Nonrelativistic Shocks of Young Supernova Remnants in Kinetic Simulations

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Setting the stage…

Diffusive Shock Acceleration (DSA) process at young SNR shocks assumed to provide the main part of Galactic cosmic-ray flux. Possibly relevant for mildly-relativistic flows in AGN jets.

Attributes relevant for DSA:

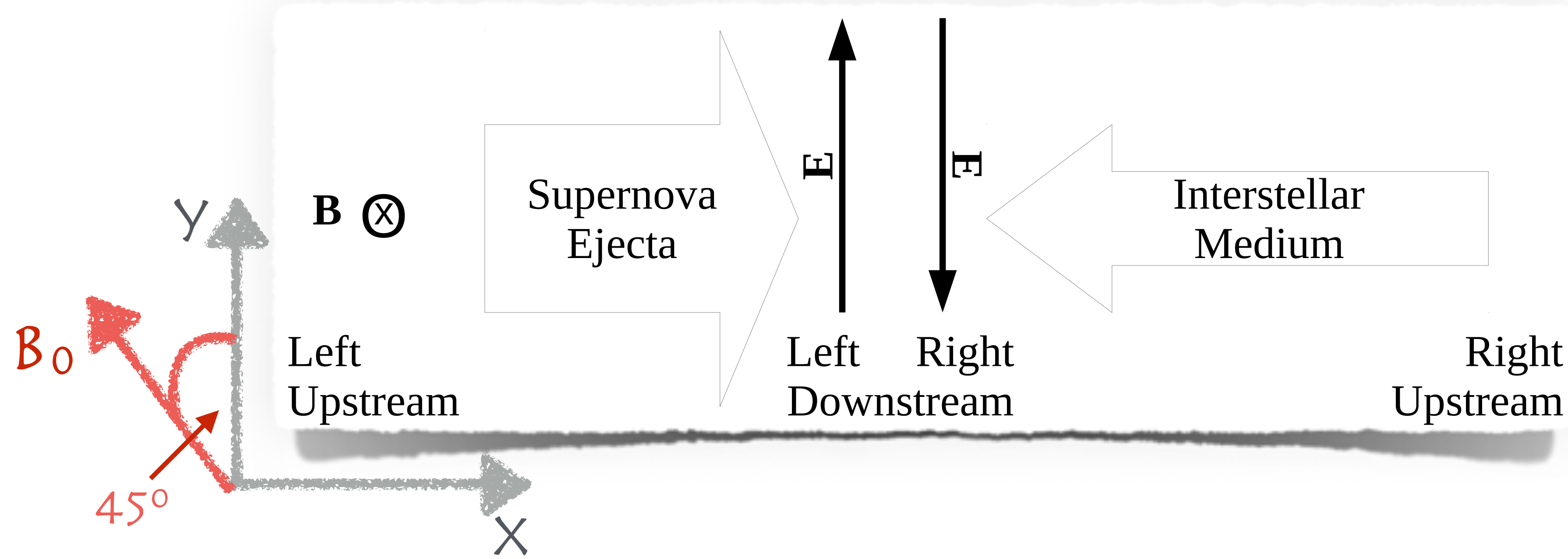
- shock structure: ion driven but electron dynamics important
- EM field amplitudes
- particle pre-acceleration processes: electron injection constitutes the central unresolved issue

Current main interest:

- high Alfven Mach number shocks: regime of weakly magnetized plasma
- high-speed nonrelativistic shocks: mediated by Weibel-type filamentation instabilities
- electron acceleration - no hybrid simulations

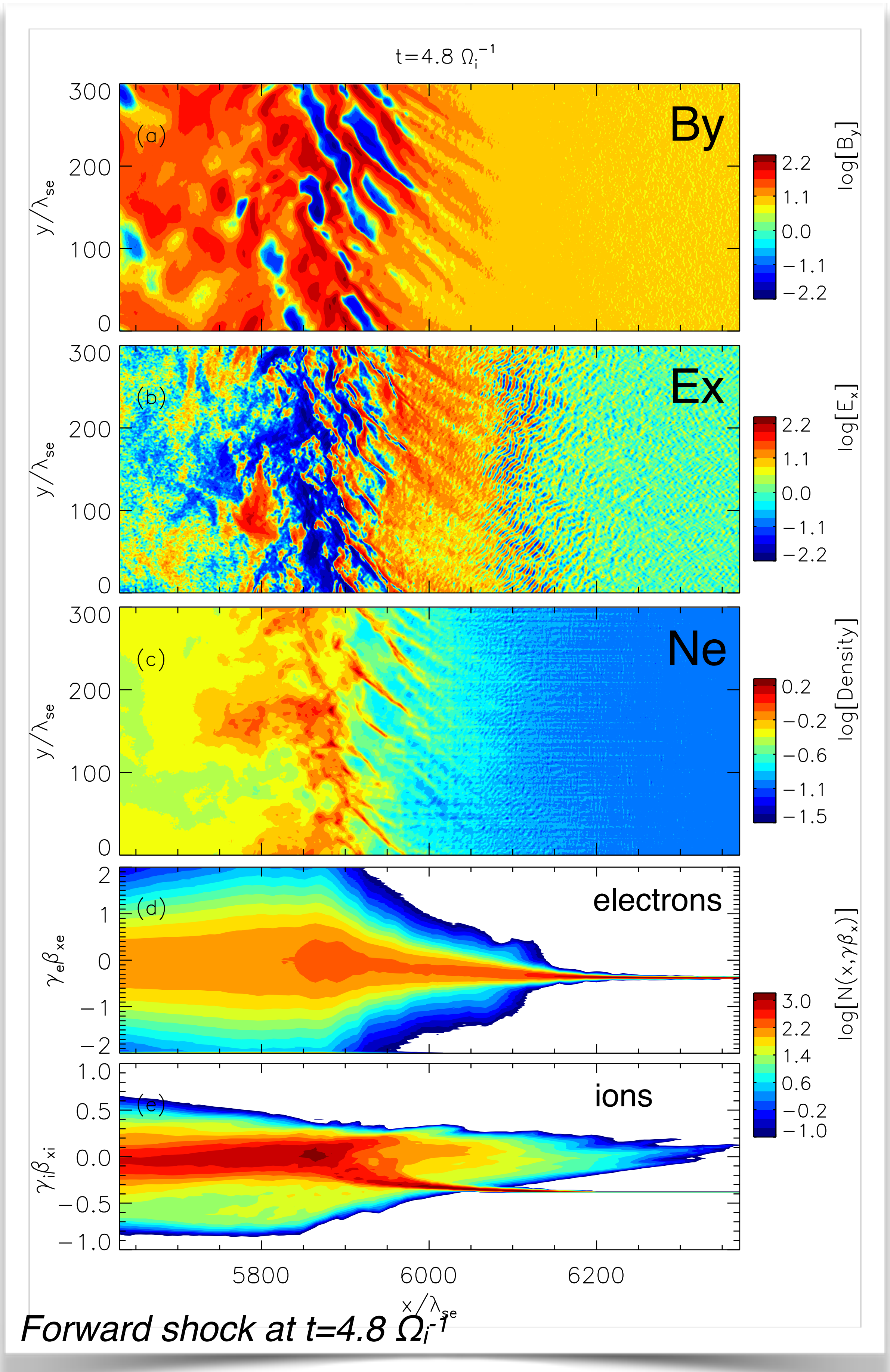
Today's topic: perpendicular shocks

Wieland et al. 2015, ApJ submitted

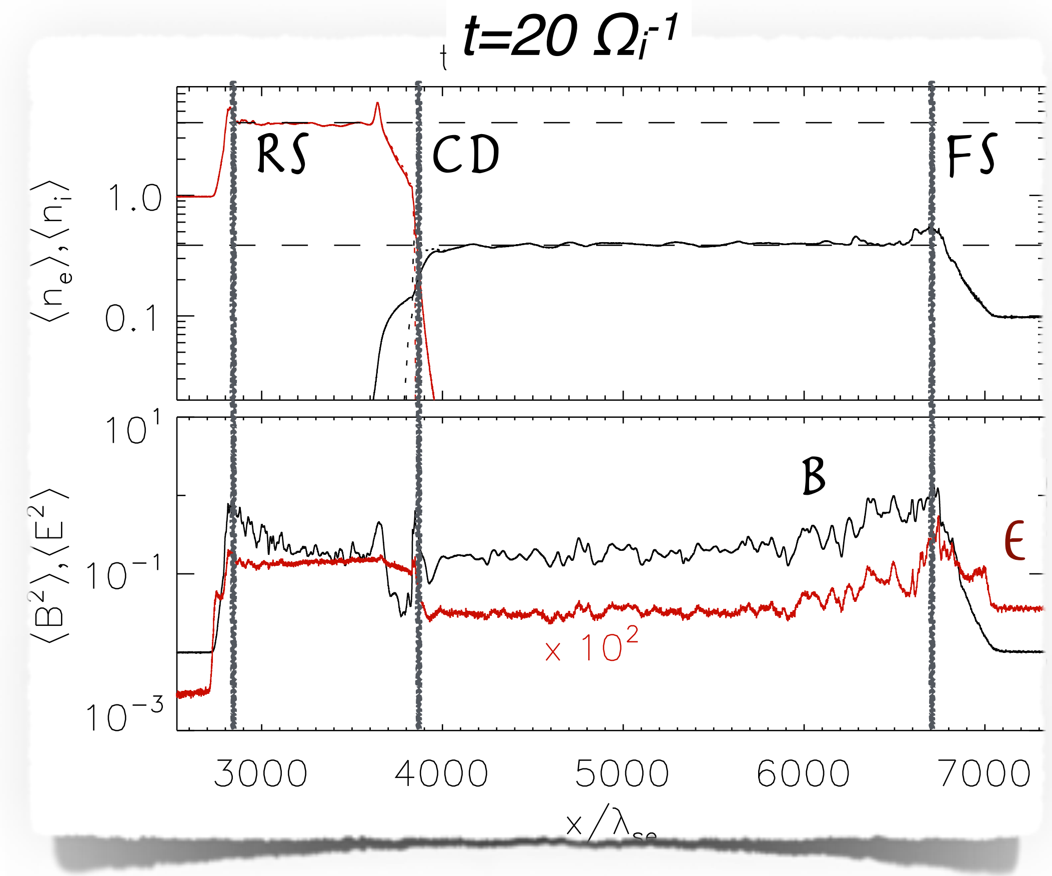


- 2D3V kinetic PIC simulations (m_i/m_e=50)
- high relative collision speed (v_{rel}=0.38c)
- stream-counterstream density asymmetry of 10: system of forward and reverse shock + CD
- Alfven Mach numbers for both shocks: M_A~ 28
- different sonic Mach numbers: M_S~ 755 (forward); M_S~ 250 (reverse)
- magnetic field at 45° to the x-y plane
- low plasma beta $\beta_e \ll 1$: initially cold plasma flows
- simulations complement recent 2D3V PIC studies of high Mach fast nonrelativistic shocks in the regime of moderate or high β_e and for strictly in-plane or out-of-plane MF orientations (*Amano, Hoshino, Kato, Matsumoto, 2009-2015*)

Shock structure

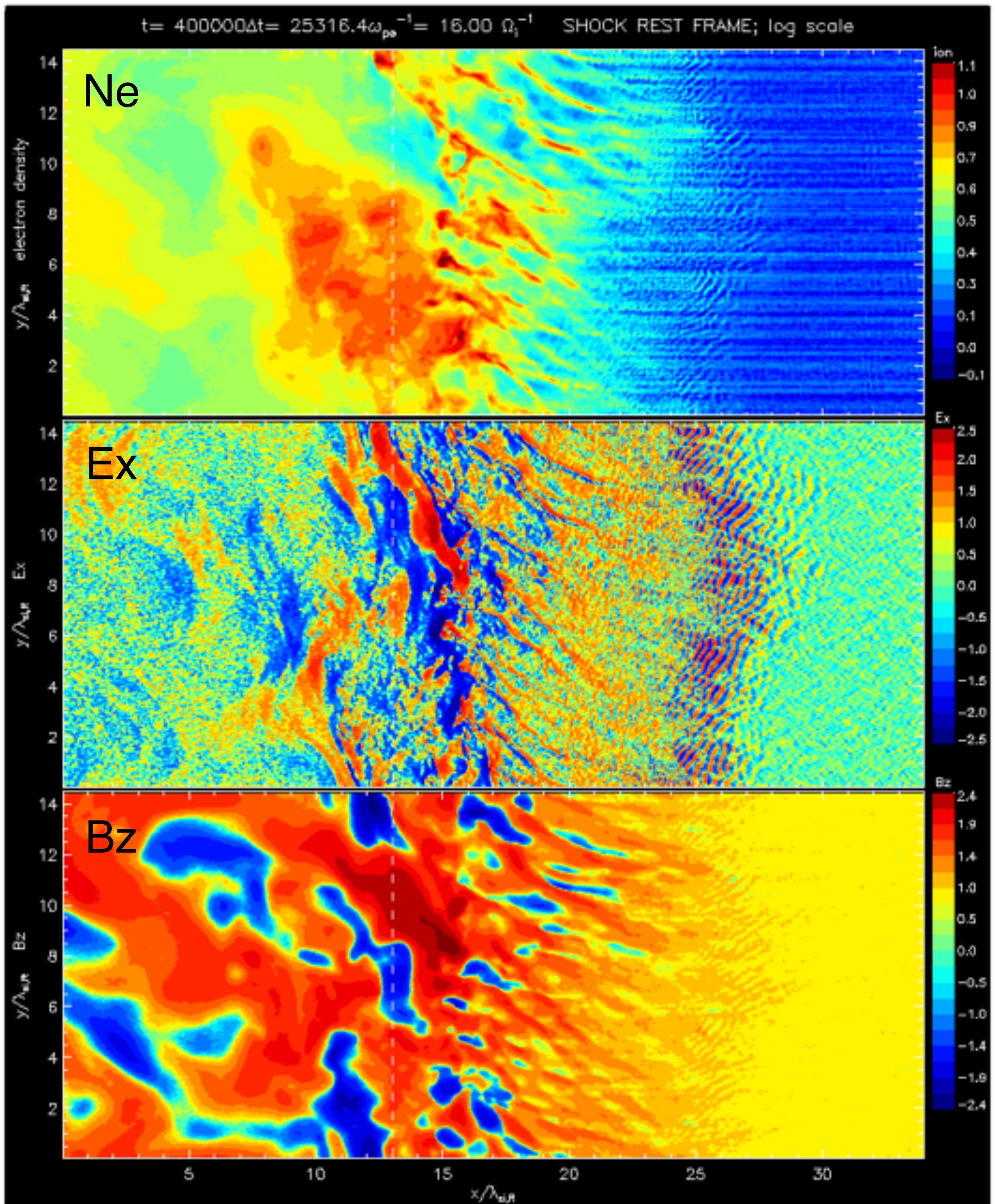


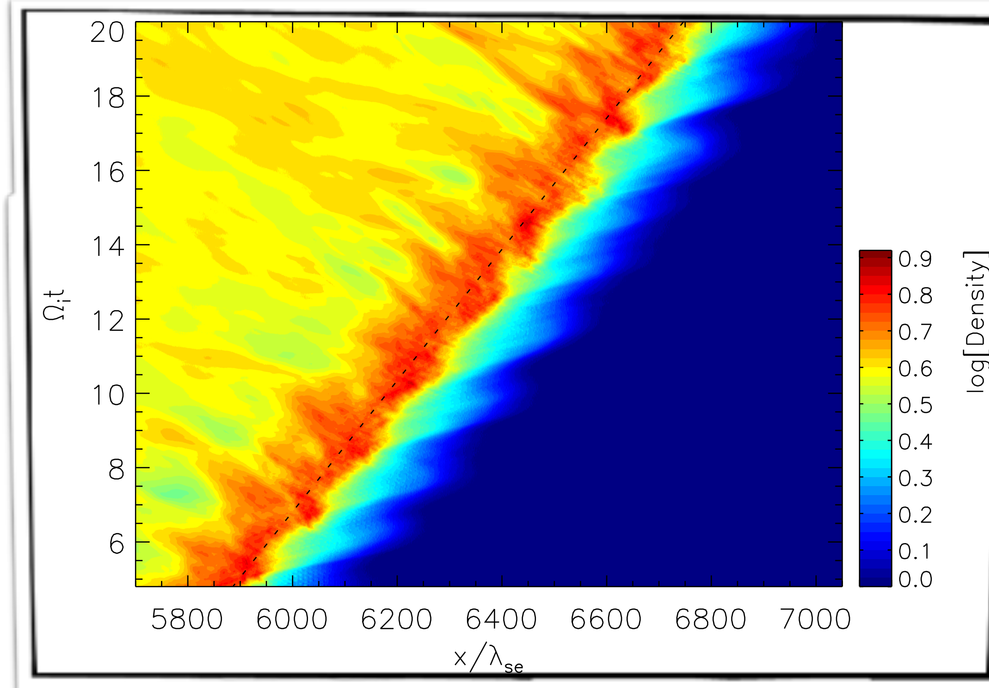
Forward shock at t=4.8 Ω_i^{-1}



- steady-state system of shocks separated by CD formed within a few ion cyclotron times
- structure governed by ion reflection
- shocks mediated by ion-beam Weibel-type filamentation instabilities that generate mainly magnetic turbulence
- strong Buneman modes in the shock foot

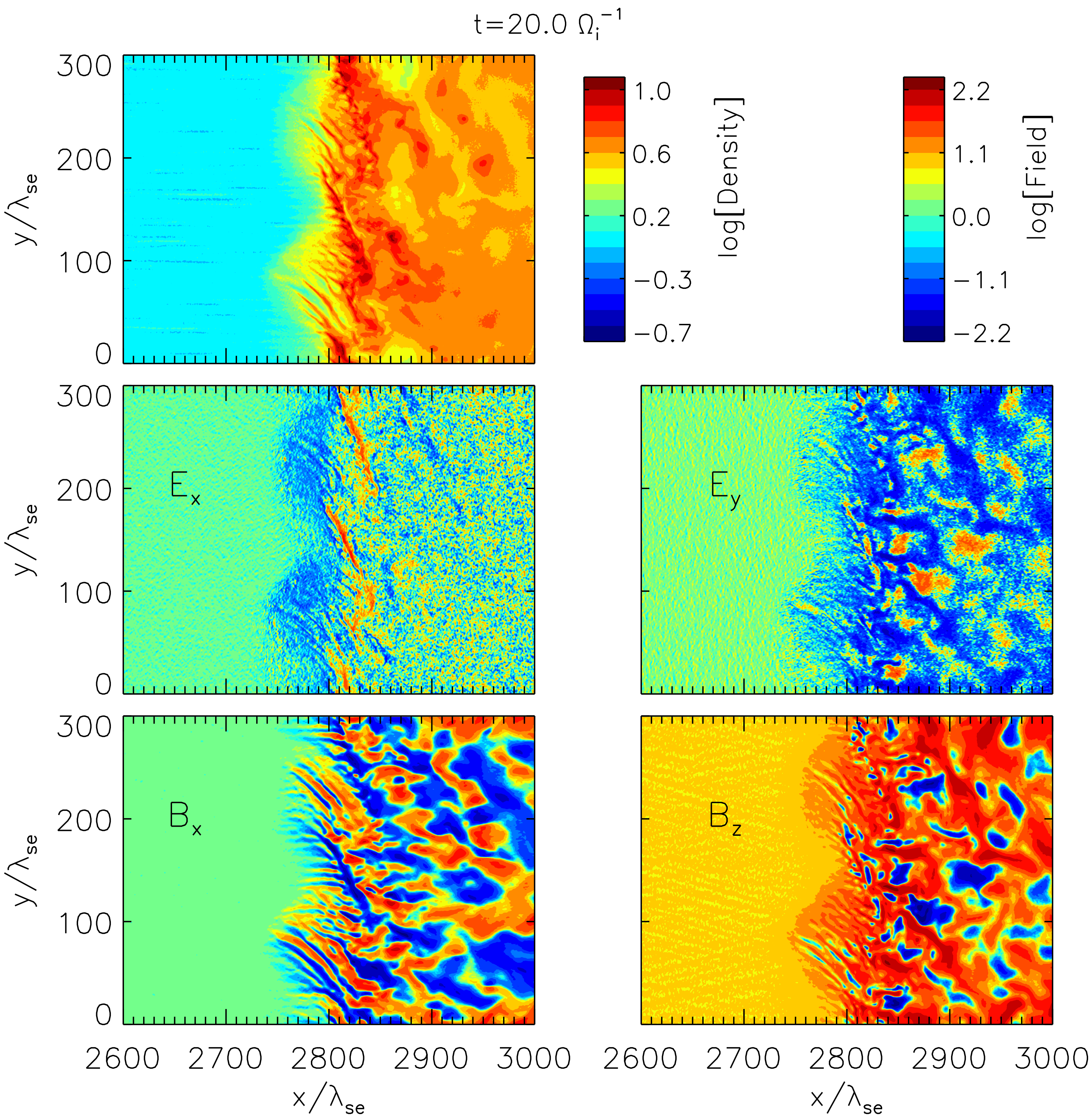
Shock reformation…





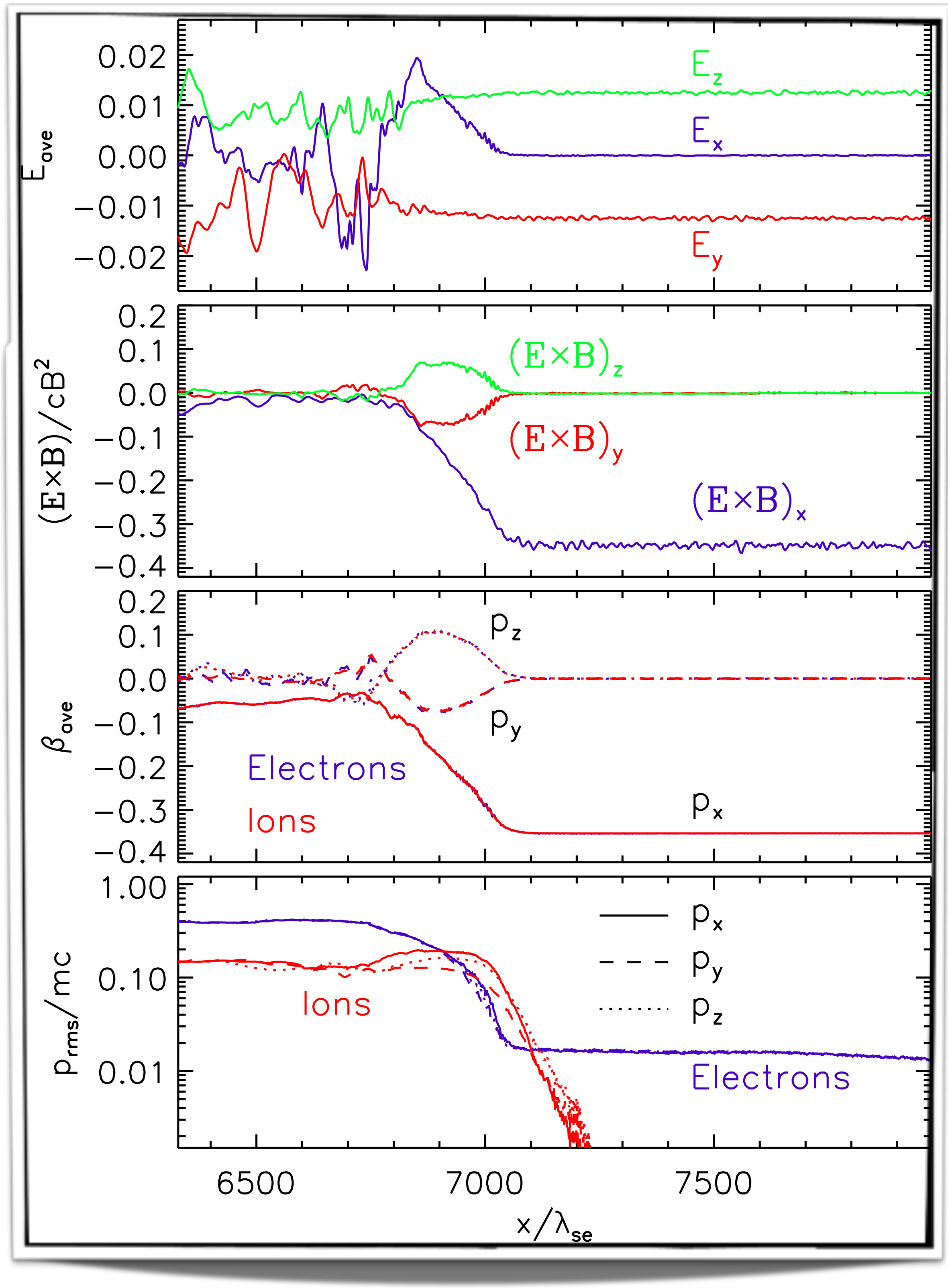
- cyclic shock self-reformation caused by dynamics of shock-reflected ions governed by the physics of current filament mergers in the shock ramp
- period of ~1.5 Ω_i^{-1} similar at both shocks and roughly constant throughout the simulation

Shock reformation… and rippling



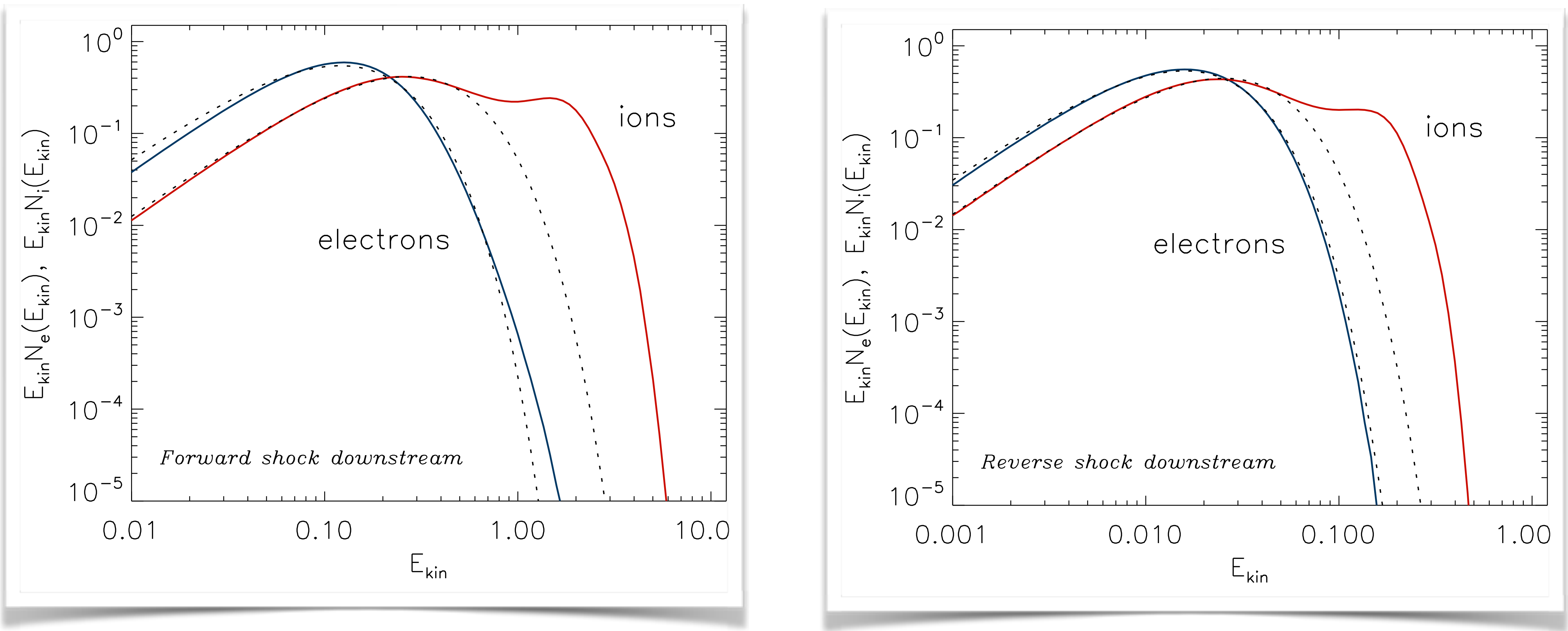
- amplitude of rippling significantly larger at reverse shock with lower sonic Mach number
- spatial scale of ~20 λ_{si}

Notes on electron injection



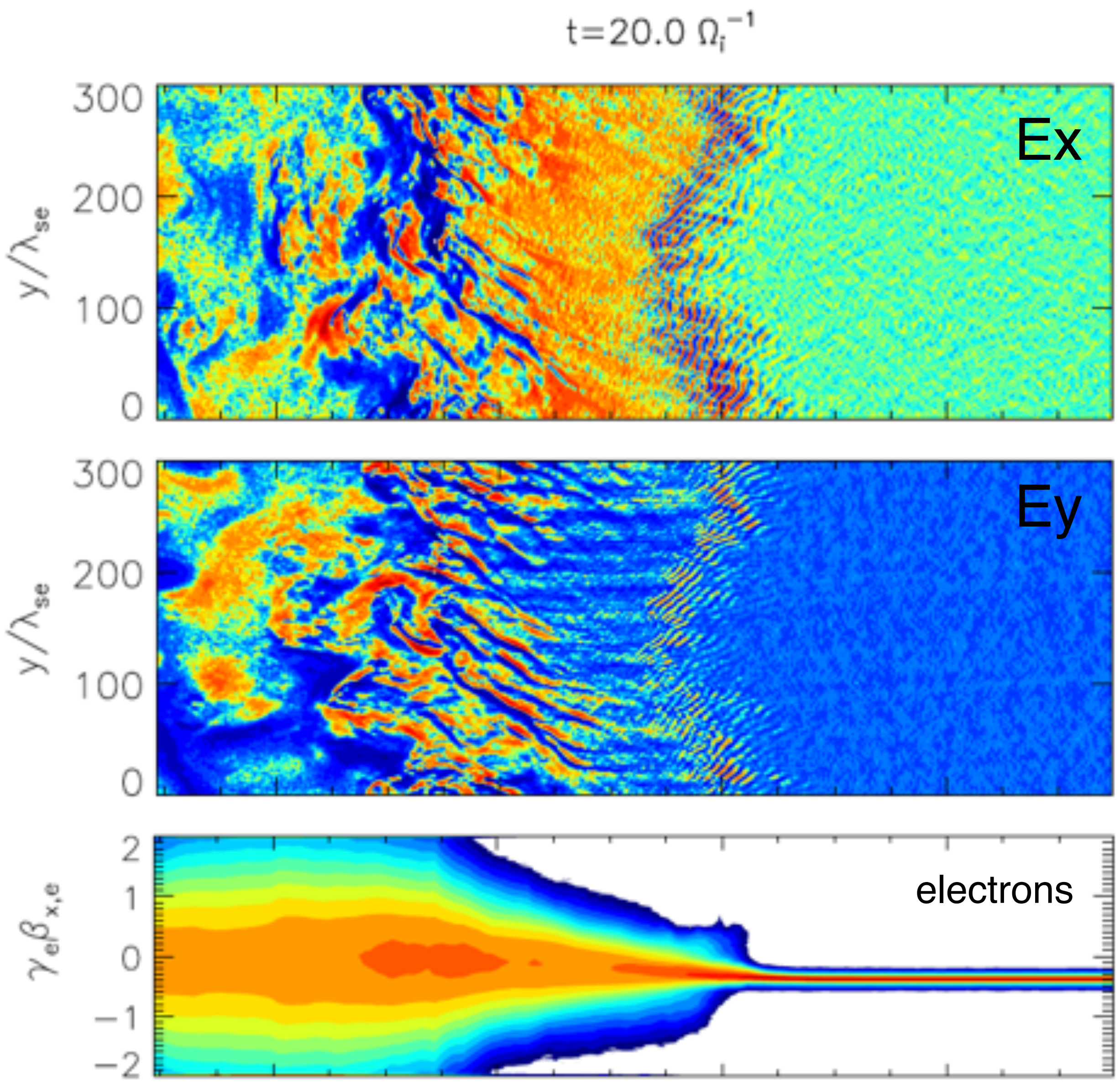
- gradient drift along elctric field - necessary for shock-drift acceleration (SDA) - is not observed: local MF gradients dominate the global gradient accross the shock
- bulk electron and ion motion commensurate with ***E***x***B*** drift in direction and amplitude

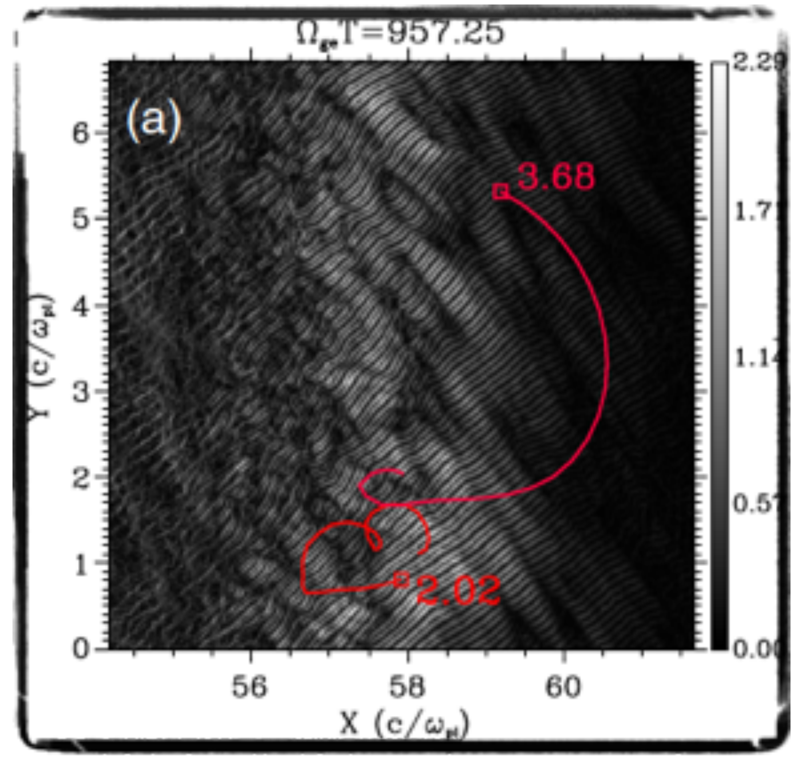
Particle pre-acceleration



- stable suprathermal tails in ion spectra resulting from shock-surfing acceleration (SSA)
- efficient electron heating; no or marginal electron acceleration to high energies

Notes on electron injection





Matsumoto et al. 2013

- despite suitable conditions exist ($M_A \gtrsim 16$; Matsumoto et al. 2012) electron SSA is not observed because the amplitude of Buneman modes in the shock foot is insufficient for trapping relativistic electrons

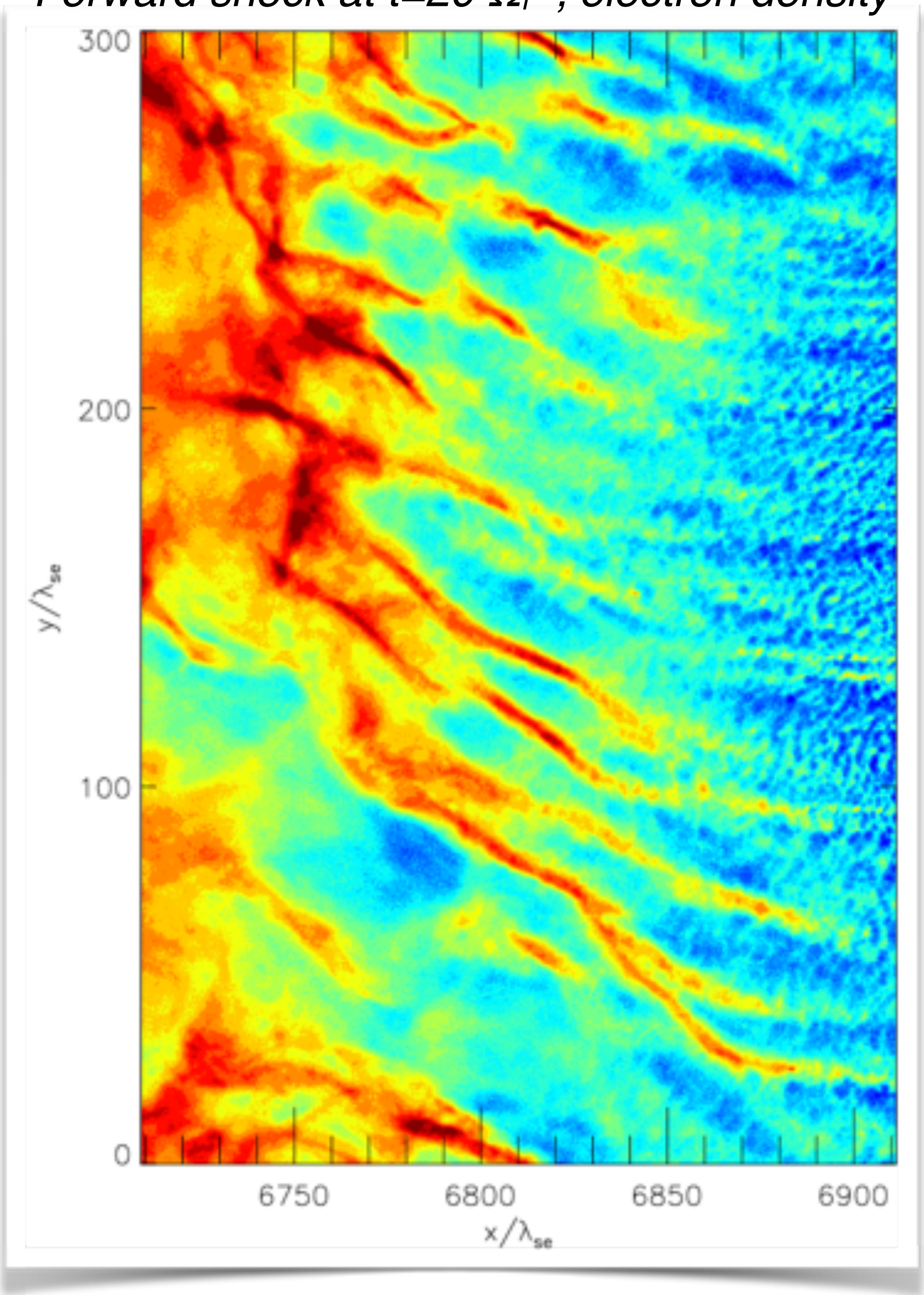
$$E_B \sim cB_0$$

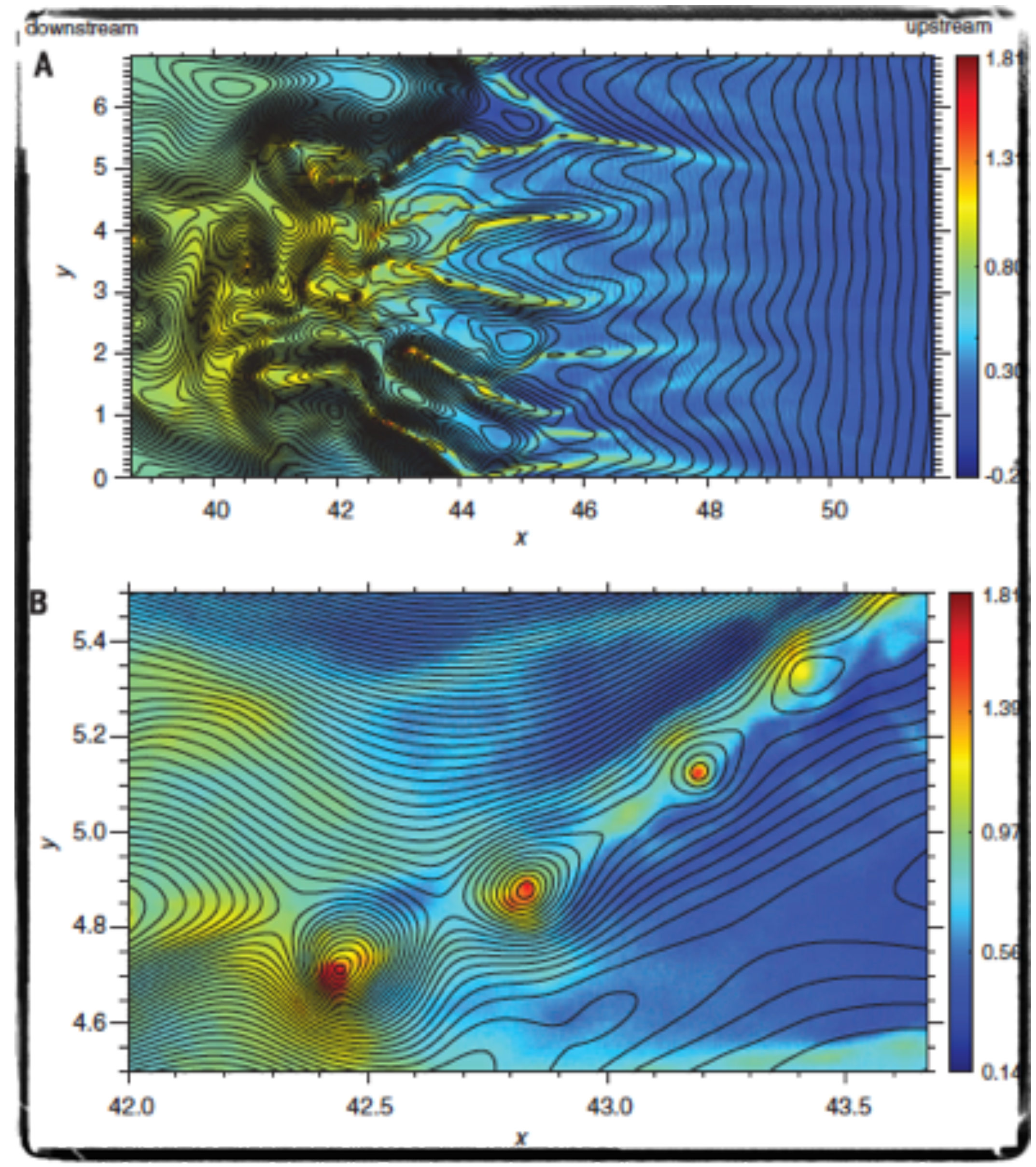
- bulk electron thermalization occurs instead

- inefficient electron acceleration observed by us (low β_e) and by Kato & Takabe (2010; high β_e) suggests that β_e is not deciding factor for the generation of non-thermal tails in the electron SSA
- possible reasons for discrepancy: MF orientation - dimensionality effects (tests suggests electron acceleration for strictly out-of-plane configuration)

Notes on electron injection

Forward shock at t=20 Ω_i^{-1}; electron density

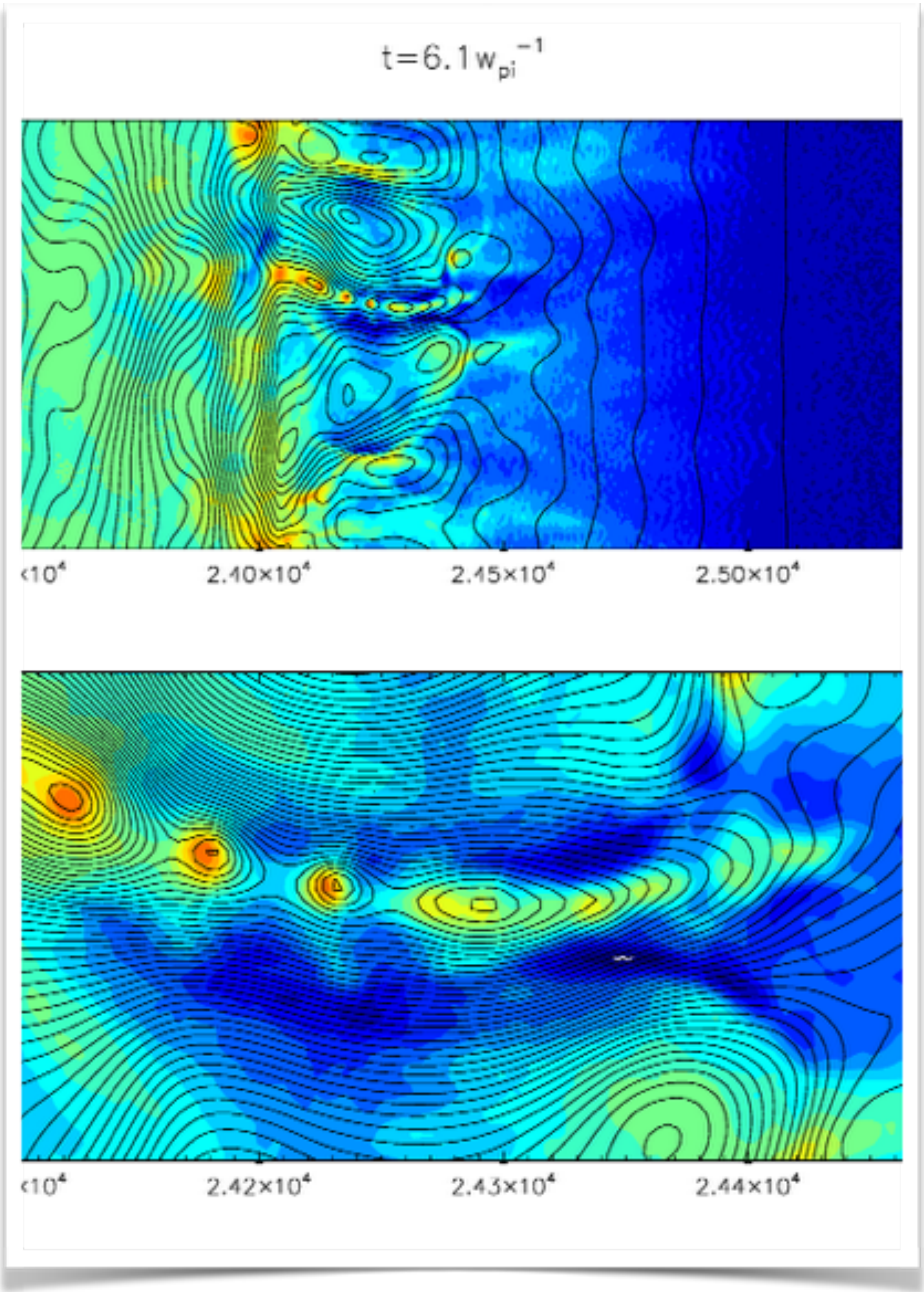


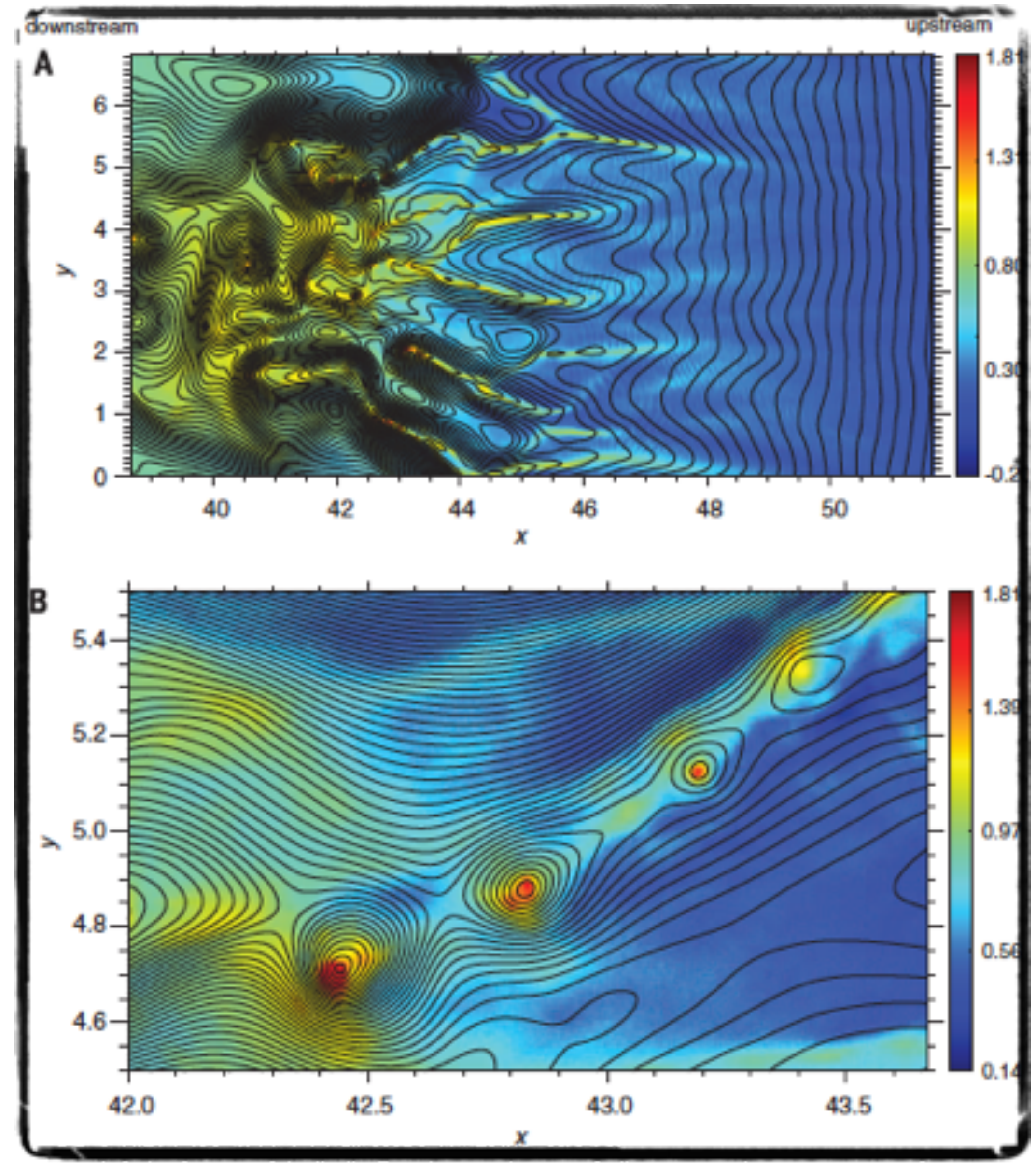


Matsumoto et al. 2015

- no evidence for turbulent reconnection at 45^{o} configuration (and also 90^{o})

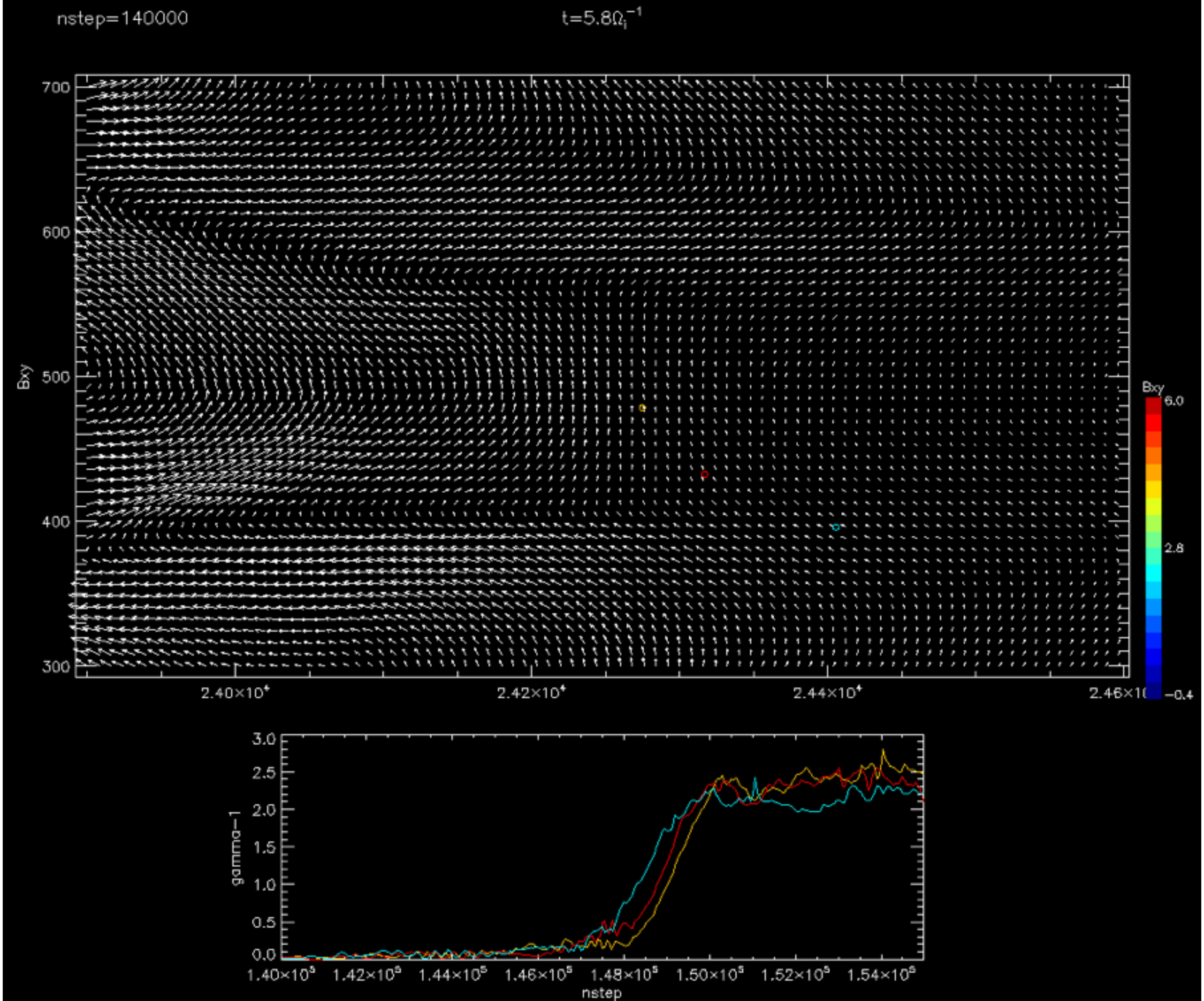
Notes on electron injection





Matsumoto et al. 2015

- turbulent reconnection observed for in-plane (0°) configuration


nstep=140000
t=5.8Ω_i^{-1}
Bxy
700
600
500
400
300
2.40×10⁴
2.42×10⁴
2.44×10⁴
Bxy
6.0
2.8
−0.4
gamma−1
3.0
2.5
2.0
1.5
1.0
0.5
0.0
1.40×10⁵
1.42×10⁵
1.44×10⁵
1.46×10⁵
1.48×10⁵
1.50×10⁵
1.52×10⁵
1.54×10⁵
nstep

Summary

- shock transition mediated by Weibel-like instabilities leading to current filaments
- rippling observed at both shocks, reformation period ~1.5 Ω_i^{-1}
- no gradient drift, only ExB drift
- electron heating, inefficient acceleration to high energy
- tests suggest electron acceleration for out-of-plane magnetic field
- 3-D simulations needed