

The H.E.S.S. Galactic Plane Survey

Ryan C. G. Chaves¹,

F. Brun, S. Carrigan, C. Deil, A. Donath, H. Gast, V. Marandon, R. Terrier

on behalf of the H.E.S.S. Collaboration



¹ Marie Curie Fellow

CNRS / IN2P3 / LUPM / Université Montpellier
Montpellier, France







The 5-tel hybrid Cherenkov telescope array

4 12-m IACTs w/ recoated mirrors + 1 28-m IACT

Multiple triggering & pointing schemes available

CT5: 2048 PMTs 614 m^2 3.2° FoV $E_{\min} \sim 20 \text{ GeV}$ $f = 38 \text{ m}$

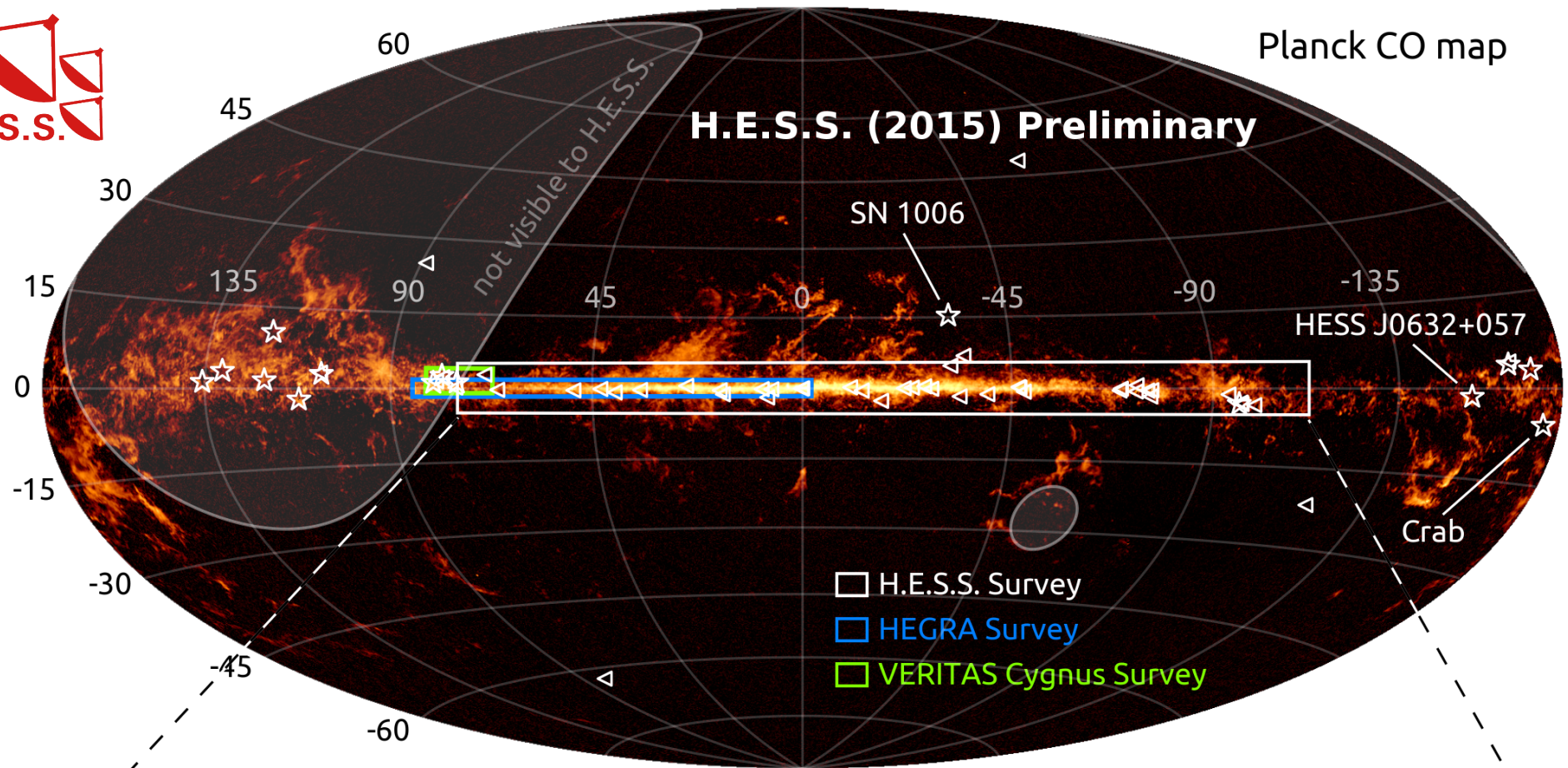
vs. 5.0° FoV (CT1-4)



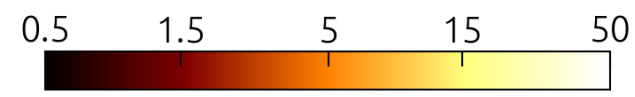


Planck CO map

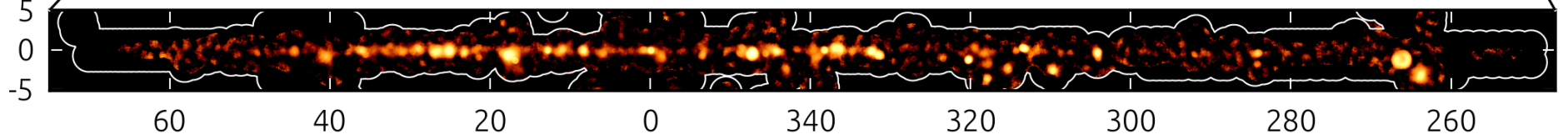
H.E.S.S. (2015) Preliminary



- H.E.S.S. Survey
- HEGRA Survey
- VERITAS Cygnus Survey



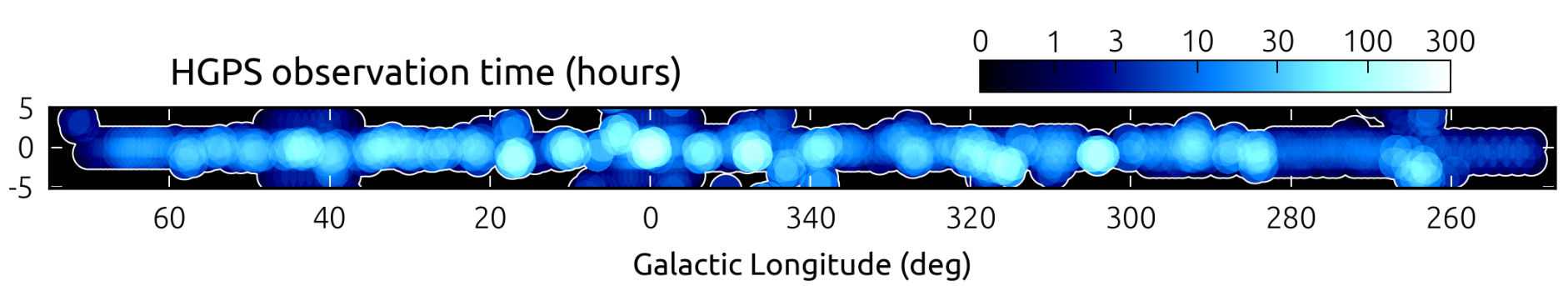
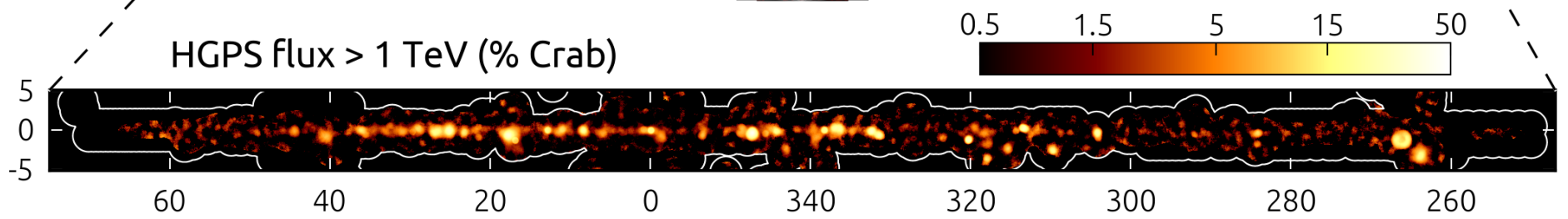
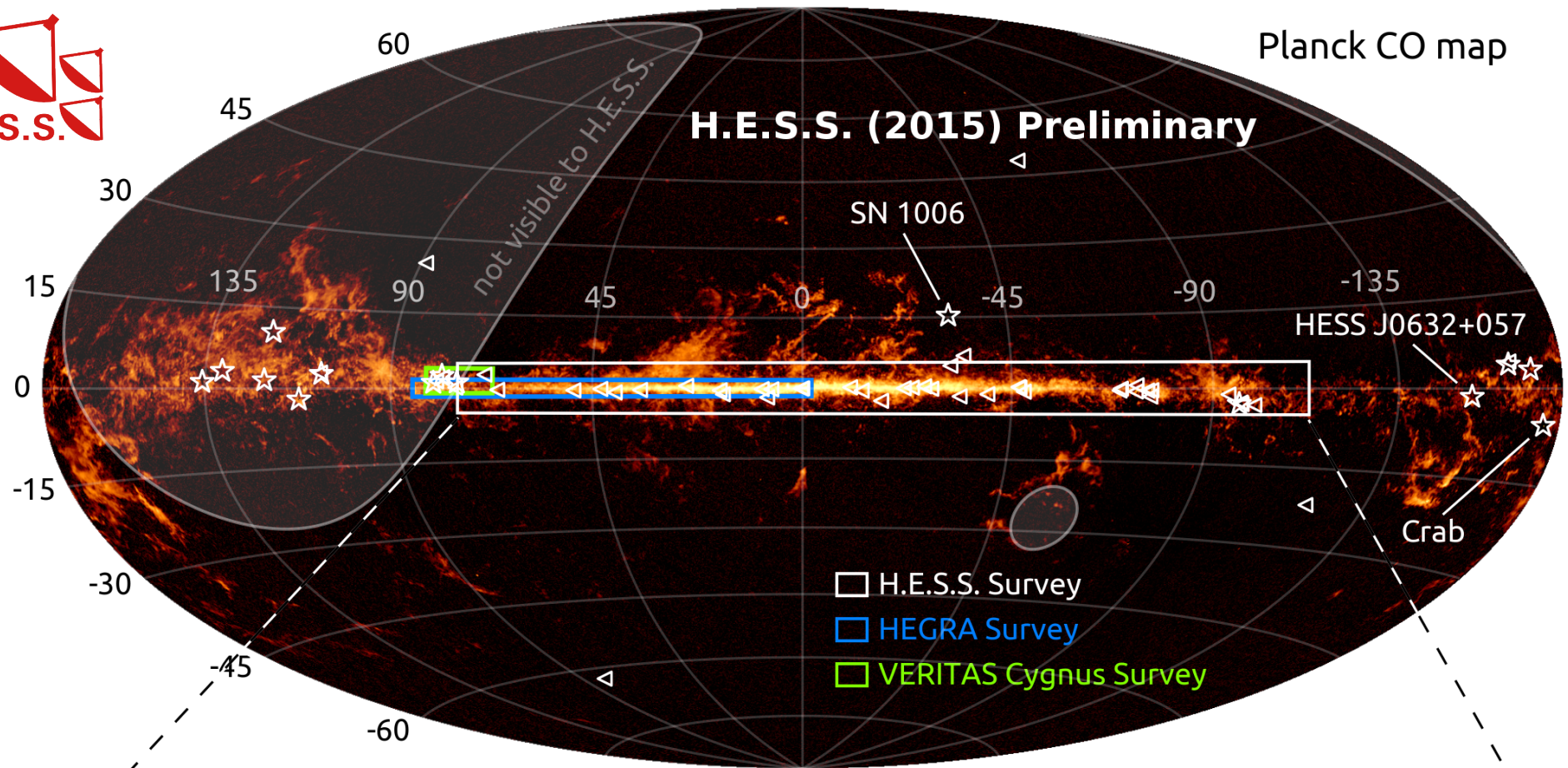
HGPS flux > 1 TeV (% Crab)



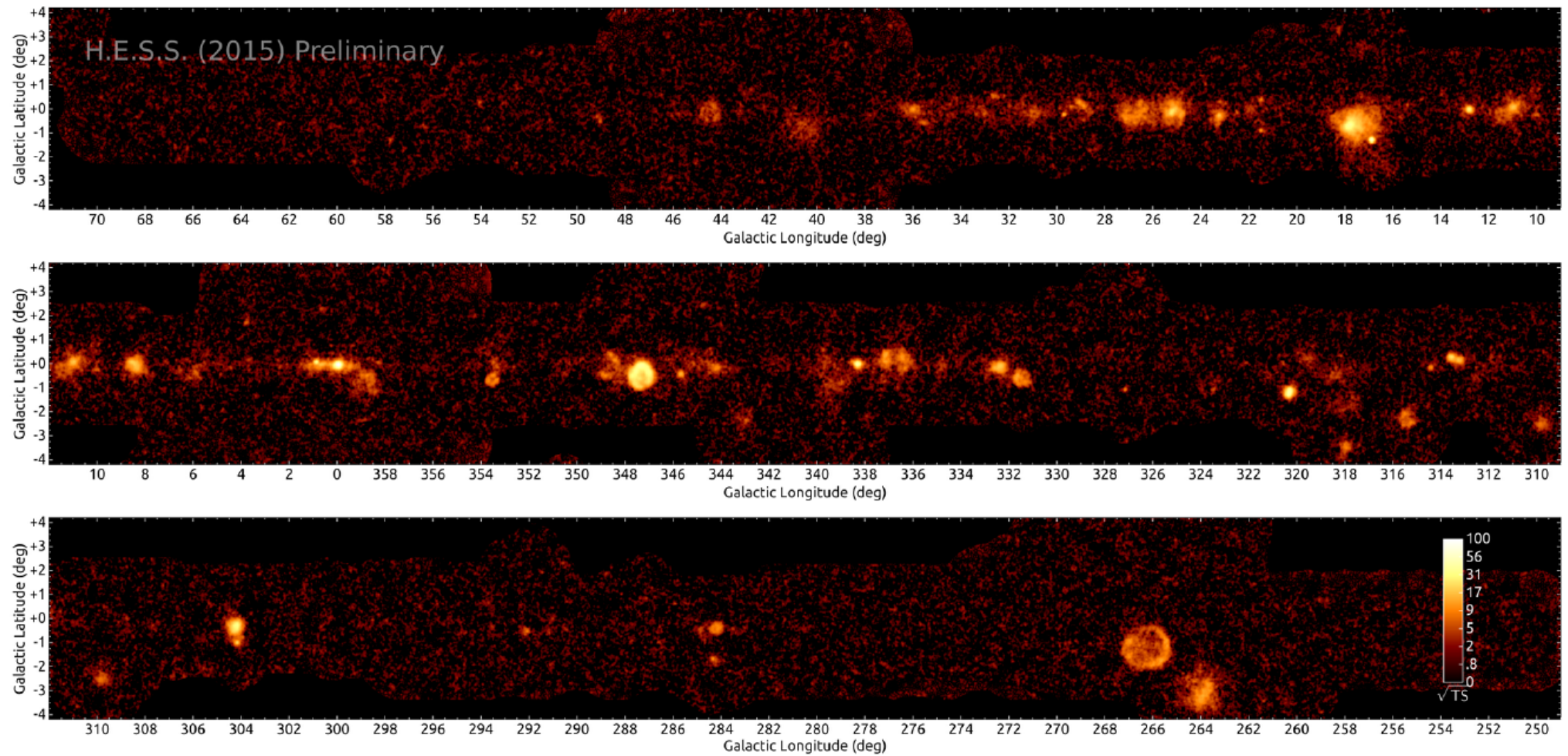


Planck CO map

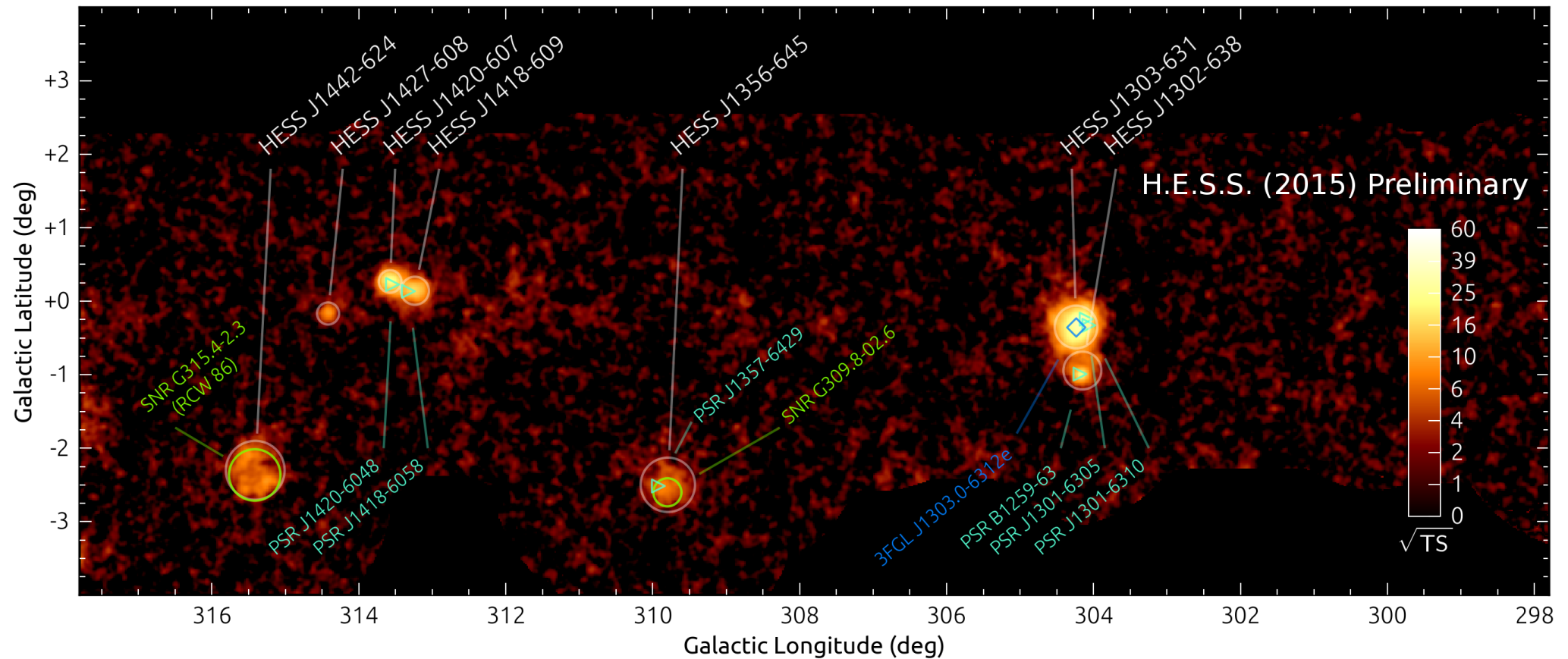
H.E.S.S. (2015) Preliminary



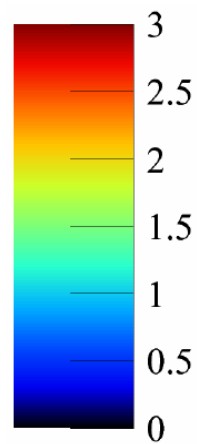
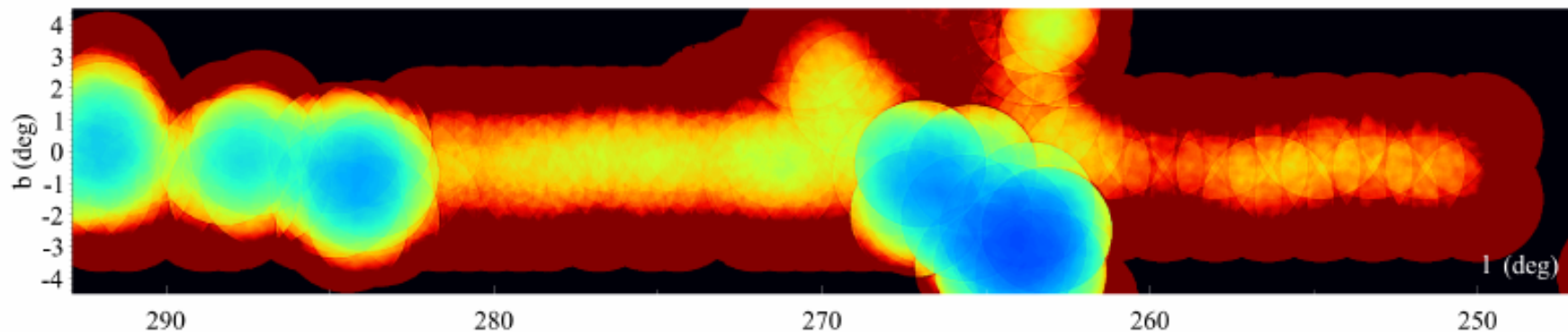
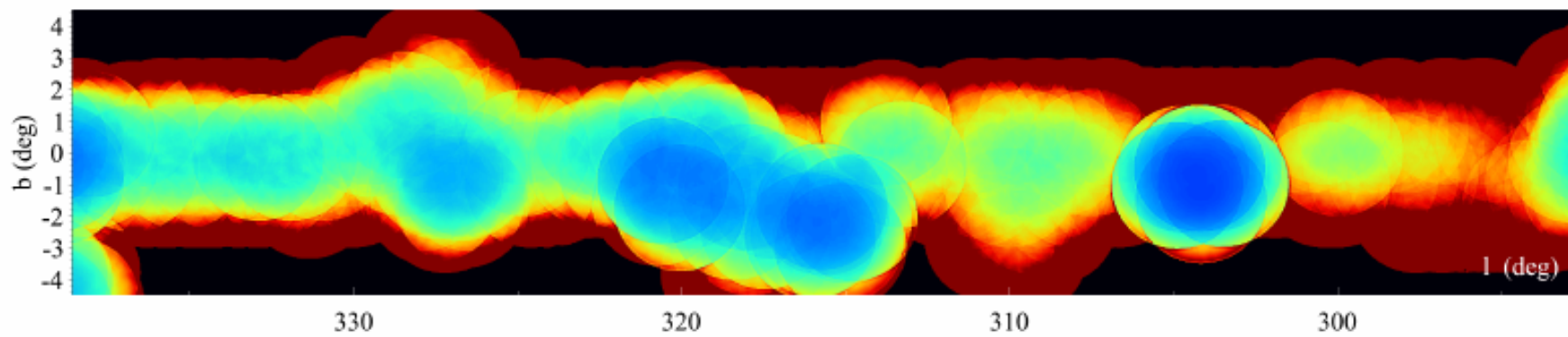
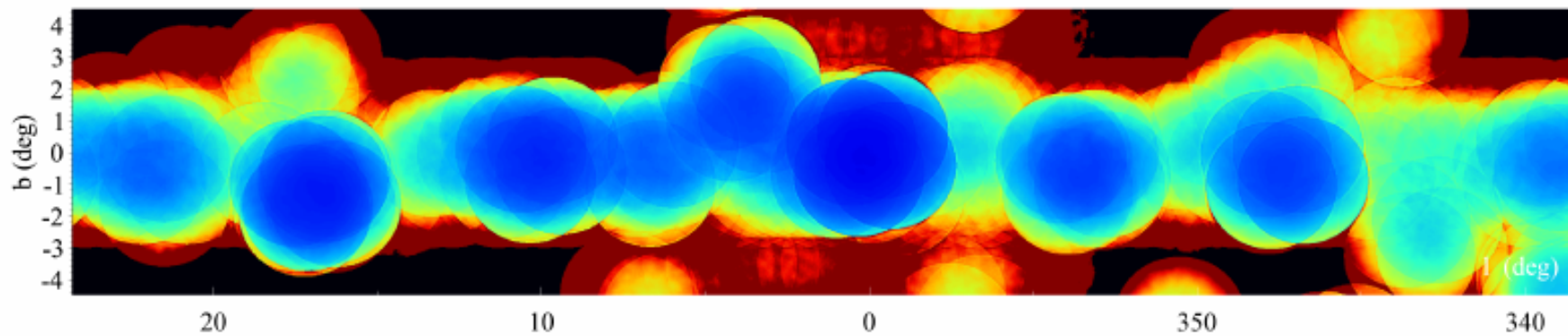
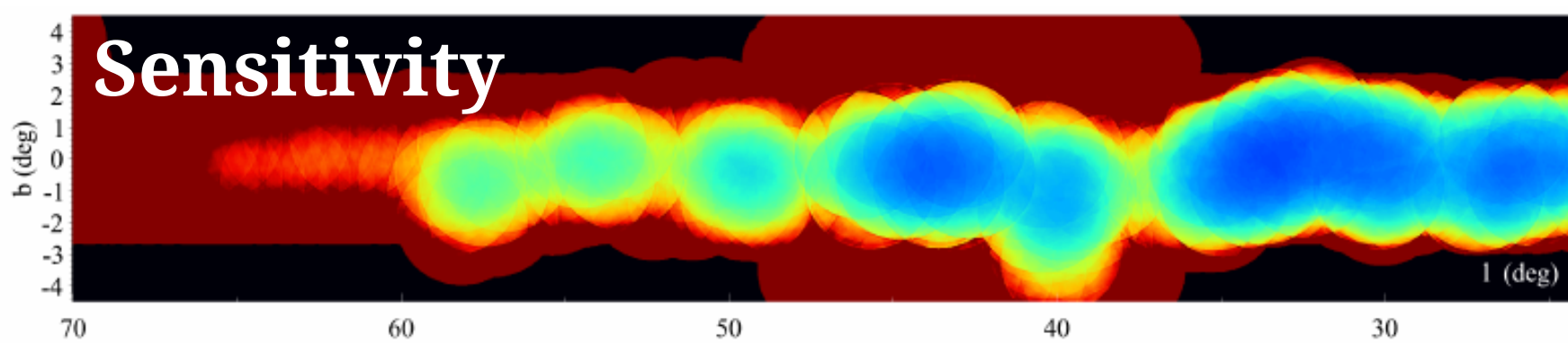
Significance, in full detail



Zoom-in



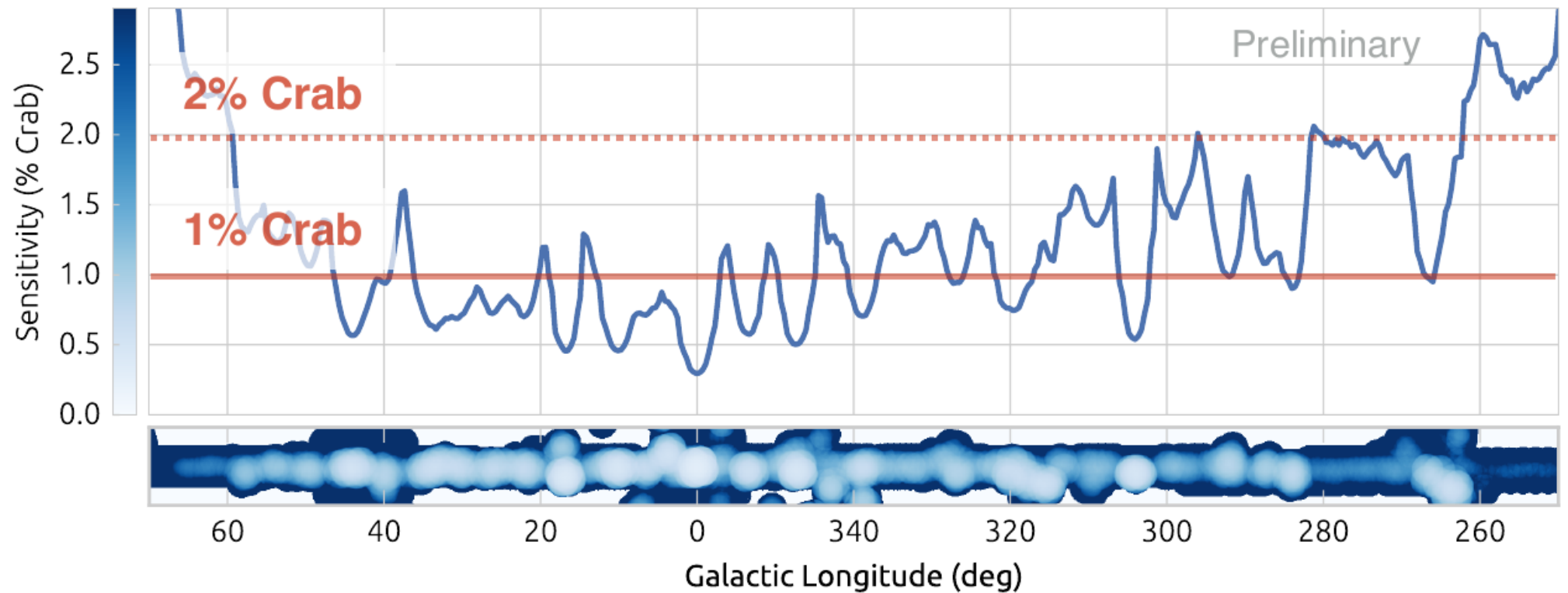
Full detail TeV sky images, catalog, & MWL associations

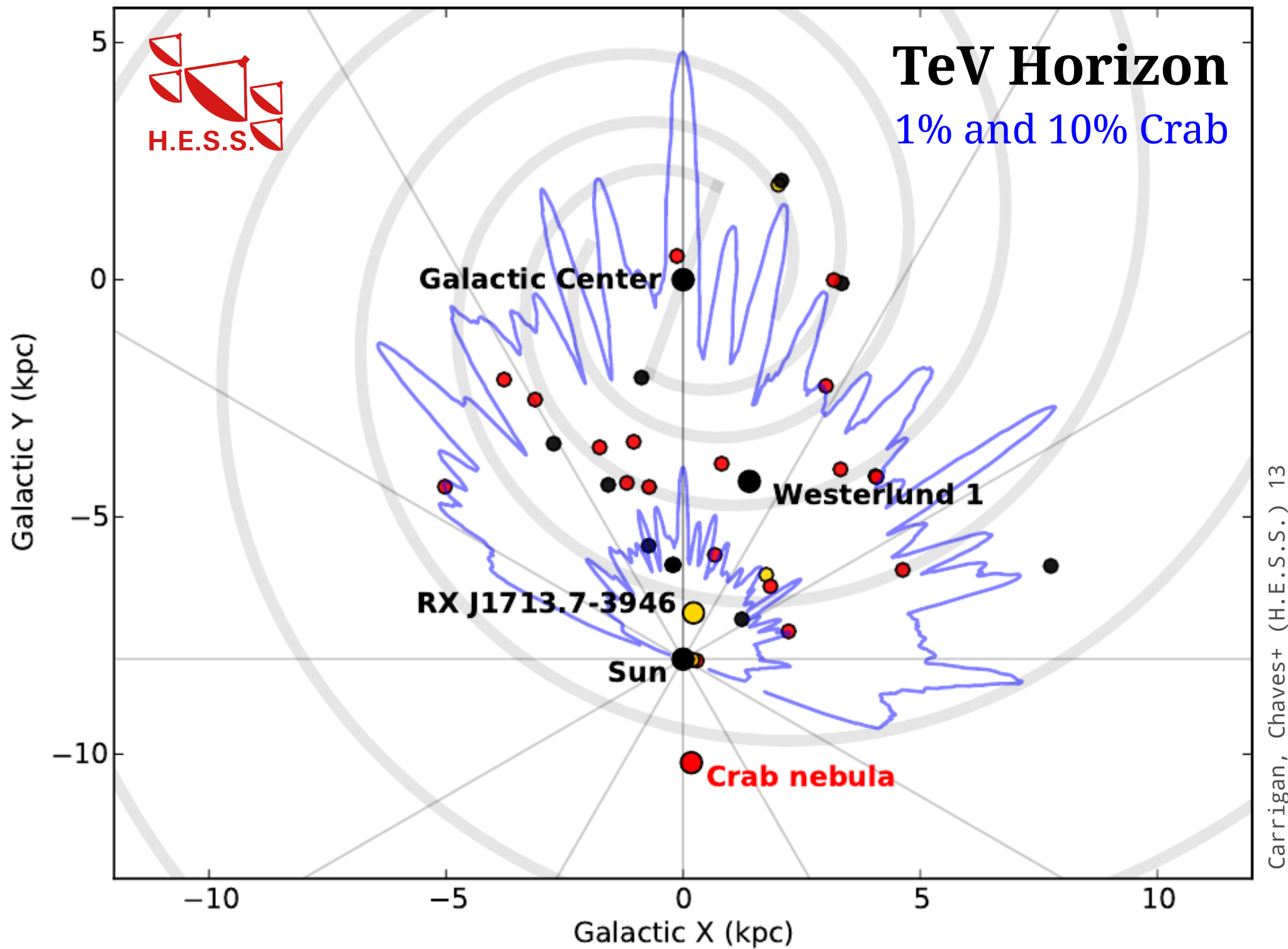


% Crab

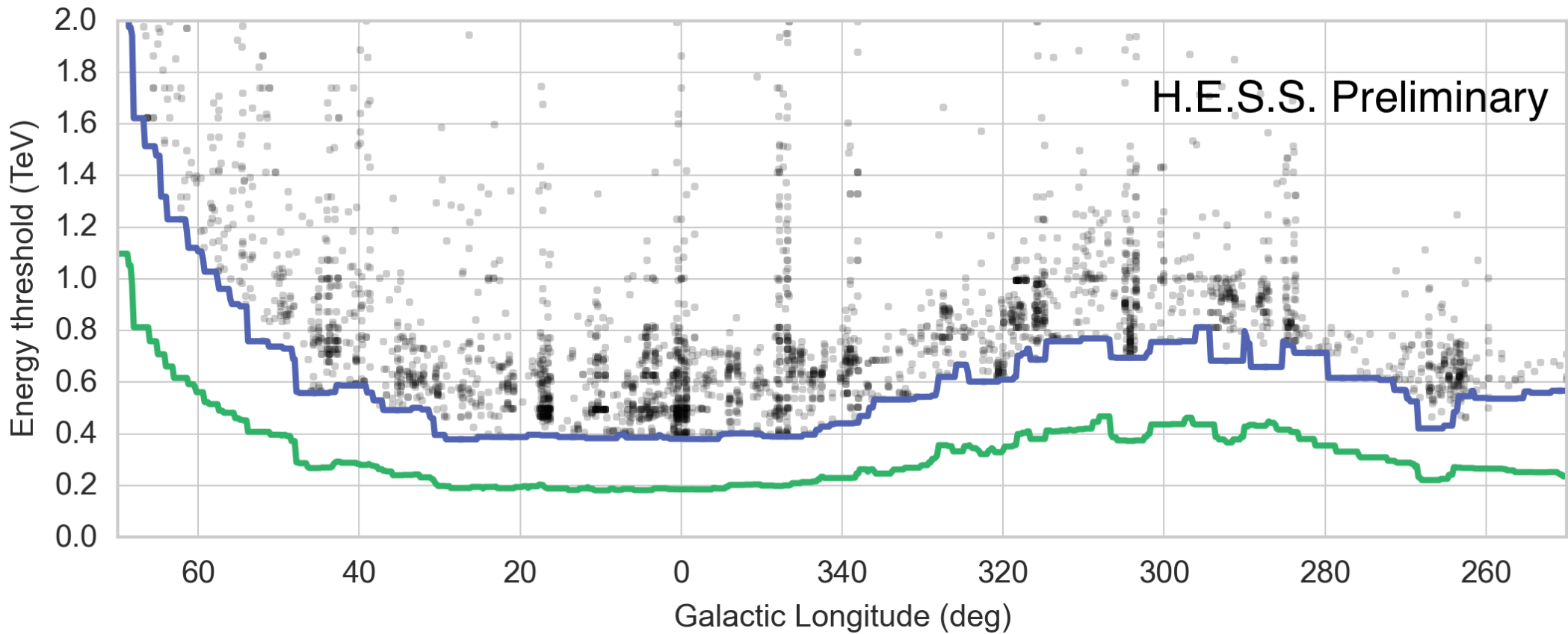
HGPS Sensitivity

along $b = 0^\circ$ for a $5\text{-}\sigma$ detection



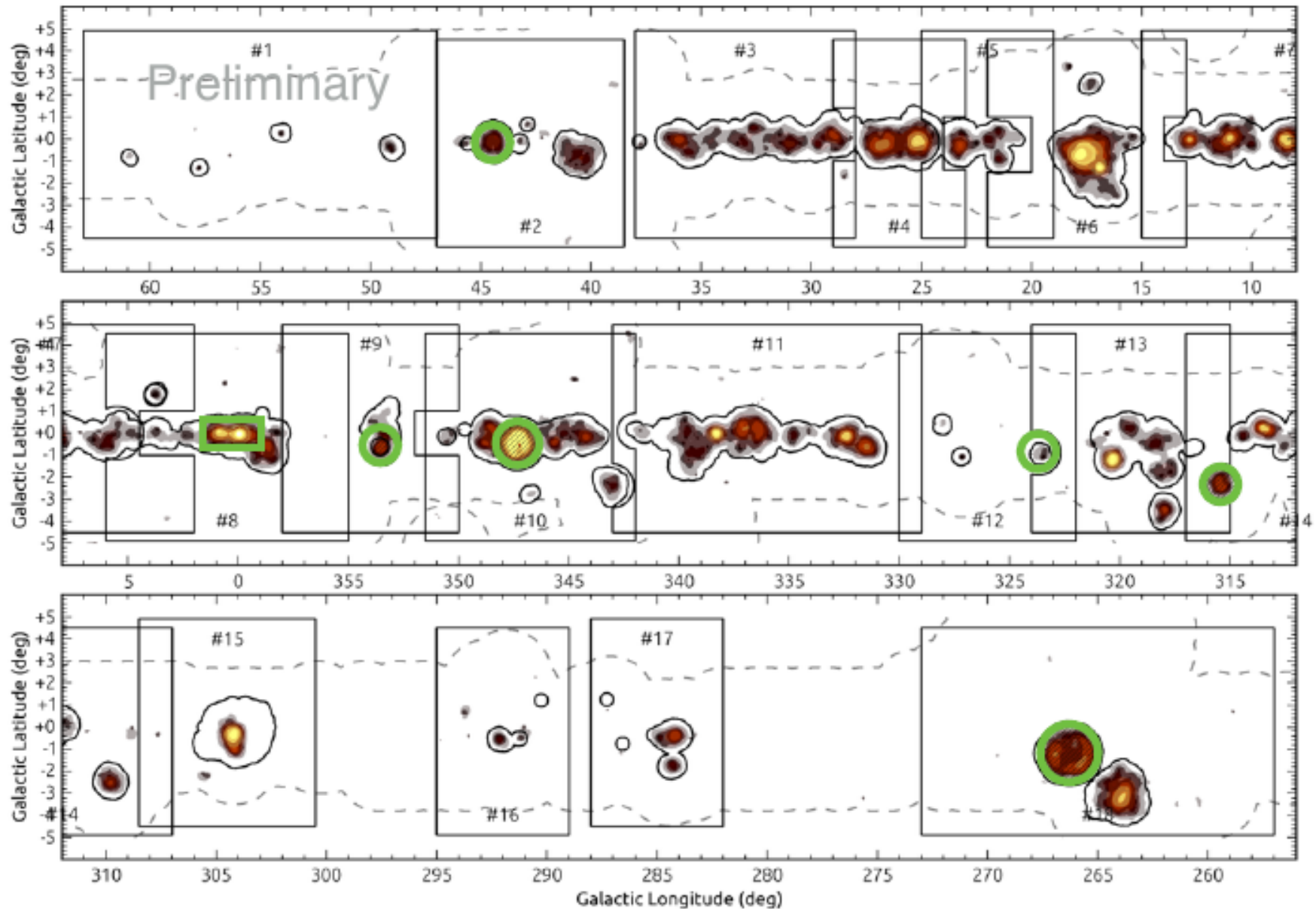


Energy threshold

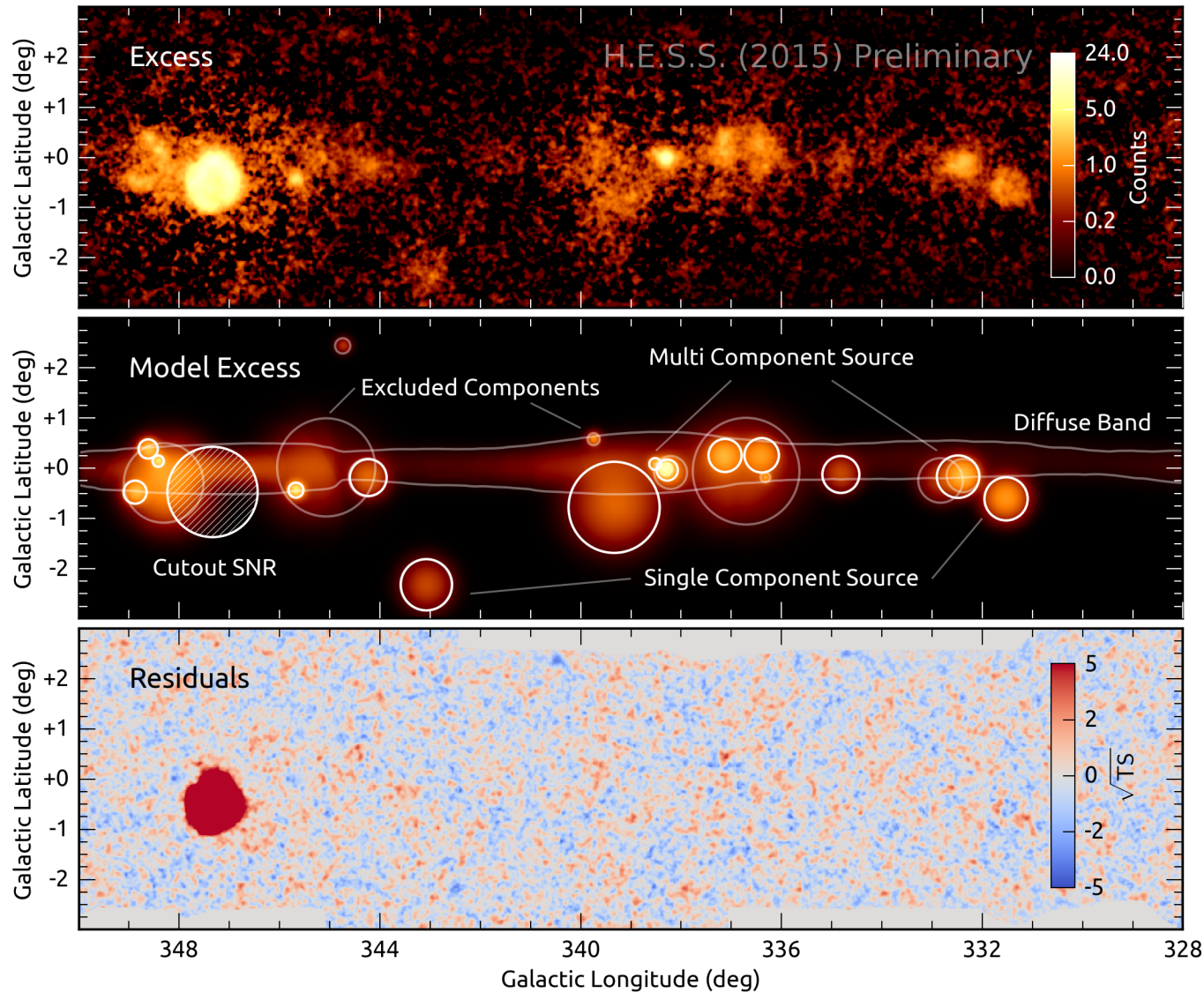


Reflects zenith angle dependency on longitude
~6000 observations $\rightarrow$ ~3000 h dataset

RoI-based analysis pipeline

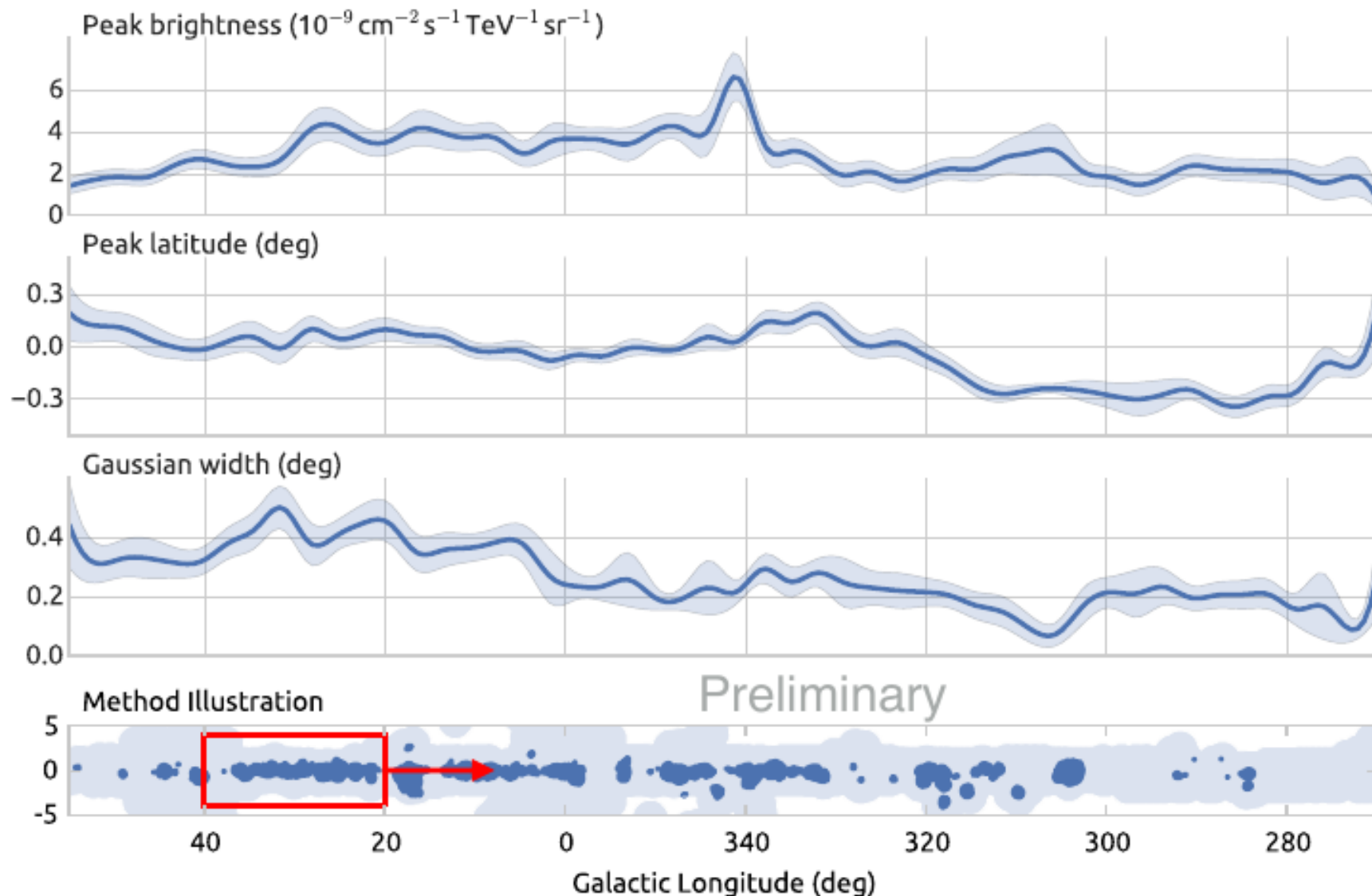


A 2D likelihood approach



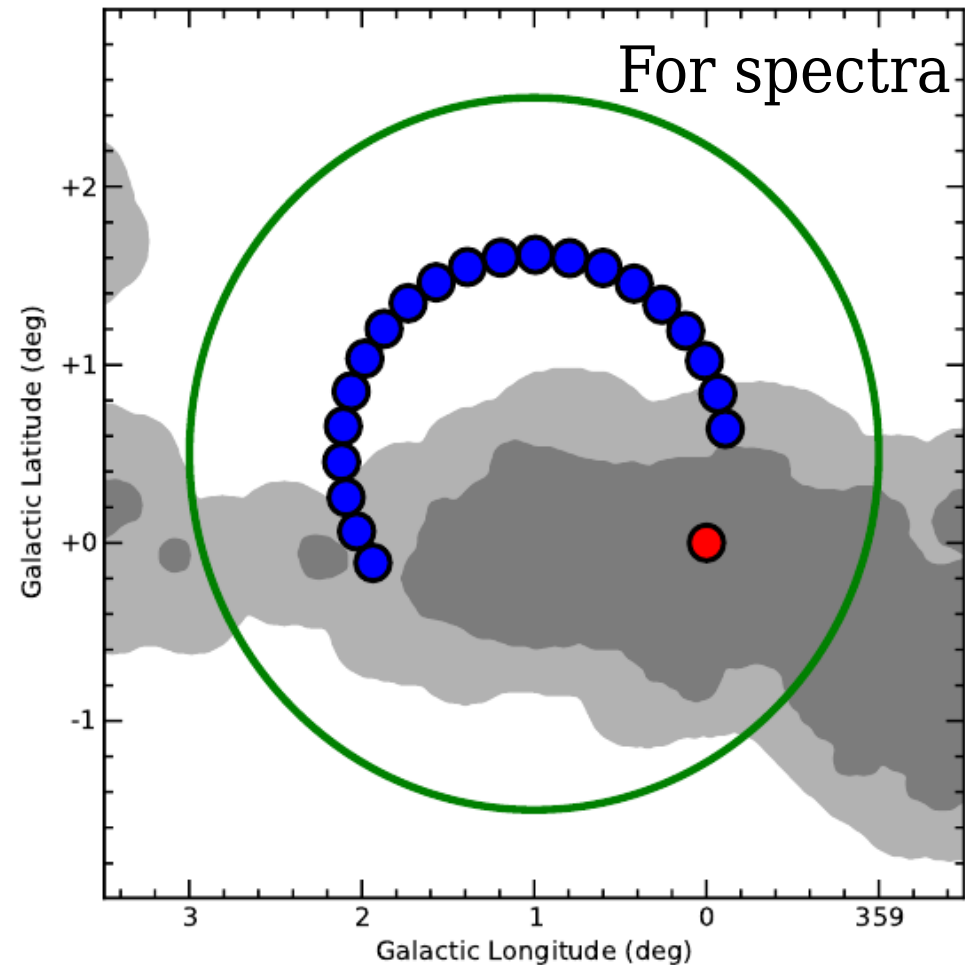
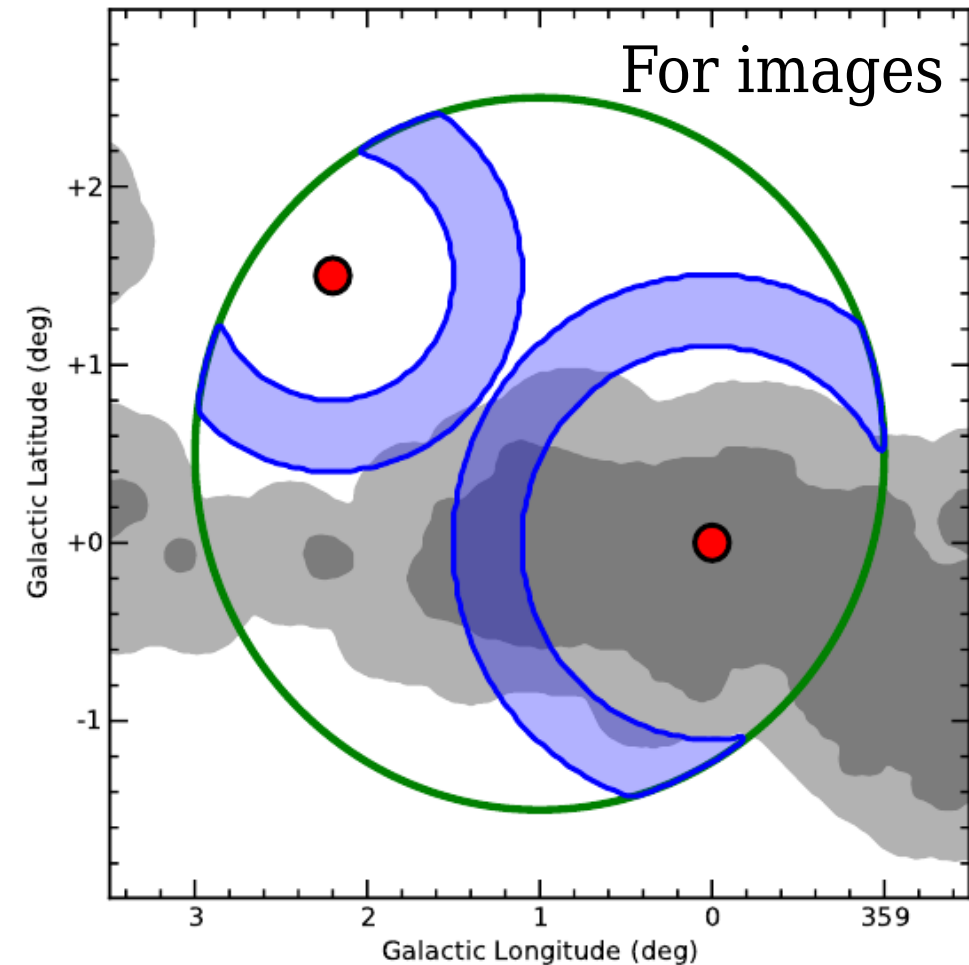
- Semi-automated, MLE-based source detection & morphology fitting
- Model sources as Gaussian plus an underlying “diffuse” component

Handling diffuse emission in the HGPS



Empirical not physical model: 20 deg sliding window w/ Gaussian in b

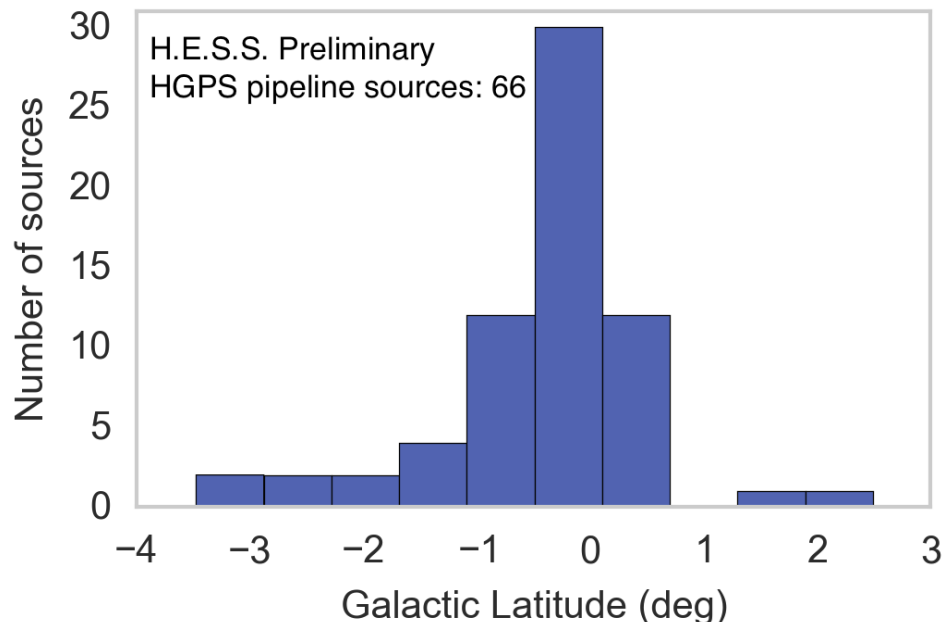
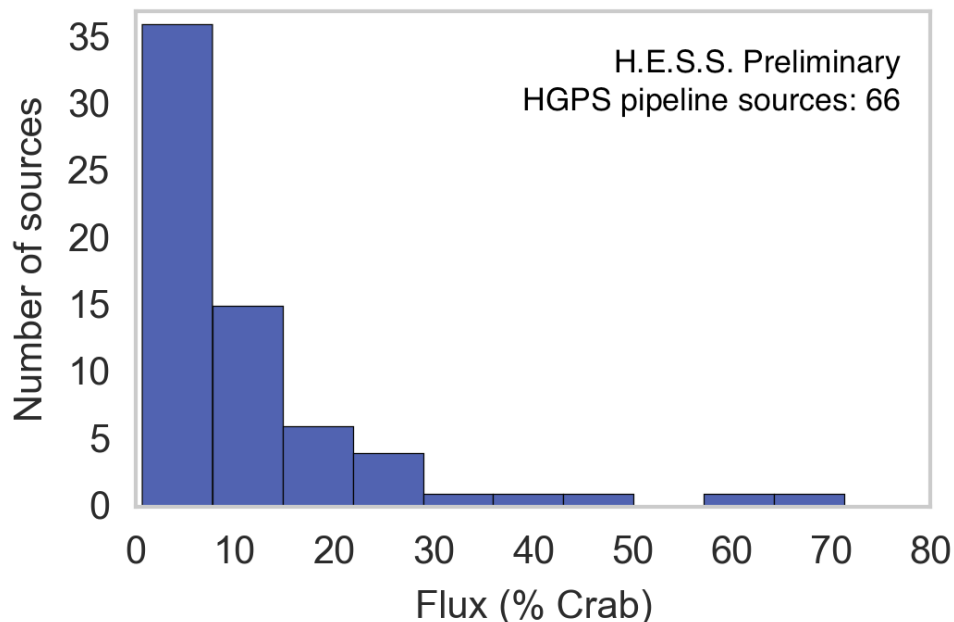
Background estimation



Adaptive background estimation with exclusion region masks



Galactic TeV sources



66 HGPS pipeline sources

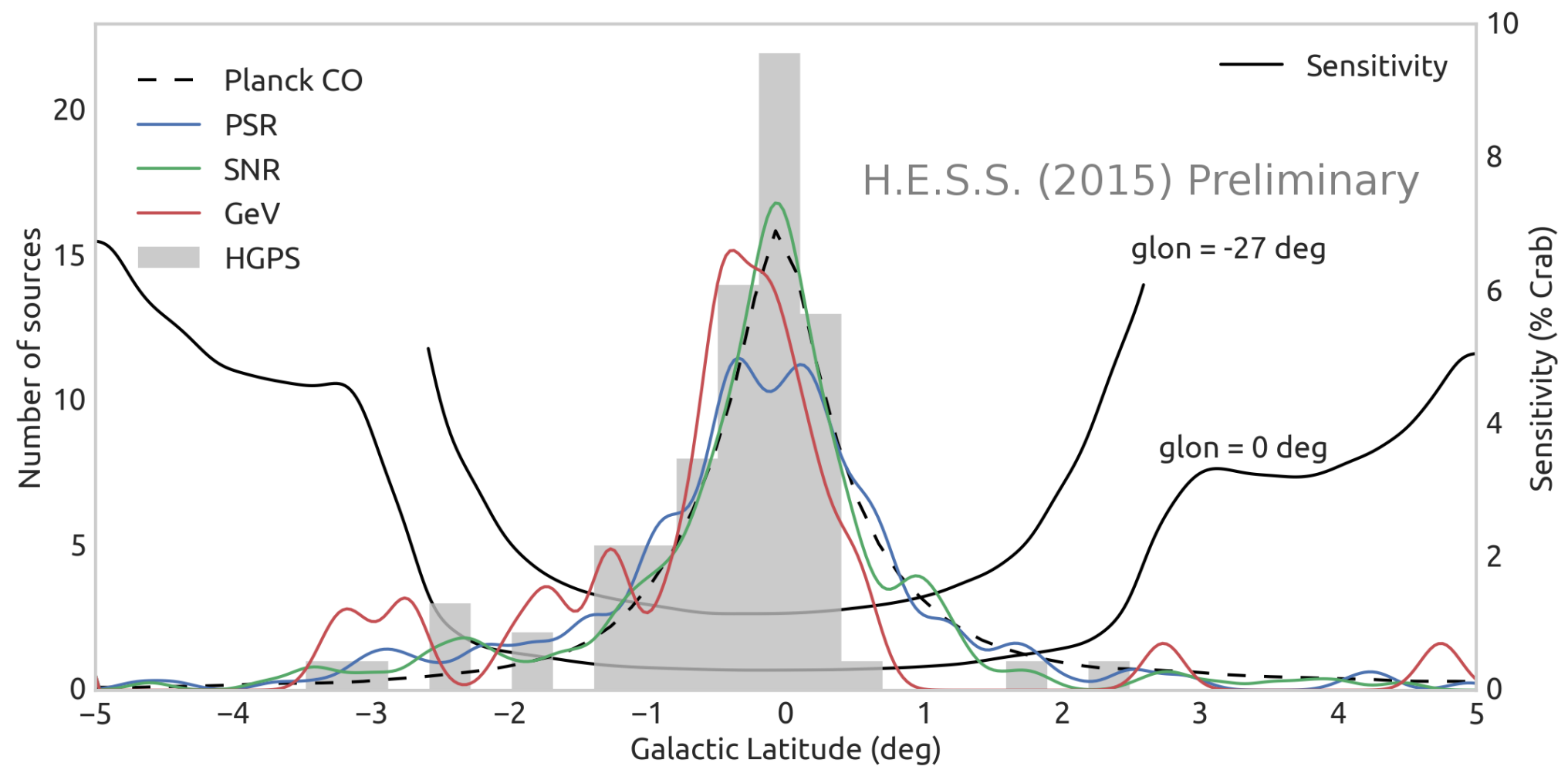
+ 11 sources treated manually (Galactic center region, TeV shells)

+ 3 sources outside HGPS region (Crab Nebula, SN 1006, HESS J0632)

= 80 Galactic sources detected by H.E.S.S.

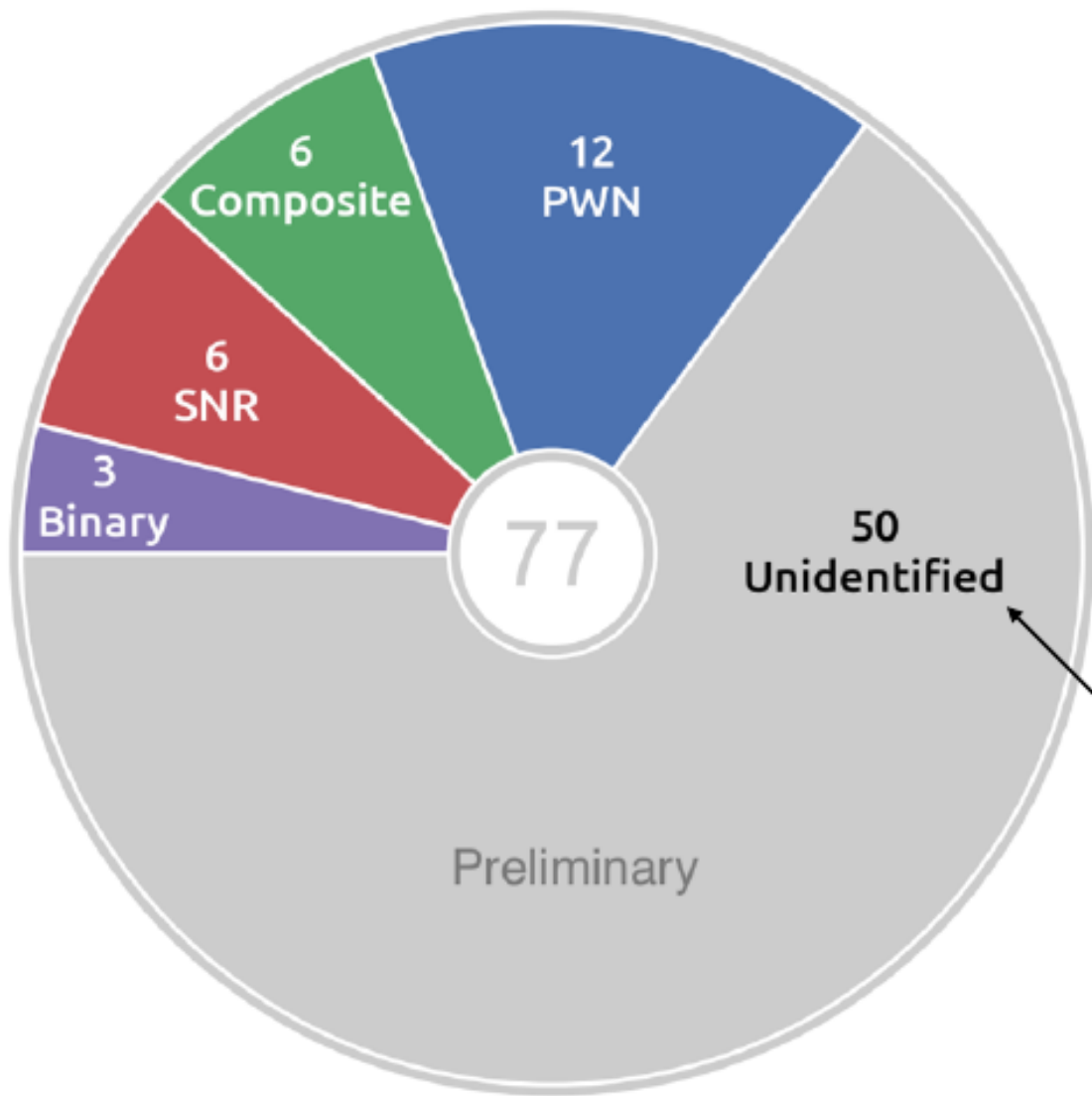


TeV source distribution





HGPS Firm identifications

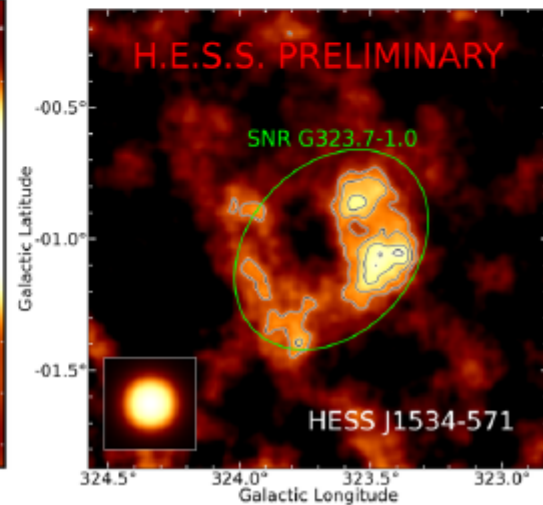
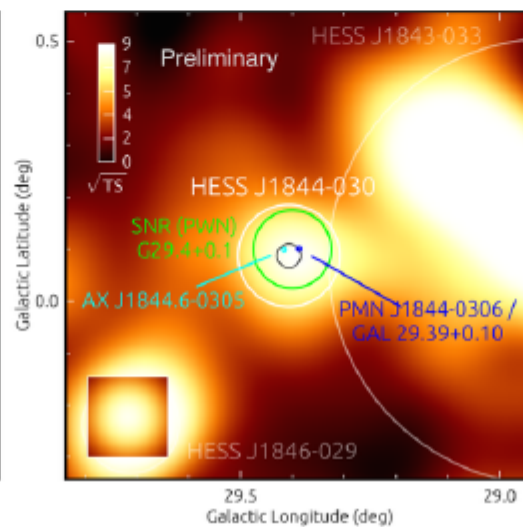
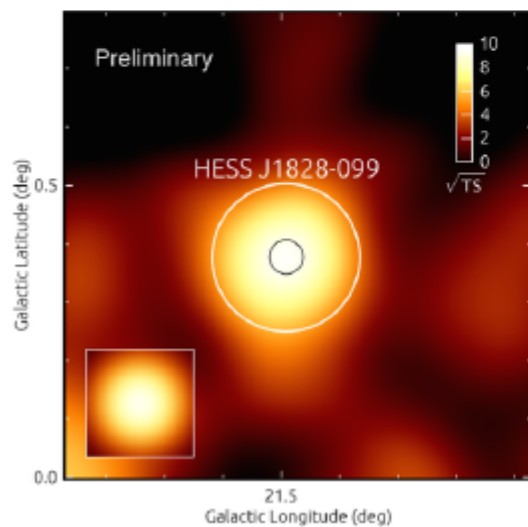
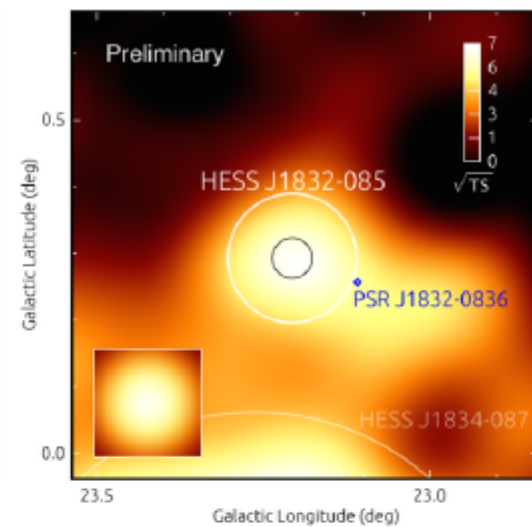
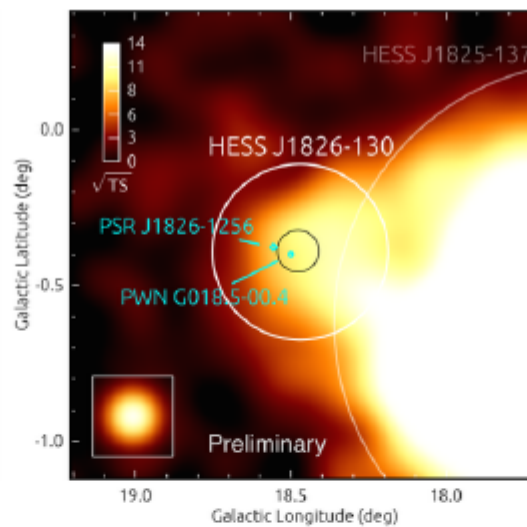
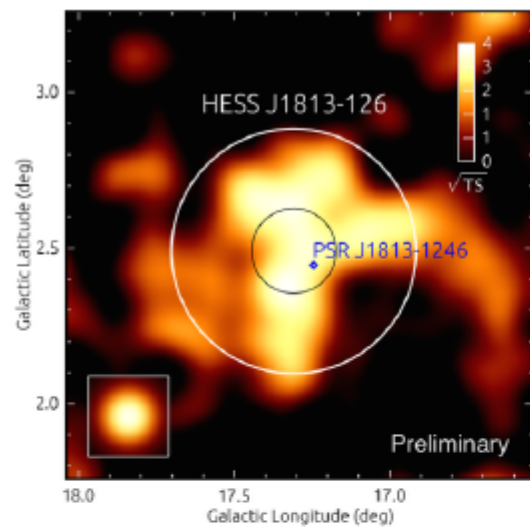


N.B.
associations
(spatial coincidences)
vs. identification

Mostly sources with
multiple associations

MWL counterparts:
Pulsars (ATNF), **PWN** (SNRcat), **SNRs** (SNRcat), **HE sources** (1FHL, 3FGL)

New source discoveries

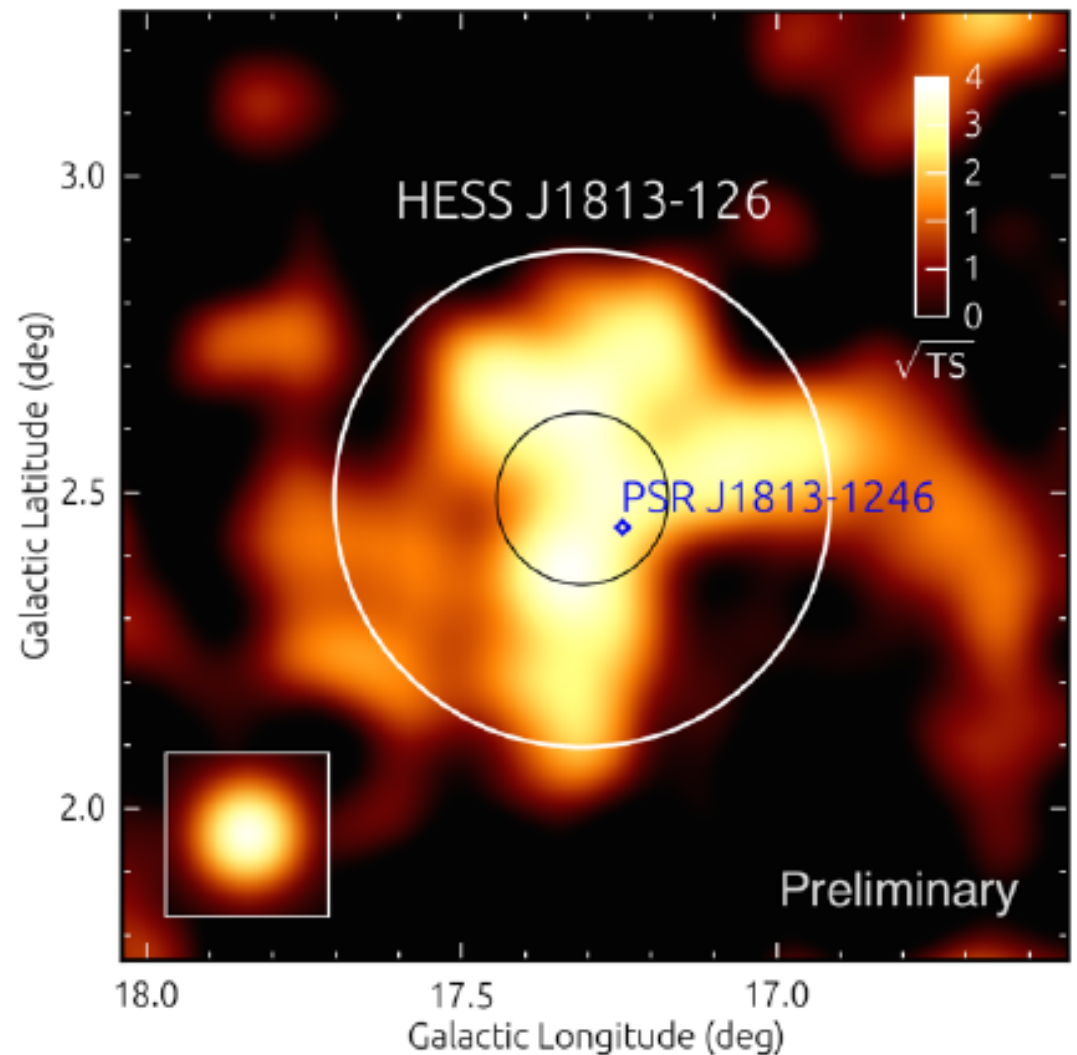




HESS J1813-126

New

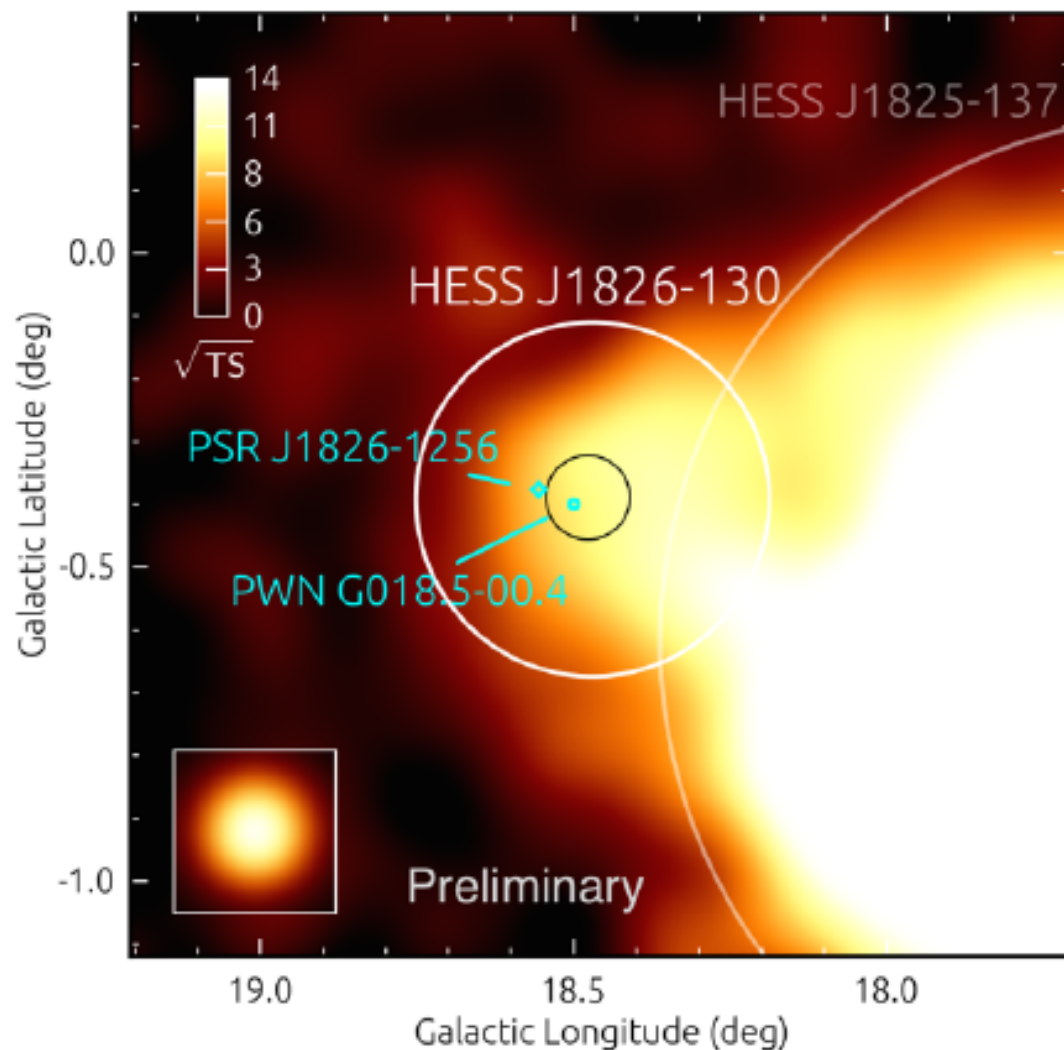
- Position:
 $(l, b) = (17.31, 2.49)$ deg
 $(\alpha, \delta) = (273.34, -12.69)$ deg
- Extension: 0.21 deg
- Flux: 4.2% Crab





HESS J1826-130

- Position:
(l, b) = (18.48, -0.39) deg
(α , δ) = (276.51, -13.02) deg
- Extension: 0.15 deg
- Flux: 3.3% Crab

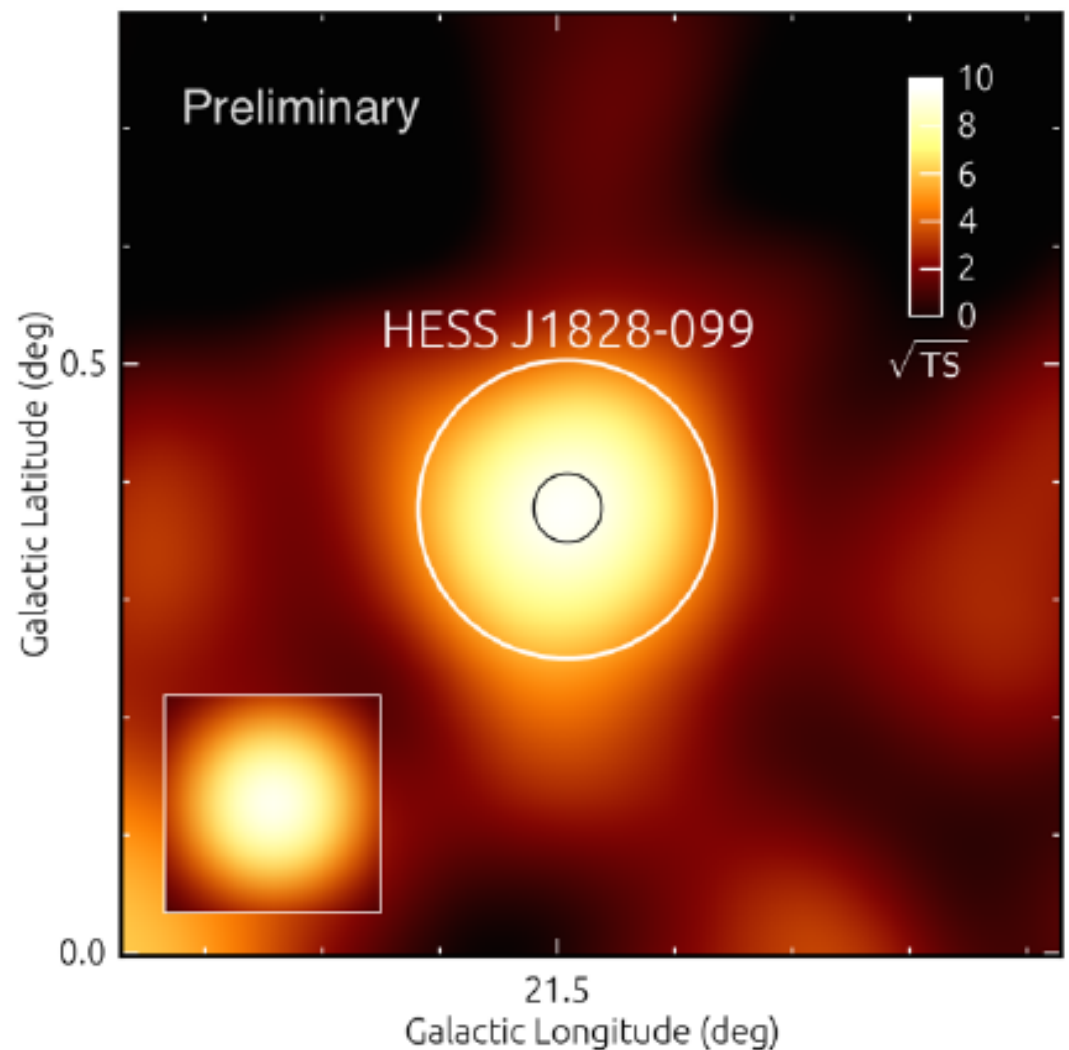




HESS J1828-099

New

- Position:
(l, b) = (21.49, 0.38) deg
(α, δ) = (277.25, -9.99) deg
- Extension upper limit:
< 0.07 deg
- Flux: 1.7% Crab

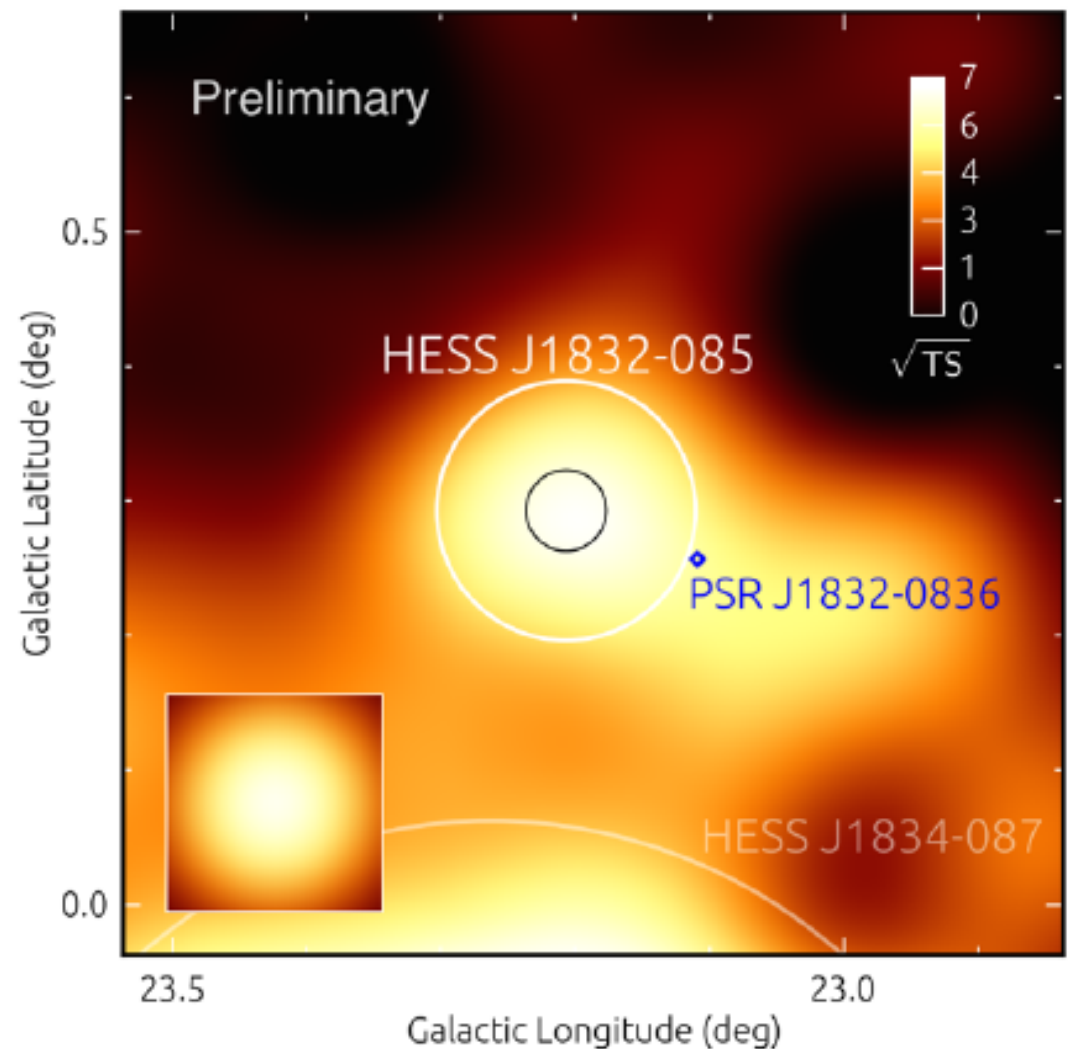




HESS J1832-085



- Position:
(l, b) = (23.21, 0.29) deg
(α , δ) = (278.13, -8.51) deg
- Extension upper limit:
< 0.05 deg
- Flux: 0.8% Crab

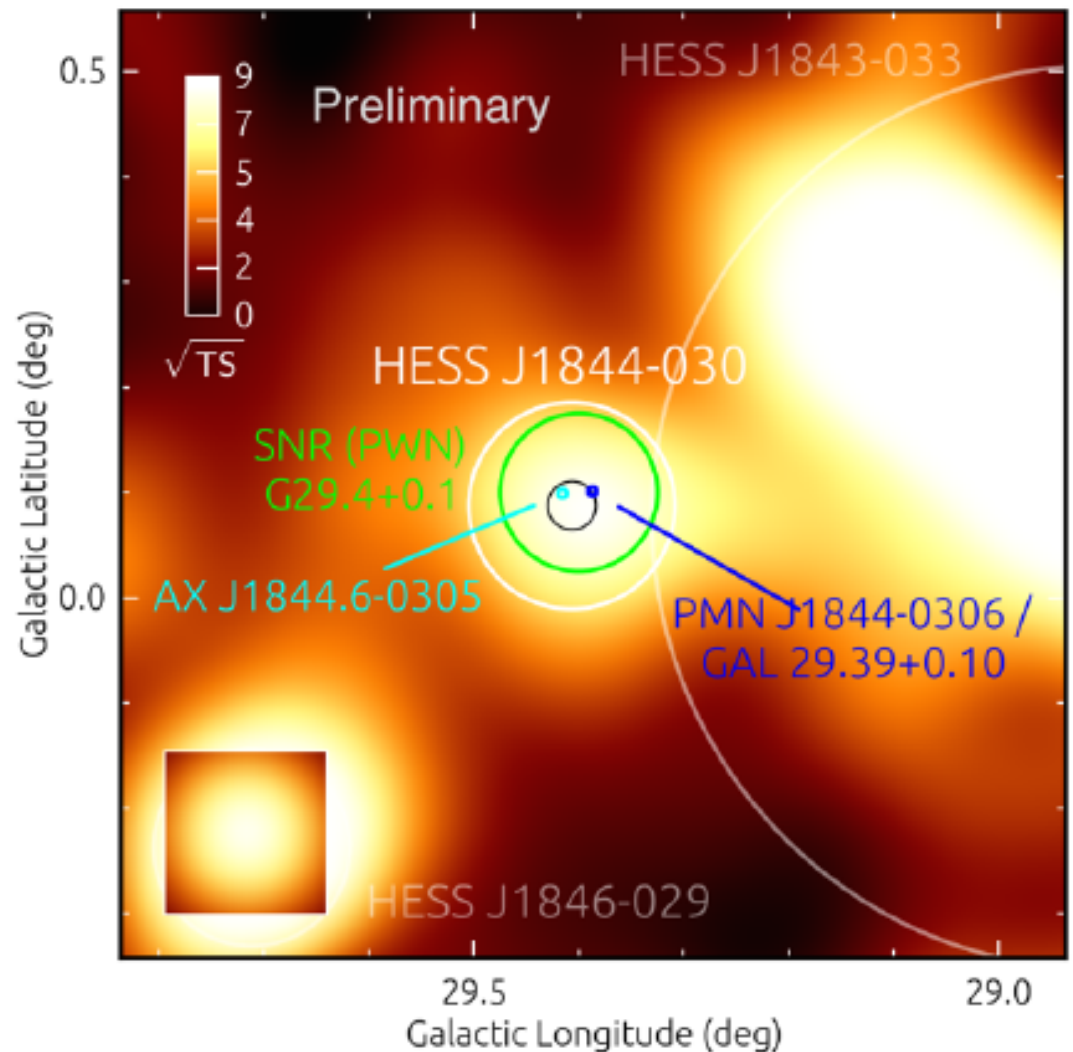




HESS J1844-030

New

- Position:
(l, b) = (29.41, 0.09) deg
(α , δ) = (281.17, -3.10) deg
- Extension upper limit:
< 0.05 deg
- Flux: 1.0% Crab





Conclusion

Demonstrated power of surveying with wide-FoV imaging atmospheric Cherenkov telescopes → **TeV gamma-ray astronomy**

Decade-long H.E.S.S. I Galactic Plane Survey completed

- Legacy paper in preparation

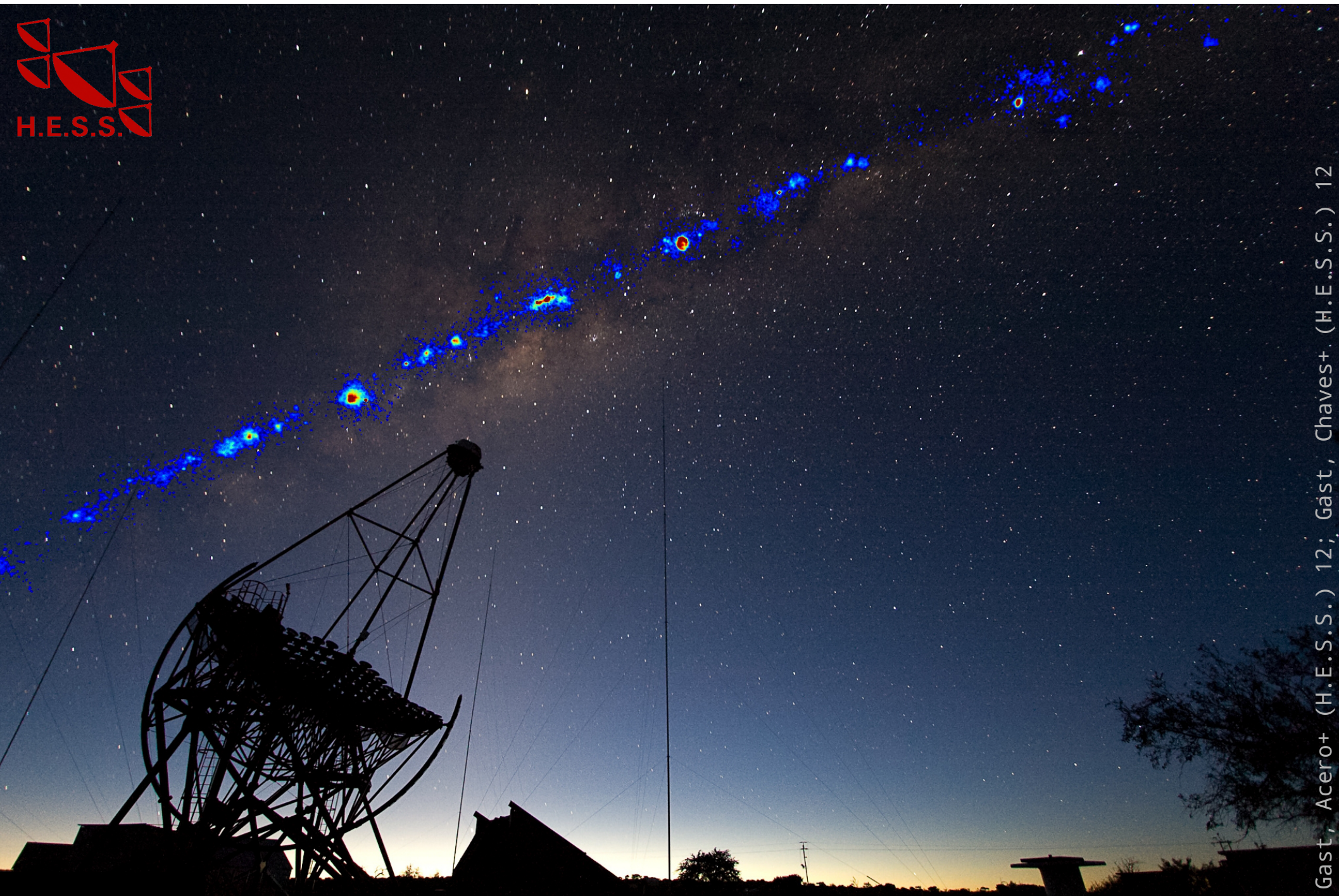
- Final catalog, high-level FITS data release, including spectra

- 17 new source discoveries

- Associations & firm identifications

- Coupled with population studies (SNRs, PWNe) & new TeV shells

The TeV Galactic Plane

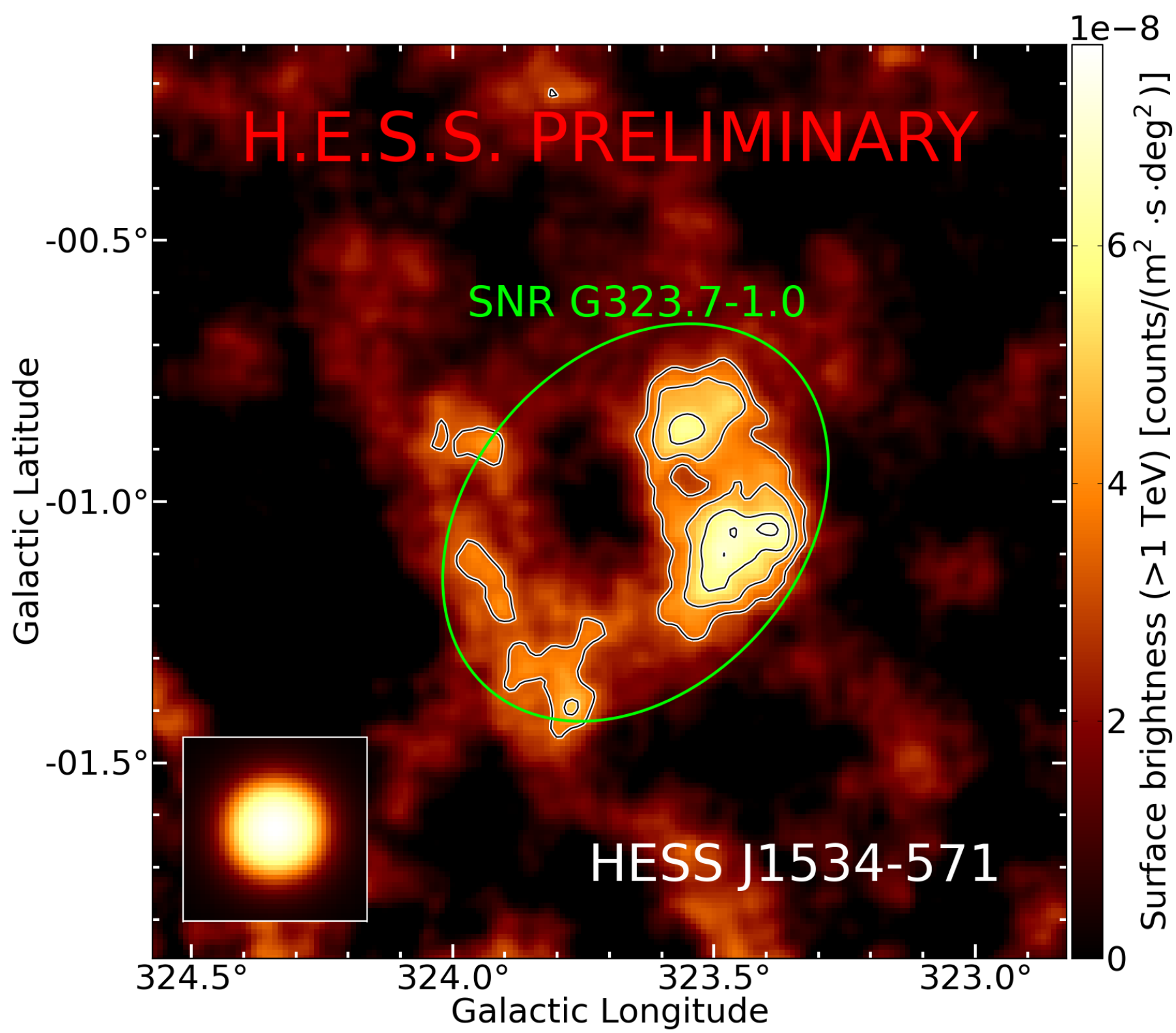


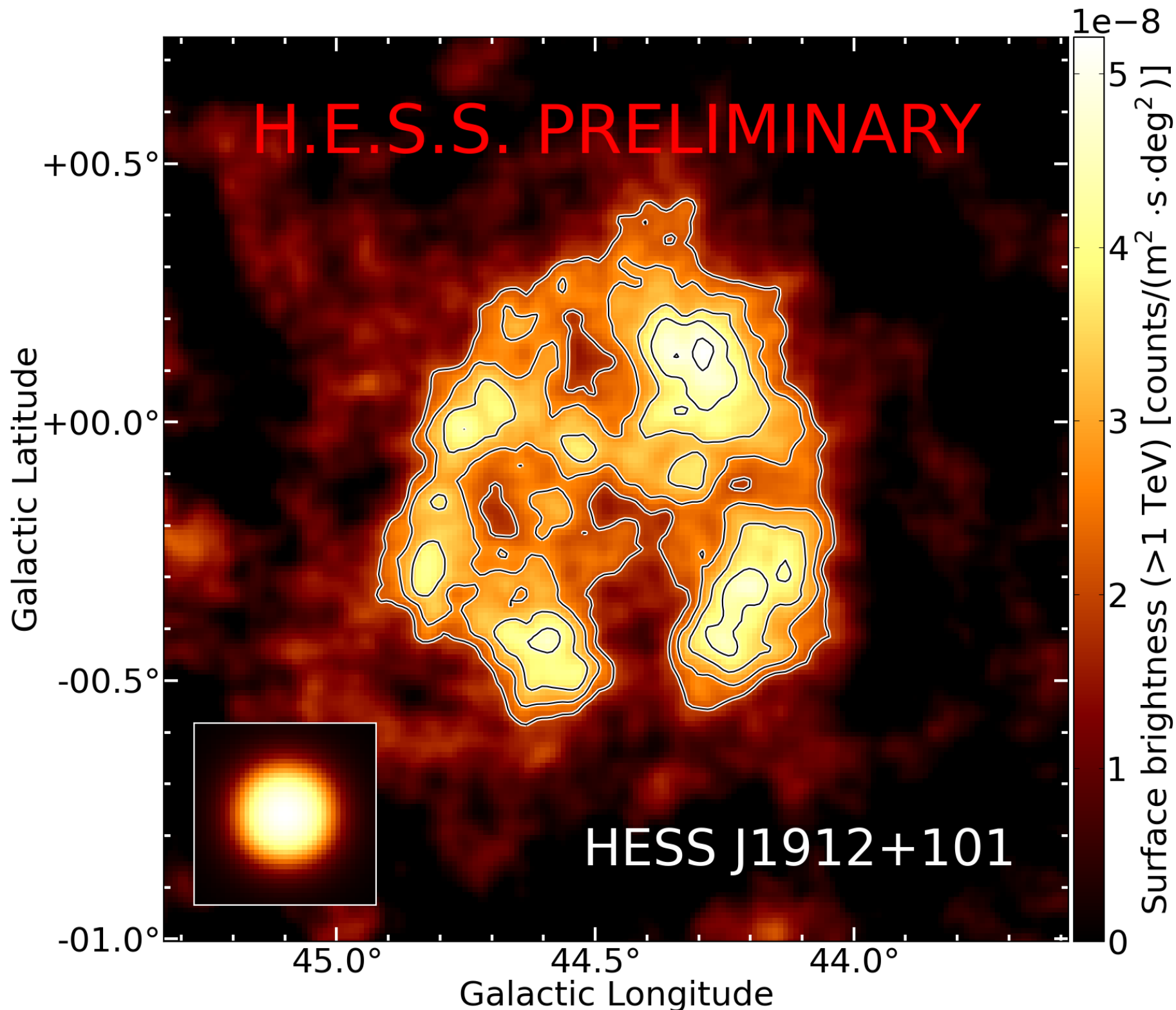


Backup slides

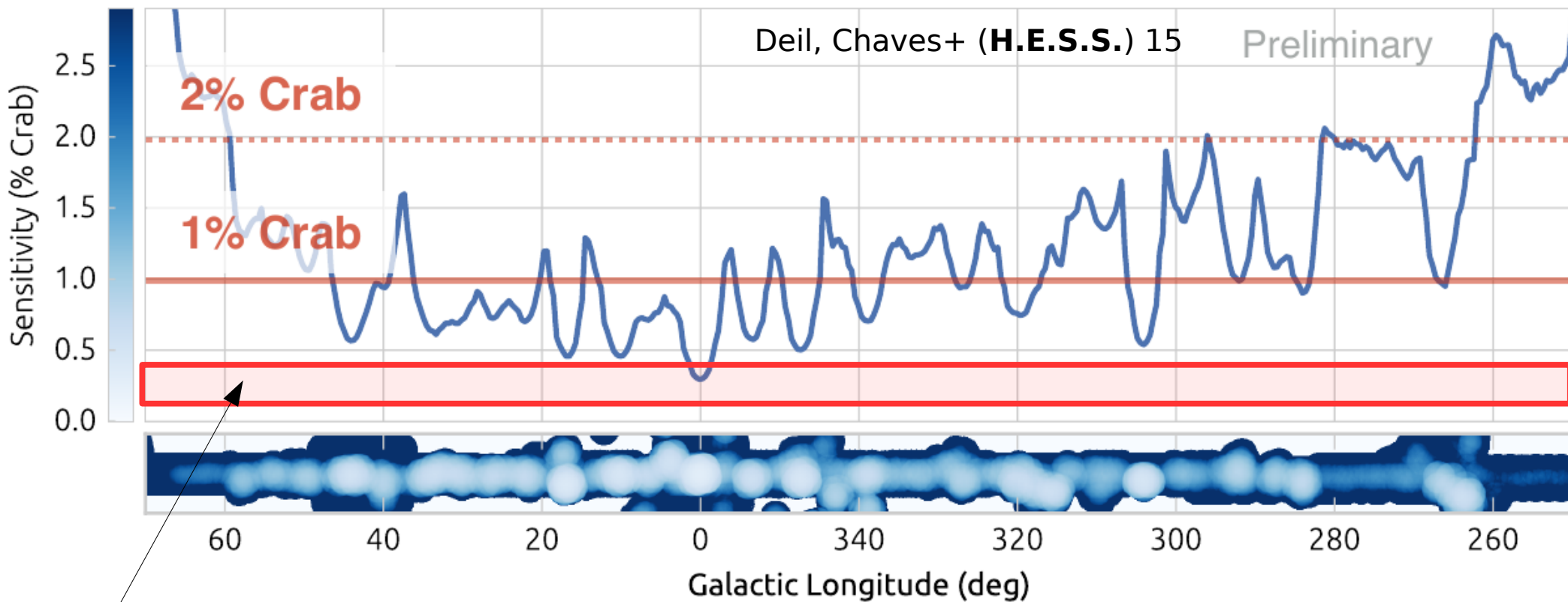
Firm identifications

Source Name	Associated object	Class	Evidence	Reference
HESS J1018–589A	1FGL J1018.6–5856	Binary	Variability	H. E. S. S. Collaboration et al. (2015a)
HESS J1302–638	PSR B1259–63	Binary	Variability	Aharonian et al. (2005a)
HESS J1826–148	LS 5039	Binary	Variability	Aharonian et al. (2006c)
HESS J0852–463	Vela Junior	SNR	Morphology	Aharonian et al. (2005b)
HESS J1442–624	RCW 86	SNR	Morphology	Aharonian et al. (2009)
HESS J1534–571	G323.7-01.0	SNR	Morphology	HESS SNR shell paper (2015)
HESS J1713–397	RX J1713.7–3946	SNR	Morphology	Aharonian et al. (2004)
HESS J1731–347	G353.6–0.7	SNR	Morphology	H.E.S.S. Collaboration et al. (2011b)
HESS J1800–240	W 28	SNR	Position	Aharonian et al. (2008)
HESS J0835–455	Vela X	PWN	Morphology	Aharonian et al. (2006a)
HESS J1303–631	PSR J1301–6305	PWN	ED Morph.	H.E.S.S. Collaboration et al. (2012)
HESS J1514–591	MSH 15–52	PWN	Morphology	Aharonian et al. (2005a)
HESS J1825–137	PSR J1826–1334	PWN	ED Morph.	Aharonian et al. (2006d)
HESS J1356–645	PSR J1357-6429	PWN	Position	H.E.S.S. Collaboration et al. (2011a)
HESS J1418–609	PSR J1418-6058	PWN	Position	Aharonian et al. (2006b)
HESS J1420–607	PSR J1420-6048	PWN	Position	Aharonian et al. (2006b)
HESS J1554–550	G327.1-01.1	PWN	Morphology	Section 5.7.5
HESS J1747–281	G0.9+0.1	PWN	Morphology	Aharonian et al. (2005b)
HESS J1818–154	G015.4+00.1	PWN	Morphology	H. E. S. S. Collaboration et al. (2014)
HESS J1849–000	PSR J1849–0001	PWN	Position	Section 5.7.15
HESS J1837–069	PSR J1838–0655	PWN?	Morphology	Marandon et al. (2008)
HESS J1640–465	G338.3-0.0	Composite?	Position	Abramowski et al. (2014b), Gotthelf et al. (2014)
HESS J1119–614	PSR J1119-6127	Composite	Position	Section 5.7.1
HESS J1813–178	PSR J1813-1749	Composite	Position	Funk et al. (2007), Gotthelf & Halpern (2009)
HESS J1833–105	G21.5–0.9	Composite	Position	Section 5.7.10
HESS J1846–029	PSR J1846-0258	Composite	Position	Section 5.7.13
HESS J1930+186	G54.1+0.3	Composite	Position	Acciari et al. (2010), Section 5.5



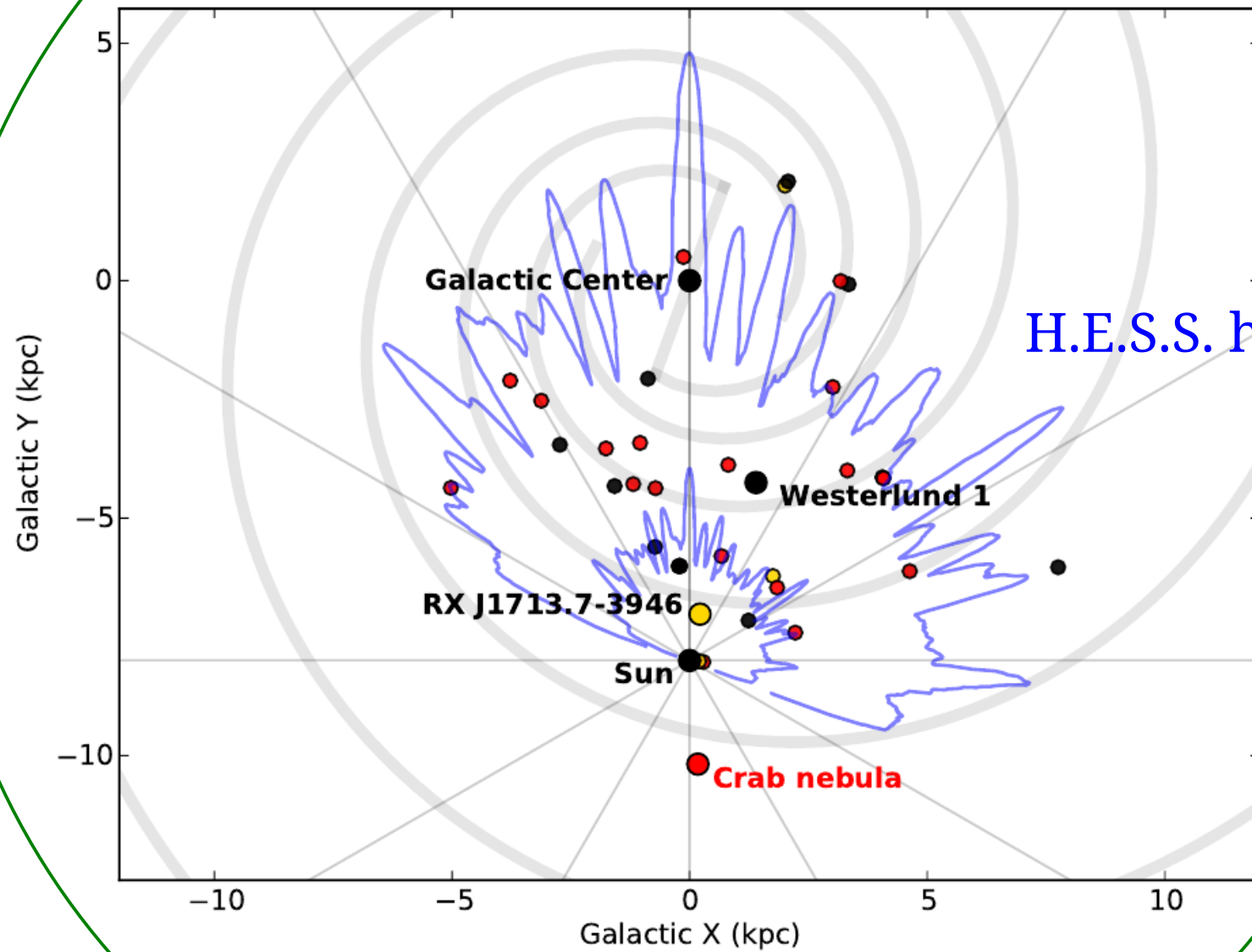


GPS IN CONTEXT



~**mCrab** and uniform sensitivity with **CTA GPS** in just 2 years

CTA horizon



H.E.S.S. horizon