

宇宙の進化と素粒子模型

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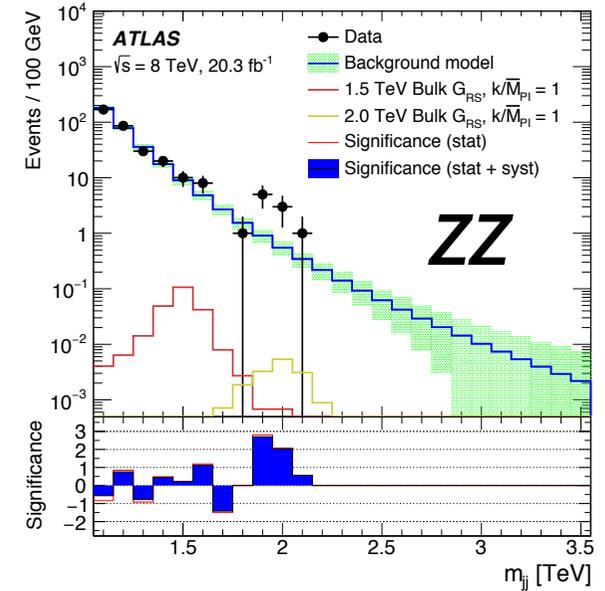
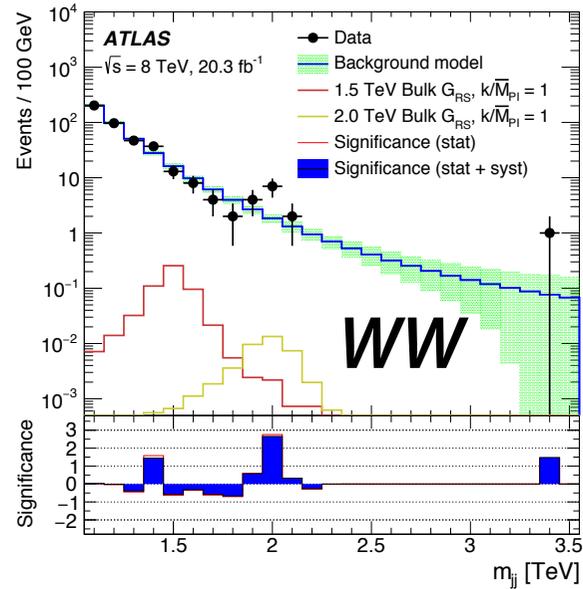
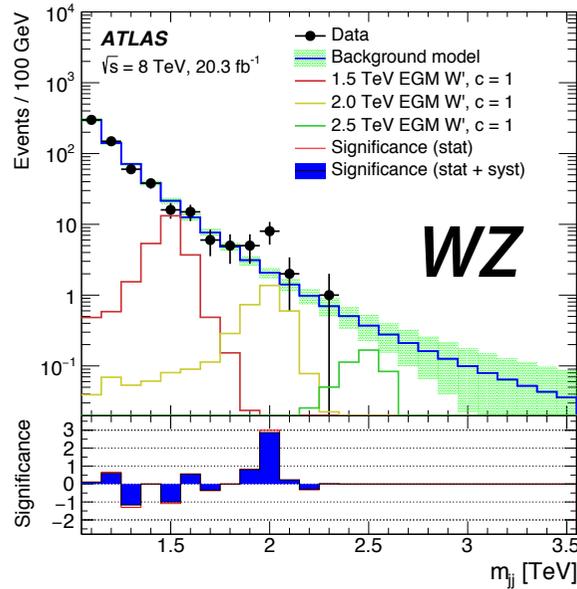
Diboson Resonance as a Portal to Hidden Strong Dynamics

with C.W. Chiang (TNCU), H. Fukuda (IPMU), K. Harigaya(UC Berkeley), T.T.Yanagida (IPMU)

JHEP 1511 (2015) 015 (arXiv:1507.02483)

Diboson Excesses !

[ATLAS arXiv:1506.00962]

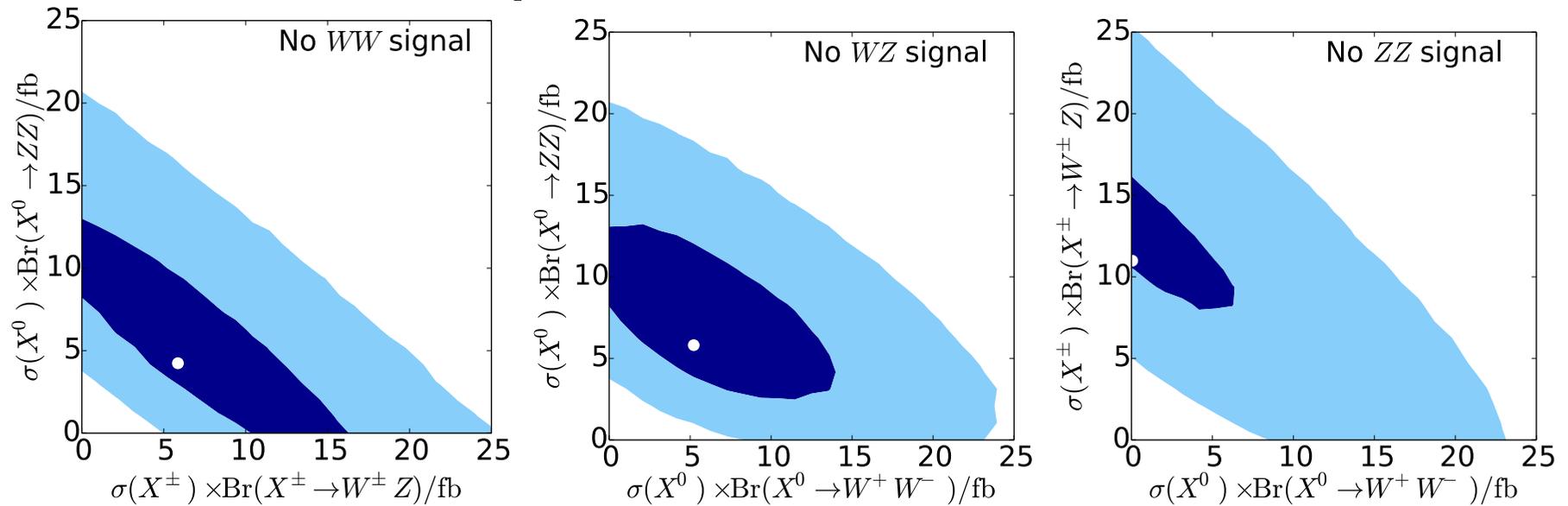


$pp \rightarrow X \rightarrow ZZ/WW/WZ \rightarrow \text{Two Fat Jets} \quad (|m_J - m_V| < 13 \text{ GeV})$

- ✓ local significances : $ZZ(2.6\sigma)/WW(2.9\sigma)/WZ(3.4\sigma)$
- ✓ $m_X \sim 2 \text{ TeV}$
- ✓ $\Gamma_X \lesssim 100 \text{ GeV}$ (narrow width)

Who is X ?

B.C. Allanach, B. Gripaios and D. Sutherland, [arXiv:1507.01638]



[preferred region : dark blue 70%CL, light blue 95%CL]

✓ ***WZ signals are not well separated from ZZ, WW***

For example, the signal can be explained without **WZ** channel by

$$\sigma_{WW} + \sigma_{ZZ} \sim 5 - 15 \text{ fb}$$

✓ **No excesses in leptonic modes @ 2TeV...**

(CMS 1405.3447, ATLAS 1503.04677)

$$\sigma_{WZ} (WZ \rightarrow lv + jets) \lesssim 10 \text{ fb}$$

$$\sigma_{WW} (WW \rightarrow lepton + jets) \lesssim 3-5 \text{ fb}$$

$$\sigma_{ZZ} (ZZ \rightarrow lepton + jets) \lesssim 10 \text{ fb}$$

[No BR's of Z,W are multiplied]

✓ **No excesses in dijet modes @ CMS (CMS 1405.1994)**

$$\sigma_{ZZ,WW,WZ} (ZZ,WW,WZ \rightarrow dijet) \lesssim 10 \text{ fb}$$

✓ These constraints do not immediately conflict with the ATLAS dijet excesses.

→ We need to wait for Run 2 results !

Spin-0 from Hidden Strong Dynamics

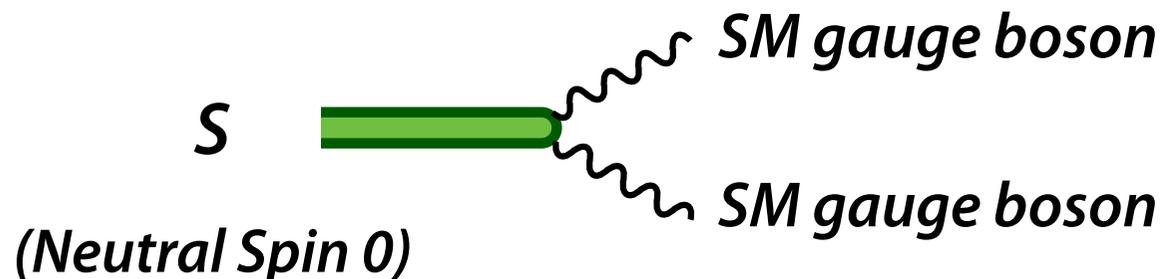
If the *spin-0* boson is a composite particle resulting from strong dynamics *at the TeV scale*, the production cross section can be sizable !

Effective theory (*S* : composite *spin-0* neutral boson)

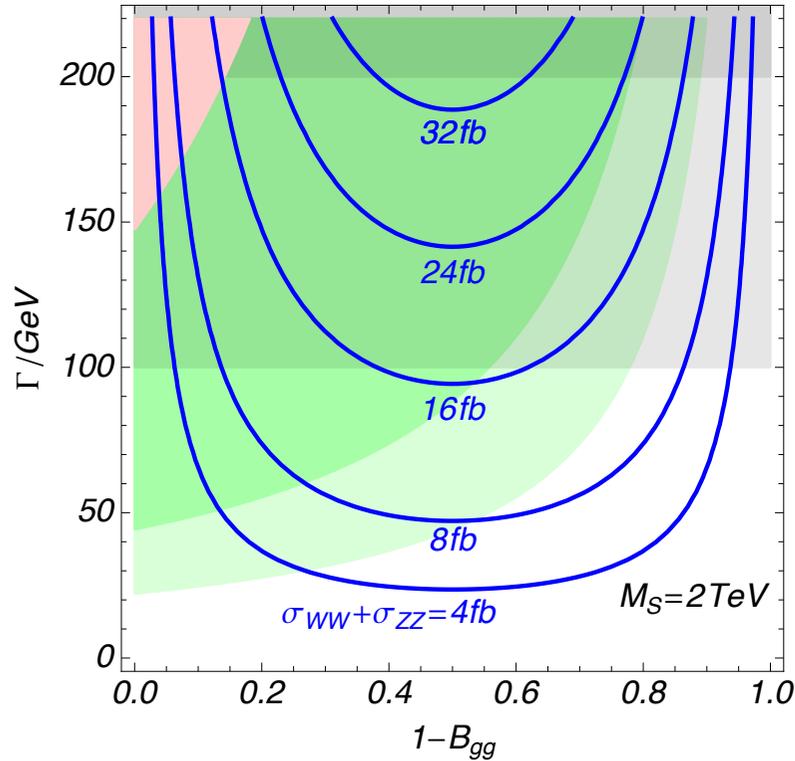
$$\mathcal{L}_{\text{eff}} = \frac{\kappa_3}{\Lambda} S G_{\mu\nu}^a G^{a\mu\nu} + \frac{\kappa_2}{\Lambda} S W_{\mu\nu}^i W^{i\mu\nu} + \frac{5}{3} \frac{\kappa_1}{\Lambda} S B_{\mu\nu} B^{\mu\nu}$$

$$\text{(Normalization : } \mathcal{L} = -\frac{1}{4g_s^2} G_{\mu\nu}^a G^{a\mu\nu} - \frac{1}{4g^2} W_{\mu\nu}^i W^{i\mu\nu} - \frac{1}{4g'^2} B_{\mu\nu} B^{\mu\nu} \text{)}$$

($\Lambda = O(1)\text{TeV}$: dynamical scale)



Predicted cross section (8TeV)



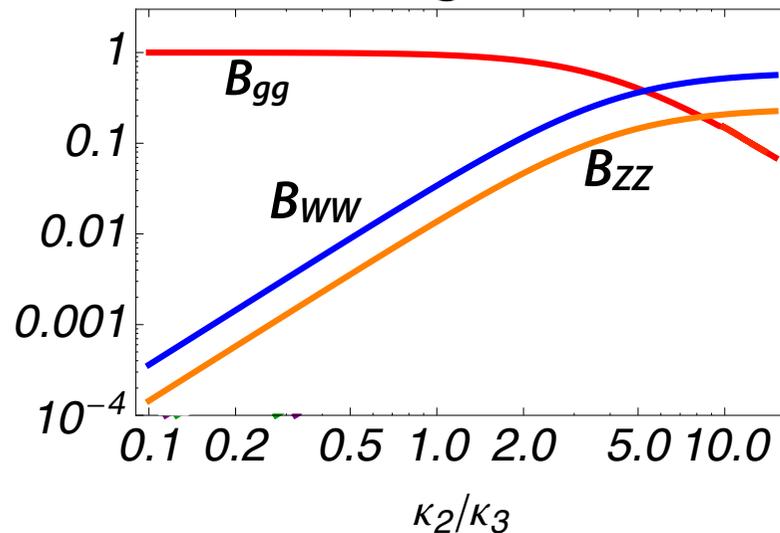
- ✓ **Gray shaded region $\Gamma_S > 100$ GeV**
- ✓ **(light) Green shaded region $\sigma_{\gamma\gamma} > 0.3$ fb for $B_{\gamma\gamma} = 1\%$ (2%) [ATLAS, 1504.05511]**
- ✓ **Pink shaded region $\sigma_{gg} > 100$ fb [ATLAS, 1407.1376, CMS, 1501.04198]**

- ✓ **Allowed region with $\sigma_{WW+ZZ} \sim 5 - 10$ fb**

$$\Gamma_S = O(10)\text{GeV}$$

$$B_{WW} + B_{ZZ} \gtrsim 0.4$$

Branching Fractions



- ✓ **$B_{WW} + B_{ZZ} > 0.4$ is achieved for $\kappa_2/\kappa_3 > 2 - 3$.**

$$\left(\mathcal{L}_{\text{eff}} = \frac{\kappa_3}{\Lambda} S G_{\mu\nu}^a G^{a\mu\nu} + \frac{\kappa_2}{\Lambda} S W_{\mu\nu}^i W^{i\mu\nu} \right)$$

Model of Hidden Dynamics

- ✓ Gauge Group : $SU(N)_H \times SM$ Gauge group
- ✓ Matter field : 5 scalars in fundamental representations of $SU(N)_H$

	$SU(N)_H$	$SU(3)_C$	$SU(2)_L$	$U(1)_Y$
Q_L	N	1	2	$1/2$
Q_D	N	3	1	$1/3$

- ✓ 5 scalars are in $(3,1)_{1/3}, (1,2)_{1/2}$ reps. of $SU(3)_C \times SU(2)_L \times U(1)_Y$
- ✓ Explicit Mass term of $O(1)$ TeV ($m_L^2 < m_D^2$)

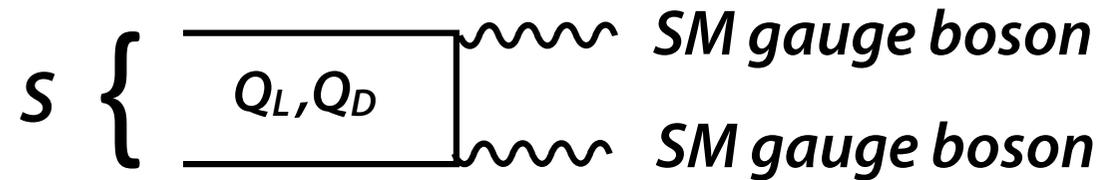
$$\mathcal{L} \supset -m_D^2 Q_D^\dagger Q_D - m_L^2 Q_L^\dagger Q_L$$

- ✓ Dynamical Scale : $\Lambda_{dyn} = O(1) \text{TeV}$

✓ By taking $m_{D,L} \approx \Lambda_{\text{dyn}}$, we assume that $Q_L^\dagger Q_L$ and $Q_D^\dagger Q_D$ dominate S .

$$S \propto \cos \theta_Q \times [Q_L^\dagger Q_L] + \sin \theta_Q \times [Q_D^\dagger Q_D]$$

(θ_Q decreases for $m_D \gg m_L$)



✓ Effective Lagrangian is given by

$$\mathcal{L}_{\text{eff}} = \frac{\kappa_3}{\Lambda} S G_{\mu\nu}^a G^{a\mu\nu} + \frac{\kappa_2}{\Lambda} S W_{\mu\nu}^i W^{i\mu\nu} + \frac{5}{3} \frac{\kappa_1}{\Lambda} S B_{\mu\nu} B^{\mu\nu}$$

$$\frac{\kappa_3}{\Lambda} = \frac{\kappa \sin \theta_Q}{4\pi \Lambda_{\text{dyn}}}, \quad \frac{\kappa_2}{\Lambda} = \frac{\kappa \cos \theta_Q}{4\pi \Lambda_{\text{dyn}}}, \quad \frac{\kappa_1}{\Lambda} = \frac{\kappa}{4\pi \Lambda_{\text{dyn}}} \frac{6}{5} \left(\frac{\sin \theta_Q}{3} + \frac{\cos \theta_Q}{2} \right)$$

[Naive dimensional analysis (NDA) ('97 Luty, '97 Cohen, Kaplan, Nelson)]

$\Lambda/\kappa = O(1)\text{TeV}$ is possible within uncertainties !

$B_{WW} + B_{ZZ} > 0.4$ is possible for a small θ_Q ($m_D \gg m_L$) !

[If we do not use, NDA, we don't have 4π suppressions.]

Dark Matter Candidate ?

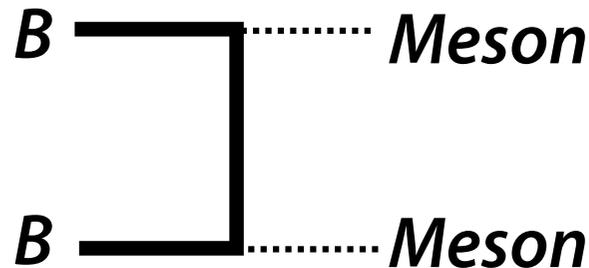
- ✓ Hidden sector so far has a global $U(1)$ symmetry.
- ✓ The lightest $U(1)$ charged composite field = **Baryon** !

For $N = 5$, the lightest baryon is a neutral scalar !

$$B \propto QQQQQ$$

→ Dark matter candidate !

Abundance ?



- ✓ If $\sigma_{ann} \sim 4\pi/m_B^2$ (unitarity limit)
 $\Omega h^2 \sim 10^{-3} (10\text{TeV}/m_B)^2$

- ✓ However, since $m_B > \Lambda_{dyn}$, σ_{ann} can be suppressed by form factors $F(m_B)^4$:

$$\Omega h^2 \propto F(m_B)^{-4}$$

The thermal relic density of **B** can be consistent with the observed relic density !

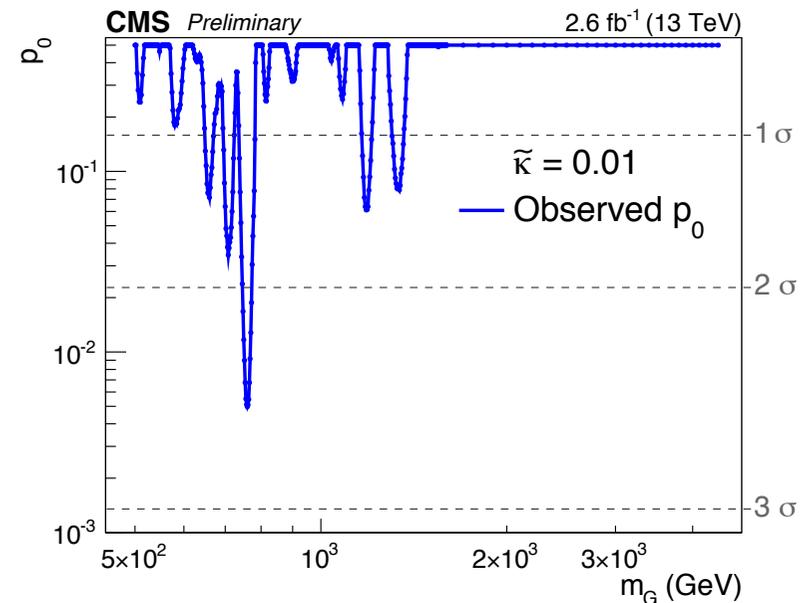
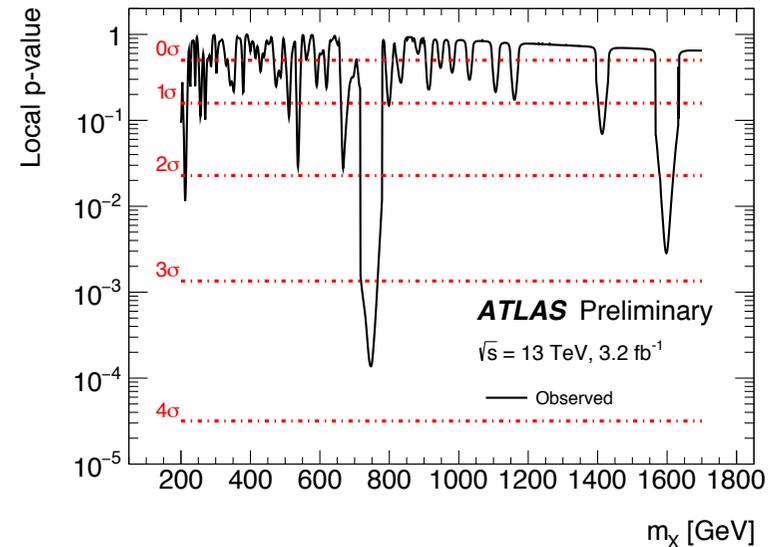
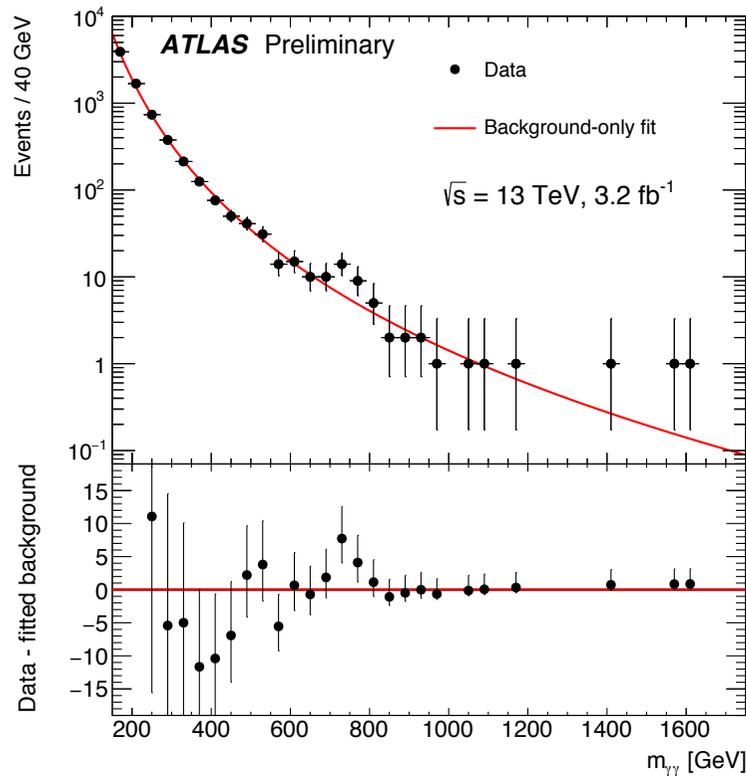
(Detection is very difficult ... heavy and suppressed interactions)

Summary

- ✓ ATLAS reported excesses in diboson resonance search.
(No excesses in leptonic modes nor dijet modes at CMS)
- ✓ Standard Model *neutral Spin-0* resonance requires dynamical scale in the TeV range.
- ✓ Simple dynamical model can provide the desired *spin-0* particle for $m_{L,D} < \Lambda_{dyn}$.
- ✓ The model predicts a lot of new charged particles in the TeV region!
- ✓ Dark Matter can be provided as a lightest neutral baryon in the dynamical sector.

The diboson resonance can be a portal to a hidden strong dynamics !

News!  Apart from the 2TeV excess, **both** ATLAS and CMS reported an excess at 750GeV in di-photon search (2015/12/15)!



We could have some dynamics within a TeV range?

(Our model can be tuned to explain this 750GeV signal.)

Backup Slides

Partial Decay Widths

$$\Gamma(S \rightarrow g + g) = \frac{2}{\pi} \left(\frac{g_s^2 \kappa_3}{\Lambda} \right)^2 M_S^3$$

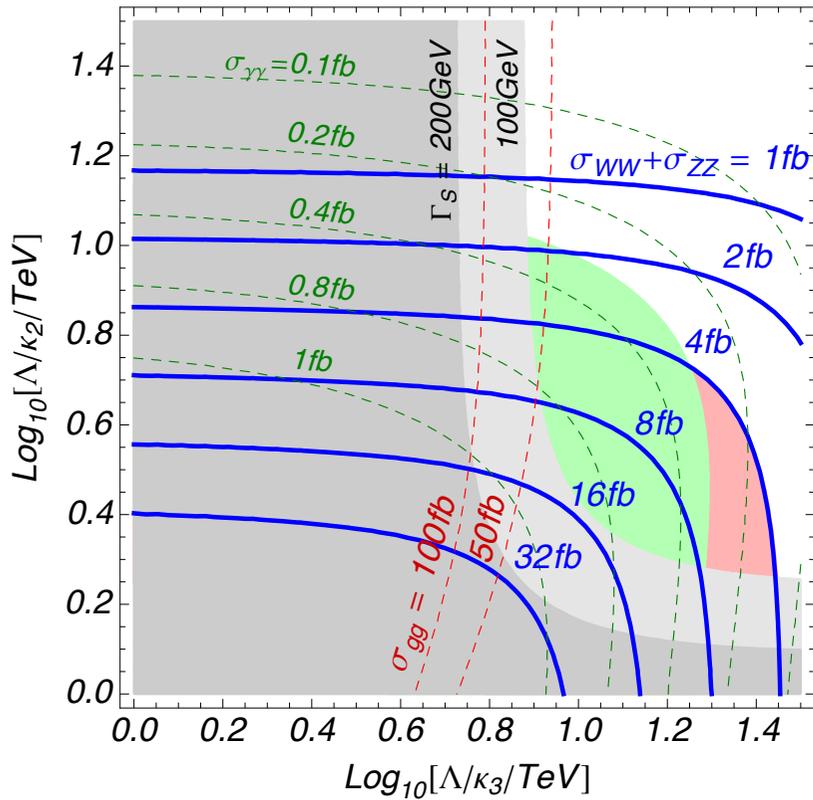
$$\Gamma(S \rightarrow W^+ + W^-) = \frac{1}{2} \frac{1}{\pi} \left(\frac{g^2 \kappa_2}{\Lambda} \right)^2 M_S^3 ,$$

$$\Gamma(S \rightarrow Z + Z) = \frac{1}{4} \frac{1}{\pi} \left[\left(\frac{g^2 \kappa_2}{\Lambda} \right) c_W^2 + \frac{3}{5} \left(\frac{g'^2 \kappa_1}{\Lambda} \right) s_W^2 \right]^2 M_S^3 ,$$

$$\Gamma(S \rightarrow \gamma + \gamma) = \frac{1}{4} \frac{1}{\pi} \left[\left(\frac{g^2 \kappa_2}{\Lambda} \right) s_W^2 + \frac{3}{5} \left(\frac{g'^2 \kappa_1}{\Lambda} \right) c_W^2 \right]^2 M_S^3 ,$$

$$\Gamma(S \rightarrow Z + \gamma) = \frac{1}{2} \frac{1}{\pi} \left[\left(\frac{g^2 \kappa_2}{\Lambda} \right) - \frac{3}{5} \left(\frac{g'^2 \kappa_1}{\Lambda} \right) \right]^2 c_W^2 s_W^2 M_S^3 ,$$

Cross section in terms of the model parameters



- ✓ Gray shaded region $\Gamma_s > 100 \text{ GeV}$
- ✓ (light) Green shaded region $\sigma_{\gamma\gamma} > 0.3 \text{ fb}$ for $B_{\gamma\gamma} = 1\%$ (2%) [ATLAS, 1504.05511]
- ✓ Pink shaded region $\sigma_{gg} > 100 \text{ fb}$

✓ $\sigma_{WW} + \sigma_{ZZ} \sim 5 - 10 \text{ fb}$ is allowed $\Lambda/\kappa \sim 1 - 1.4 \text{ TeV}$

Naive Dimensional Counting ('97 Luty, '97 Cohen, Kaplan, Nelson)

$$S \propto \cos \theta_Q \times [Q_L^\dagger Q_L] + \sin \theta_Q \times [Q_D^\dagger Q_D]$$

(θ_Q decreases for $m_D \gg m_L$)

Operator Matching

$$S \simeq \frac{4\pi}{\kappa\Lambda_{\text{dyn}}} \cos \theta_Q \times [Q_L^\dagger Q_L] + \frac{4\pi}{\kappa\Lambda_{\text{dyn}}} \sin \theta_Q \times [Q_D^\dagger Q_D]$$

Effective Lagrangian is given by

$$\mathcal{L}_{\text{eff}} = \frac{\kappa_3}{\Lambda} S G_{\mu\nu}^a G^{a\mu\nu} + \frac{\kappa_2}{\Lambda} S W_{\mu\nu}^i W^{i\mu\nu} + \frac{5}{3} \frac{\kappa_1}{\Lambda} S B_{\mu\nu} B^{\mu\nu}$$
$$\frac{\kappa_3}{\Lambda} = \frac{\kappa \sin \theta_Q}{4\pi\Lambda_{\text{dyn}}}, \quad \frac{\kappa_2}{\Lambda} = \frac{\kappa \cos \theta_Q}{4\pi\Lambda_{\text{dyn}}}, \quad \frac{\kappa_1}{\Lambda} = \frac{\kappa}{4\pi\Lambda_{\text{dyn}}} \frac{6}{5} \left(\frac{\sin \theta_Q}{3} + \frac{\cos \theta_Q}{2} \right)$$

S is accompanied by charged spin-0 particles

$$Q_{L,D}^\dagger Q_{L,D} \sim (1,1)_0 \times 2, (1,3)_0, (8,1)_0, (3,2)_{5/6} \text{ [25 states]}$$

- ✓ Heavier singlets : $Q_D^\dagger Q_D$ dominated but decaying to S .
- ✓ Octet : decaying to two gluons [$\sigma_{gg} < 30\text{fb}$ for $M_8 > 3\text{TeV}$]
- ✓ Triplet : decaying to HH, WZ via $L = (H^\dagger \sigma^i H)(Q_L^\dagger \sigma^i Q_L)$
(produced via Drell-Yan process at the LHC)
- ✓ $(3,2)_{5/6}$: we need new particle to make $(3,2)_{5/6}$ decay

$$(3,2)_{5/6} \left\{ \begin{array}{l} Q_L \\ Q_D \end{array} \right. \begin{array}{c} \text{---} \\ \text{---} \\ \text{---} \end{array} \begin{array}{c} \diagup \\ \psi \\ \diagdown \end{array} \begin{array}{l} \text{Lepton} \\ \text{quark} \end{array} \quad \mathcal{L} \supset y Q_D^\dagger \psi_Q \bar{d}_R + y Q_L^\dagger \psi_Q \ell_L + M \psi \bar{\psi}$$

(ψ : $SU(N)_H$ fundamental fermion)

Higher spin modes are heavier and decay immediately .

Electroweak Precision Constraints 1

The decay operator

$$L = (H^\dagger \sigma^i H)(Q_L^\dagger \sigma^i Q_L)$$

leads to a tadpole to the triplet scalar

$$\mathcal{L} \simeq \frac{\lambda}{4\pi} \Lambda_{\text{dyn}} H^\dagger \sigma^i H T^i,$$

Thus, the triplet obtains a small VEV,

$$\langle T^3 \rangle \simeq \frac{\lambda v_{\text{EW}}^2 \Lambda_{\text{dyn}}}{4\pi M_T^2} = 0.6 \text{ GeV} \times \lambda \frac{\Lambda_{\text{dyn}}}{1 \text{ TeV}} \left(\frac{M_T}{2 \text{ TeV}} \right)^{-2}$$

whose contribution to T -parameter is small enough.

Electroweak Precision Constraints 2

The decay operator

$$L = (H^\dagger \sigma^i H)(Q_L^\dagger \sigma^i Q_L)$$

also induces

$$\frac{g^2}{16\pi^2} \frac{1}{m_L^2} (H^\dagger W_{\mu\nu}^i \sigma^i H) B^{\mu\nu}$$

which contributes to **S**-parameter.

Again, it is very small.

Electroweak Precision Constraints 3

Spin-1 composite fields may have kinetic mixing to the SM gauge boson

e.g. $\epsilon F'_{\mu\nu} F^{\mu\nu}$

After removing the kinetic mixing and integrating out F' , we end up with

$$\frac{\epsilon^2}{M_F^2} (H^\dagger D H)^2 \quad \frac{\epsilon^2}{M_F^2} J J$$

Both contributes to the electroweak precision measurements but they are small enough for $M_F > \mathbf{a\ few\ TeV}$.