ resolution

arXiv:1611.06118



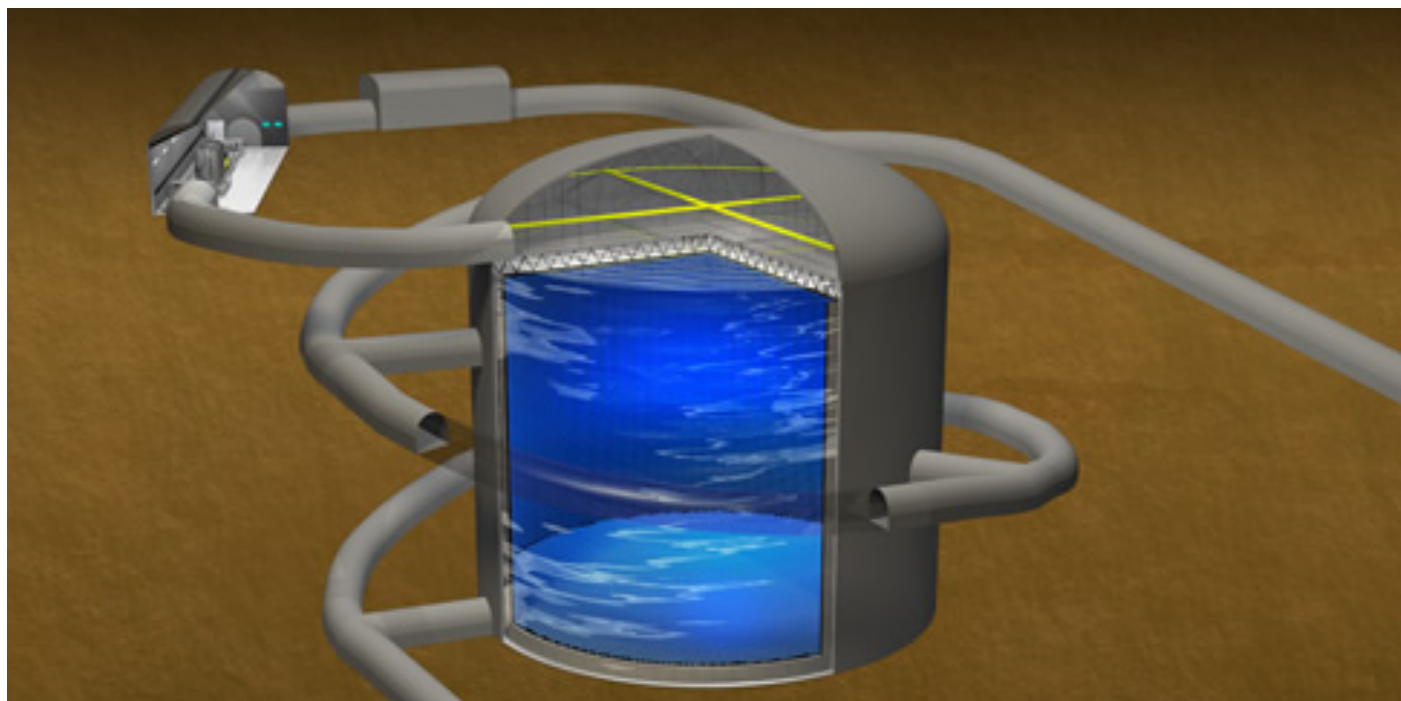
Band indicates range  
 $0.4 < \sin^2 \theta_{23} < 0.6$



The second detector in Korea allows us to better measure the CP-phase, compared with both detectors in Japan

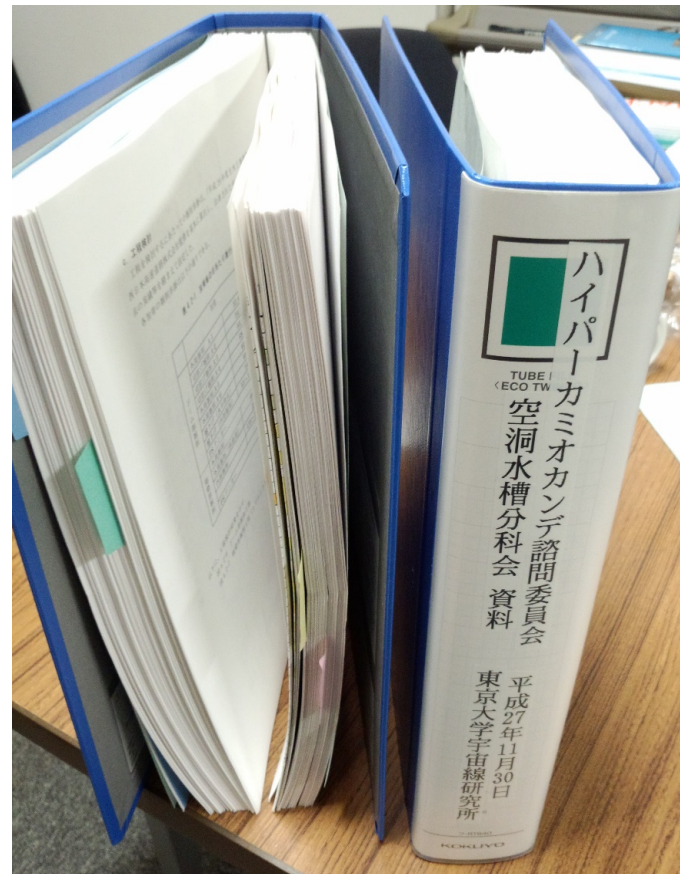
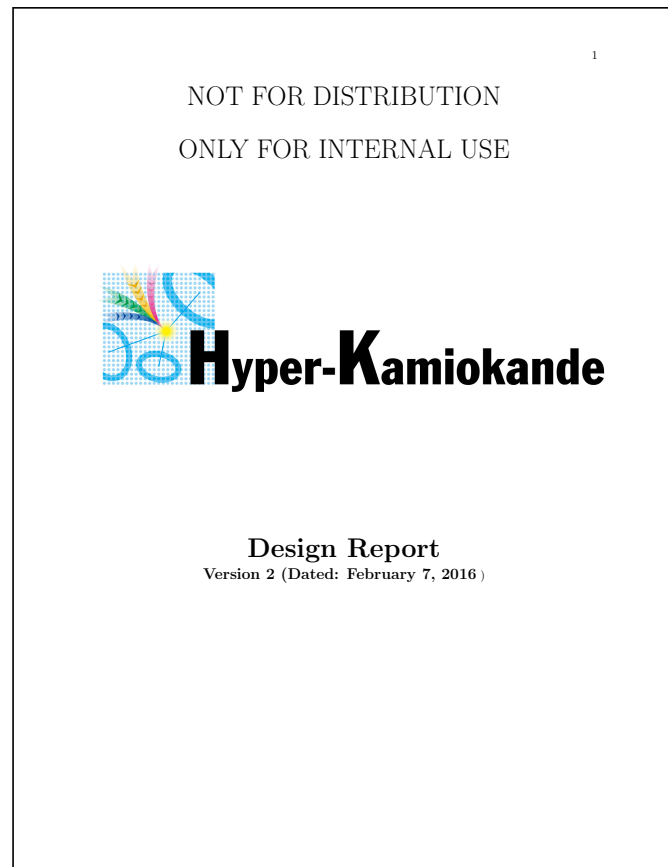
# Project status in Japan

- ‘Hyper-K Design Report’ released
  - KEK preprint 2016-21, ICRR-Report-701-2016-1
- Strong commitment from host institutes: ICRR, U.Tokyo and KEK (MoU for Hyper-K)
- Strong support from Japanese communities
  - Cosmic-ray (CRC) and high-energy (JAHEP)
- Science Council of Japan selected Hyper-K as one of the top priority large-scale projects in ‘Master Plan 2017’
- Hyper-K listed in ‘Roadmap 2017’ of MEXT (funding agency)
- Aim to begin operation in JFY 2026





# Hyper-K Design Report



ICRR-Report-701-2016-1  
KEK Preprint 2016-21

- Hyper-K baseline design and physics sensitivities in KEK preprint and ICRR report
- Being updated by adopting recent progress of detector design and physics sensitivity studies

# Summary

- **Wide Physics topics, many discovery potentials**
  - $\nu$  CPV: 76% of  $\delta$  space w/  $3\sigma$ ,  $\delta$  resolution  $<20^\circ$
  - Proton decay discovery sensitivity reaches  $10^{35}$  years
  - SN burst, SN relic  $\nu$ , indirect WIMP search, etc
  - Physics sensitivity enhanced with new photosensor
- **Project is boosted toward an early realization**
  - International proto-collaboration formed
  - Hyper-K Design Report released
  - Strong support from Japanese communities and host institutes
  - Selected in 'Master Plan' of Science Council of Japan
  - Listed in 'Roadmap 2017' of MEXT
  - Aim to begin the operation in JFY 2026
- **Open for new collaborators**