

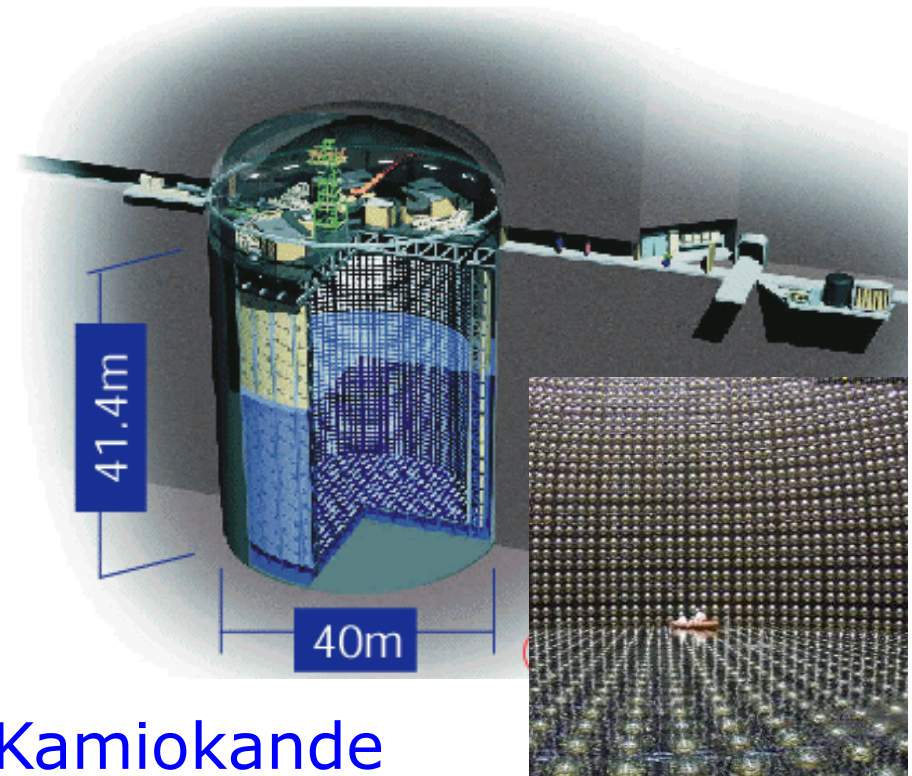
# スーパーカミオカンデにおける 最新結果 (大気ニュートリノ、核子崩壊探索等)

東京大学宇宙線研究所  
神岡宇宙素粒子研究施設  
三浦 真

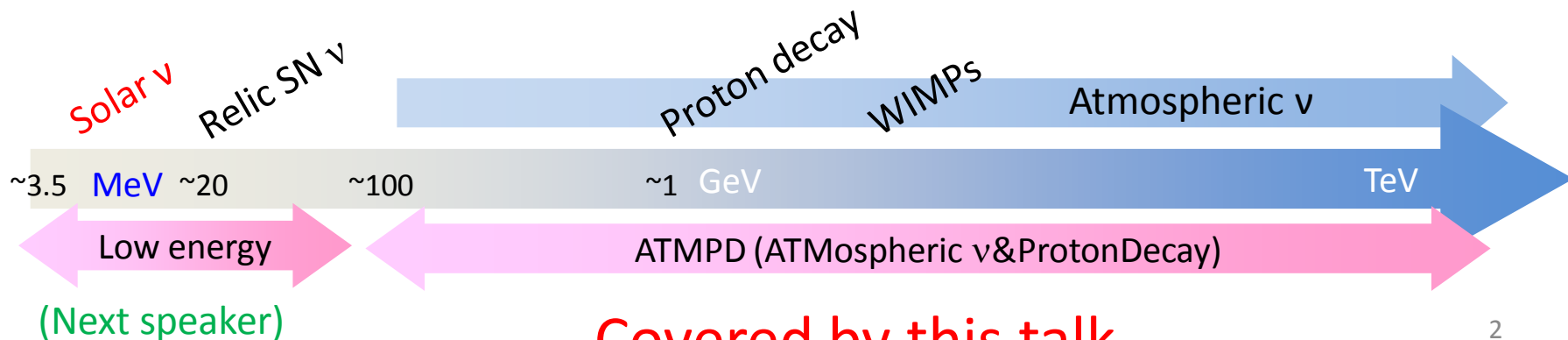
2015/12/19 共同利用研究成果発表会

# 1. Introduction: Super-Kamiokande

- 50kton pure water Cherenkov detector
- 1km (2.7km w.e) underground in Kamioka
- 11129 50cm PMTs  
in Inner Detector
- 1885 20cm PMTs  
in Outer Detector



## Physics targets of Super-Kamiokande



Covered by this talk

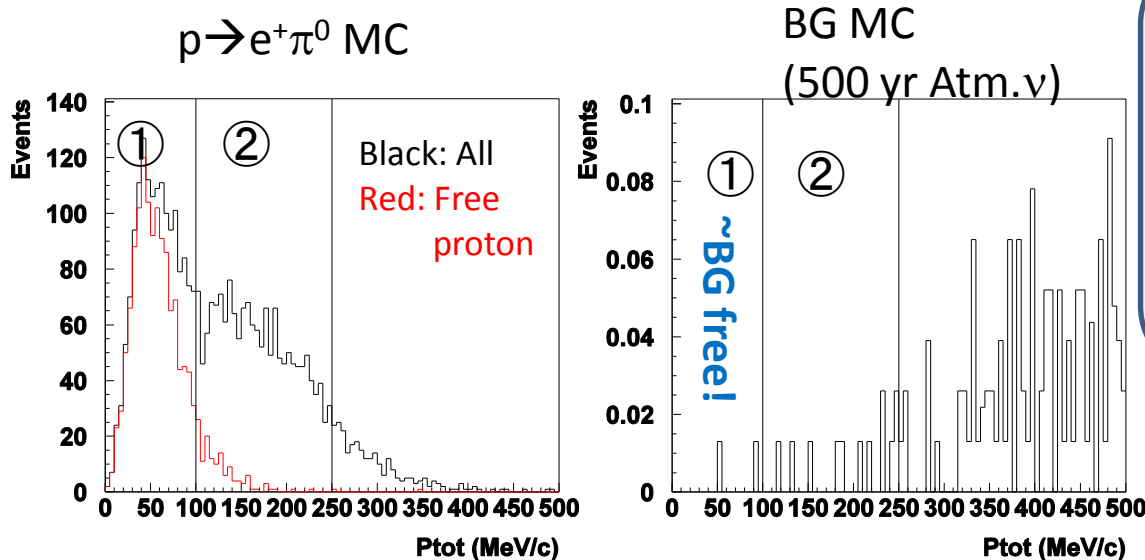
## 2. Publications in this year (ATMPD)

- Test of Lorentz invariance with atmospheric neutrinos: Phys. Rev. D 91, 052003 (2015)
- Limits on sterile neutrino mixing using atmospheric neutrinos in Super-Kamiokande: Phys. Rev. D91, 052019 (2015)
- Search for Neutrinos from Annihilation of Captured Low-Mass Dark Matter Particles in the Sun by Super-Kamiokande: Phys. Rev. Lett. 114, 141301 (2015)
- Search for  $n$ - $\bar{n}$  oscillation in Super-Kamiokande: Phys. Rev. D91, 072006 (2015)
- Search for dinucleon decay into pions at Super-Kamiokande: Phys. Rev. D91, 072009 (2015)
- Search for Nucleon and DiNucleon Decays with an Invisible Particle and a Charged Lepton in the Final State at the Super-Kamiokande Experiment: Phys. Rev. Lett. 115, 121803 (2015)

2 neutrino oscillation papers,  
1 astrophysical paper,  
3 nucleon decay papers.

# 3. Nucleon decay analysis

- Update  $p \rightarrow e^+\pi^0$ ,  $\mu^+\pi^0$ , and  $\nu K^+$  by data until March, 2015 (total exposure: **306 kt·year**).
- Neutron tagging is applied for SK4 analysis. Backgrounds are reduced by almost half (reported in the previous year).
- In  $p \rightarrow e^+\pi^0$ ,  $\mu^+\pi^0$ , signal region is divided into two parts;
  - ①  $P_{tot} < 100$  MeV/c: **Free proton** dominant, almost **BG free!**
  - ②  $100 < P_{tot} < 250$  MeV/c: Bound proton dominant.



• Expected BG: 0.07 events  
 Obs.in ①  $\geq 2 \rightarrow 3\sigma$   
                    $\geq 3 \rightarrow 5\sigma$   
 • BG increases very slowly:  
 0.003evts/year



**Keep discovery potential!**

# Results with 306kt·year exposure

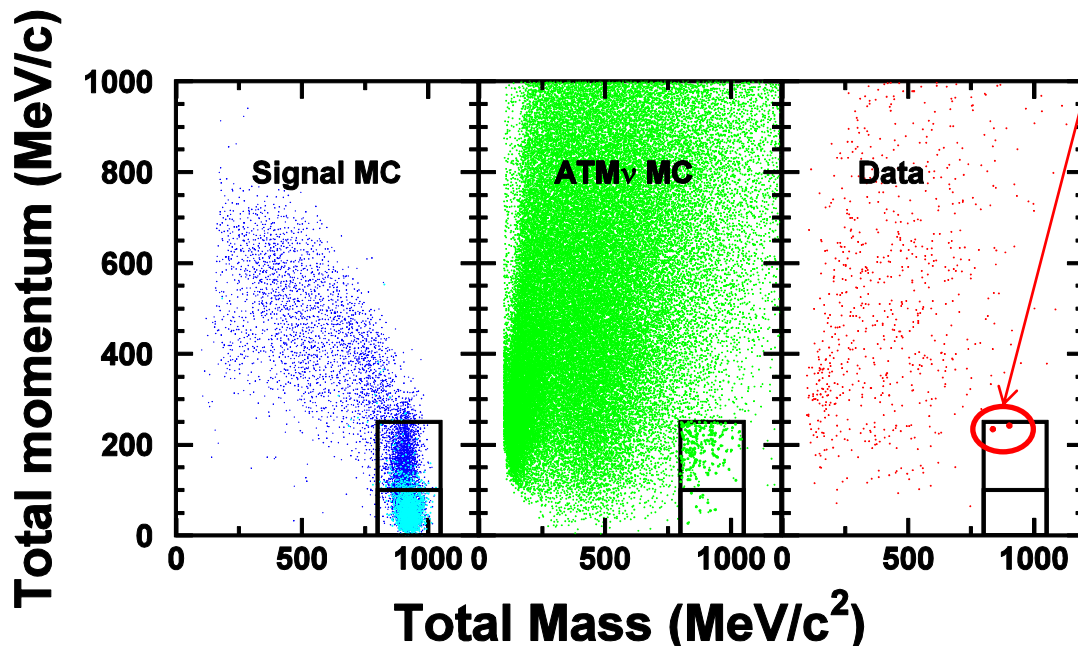
$p \rightarrow e^+ \pi^0$	0~100MeV/c	100~250MeV/c	Total
Observed	0	0	0
BKG	0.07	0.54	0.61

$p \rightarrow \mu^+ \pi^0$	0~100MeV/c	100~250MeV/c	Total
Observed	0	2	2
BKG	0.05	0.82 (SK4:0.23)	0.87

2 events remained  
in  $p \rightarrow \mu \pi^0$ !

Both events are

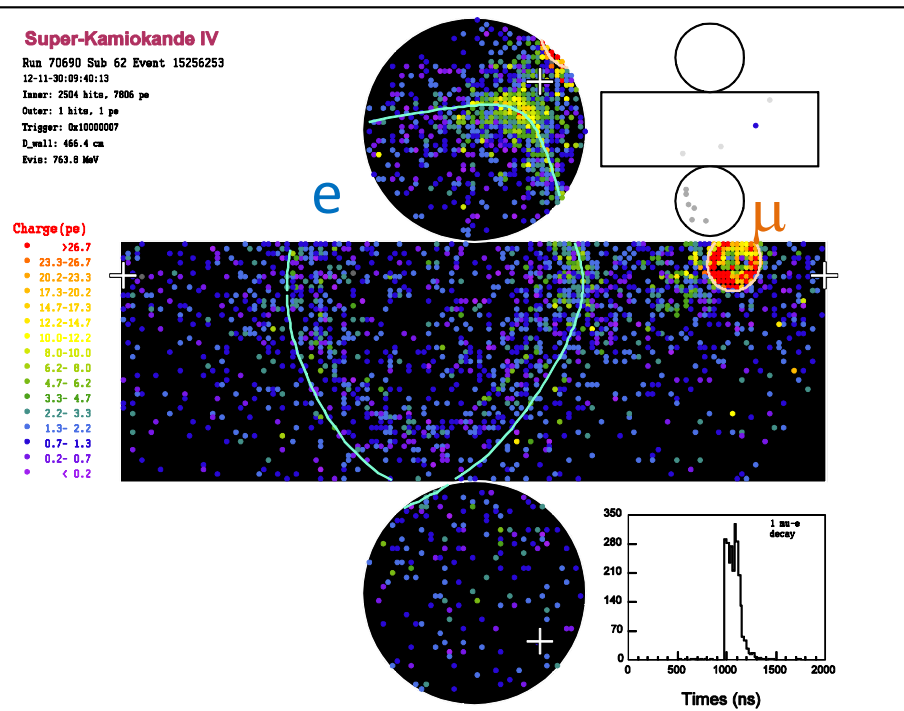
- in SK4 period,
- near by boarder ( $P_{\text{tot}} \sim 250 \text{ MeV/c}$ ),
- 2 ring (e-like+ $\mu$ -like)



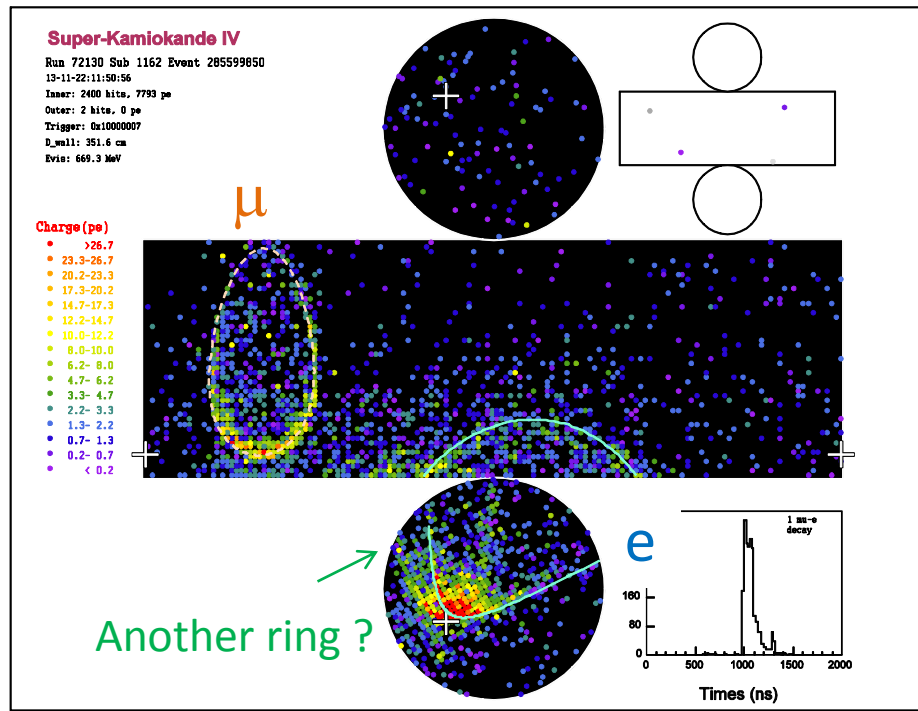
Probability assuming BG;  
 $\text{Poisson}(\geq 2, 0.82) = 20\%$   
 (SK4 only: 2.3 %)

# Observed events

## 1<sup>st</sup> event



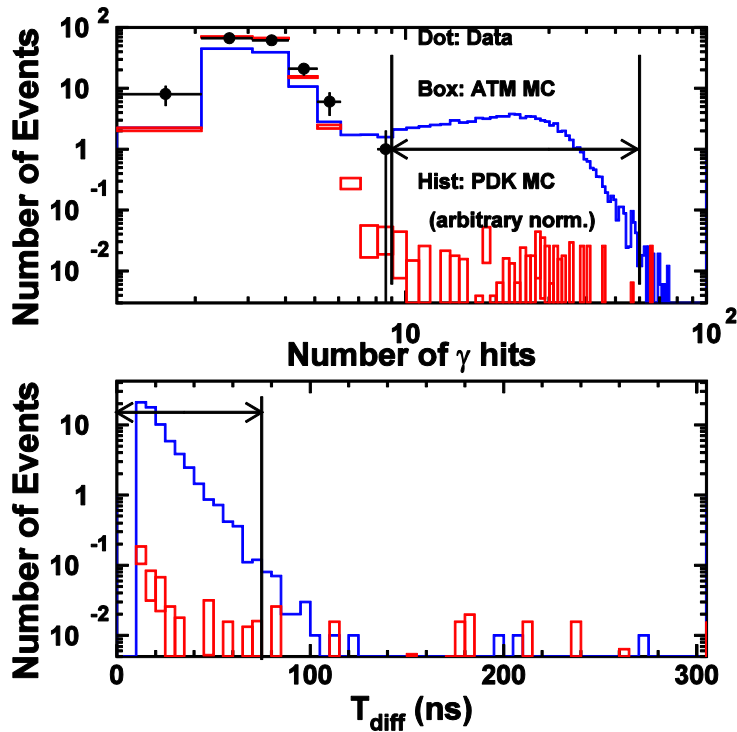
## 2<sup>nd</sup> event



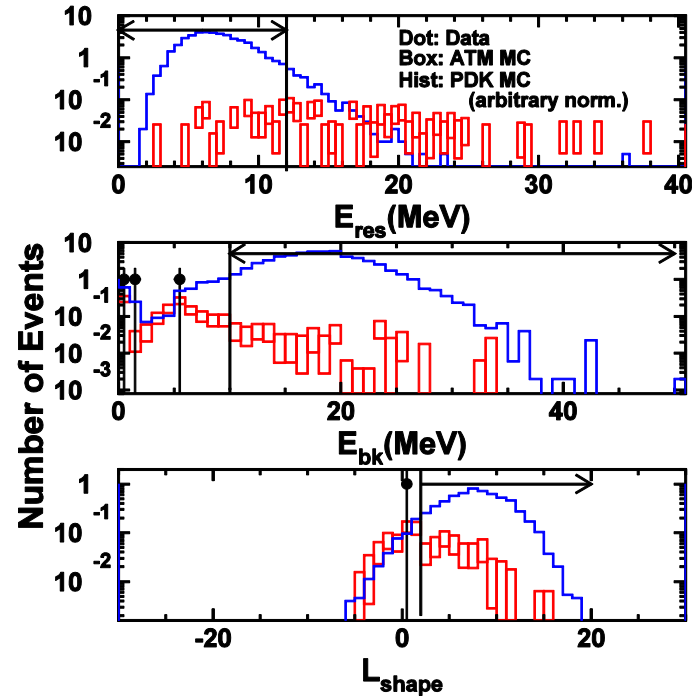
	1 <sup>st</sup> event	2 <sup>nd</sup> event
Mtot (MeV/c <sup>2</sup> )	902.5	832.4
Ptot (MeV/c)	248.0	237.9
Pe (MeV/c)	374.9	460.5
Pμ(MeV/c)	551.1	391.3
Opening ang (deg)	157.9	148.9

# $p \rightarrow \nu K^+$ : No candidates

$K^+ \rightarrow \nu \mu$  with nuclear  $\gamma$



$K^+ \rightarrow \pi^+ \pi^0$



Updated lower limit of proton lifetime

$$p \rightarrow e^+ \pi^0: > 1.7 \times 10^{34} \text{ years}$$

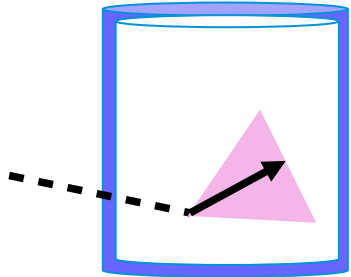
$$p \rightarrow \mu^+ \pi^0: > 7.8 \times 10^{33} \text{ years}$$

$$p \rightarrow \nu K^+: > 6.6 \times 10^{33} \text{ years}$$

# 4. Neutrino oscillation analysis

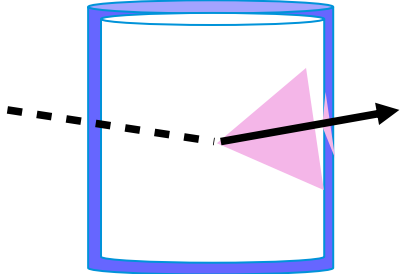
Fully Contained (FC)

~1GeV



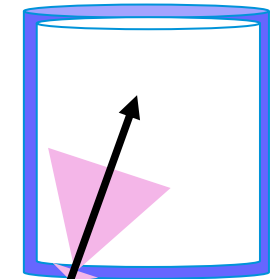
Partially Contained (PC)

~10GeV

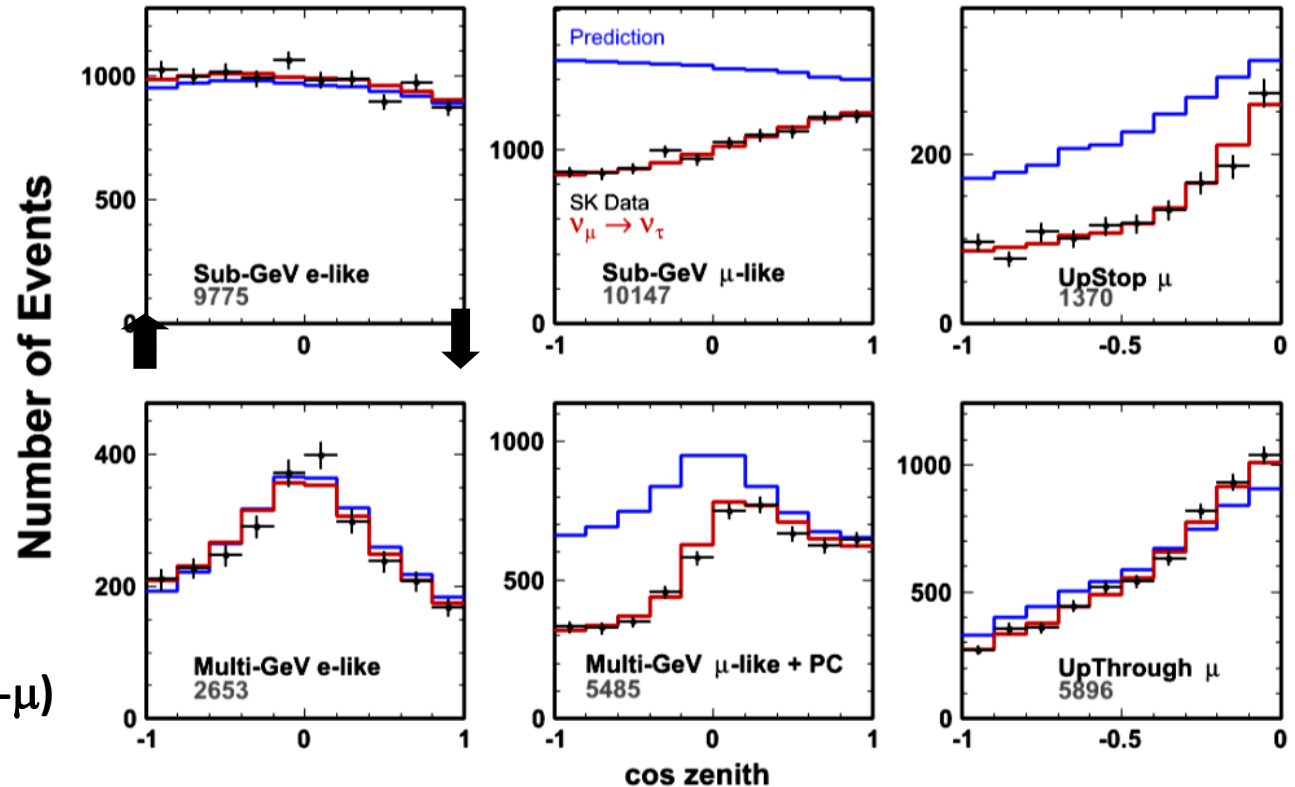


Upward-going Muons (Up- $\mu$ )

~ 100GeV



*Zenith angle distribution of each sample*



- In total 19 analysis samples  
(classified by  $\nu$  flavors, event topologies, energies, ...)
- *Fit to the data in bins of  $\cos\theta_{zenith}$  and momentum*
- Dominated by  $\nu_{\mu} \rightarrow \nu_{\tau}$  oscillations
- Interested in sub-dominant contributions.



# Neutrino oscillation studies using atmospheric $\nu$

High statistics atmospheric neutrino data : FC  $\sim 48k$  ev (98': 4353ev)

$\sim$  Possibility in observing small distortion in  $\nu_e$

- Matter effect  $\sim$  from mass hierarchy

Possible  $\nu_e$  enhancement in several GeV  
passed through the earth core

- Solar term  $\sim$  from  $\theta_{23}$  octant degeneracy

Possible  $\nu_e$  enhancement  
in sub-GeV

- Interference

CP phase could be studied.

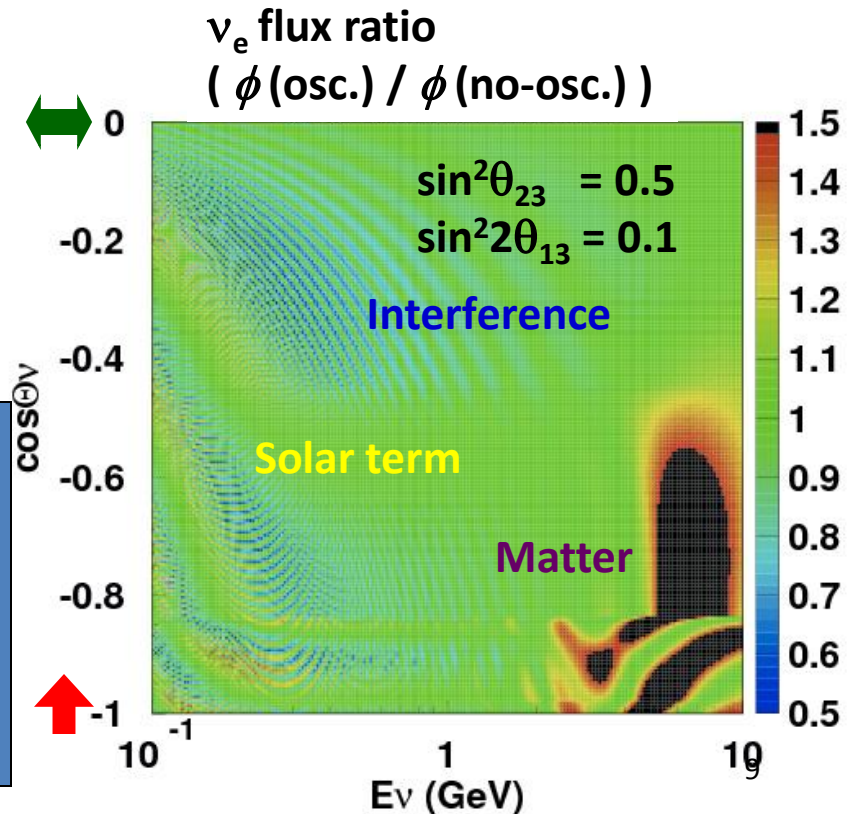
*Difference in # of electron events:*

$$\Delta_e \equiv \frac{N_e}{N_e^0} \cong \Delta_1(\theta_{13}) + \Delta_2(\Delta m_{12}^2) + \Delta_3(\theta_{13}, \Delta m_{12}^2, \delta)$$

← Matter effect

← Solar term

← Interference

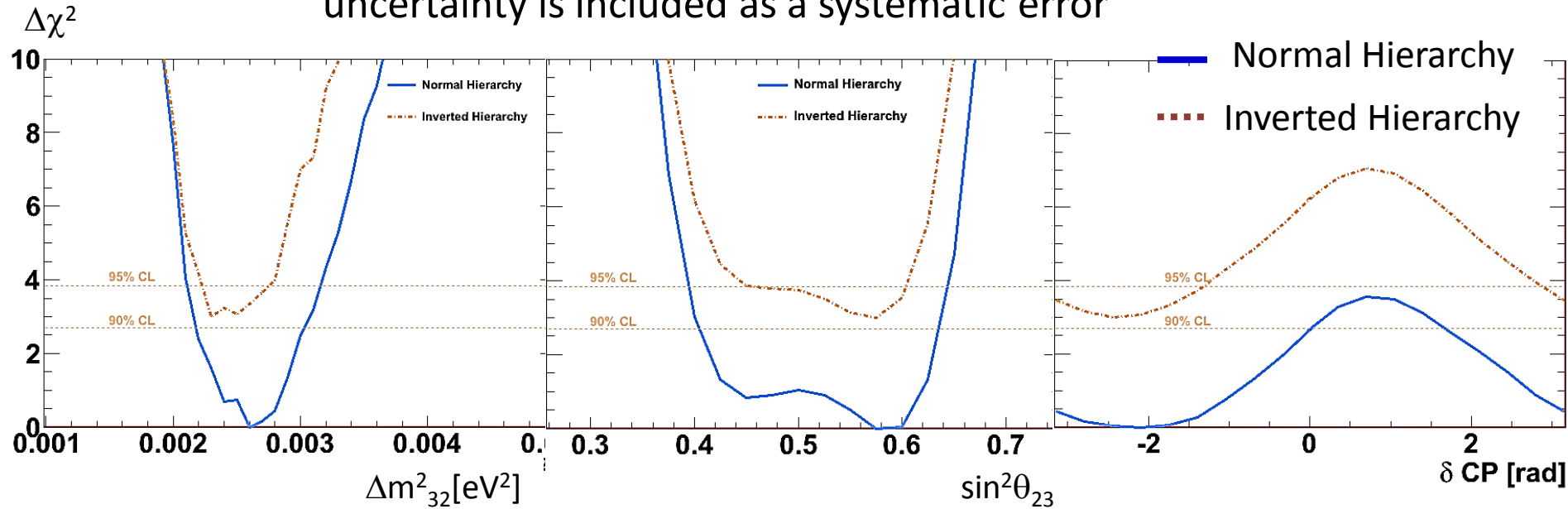


# Neutrino oscillation studies using atmospheric $\nu$

## $\theta_{13}$ Fixed Analysis (NH+IH) SK Only

$\theta_{13}$  fixed to PDG2014 value (  $\sin^2 2\theta_{13} = 0.093 \pm 0.08$  )  
uncertainty is included as a systematic error

*Preliminary*



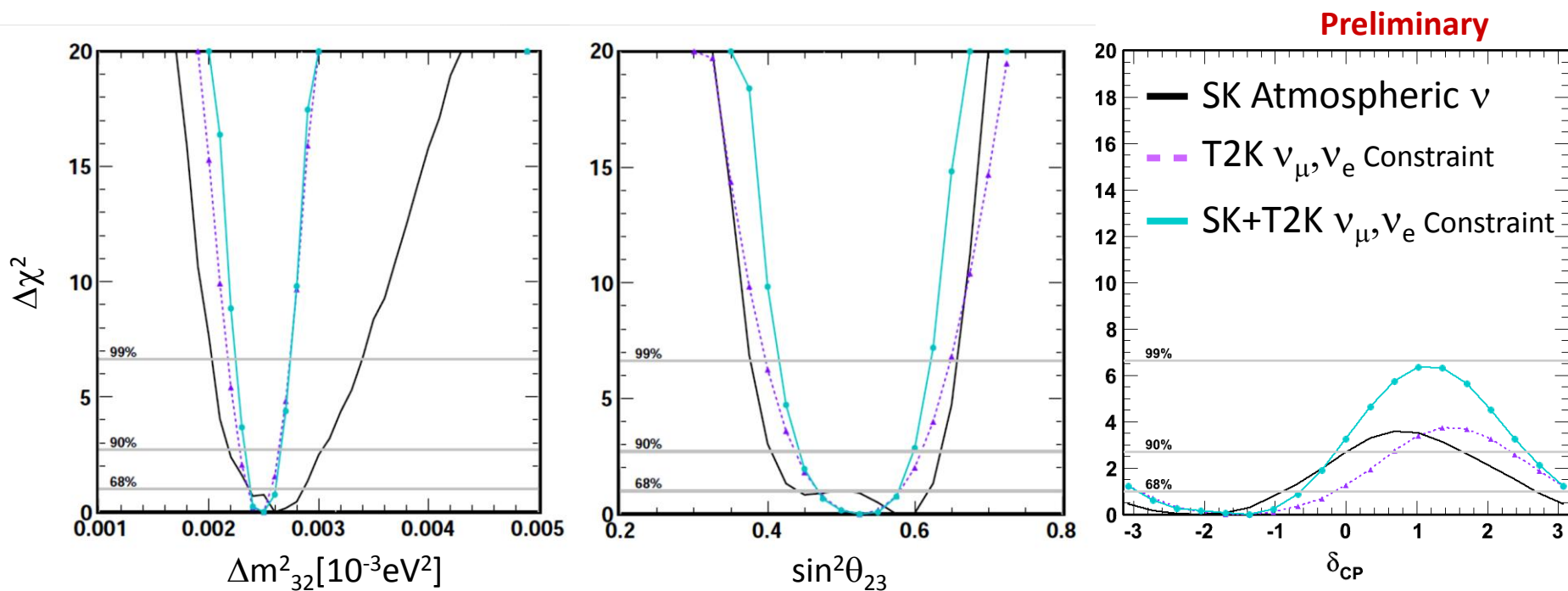
Fit (517 dof)	$\chi^2$	$\delta_{cp}$ (rad)	$\sin^2\theta_{23}$	$\Delta m^2_{32/23}$ ( $\times 10^{-3} eV^2$ )
SK (NH)	<b>582.4</b>	-2.09 (-0.67 $\pi$ )	0.575	2.6
SK (IH)	<b>585.4</b>	-2.44 (-0.78 $\pi$ )	0.575	2.3

Offsets in these curves show the absolute  $\chi^2$  diff. in the hierarchies.

*Normal hierarchy is favored at  $\chi^2_{NH} - \chi^2_{IH} = -3.0$  ( not significant still. )*

# Neutrino oscillation studies using atmospheric $\nu$

$\theta_{13}$  **Fixed** SK + T2K  $\nu_{\mu}, \nu_e$  (External Constraint) Normal Hierarchy



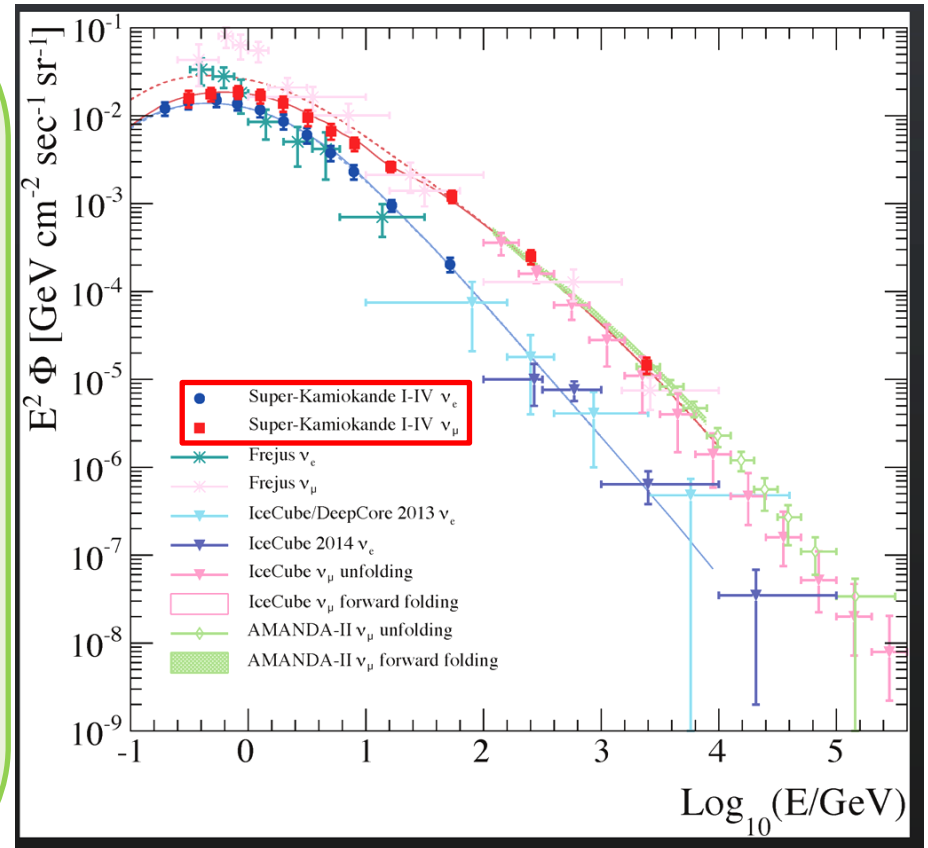
Fit (543 dof)	$\chi^2$	$\theta_{13}$	$\delta_{\text{CP}}$	$\theta_{23}$	$\Delta m^2_{32/23}$ ( $\times 10^{-3} \text{eV}^2$ )
SK + T2K (NH)	651.5	0.0238	-1.39	0.525	2.5
SK + T2K (IH)	654.7	0.0238	-2.09	0.550	2.4

$$\chi^2_{\text{NH}} - \chi^2_{\text{IH}} = \mathbf{-3.2} \text{ (-3.0 SK only)}$$

$\sin\delta_{\text{CP}} = 0$  is still allowed at (at least) 90% C.L. for both hierarchies.

# 5. Atmospheric $\nu$ flux measurement

- Our measurement can be translated to  $\nu$  flux via detector response matrix estimated by MC.
- Our data provide improved precision, extending **up to 100 GeV for  $\nu_e$**  and **10 TeV for  $\nu_\mu$** .
- First data **below 320 MeV**.
- Overlap with km<sup>3</sup> detectors.
- Consistent with HKKM11, Bartol, and Fluka Models.



Paper has been submitted (<http://arxiv.org/abs/1510.08127>).

# 6. Summary

- 6 papers related to ATMPD have been published in 2015.
- Updated major proton decay modes.
  - 2 box analysis introduced to  $p \rightarrow e\pi^0, \mu\pi^0$  to keep high sensitivity and discovery potential even for longer exposure.
  - $p \rightarrow e\pi^0$ : 0 event observed,  $> 1.7 \times 10^{34}$  years
  - $p \rightarrow \mu\pi^0$ : 2 events observed,  $> 7.8 \times 10^{33}$  years
  - $p \rightarrow \nu K^+$ : 0 event observed,  $> 6.6 \times 10^{33}$  years
- Atmospheric  $\nu$  data favors Normal hierarchy slightly ( $\chi^2_{\text{NH}} - \chi^2_{\text{IH}} = -3.0$ , SK only).
- SK provides precise data points of  $\nu$  flux. Consistent with existing models.