

Neutrino physics

- from Super-K to Hyper-K -

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3-flavor oscillation scheme

$$\text{Flavor eigenstate} \quad \text{Mass eigenstate}$$

$${}^t(\nu_e, \nu_\mu, \nu_\tau) = U_{\alpha i}^{MNS} {}^t(\nu_1, \nu_2, \nu_3)$$

U^{MNS} : Maki-Nakagawa-Sakata Matrix

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} \cos \vartheta_{12} & \sin \vartheta_{12} & 0 \\ -\sin \vartheta_{12} & \cos \vartheta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \cos \vartheta_{13} & 0 & \sin \vartheta_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin \vartheta_{13} e^{i\delta} & 0 & \cos \vartheta_{13} \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \vartheta_{23} & \sin \vartheta_{23} \\ 0 & -\sin \vartheta_{23} & \cos \vartheta_{23} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$P(\bar{\nu}_\alpha \rightarrow \bar{\nu}_\beta) = \delta_{\alpha\beta} - 4 \sum_{j>i} \operatorname{Re}(U_{\alpha i}^* U_{\beta i} U_{\alpha i}^* U_{\beta j}) \sin^2 \frac{(m_j^2 - m_i^2)L}{4E_\nu}$$

$$(\mp 2 \sum_{j>i} \operatorname{Im}(U_{\alpha i}^* U_{\beta i} U_{\alpha i}^* U_{\beta j}) \sin \frac{(m_j^2 - m_i^2)L}{2E_\nu}$$

Parameterized by 4 (mixing matrix) and 2 (difference of squared masses)

$\theta_{23} \sim 45 \pm 5^\circ$
$\Delta m^2_{23} = 2.4 \times 10^{-3} \text{ eV}^2$
Atmospheric ν , Accelerator ν

$\theta_{12} \sim 34 \pm 3^\circ$
$\Delta m^2_{21} = 7.6 \times 10^{-5} \text{ eV}^2$
Solar ν , Reactor ν

$\theta_{13} \sim 9^\circ$
Accelerator ν , Reactor ν

$\delta = \text{unknown}$
Accelerator ν , Atmospheric ν

$\Delta m^2_{32} = m^2_3 - m^2_2 > 0$ (Normal mass hierarchy) or $\Delta m^2_{32} < 0$ (Inverted mass hierarchy) is also unknown
(key parameter for ν less $\beta\beta$ decay searches)

Majorana nature is also unknown

Super-K detector

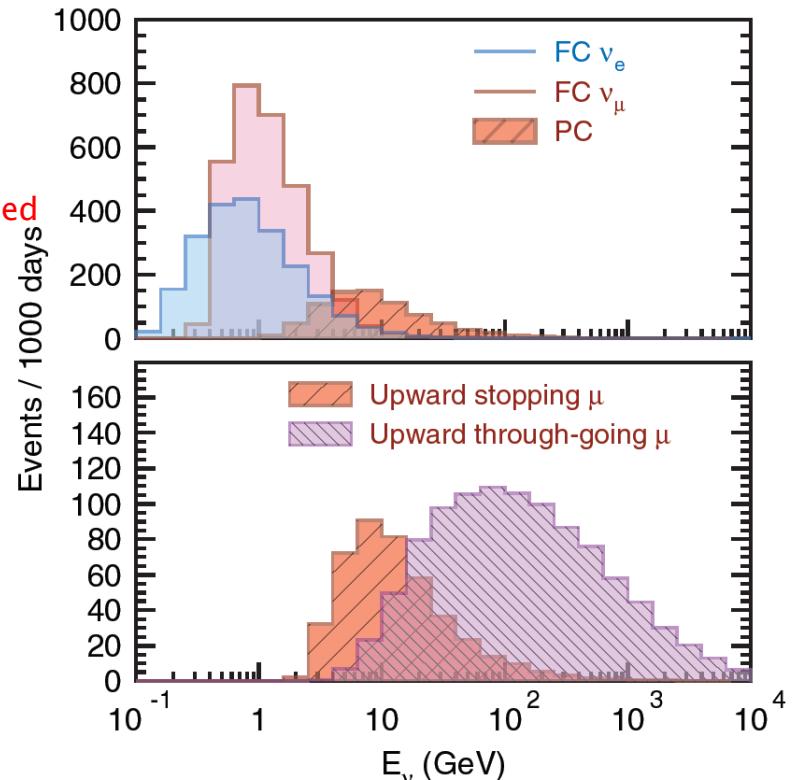
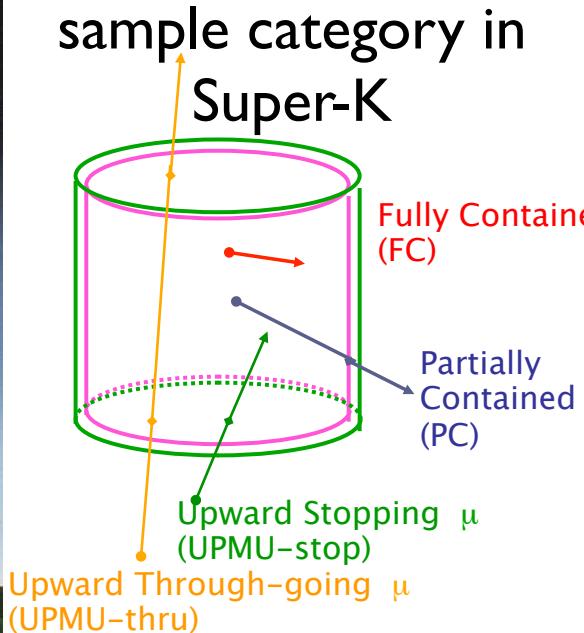
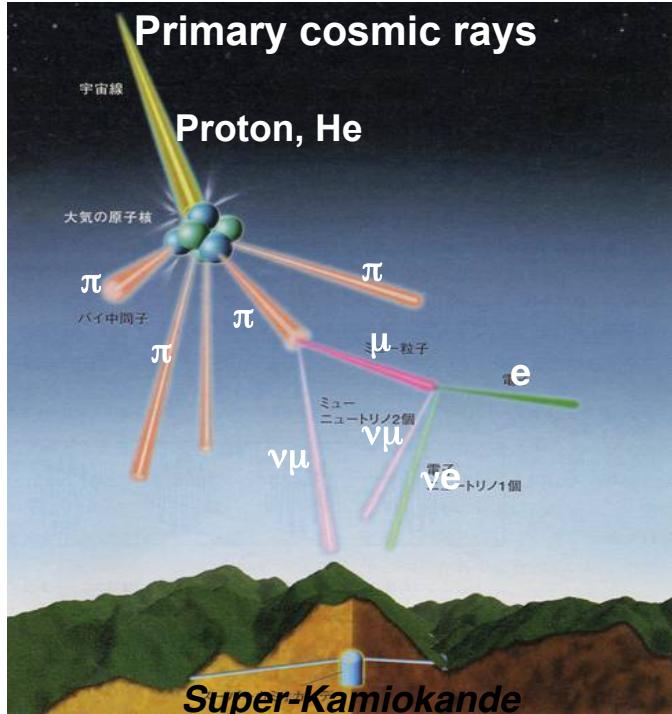


11,129 x 20inch PMTs (inner detector, ID)

- Ring-imaging water Cherenkov detector
 - Fiducial volume 22kton (Total volume 50 kton)
 - Photon yield ~ 10 p.e. / MeV
 - Atmospheric $\nu \sim 10$ events/day
 - Solar $\nu \sim 15$ events/day
 - Accelerator $\nu \sim 1$ events/day (depends on the accelerator power)
 - always ready for Supernova ν and nucleon decays
- Observables
 - Direction of recoiled charged particles (leptons, pions, γ) by neutrinos
 - Particle species (neutrino flavor)
 - Energy
 - Time

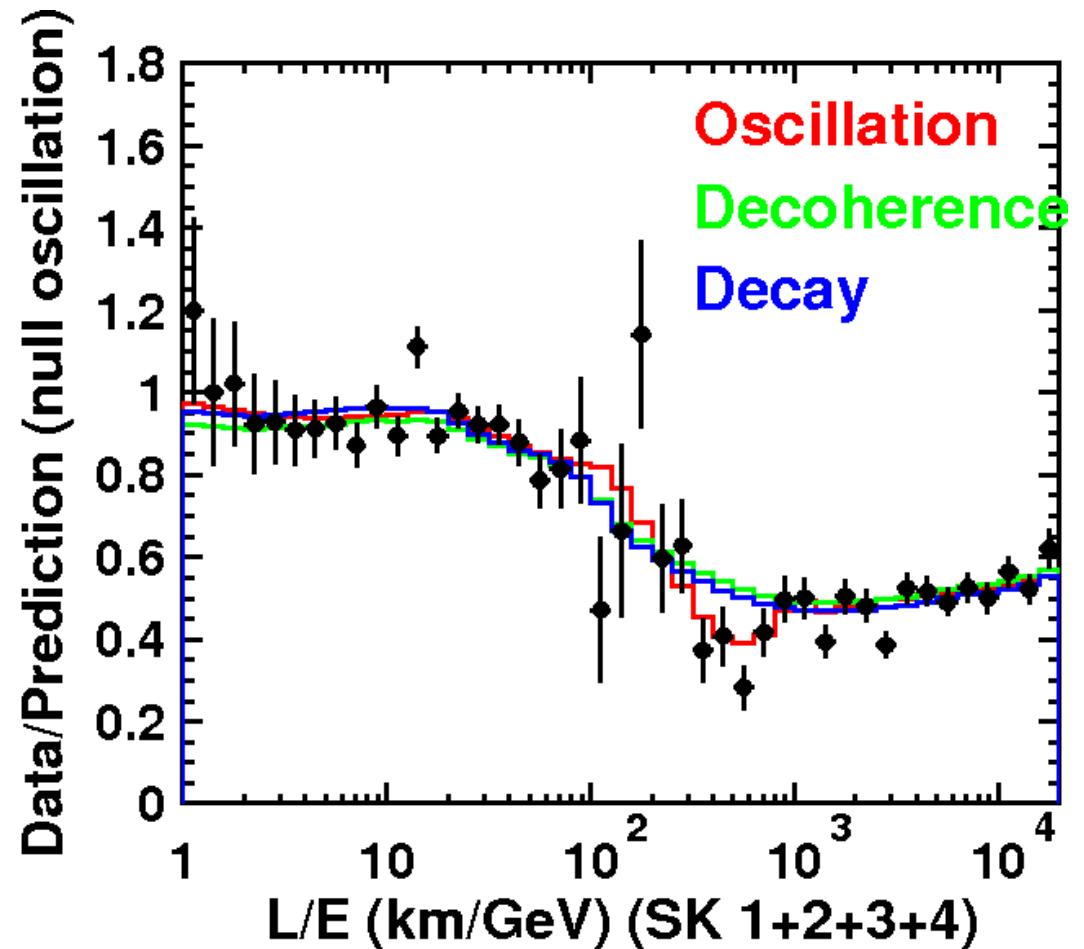
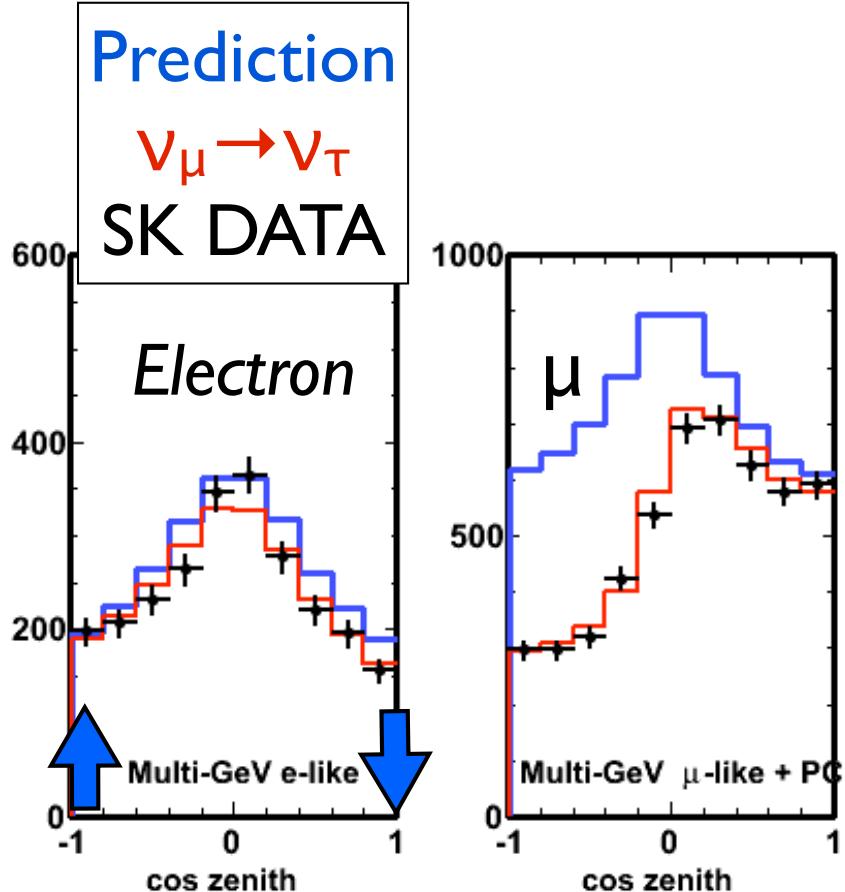


Atmospheric Neutrinos in Super-K



- Unique and complementary studies to solar/reactor/accelerator ν experiments
 - High statistical data (>40,000 events in Super-K)
 - Wide range of ν energy ($0.1\text{GeV} \sim 10^4\text{GeV}$)
 - Wide range of ν baseline (10km downward \sim 13,000km upward)
 - $\nu_\mu:\nu_e \sim 2:1$ at production
 - ν oscillation study by all three flavors (ν_e, ν_μ, ν_τ)
 - test of ν 's exotic property (4th ν , Lorentz violation, etc)

Early history of atm ν 's oscillations



- Dominant effect is ν_μ disappearance (**discovered in 1998**)
- Oscillatory signature (**evidence in 2004**)

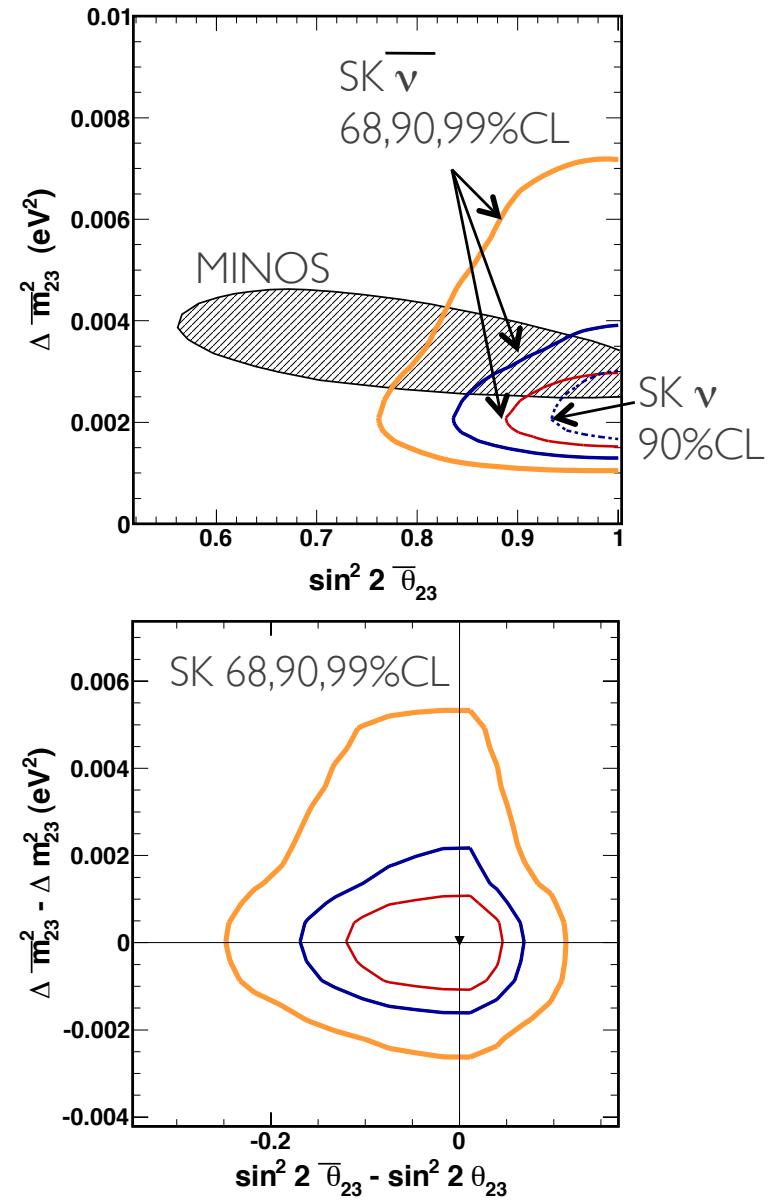
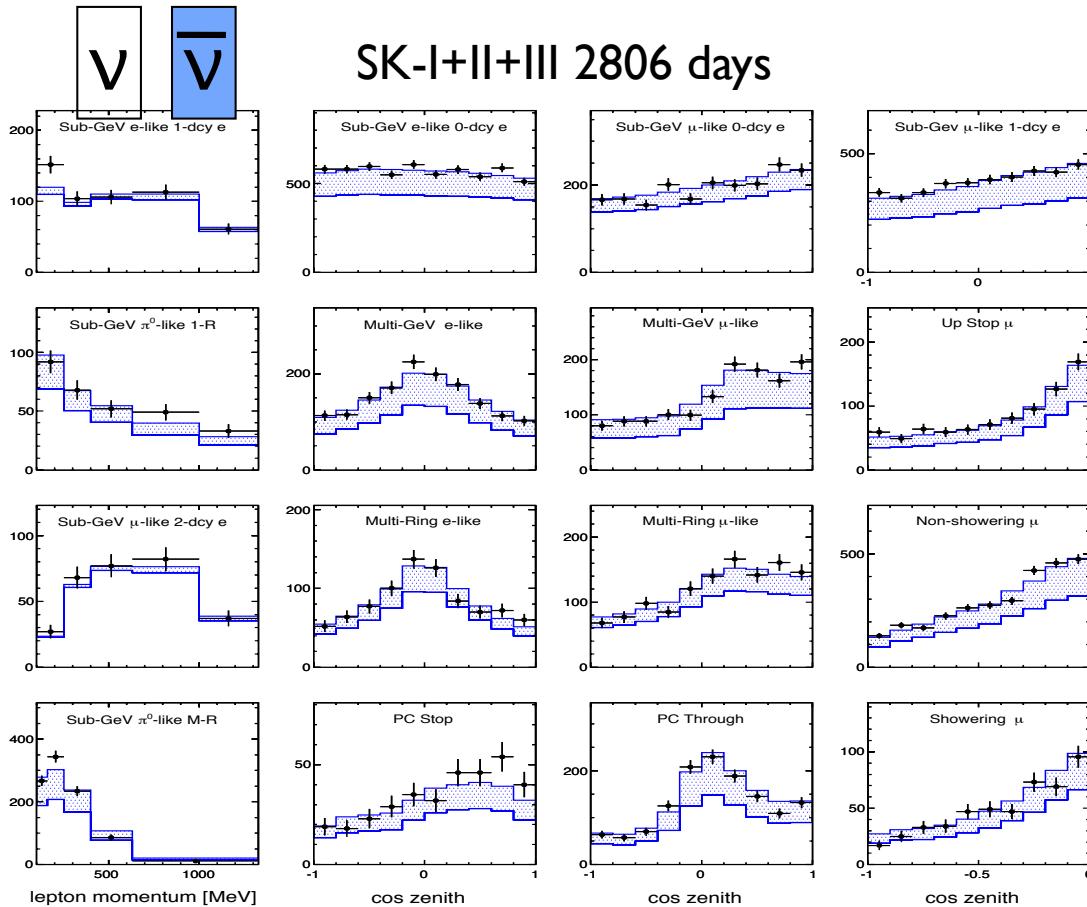


$\bar{\nu}_\mu$ oscillation study

$$P(\nu_\mu \rightarrow \nu_\tau) = 1 - \sin^2 2\vartheta \sin\left(\frac{\Delta m^2 L}{E}\right)$$

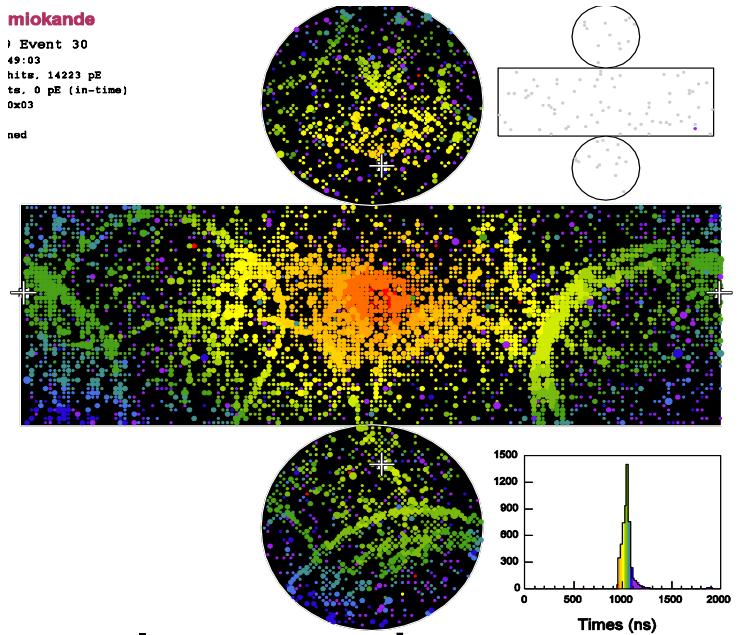
$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_\tau) = 1 - \sin^2 2\bar{\vartheta} \sin\left(\frac{\Delta \bar{m}^2 L}{E}\right)$$

- ν and $\bar{\nu}$ differ
- cross sections
- flux
- event topology



Good constraints on anti-neutrino parameters.
Data is consistent with CPT conservation.

Evidence for τ neutrino appearance



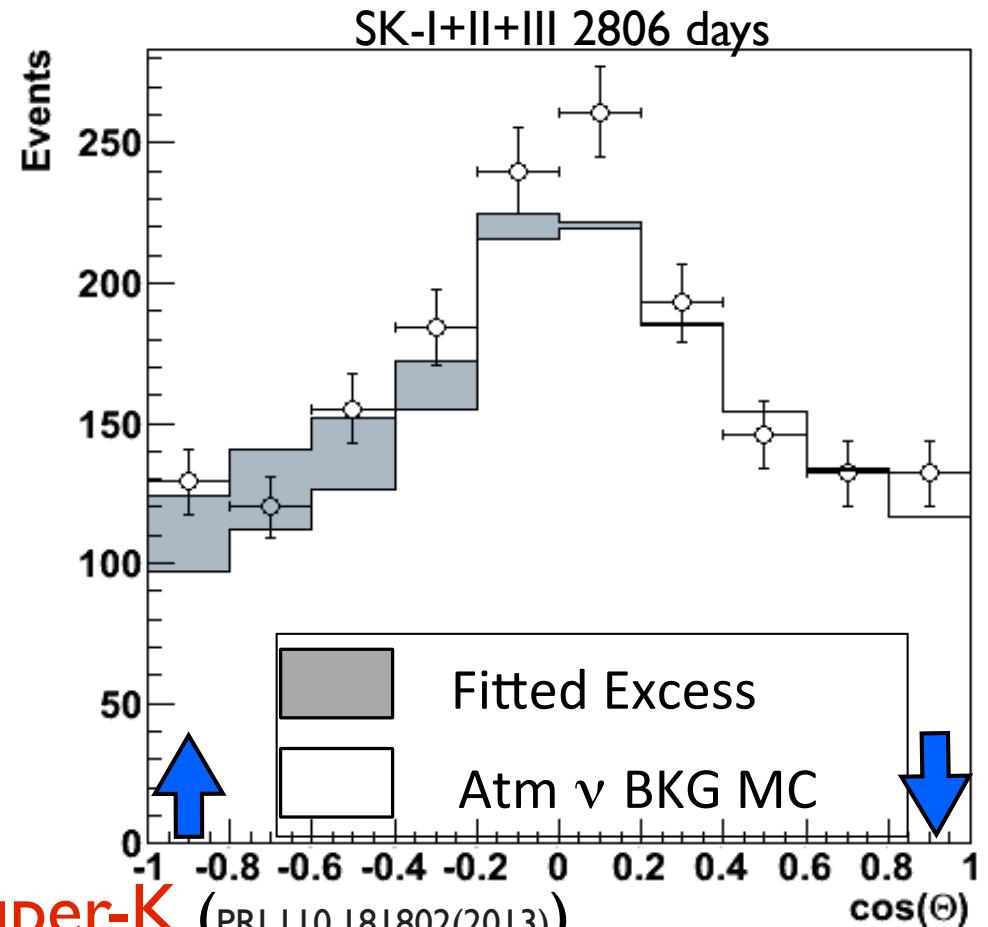
Neural network to enhance events consistent with hadronic decays of τ

$$N_{\tau}^{\text{DATA}}/N_{\tau}^{\text{exp}} = 1.42 \pm 0.35(\text{stat})^{+0.14}_{-0.12}(\text{syst})$$

Evidence(3.8σ significance)@Super-K (PRL 110, 181802(2013))
5 τ observation ($>5\sigma$) @ OPERA (PRL 115, 121802(2015))

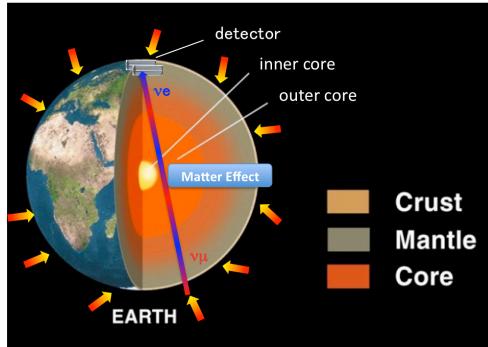
PRL 110, 181802 (2013)

Zenith Distribution of τ -like events



Atmospheric ν anomaly (problem) is finally concluded.
More precise measurement will be conducted in future.

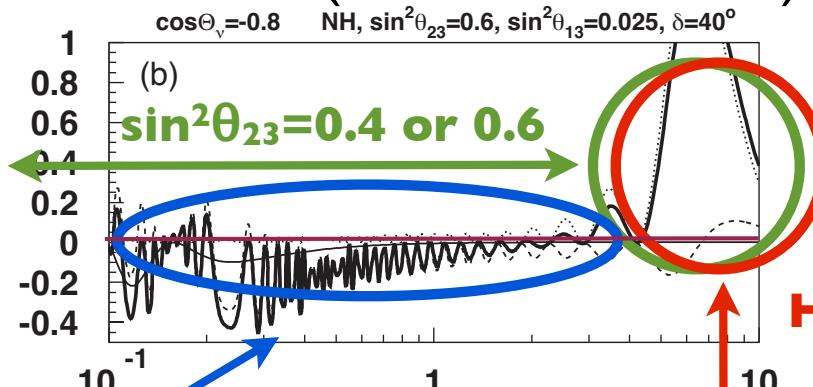
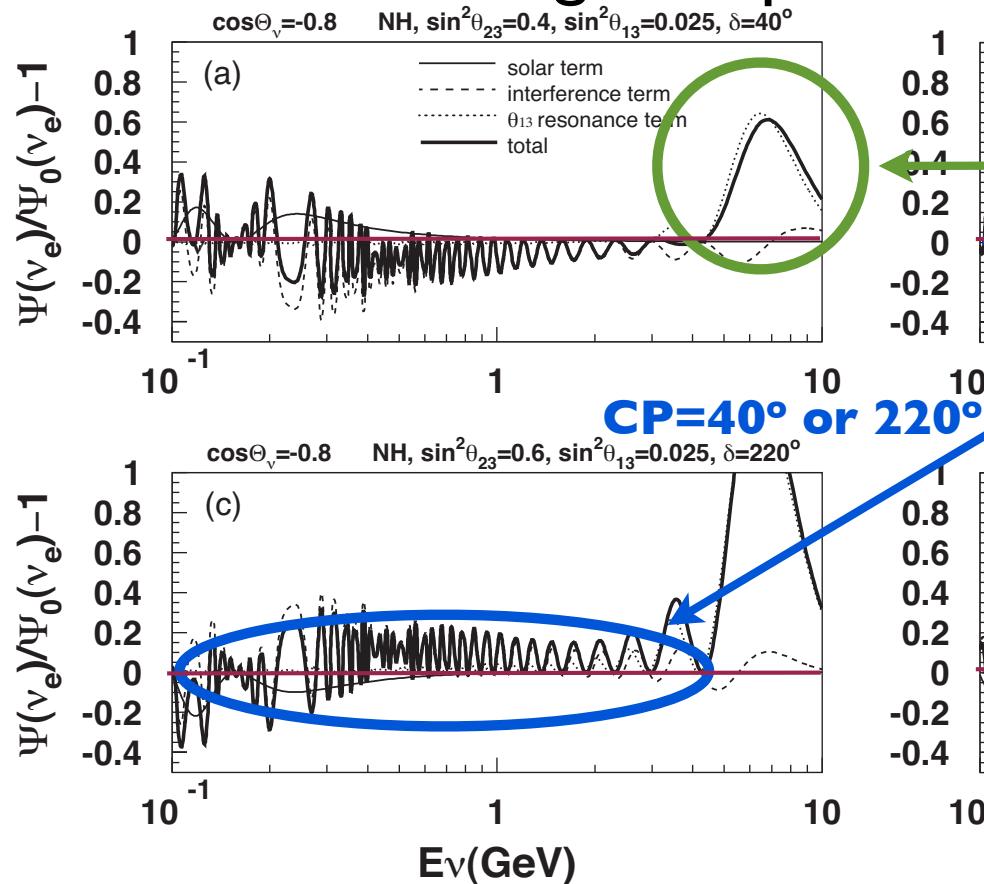
Atmospheric $\bar{\nu}_e$ oscillation



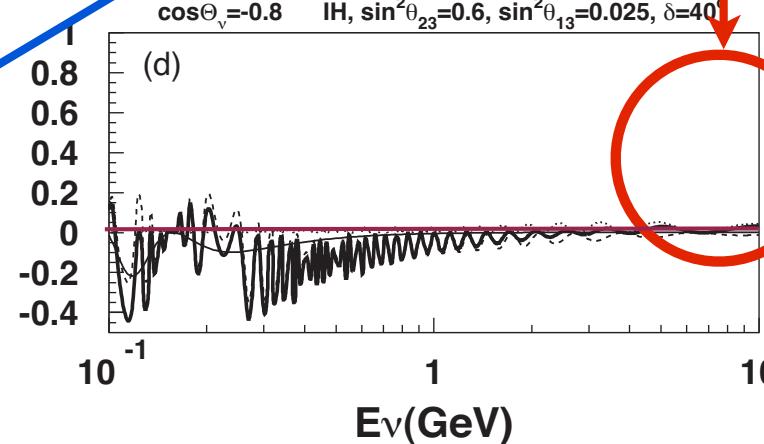
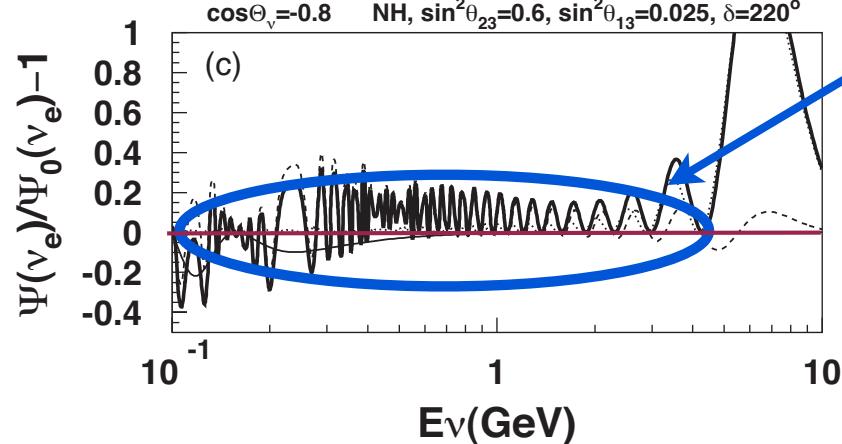
Through the matter effect in the Earth, we study on

- **Mass hierarchy** : resonance in multi-GeV $\bar{\nu}_e$ or ν_e
- **CP δ** : interference btw two Δm^2 driven oscill.
- **θ_{23} octant** : magnitude of the resonance

“Fractional change of upward ν_e flux ($\cos\Theta_{\text{zenith}}=-0.8$)”

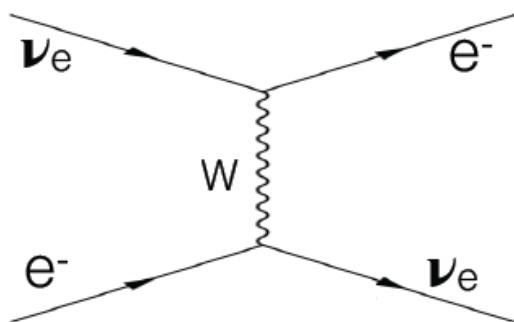


Hierarchy is NH or IH



Resonance in $\bar{\nu}_e$ (not shown) in the case of IH.

Test of Earth's matter effect

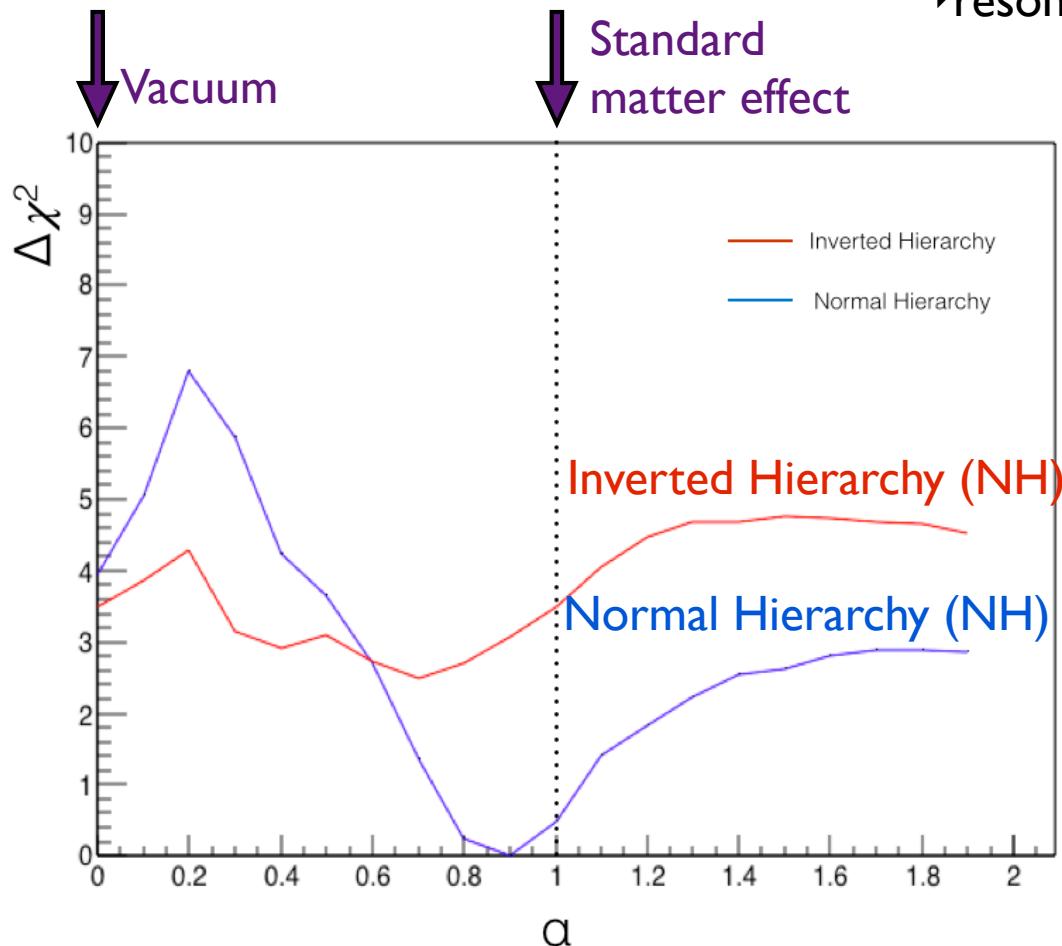


$$V_{cc}^{\nu_e} = \pm \sqrt{2} G_F N_e = \pm 7.56 \times 10^{-14} \left(\frac{\rho}{g/cm^3} \right) Y_e eV$$

($V>0$ for ν , $V<0$ for anti- ν)

$$\sin 2\theta_m = \frac{\sin 2\theta}{\sqrt{\sin^2 2\theta + (\cos 2\theta - \frac{2EV}{\Delta m^2})^2}}$$

- $\Delta m^2_{32} > 0$ for Normal Mass Hierarchy (NH)
→ resonant oscillation in ν ($V>0$)
- $\Delta m^2_{32} < 0$ for Inverted Mass Hierarchy (IH)
→ resonant oscillation in anti- ν ($V<0$)



α : scaling factor of the matter effect
($\alpha=1$: standard, $\alpha=0$: vacuum)

χ^2 minimum at $\alpha=0.9$

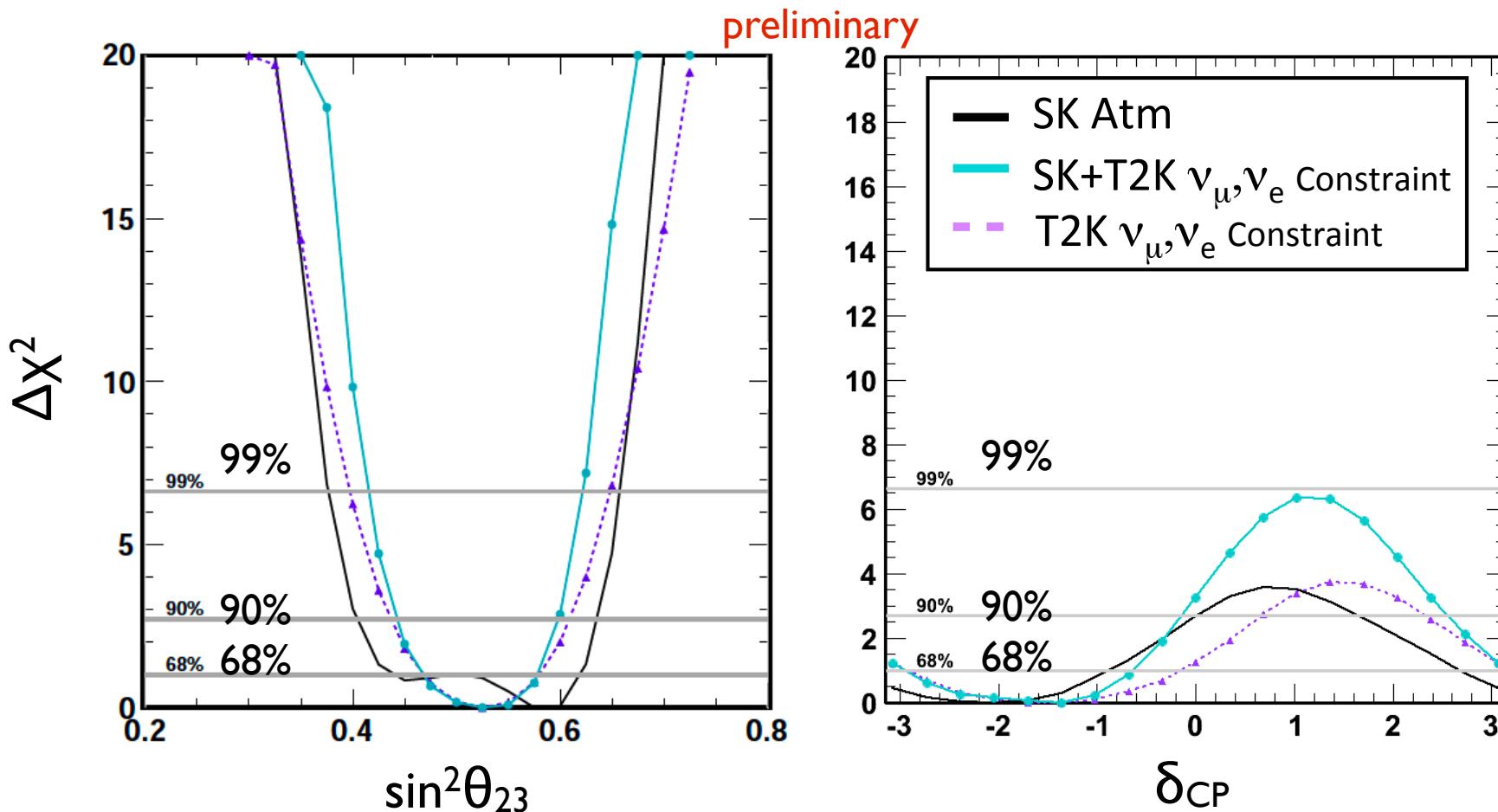
$$\Delta\chi^2_{\text{vac}} = 3.5$$

$$\Delta\chi^2 \text{ "standard" matter} = 0.5$$

$$\chi^2_{\text{IH}} - \chi^2_{\text{NH}} = 3 \quad (\alpha=1 \text{ fixed})$$

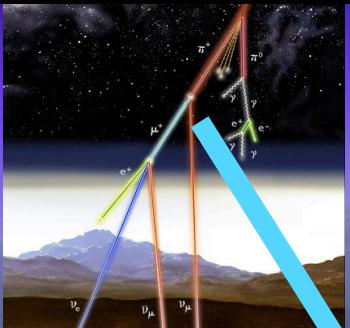
Normal Hierarchy is favored.
More data is necessary, analysis improvement is also going on.

Results w/ T2K's constraints



- Preference of normal hierarchy is slightly strengthened, $\chi^2_{IH} - \chi^2_{NH} = 3.2$ (3.0 for SK only)
- T2K/SK are getting constraint on CP δ
 - Preference of δ_{CP} near $-\pi/2$ is strengthened by T2K
 - CP conservation ($\sin \delta_{CP} = 0$) is still allowed at 90% CL

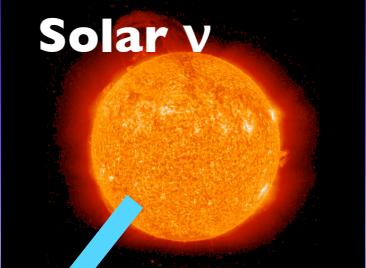
Atmospheric ν



Supernova ν

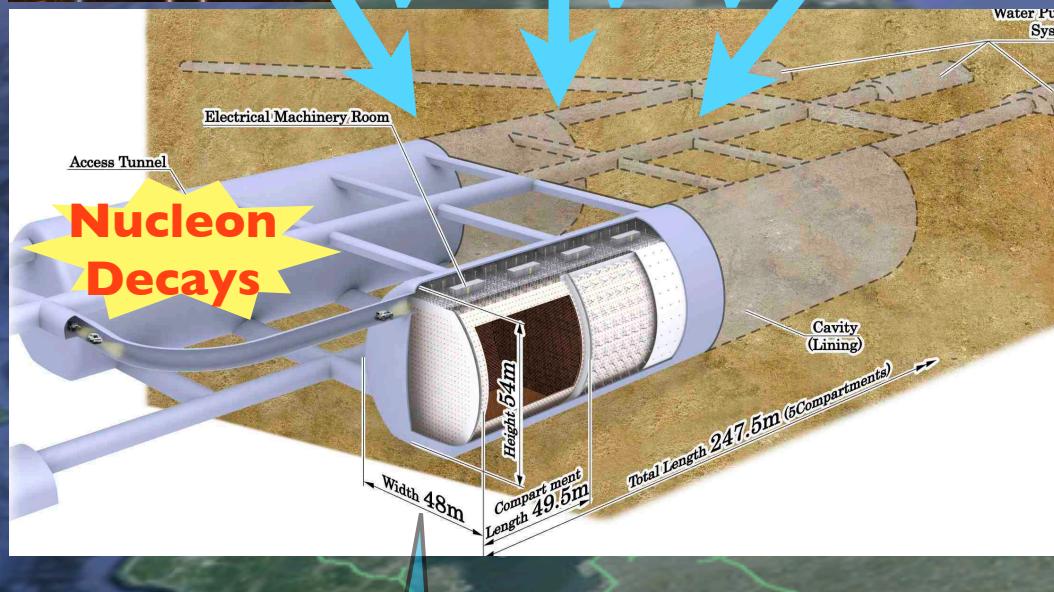
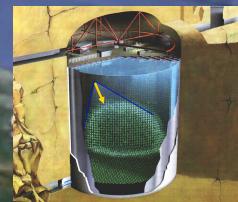


Solar ν



V
V
V

Super-Kamiokande



Nucleon
Decays

Hyper-Kamiokande

25 × Super-K fiducial mass
as neutrino target and
proton decay source

J-PARC

High intensity neutrino and
anti-neutrino beam

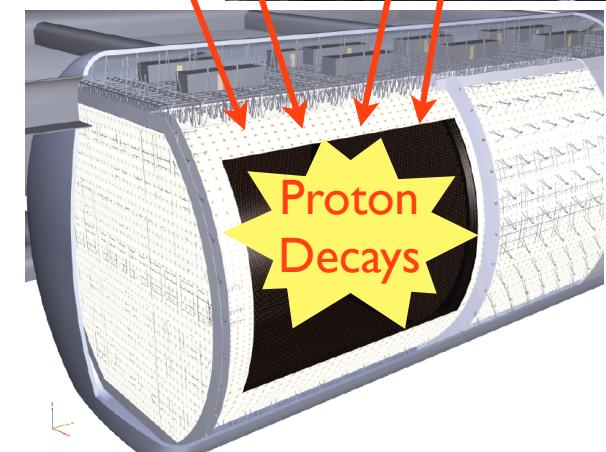
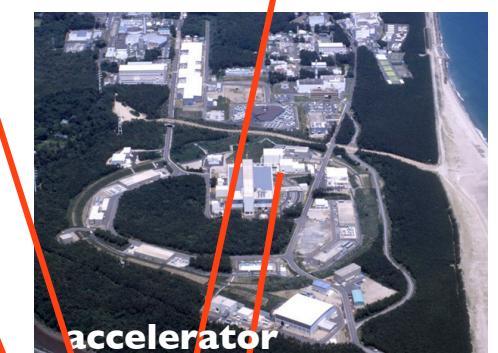
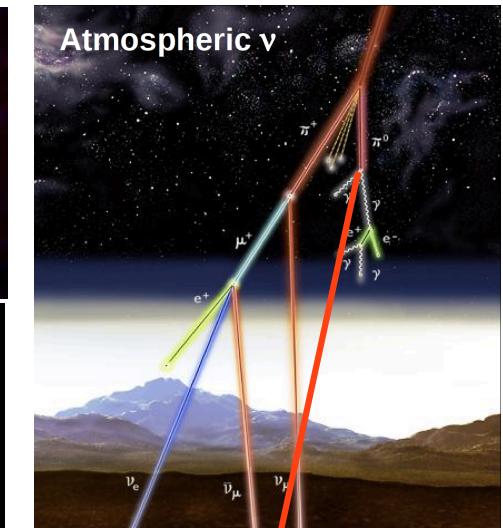
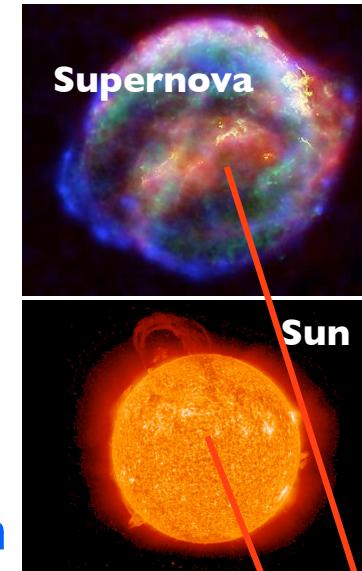


Multi-purpose detector, Hyper-K

Letter of Intent, Hyper-K WG,
arXiv:1109.3262 [hep-ex]

LBL study, Hyper-K WG,
arXiv:1502.05199 and
published in PTEP

- Proton decay 3σ discovery potential
 - 5×10^{34} years for $p \rightarrow e^+ \pi^0$
 - 1×10^{34} years for $p \rightarrow \nu K^+$
- Comprehensive study on ν oscillations
 - CPV (76% of δ space at 3σ), $< 20^\circ$ precision
 - MH determination for all δ by J-PARC/Atm ν
 - θ_{23} octant: $\sin^2 \theta_{23} < 0.47$ or $\sin^2 \theta_{23} > 0.53$
 - $< 1\%$ precision of Δm^2_{32}
 - test of exotic scenarios by J-PARC/Atm ν
- Astrophysical neutrino observatory
 - Supernova up to 2Mpc distance, $\sim 1SN / 10$ years
 - Supernova relic ν signal ($\sim 200\nu$ events/10yrs)
 - Dark matter neutrinos from Sun, Galaxy, and Earth
 - Solar neutrino $\sim 200\nu$ events/day



Hyper-K proto-collaboration w/ cooperation of KEK-IPNS and UTokyo-ICRR

Inaugural Symposium on 1/31, 2015



Hyper-K Proto-Collaboration has been formed

MoU signing by KEK/ICRR

- KEK-IPNS and Tokyo-ICRR signed the MOU of the cooperation in promoting the Hyper-Kamiokande.



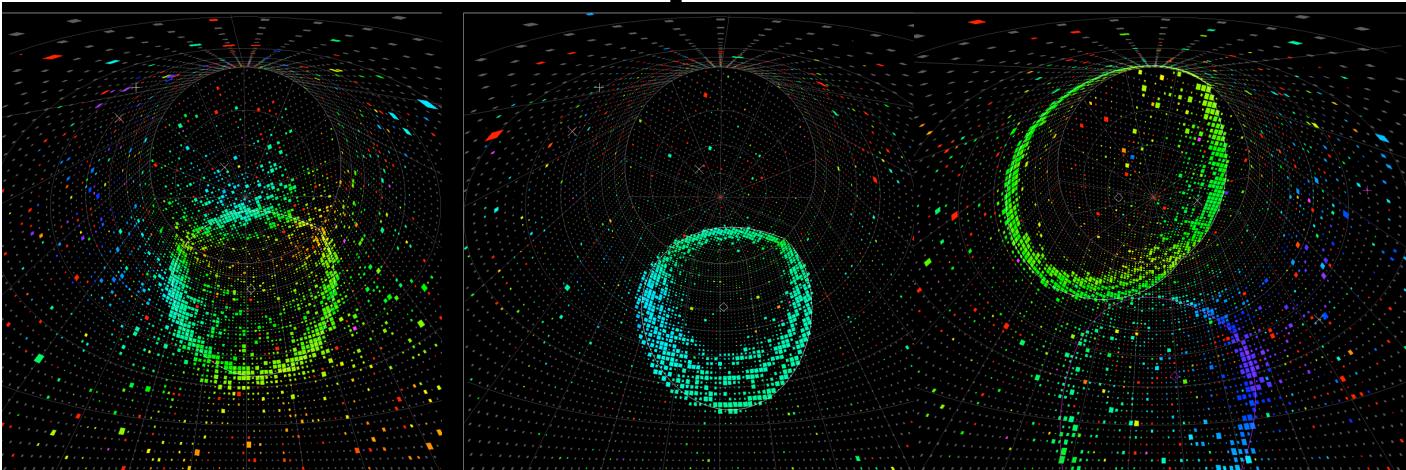
What's next?: Design Report

- **Design Report** to be prepared in 2015
 - Optimum design, Construction cost/period, Beam&Near detectors, International responsibility
- **International Advisory Committee** under **KEK-IPNS/ICRR** to develop the project
- Start **budget request**, and **Start construction** in 2018
→ start operation in ~2025

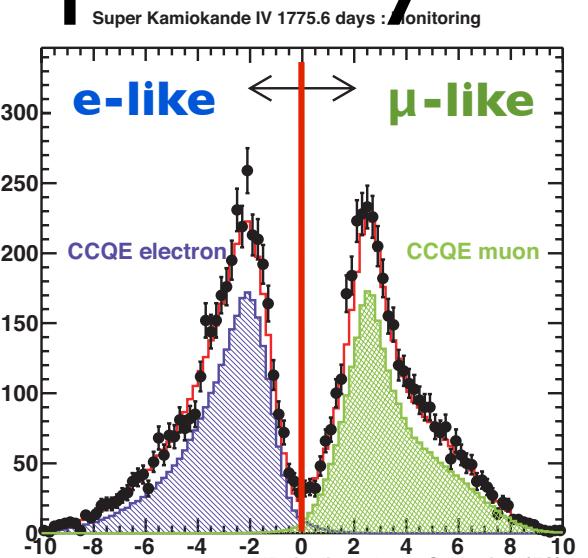
It is critical period to promote the project

Still open for new collaborators

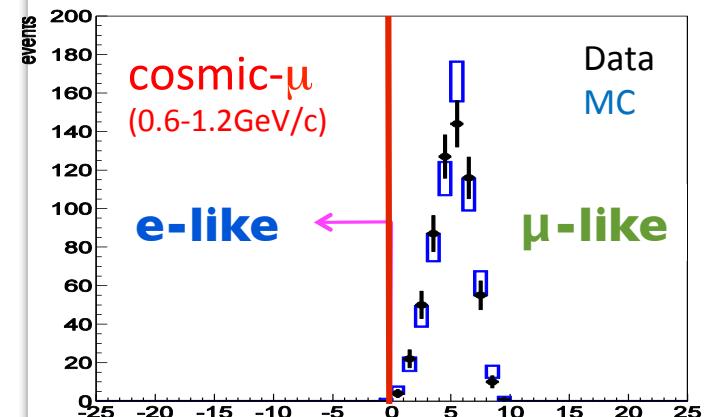
Detector performance for p-decays



- High mass (1Mton scale, 20×Super-K)
- Good ring-imaging capability at $\sim 1\text{ GeV}$
 - atmospheric ν , proton decays, accelerator ν
- Excellent particle ID (e or μ) capability $> 99\%$
- Energy resolution for e and $\mu \sim 3\%$
- opportunity to improve more
- for proton decay search via $p \rightarrow e^+ \pi^0$
 - good $\sim 5\%$ invariant proton mass resolution
 - high 40% signal efficiency
 - 99.998% atmospheric ν BG rejection



PID likelihood (atmν)

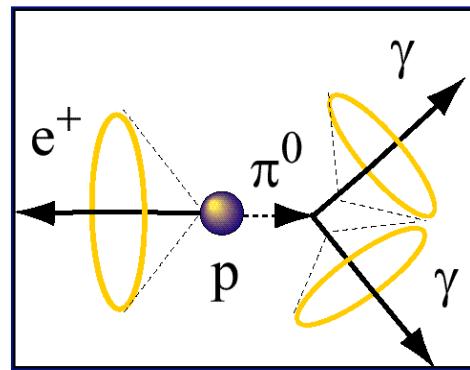


PID likelihood (CRμ)

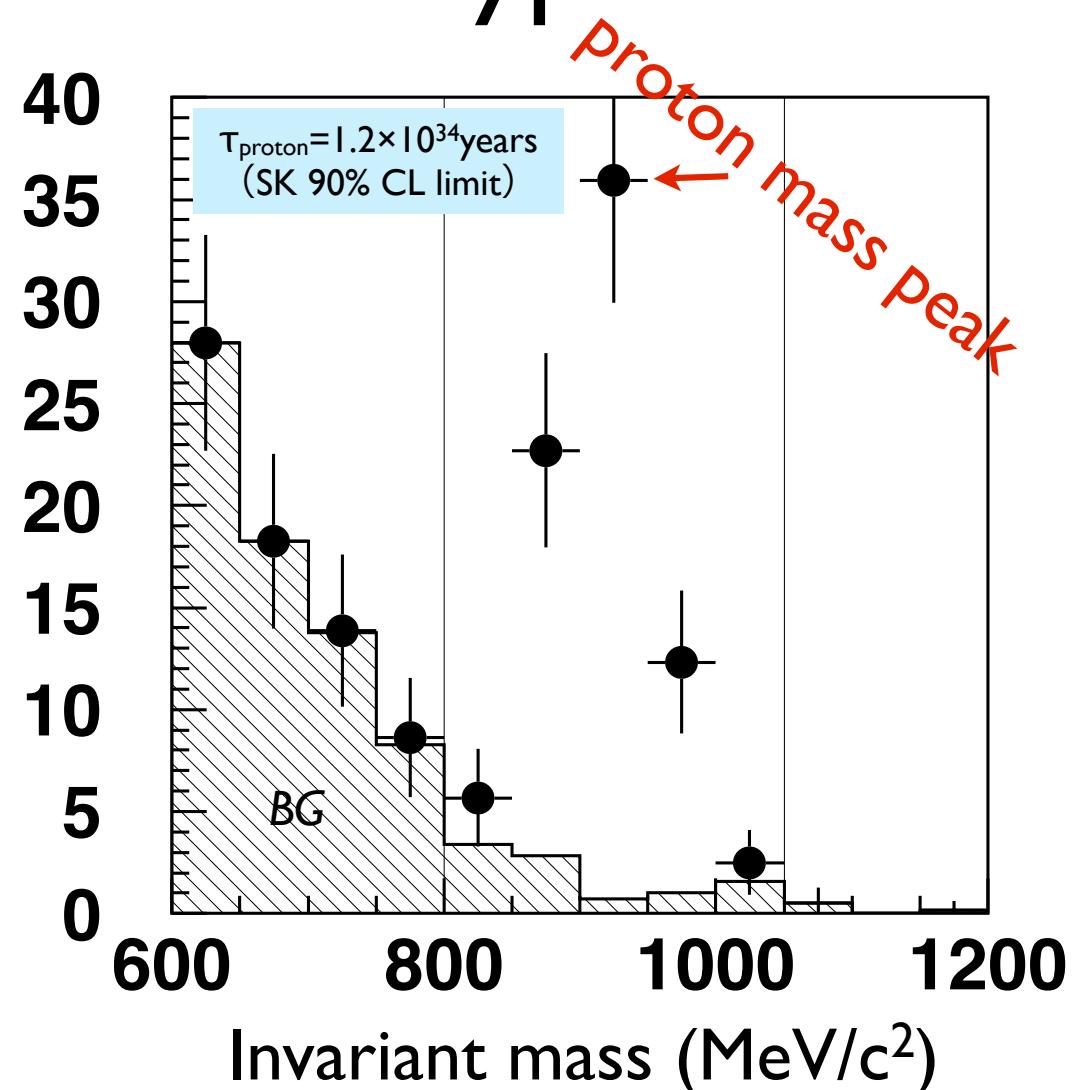
mis-PID:
 Data: $0.00 \pm 0.16(\text{stat.})\%$
 MC : $0.10 \pm 0.10(\text{stat.})\%$

Discovery potential in Hyper-K

$$p \rightarrow e^+ \pi^0$$

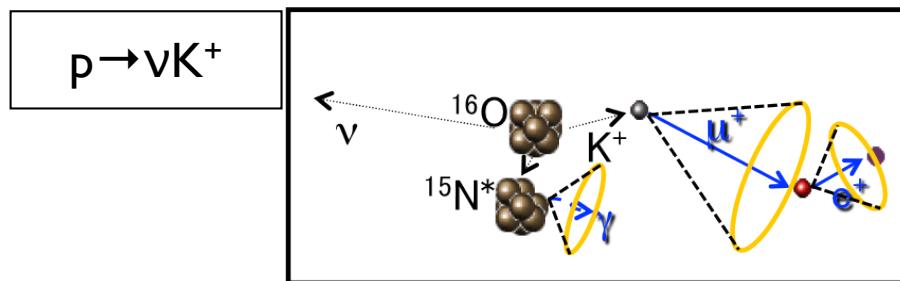


- ▶ Discovery reach (3σ)
▶ $\tau(p \rightarrow e^+ \pi^0) \sim 5 \times 10^{34} \text{ years}$ (HK 10 yrs)
- ▶ Limit (90%CL)
▶ $\tau(p \rightarrow e^+ \pi^0) > 1 \times 10^{35} \text{ years}$ (HK 10 yrs)

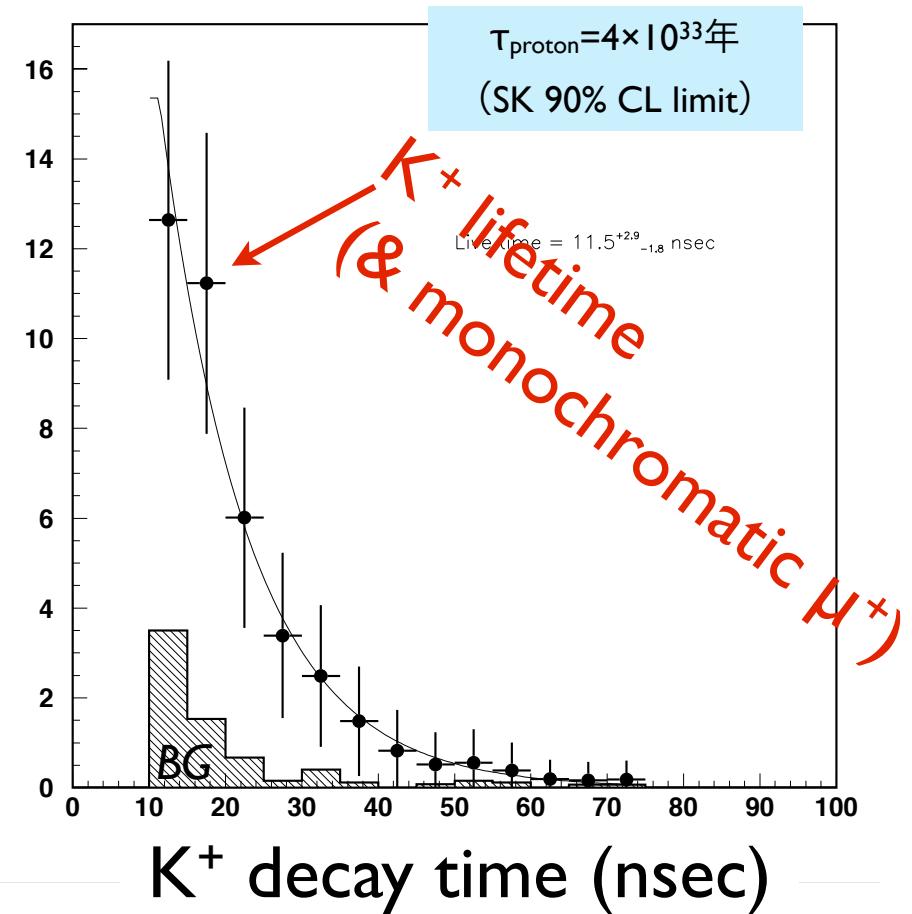
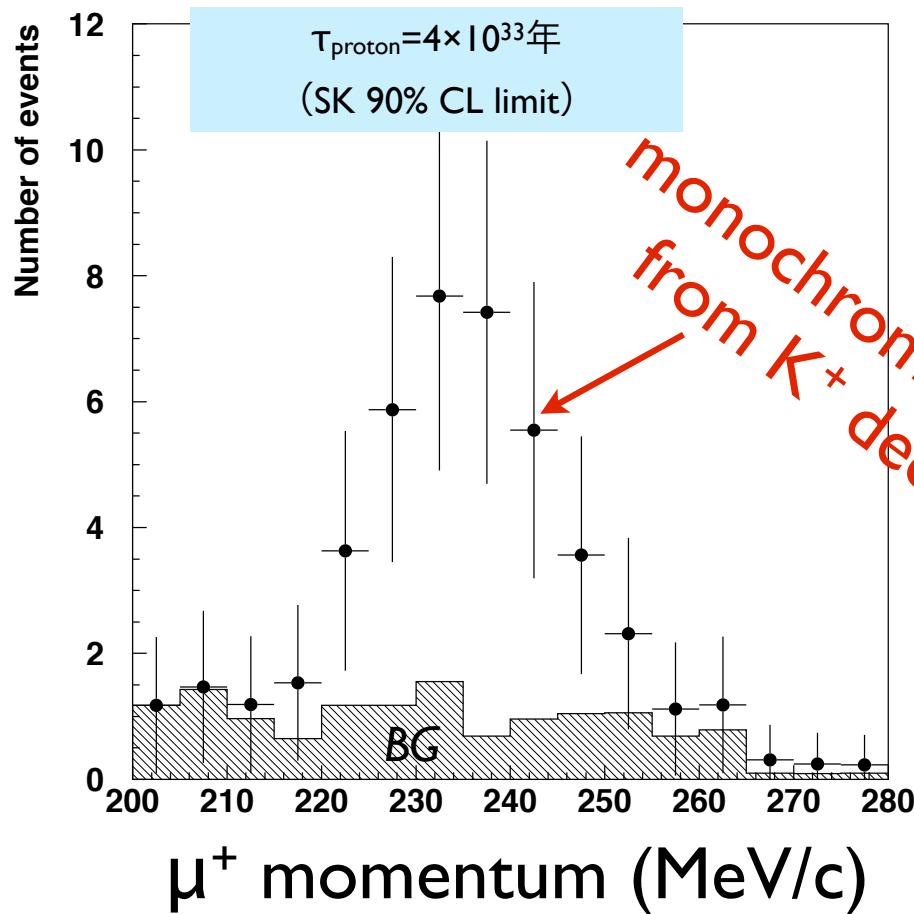


Only realistic proposal to reach the lifetime of 10^{35} years
for $p \rightarrow e^+ \pi^0$

Discovery potential (2)

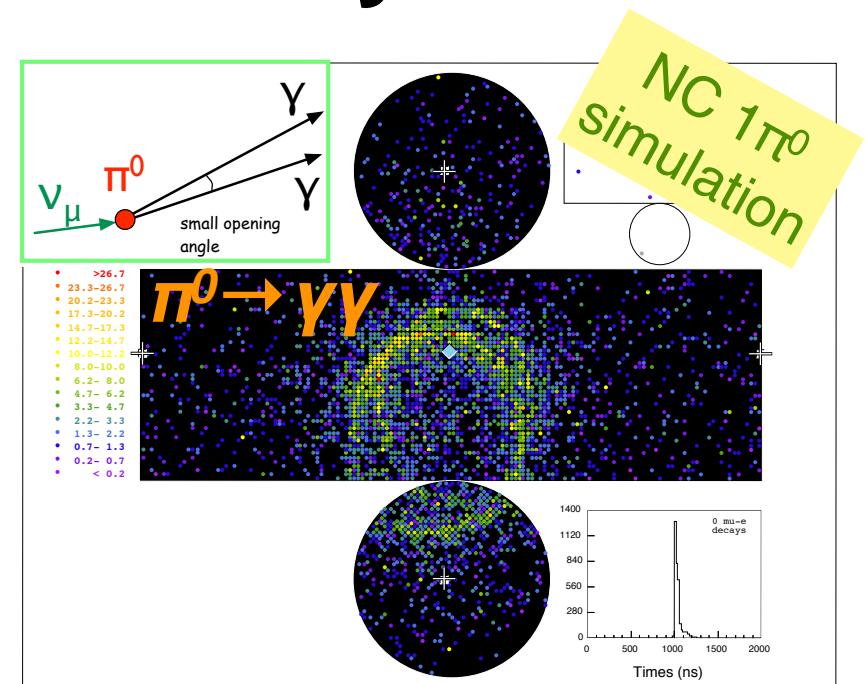
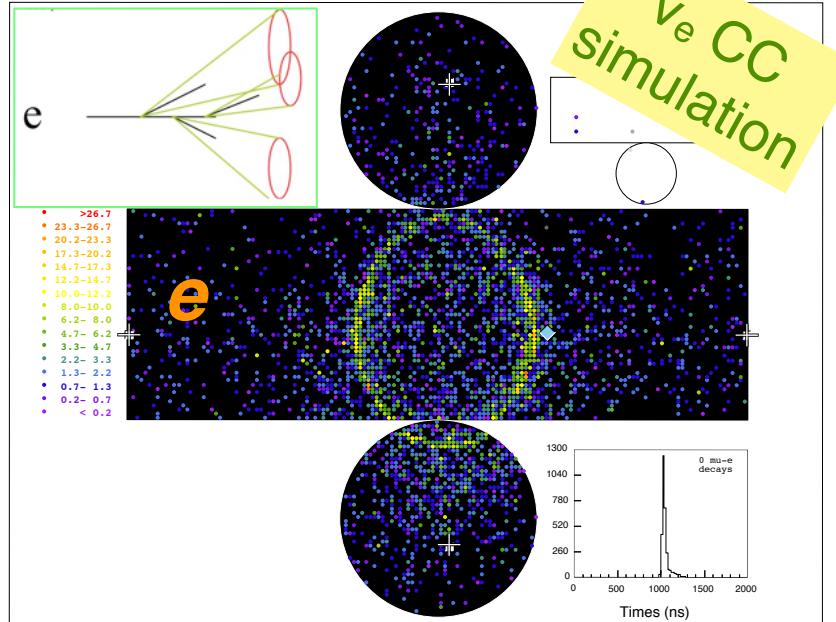


- ▶ Discovery reach (3σ)
 - ▶ $\tau(p \rightarrow v K^+) \sim 1 \times 10^{34} \text{ years}$ (HK 10 yrs)
- ▶ Limit (90% CL)
 - ▶ $\tau(p \rightarrow v K^+) > 3 \times 10^{34} \text{ years}$ (HK 10 yrs)



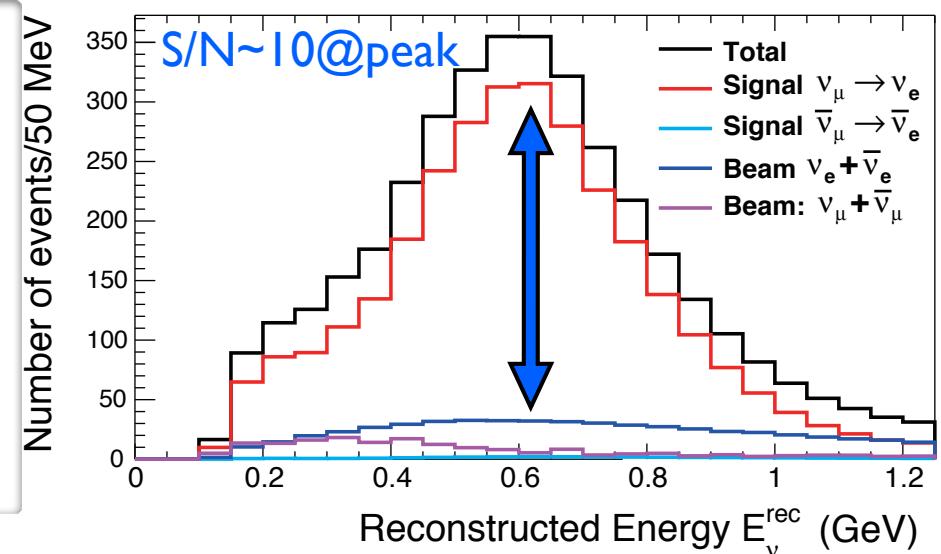
Experimental test on Supersymmetry

Detector performance for J-PARC ν



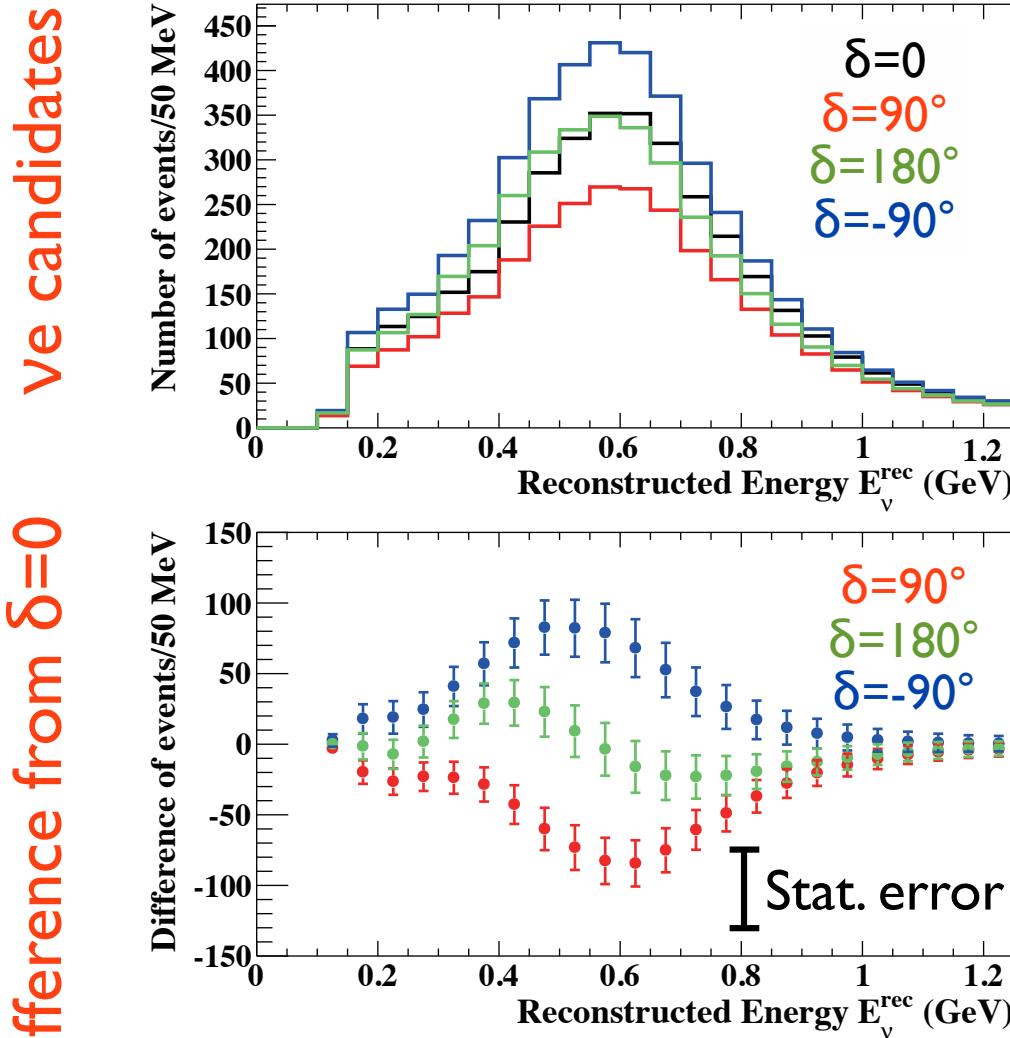
Appearance ν mode

- For ν_e appearance in J-PARC ν_μ beam
 - high 60% ν_e signal efficiency
 - >99.9% ν_μ CC rejection, 99% NC π^0 rejection
- opportunity to improve more

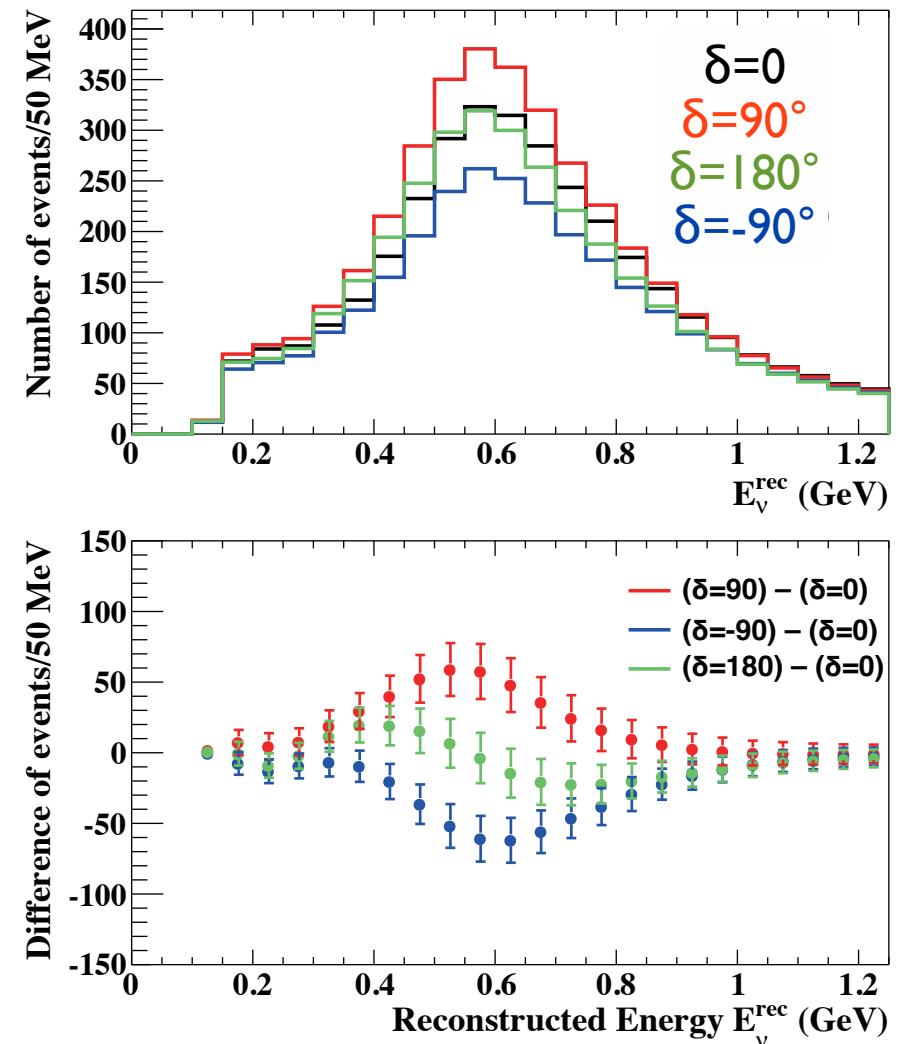


δ_{CP} dependence of observables

Neutrino mode: Appearance



7.5MW $\times 10^7$ s (1.56 $\times 10^{22}$ POT)
Antineutrino mode: Appearance

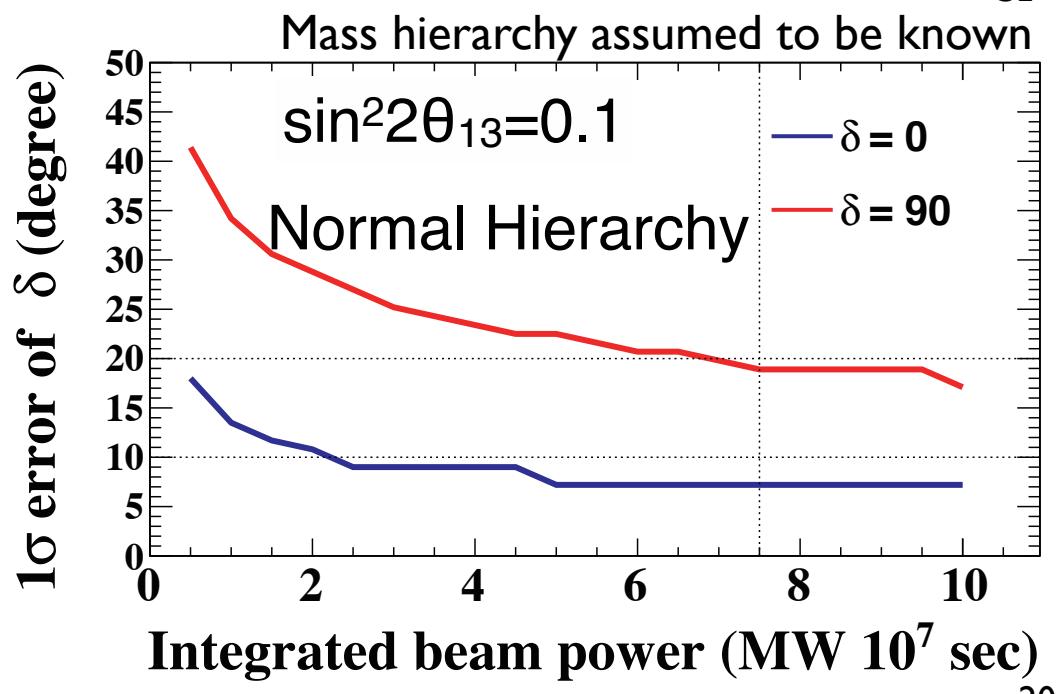
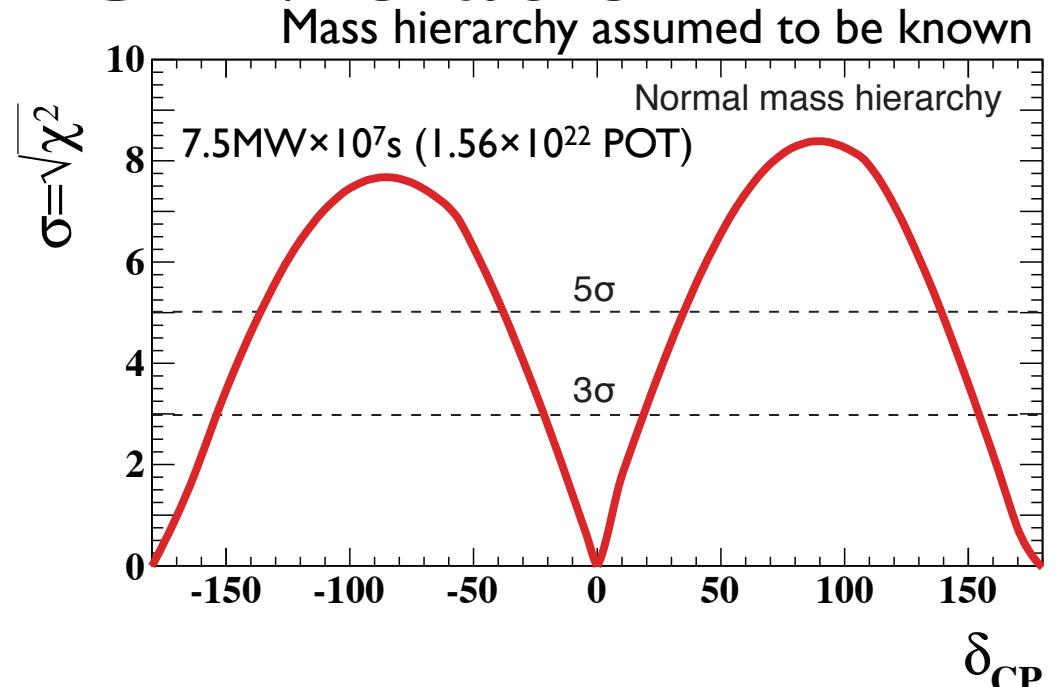


Sensitive to all values of δ with numbers + shape

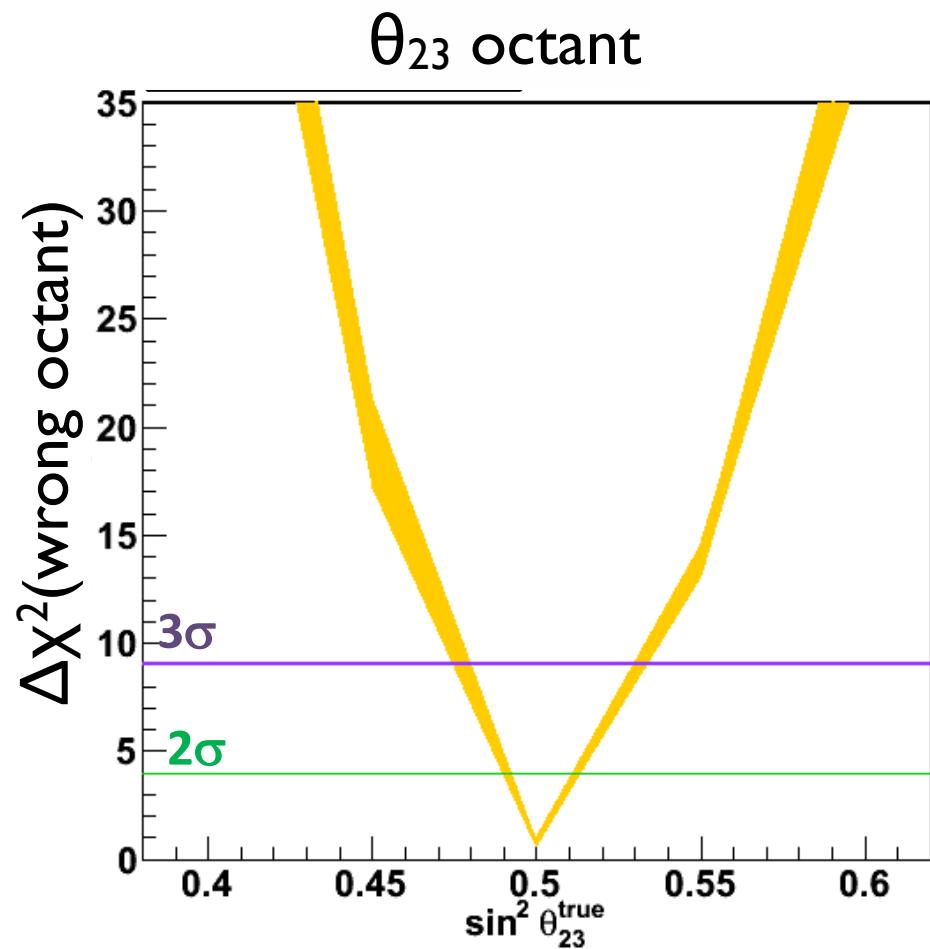
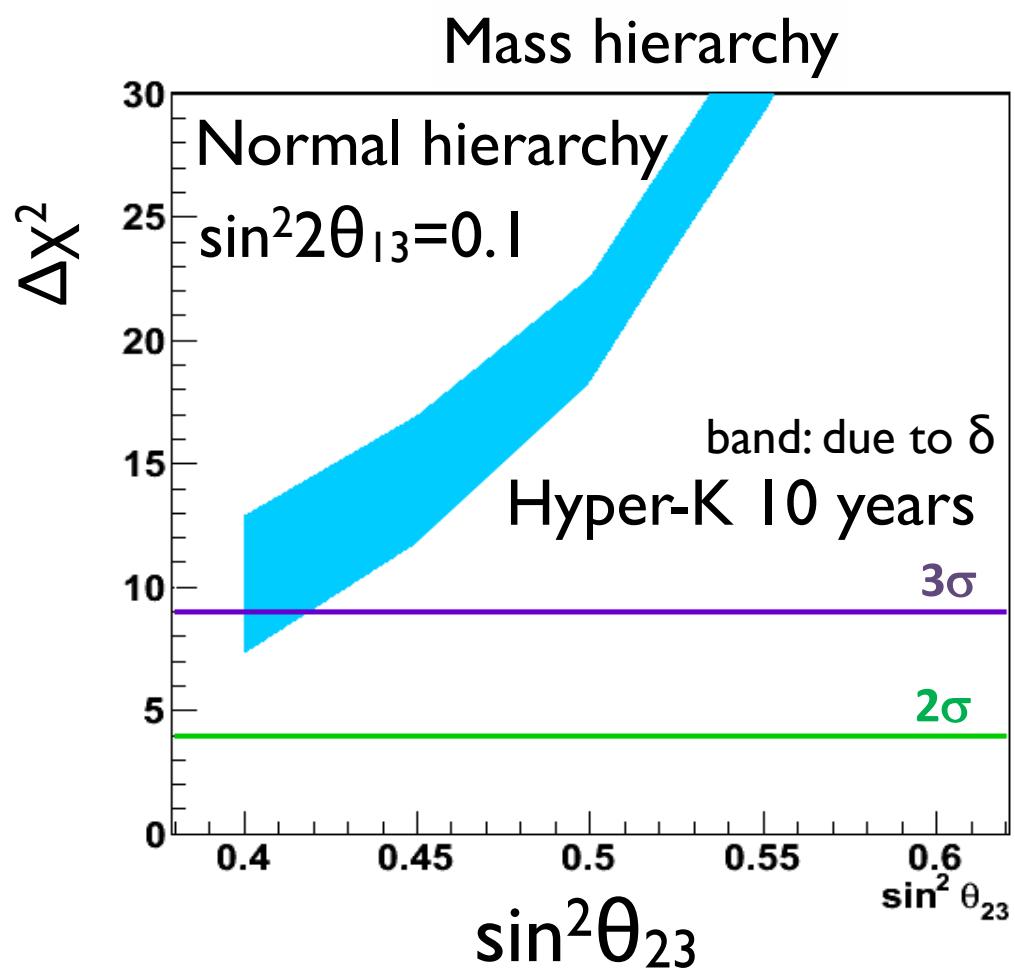
Sensitivity to CP violation

arXiv:1502.05199 and
published in PTEP

- Exclusion of $\sin\delta=0$
 - >3 σ for 76% of δ
 - >5 σ for 58% of δ
- 8°-19° precision depending on the true value of δ



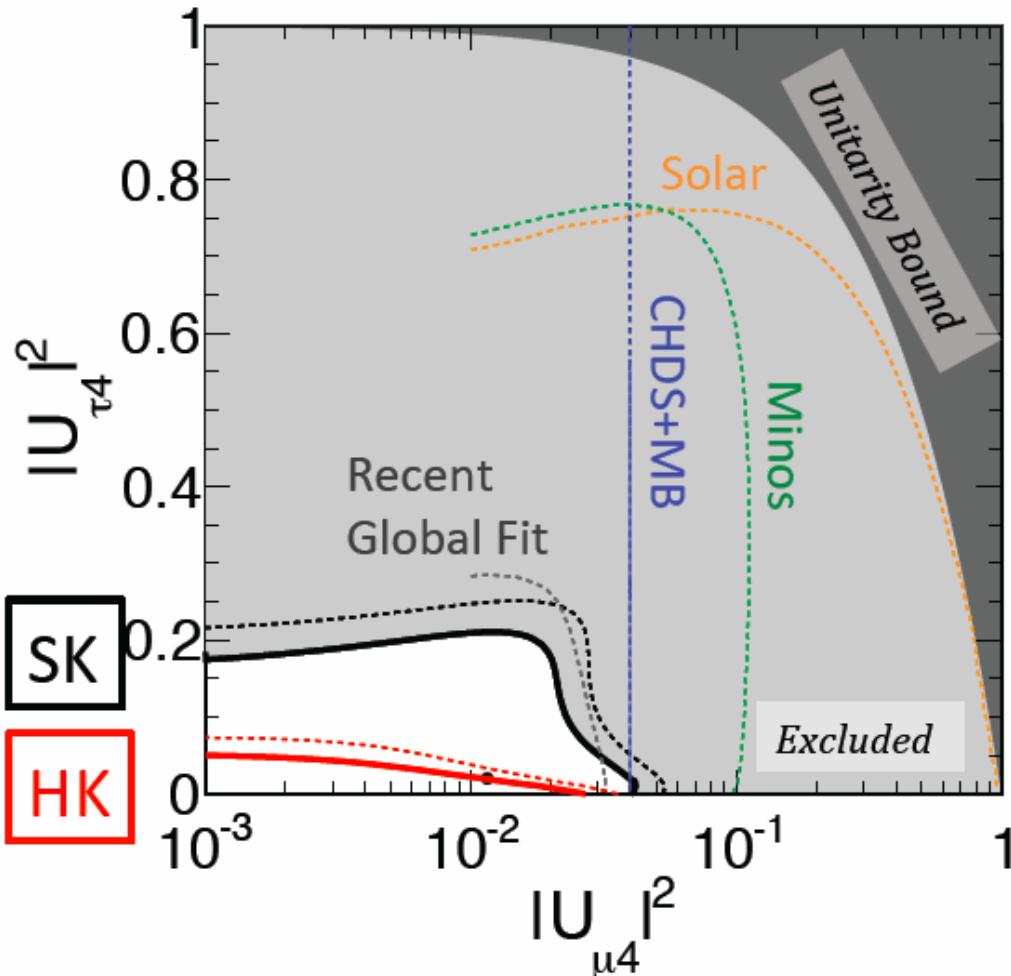
Atmospheric ν



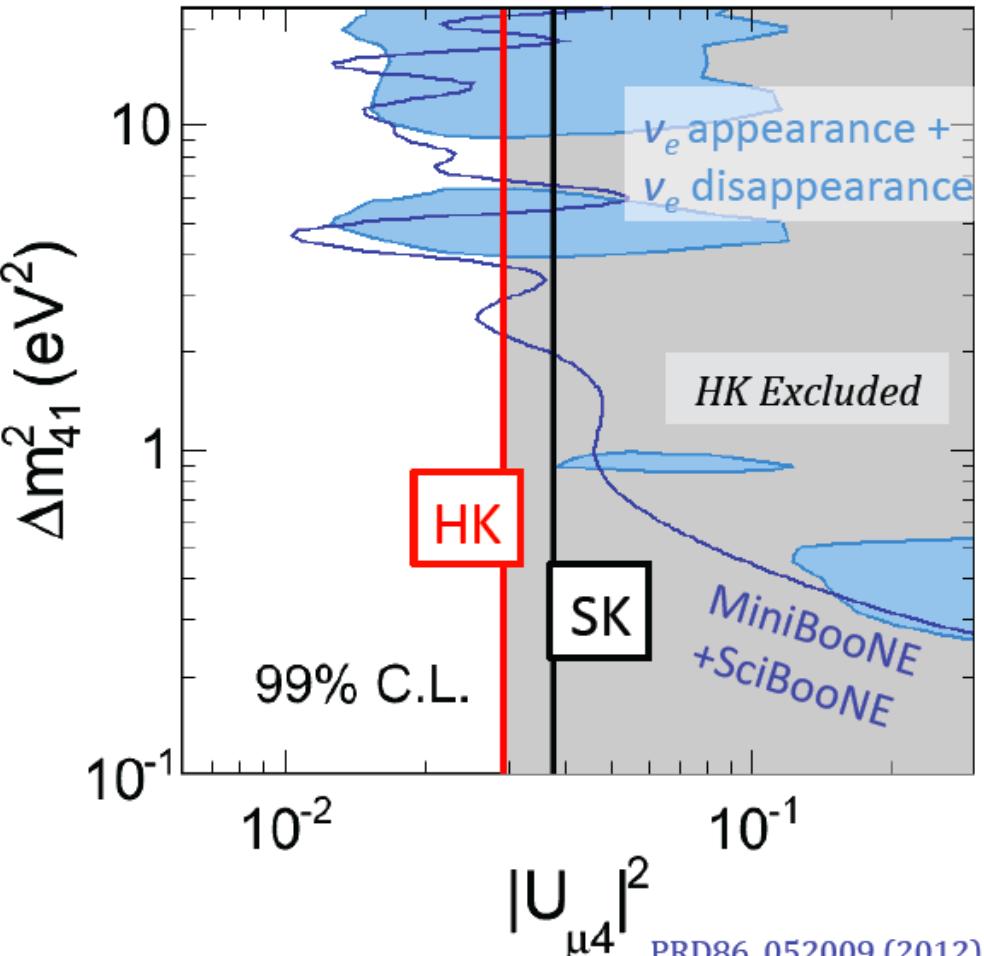
Complementary measurements to accelerator ν
Combined analysis of acc + atm ν will enhance capability

Test of Sterile ν by atmospheric ν

Look for **extra overall muon deficit or shape distortion**



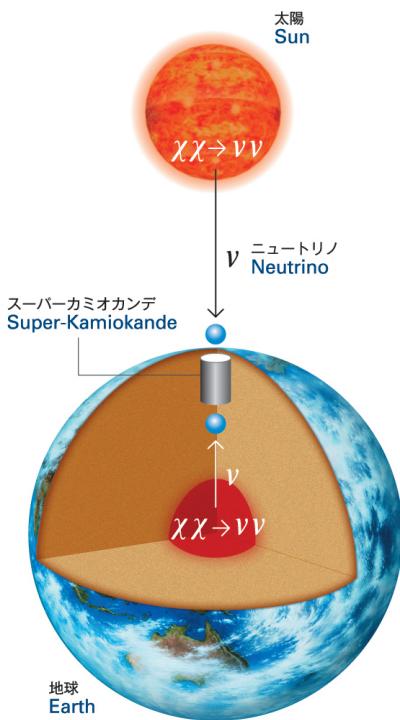
$|U_{\tau 4}|^2 < 0.066$ @99%CL
(0.164 in Super-K)



$|U_{\mu 4}|^2 < 0.029$ @99%CL
(0.038 in Super-K)

PRD86, 052009 (2012)
JHEP1305 (2013) 050

Complementary to other experiments

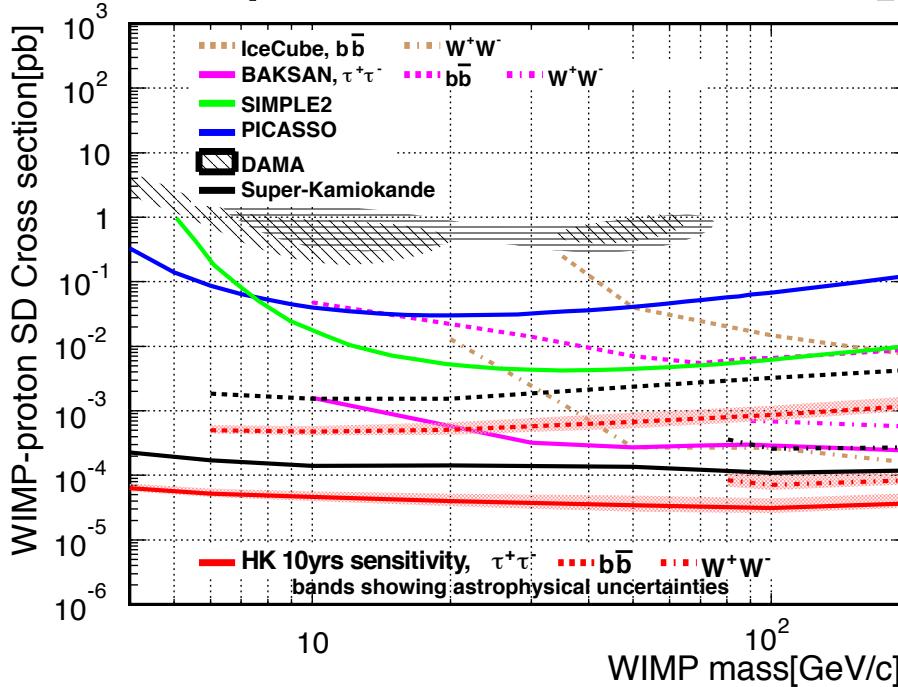


Search for ν's induced by dark matters

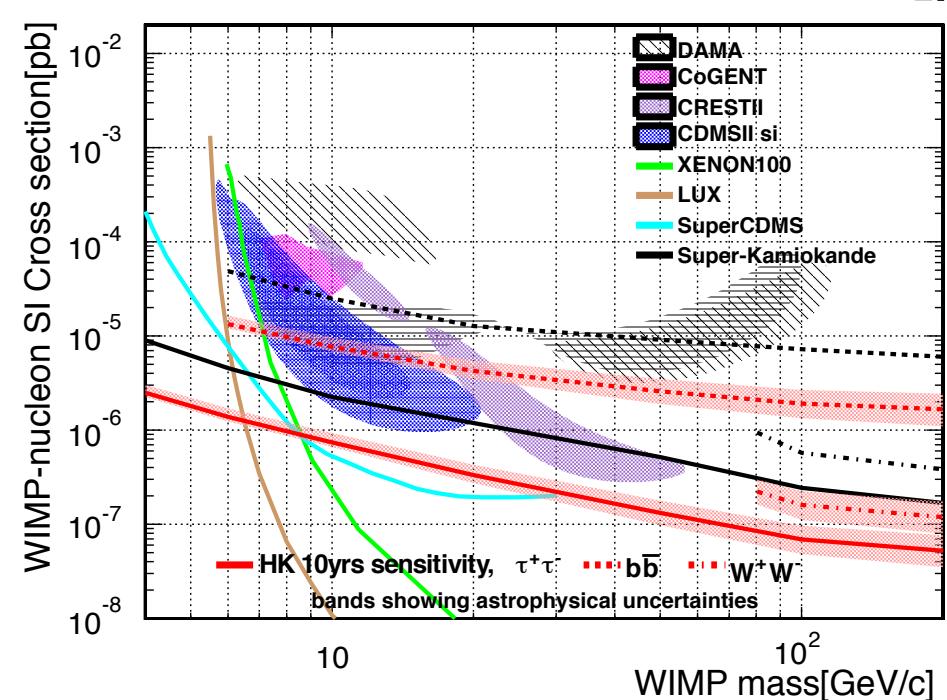
- provide complementary information w/ direct detection experiments
- Sensitive to low mass (GeV/c^2) WIMPs

Expected sensitivity for Solar WIMPs

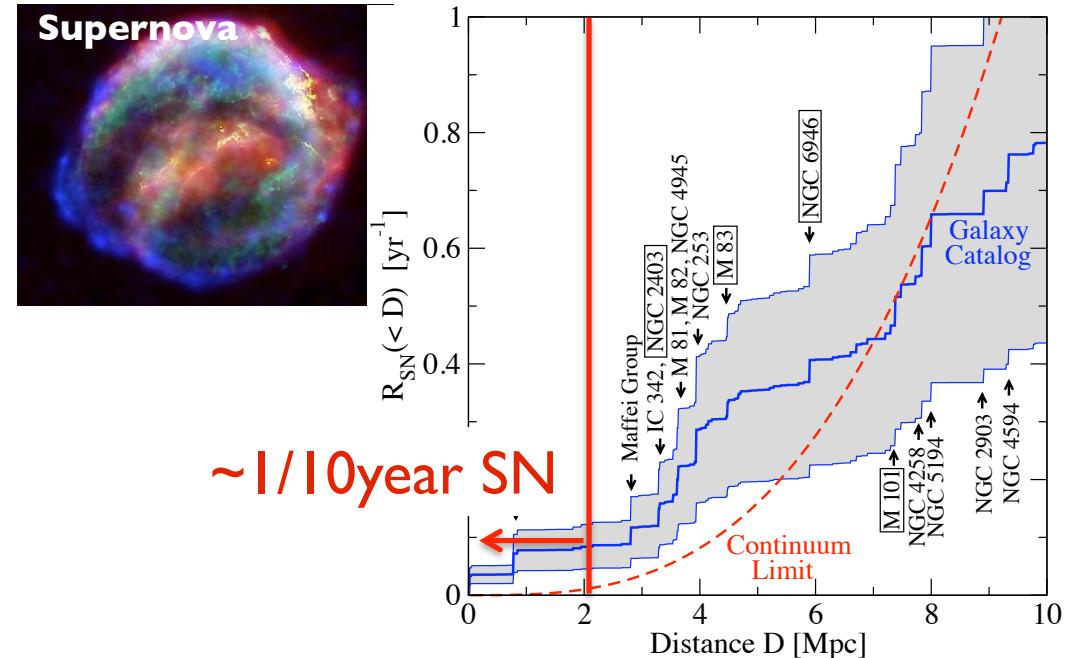
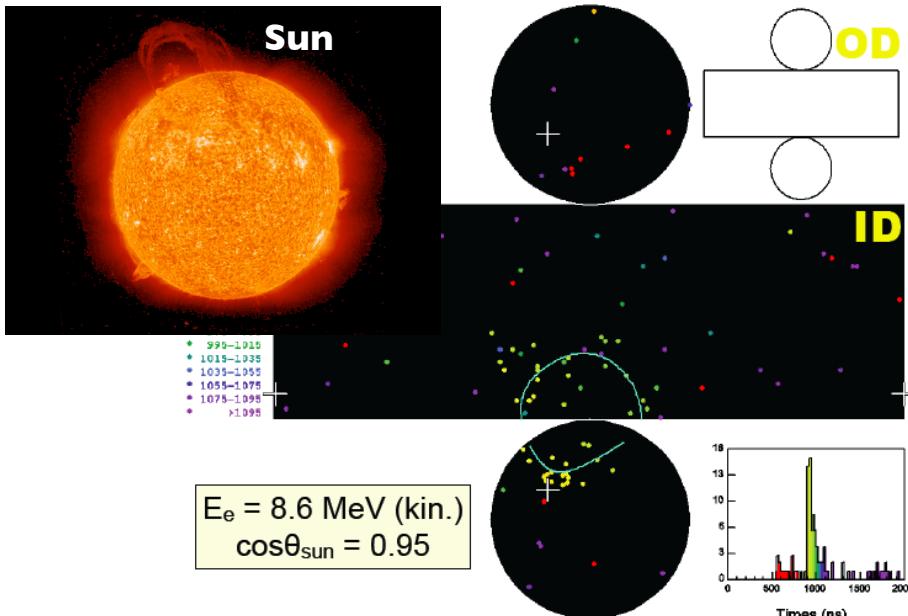
WIMP-proton cross section[pb]



WIMP-nucleon cross section[pb]



Det. performance for astrophysical ν



- Astrophysical neutrinos such as **galactic supernova, supernova in nearby galaxies, relic SN vs**, solar vs
 - Energy threshold $\sim 5 \text{ MeV}$ by established techniques of water purification, triggering system, analysis algorithms
 - tagging capability of $\mu \rightarrow e \nu \nu$ and nuclear de-excitation γ in $p \rightarrow \nu K^+$
- energy scale stability $\sim 1\%$
- stable operation (small $< 1\%$ deadtime for Supernova observation)

Summary

- **Neutrinos play critical roles in various physics field**
 - Particle physics, Cosmology, Astroparticle physics
 - Challenges are being made for mass hierarchy, δ_{CP} , θ_{23} octant
- **Hyper-K would have wide physics topics, many discovery potentials**
 - Proton decay discovery
 - CPV (76% of δ space at 3σ), δ precision of $<20^\circ$
 - SN bursts, relic SN v, WIMP annihilation v ...
- **Boost promoting the project**
 - International proto-collaboration has been formed
 - Cooperation with KEK-IPNS/ICRR to develop the project
 - Design Report to be prepared in 2015