# Disappearing track searches at LHC and future colliders

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Dark Matter searches in the 2020s At the crossroads of the WIMP 11—13 Nov, 2019, U-Tokyo, Kashiwa Campus



## Neutralino dark matter

- The lightest neutralino is a good candidate of dark matter.
- Bino DM would be overproduced, while Wino and Higgsino could have the right relic density.
- If thermally produced, the dark matter mass should be < ~1 TeV for Higgsino and < -3 TeV for Wino.
- Small mass differen between the lightest change and neutralino(= DM) is favoured

 $\Delta m_{\chi_{2}^{0},\chi_{1}^{0}} = \bullet <2$ 

•10



 $\mu$ [TeV]

### Viable parameter space

#### Wino DM

- The lightest SUSY particle in AMSB and PGM models
- It is not excluded (except for narrow region around ~2.4 TeV) by indirect DM searches when the DM-halo densities in galaxies have flat central profiles ("cores").
- It is not excluded by direct searches due to the extremely small directdetection cross-section.

#### <u>Higgsino DM</u>

- Very pure higgsino DM is excluded due to too high Z-boson-mediated nucleon scattering cross section. Gaugino mass should be < O(10) PeV.</li>
- The electron EDM limit and direct DM searches require gaugino masses to be > O(TeV).

### Can wino/higgsino DM be discovered at the LHC ?

- In general DM searches, mono-X (X = jet, higgs, Z...) signatures are used \*.
- Low sensitivity by mono-jet for wino and higgsino DM due to the small production cross section.
- For suppression BG, we need additional object.

**E**<sub>T</sub>miss





 $\pi$ iou ui $\pi$  ( $\pi$ - jet etc.) signature norm mediator-decay is also used

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\*Also di-X (X= jet etc.) signature from mediator-decay is also used.

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### Can wino/higgsino DM be discovered at the LHC?

- In general DM searches, mono-X (X = jet, higgs, Z...) signatures are used \*.
- Low sensitivity by mono-jet for wino and higgsino DM due to the small production cross section.
- For suppression BG, we need additional object.
  - If Δm (chargino-neutralino splitting) is not too small
    - → Low momentum leptons from offshell W and Z
  - If ∆m is tiny, chargino's lifetime is long
    → Disappearing track



## Charged wino lifetime

JHEP 06 (2018) 022, auxiliary material

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- When wino is the lightest SUSY particle, the lightest neutralino is "pure" wino in a wide parameter region (i.e. higgsino is heavier than wino by ~ 1 TeV)
  - Δm ~ 160 MeV
  - Chargino ст<sub>0</sub> ~ 6 ст

## Charged higgsino lifetime

- Lifetime of charged higgsino depends on gaugino masses (due to mixing).
- EDM and direct-DM searches favour smaller mixing, so smaller  $\Delta m$ .
  - Δm < O(1GeV)
  - Higgsino :  $c\tau_0 < \sim 2$  cm



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### Long-lived chargino : history

#### 1997: C.-H. Chen, M. Drees, and J. F. Gunion mention ideas for the search \*.

"We will also find that the  $\tilde{\chi}_1^{\pm}$  's lifetime is not likely to be sufficiently long that it will appear as a stable particle track in the detector short tracks in a vertex detector are, however, a distinct possibility."

Phys. Rev. D 55, 330 (1997)

2008: **S. Asai, T. Moroi, T. T. Yanagida** proposed a method to measure wino properties using ATLAS detector and the disappearing tracks.

- mass measurement using the velocity (ToF) and momentum
- Iifetime measurement from decay radius distribution.

Phys. Lett. B 664, 185 (2008)

2011: The first ATLAS result

2015: The first CMS result

\* This paper mention also about heavily-ionizing particles and soft tracks from the chargino decay

## Inner tracking detector

**ATLAS Inner Detector** 



The new inner most layer from 2015

Figure 1 3D visualisation of the structure of

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### Disappearing track search @ LHC



- Long lived chargino searches using disappearing track
  - Short tracks which do not have associated hits in the outer part of the tracker and calorimeters.
- Pions, with the transverse momentum about the mass difference (i.e. ~200 MeV from wino, ~300 MeV from higgsino), are not reconstructed in standard tracking algorithms. (e.g. Track threshold in ATLAS is ~500 MeV)



Invisible

## Wino search by ATLAS/CMS 2015-2016 data



I he difference in peak sensitivity in lifetime is due that in the track length (shorter in ATLAS).

JHEP 08 (2018) 016

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### Higgsino direct production



### Higgsino direct production



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### Higgsino: Filling the gap 1

*ATL-PHYS-PUB-2019-011* 

Improving the disappearing track search for shorter lifetimes

- Currently, 4-layer pixel tracks (R~12 cm) are used in ATLAS
- We could use 3-layer pixel tracks (R-scm) for better sensitivity, but it is technically challenging.
  - Poor momentum resolution
    - Could be mitigated by using the collision vertex as additional point in tracking (vertexconstraint).
  - High combinatorial background rate
    - Suppress BG by adding low momentum-pion tracking (pT > 300 MeV) to the signature.
    - For reducing CPU and disk, soft-pion tracks are seeded only around the disappearing track.



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### Higgsino: Filling the gap 2



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## LHC / HL-LHC plan

#### LHC / HL-LHC Plan

![](_page_16_Picture_2.jpeg)

![](_page_16_Figure_3.jpeg)

Comput. Softw. Big Sci. 3 (2019) 7

#### 20 time more data will be collected at the HL-LHC

### HL-LHC: Inner detector

#### ATLAS (scaled)

ATL-PHYS-PUB-2019-014

![](_page_17_Figure_3.jpeg)

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### Wino/Higgsino at the HL-LHC

ATL-PHYS-PUB-2018-031

![](_page_18_Figure_2.jpeg)

Sensitivity (95%CL exclusion): **wino: 850 GeV**, **higgsino: 250GeV**, assuming a similar analysis to the 36 fb<sup>-1</sup> would be done. **The reach may be higher by improving the analysis technique.** 

### Future Circular Collider (FCC-hh)

re Cenet

**Prealps** 

**Aravis** 

Copyright CERN 2014

LHC

Jura

- 100 km tunnel in Geneva area
- pp collider with  $\sqrt{s} = 100 \text{ TeV}$
- 200—1000 collisions per bunch crossing
- Total integrated lumi. ~ 20 ab<sup>-1</sup>
- Much higher sensitivities to various new physics than LHC

![](_page_19_Figure_6.jpeg)

### DT + Hit-time information

- Low Gain Avalanche Detectors (LGAD) have time resolution of 10-30 ps
- If the detector can be used in FCC as the inner pixel-detector (not at an additional timing-layer), we can use the hit-time for two purpose
  - 1. BG fake tracks (random-combination) decrease by requiring consistent time of pixel-hits on track.
  - 2. Measure the velocity of a particle.
    - If hit-time resolution is 20 ps, velocity resolution for charginos could be  $\sim 6\%$ .

![](_page_20_Picture_8.jpeg)

2017 JINST 12 P05003

![](_page_20_Picture_10.jpeg)

## Expected reach at FCC

- Eur. Phys. J. C 79 (2019) 469
- Modified pixel-detector layout (5 layers within 15cm from the beamline)
- BG rejection using time information ( $\chi^2$ /ndf in the fit)

![](_page_21_Figure_4.jpeg)

## FCC-hh: forward region ?

JHEP06(2017)119

- Due to the very high beam energy, signal charginos are emitted also to the forward region.
- The lifetime in the lab-frame is longer due to the boost.
- We could have significant signal events in the forward region by putting the tracker close to the target (~40 cm)

![](_page_22_Figure_5.jpeg)

![](_page_22_Figure_6.jpeg)

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### Gaugino mass measurement at the FCC-hh

JHEP 05 (2019) 179

- If we can measure the velocity and momentum of wino, the mass can be measured.
- Other gaugino masses (i.e. bino and gluino) can be reconstructed by using the measured wino mass.
- The gaugino mass measurements would imply also the next particle-mass scale (higgsino, Higgses and sfermions)

![](_page_23_Figure_4.jpeg)

## Conclusion

- Disappearing track is a unique tool to search for (nearly) pure wino and higgsino at the LHC.
- For nearly pure higgsino, there is a gap between the prompt and disappearing track searches. There are several ideas to explore the gap region.
- FCC-hh will **completely cover** the mass region having the correct DM relic abundance.

Back up

### CMS inner tracker

CMS-TDR-014

![](_page_26_Figure_2.jpeg)

### Strong production ATLAS/CMS

![](_page_27_Figure_1.jpeg)

### CMS analysis

<u>CMS (38 fb-1)</u>

![](_page_28_Figure_3.jpeg)

Run period	Estimated 1	Observed events		
	Leptons	Spurious tracks	Total	Observed events
2015	$0.1\pm0.1$	$0^{+0.1}_{-0}$	$0.1\pm0.1$	1
2016A	$2.0\pm0.4\pm0.1$	$0.4\pm0.2\pm0.4$	$2.4\pm0.5\pm0.4$	2
2016B	$3.1\pm0.6\pm0.2$	$0.9\pm0.4\pm0.9$	$4.0\pm0.7\pm0.9$	4
Total	$5.2\pm0.8\pm0.3$	$1.3\pm0.4\pm1.0$	$6.5\pm0.9\pm1.0$	7

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### ATLAS analysis

#### ATLAS (36 fb-1)

![](_page_29_Figure_3.jpeg)

	Electroweak channel			Strong channel				
Number of observed eve	ents with	$p_{\rm T} > 100$	GeV in	$high-E_T^m$	<sup>iiss</sup> regions			
	9			2				
Number of expected events with $p_{T} > 100 \text{ GeV}$ in high- $E_{T}^{miss}$ regions								
Hadron+electron background	6.1	$\pm 0.6$		1.78	$\pm 0.32$			
Muon background	0.15	$\pm 0.09$		0.05	$\pm 0.08$			
Fake background	5.5	$\pm 3.3$		0.1	$\pm 0.4$			
Total background	11.8	$\pm 3.1$		1.9	$\pm 0.4$			
$p_0$	0.5			0.47				
Observed $\sigma_{\rm vis}^{95\%}$ [fb]	0.22			0.12				
Expected $\sigma_{\rm vis}^{95\%}$ [fb]	$0.28^{+0.11}_{-0.08}$			$0.12\substack{+0.07 \\ -0.04}$				
Number of expected signal events with $p_{T} > 100 \text{ GeV}$ in high- $E_{T}^{miss}$ regions								
	13.5	$\pm 2.1$		5.6	$\pm 0.8$			

### Wino mass measurement

mass = momentum/
$$\beta \cdot \sqrt{(1 - \beta^2)}$$
,  
 $\beta = v/c$ 

- Velocity is measured by the tracker using time information
- How to measure the **momentum** ?
  - We can not use the momentum of charged-wino tracks because of the too poor resolution; the track length is too short (< 10cm) !</li>
  - Instead, we can reconstruct from *E*T<sup>miss</sup> and direction of charged winos, because pion carry little momentum (O(100 MeV))

![](_page_30_Figure_6.jpeg)

![](_page_30_Figure_7.jpeg)

### Bino mass measurement

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- Reconstruct **Bino mass** from **Wino and W** momentum
  - Wino momentum : reconstruct from the measured **velocity** and Wino **mass**.
  - W momentum : reconstruct using fat jets.

![](_page_31_Figure_5.jpeg)

![](_page_31_Figure_6.jpeg)

![](_page_31_Figure_7.jpeg)

### Gluino mass measurement

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#### Gluino mass is reconstructed by "hemisphere" analysis

- 1. Define two hemi-spheres using two disappearing-track directions
- 2. Iteratively assign jets to each hemi-sphere and update the directions
- 3. Reconstruct the gluino mass from jets and Winos.
- Gluino mass can be estimated from the cross-section.
  - Comparing the two estimates would be good test of SUSY hypothesis.

![](_page_32_Figure_8.jpeg)

Isolated lepton-veto is also applied

## Implication

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![](_page_33_Figure_2.jpeg)

Three model parameters can be constrained by the gaugino mass measurements.

The gaugino mass measurements would imply also **the next particlemass scale** (Higgsino, Higgses and sferimions)