

The CTA Project

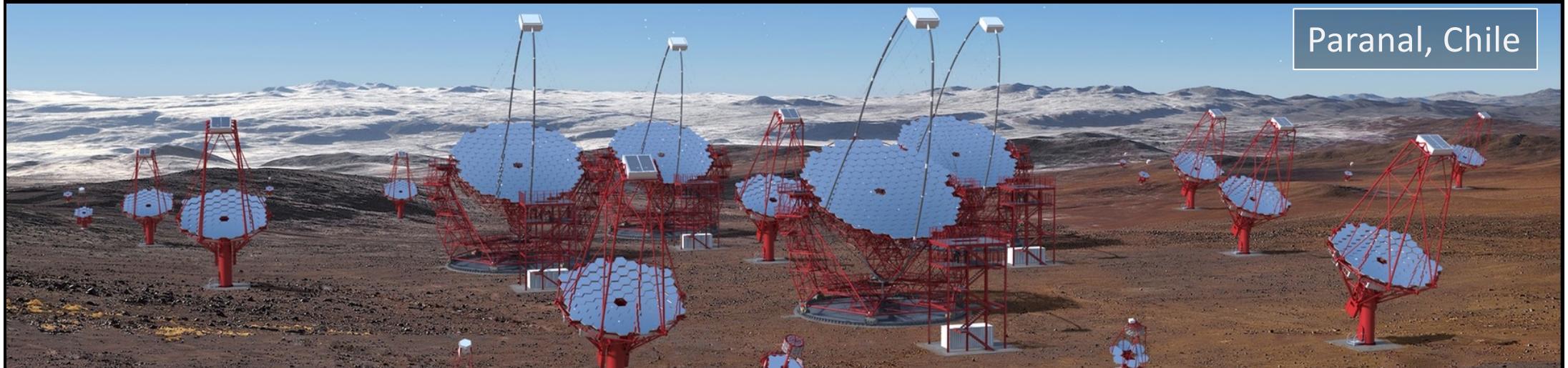
Masahiro Teshima

*for the CTA Japan Consortium
Institute for Cosmic Ray Research, The University of Tokyo*

La Palma, Spain



Paranal, Chile

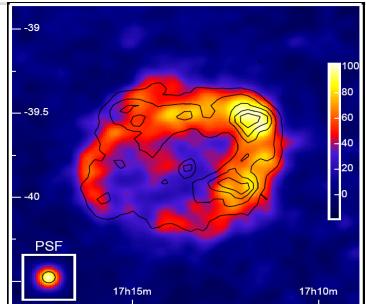


Science of CTA is very wide

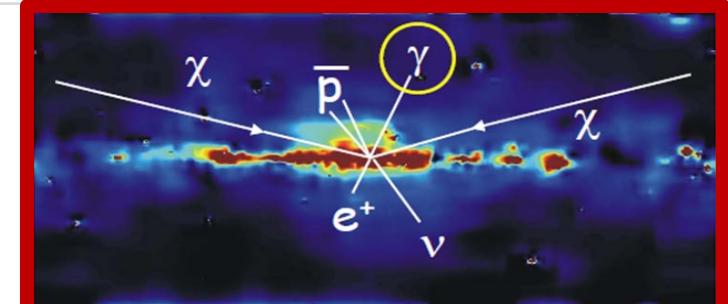
CTA-LST will cover S.M.B.H., Dark Matter, AGNs, GRBs



Cosmic Ray Origin

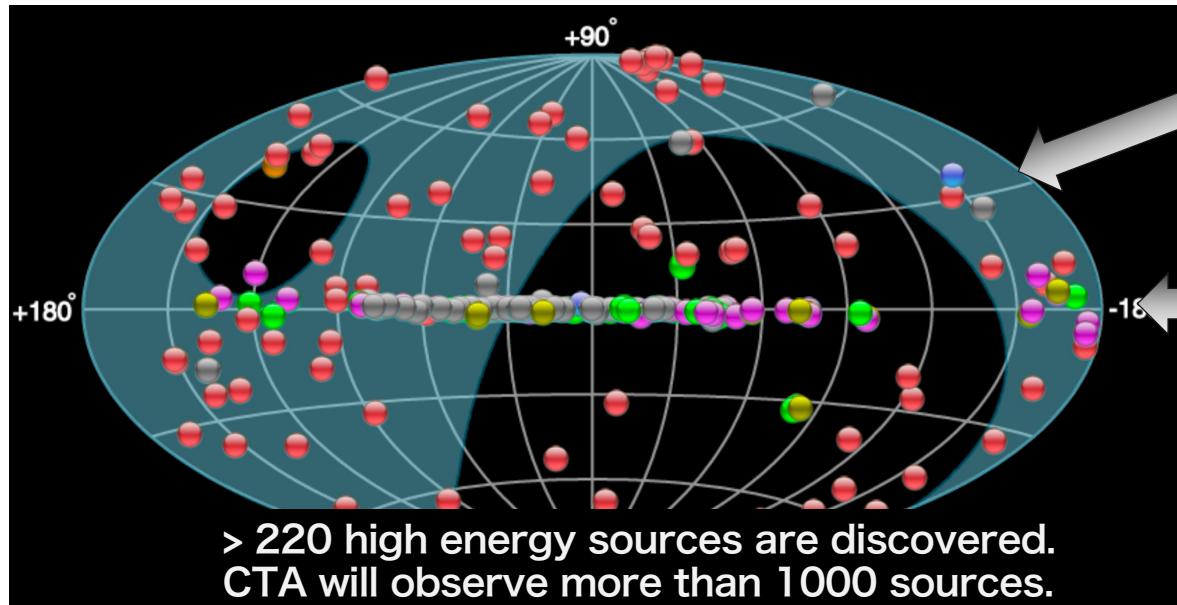


Super Massive
Black Holes



Dark Matter Search (Discovery)

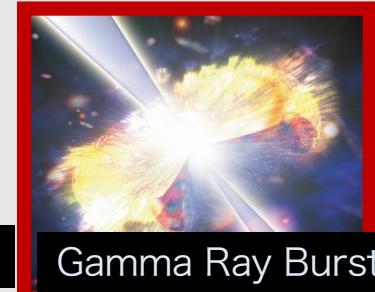
- Origin of Cosmic Rays (Big accelerators)
- Black Hole and S.M.B.H.
- Dark Matter Search



Extragalactic Sources

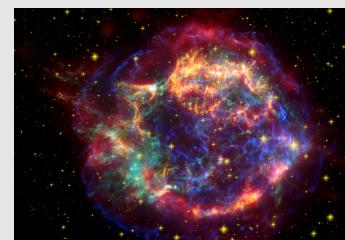


Active Galactic Nuclei

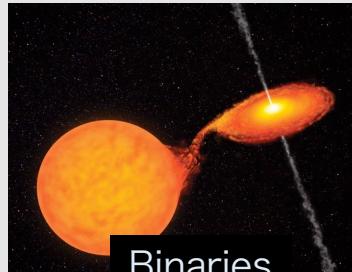


Gamma Ray Bursts

Galactic Sources



Super Nova Remnants



Binaries

共同利用体制

- 宇宙線研究所共同利用の現在
 - 14件の共同利用研究、160名の研究者
- CTAOとしての国際共同利用（公開天文台）
 - 40% Key Science Program (CTAC members)
 - 130-160 members can sign the papers among 1400 (10%)
 - 銀河中心、銀河面サーベイ、AGN、トランジェント、etc.
 - 40% Open Call for Proposals (Country/community)
 - 日本枠は 40% x 10% → 4% (50hrs)
 - 20% Host premium
 - ESO (ESO メンバー国へ還元) 、 IAC(Spain)
- CTA Southへの貢献が日本の共同利用をさらに推進する。

Key Science Programs

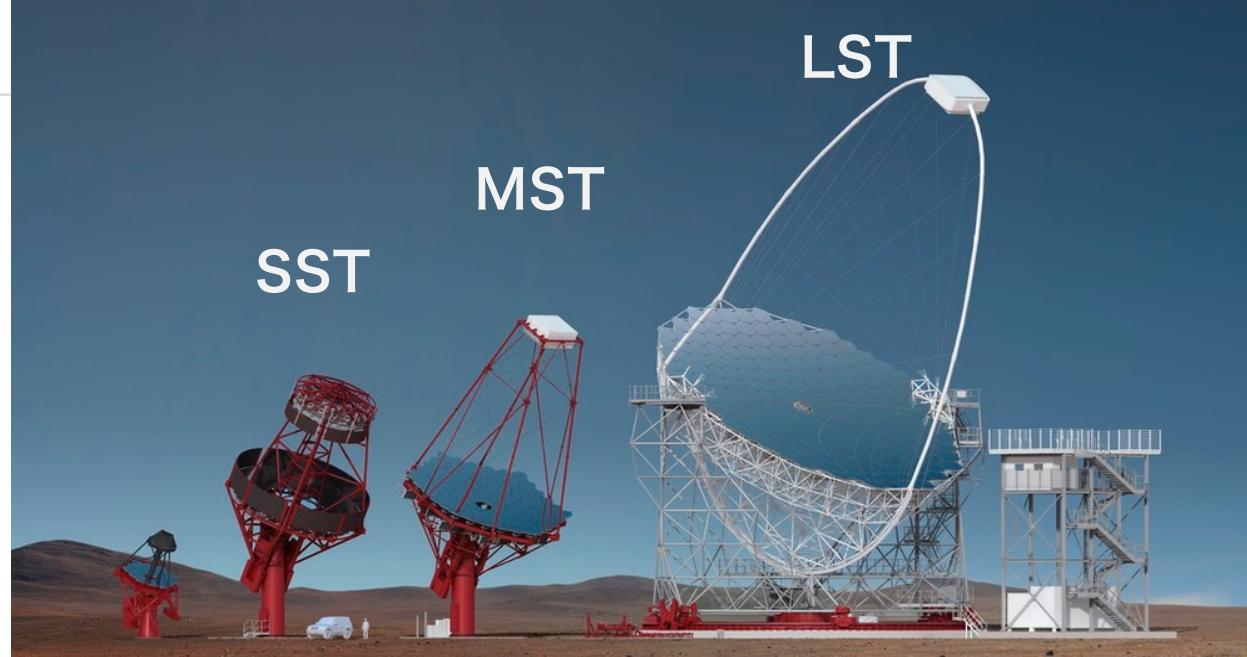
40% obs. Time in the first 10years

Theme	Question	Dark Matter Programme	Galactic Centre Survey	Galactic Plane Survey	LMC Survey	Extra- galactic Survey	Transients	Cosmic Ray PeVatrons	Star-forming Systems	Active Galactic Nuclei	Galaxy Clusters
Understanding the Origin and Role of Relativistic Cosmic Particles	1.1 What are the sites of high-energy particle acceleration in the universe?		✓	vv	vv	vv	vv	✓	✓	✓	vv
	1.2 What are the mechanisms for cosmic particle acceleration?		✓	✓	✓		vv	vv	✓	vv	✓
	1.3 What role do accelerated particles play in feedback on star formation and galaxy evolution?		✓		✓				vv	✓	✓
Probing Extreme Environments	2.1 What physical processes are at work close to neutron stars and black holes?		✓	✓	✓			vv		vv	
	2.2 What are the characteristics of relativistic jets, winds and explosions?		✓	✓	✓	✓	vv	vv		vv	
	2.3 How intense are radiation fields and magnetic fields in cosmic voids, and how do these evolve over cosmic time?					✓	✓			vv	
Exploring Frontiers in Physics	3.1 What is the nature of Dark Matter? How is it distributed?	vv	vv		✓						✓
	3.2 Are there quantum gravitational effects on photon propagation?						vv	✓		vv	
	3.3 Do Axion-like particles exist?					✓	✓			vv	

Figure 3.1 – Matrix of CTA science questions and proposed Key Science Projects (KSPs). The KSPs are sets of observations addressing multiple science questions within the CTA themes. KSPs which contribute to the programme aimed at dark matter detection are indicated in green, with the exclusively dark-matter-oriented targets described entirely within the Dark Matter Programme in Chapter 4. For KSPs simultaneously addressing dark matter and other physics/astrophysics, the motivation and context for the dark matter element is again described in Chapter 4. The order of the KSPs in this table starts with dark matter due to its importance and transversal nature and follows with surveys and more focused KSPs by increasing distance scale. The check marks are intended to give a qualitative assessment of the impact of each KSP on a particular science question.

プロジェクトの熟成度

Telescope Design Fixed



Telescope Types	SST	MST	LST
Optics	Schwarzschild-Couder	Davies-Cotton	Parabolic (Isochronous)
FoV and Camera	10.5 deg SiPM	7.5 deg PMT	4.3 deg PMT
Mirror Diameter	4.3m	11.5m	23m
Energy Range	3 TeV - 200 TeV	100GeV - 10TeV	20GeV – 2000GeV
Science Target	Galactic Sources PeVatron (UHE CR)	Galactic Sources Nearby AGNs (z<0.5) Dark Matter	Transient Sources AGNs(z<2), GRBs(z <4) Dark Matter



cherenkov
telescope
array

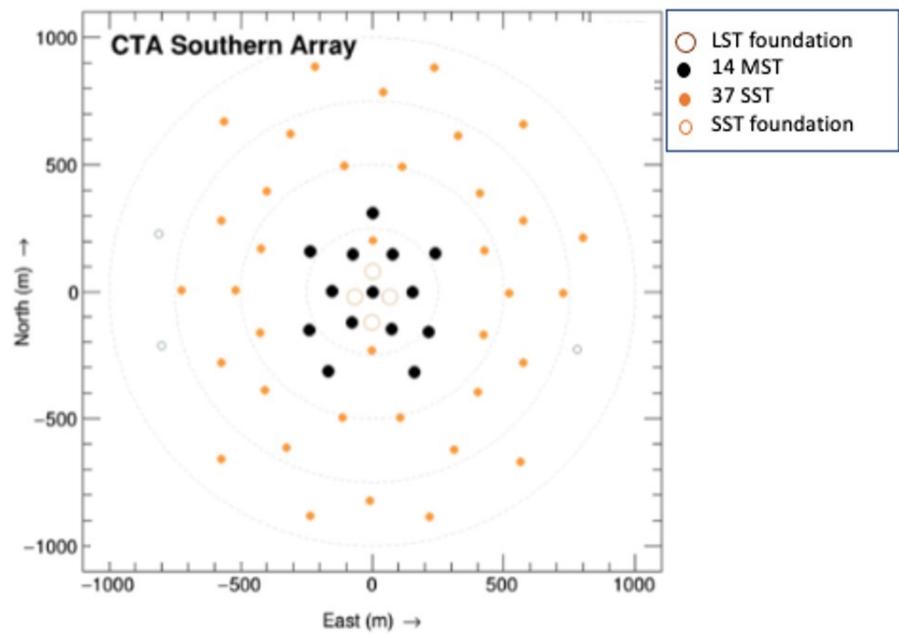
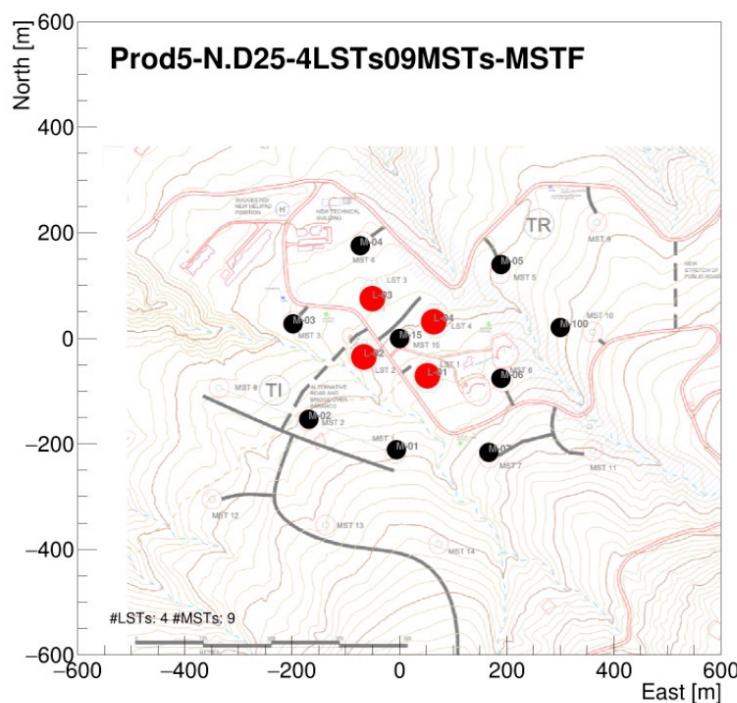
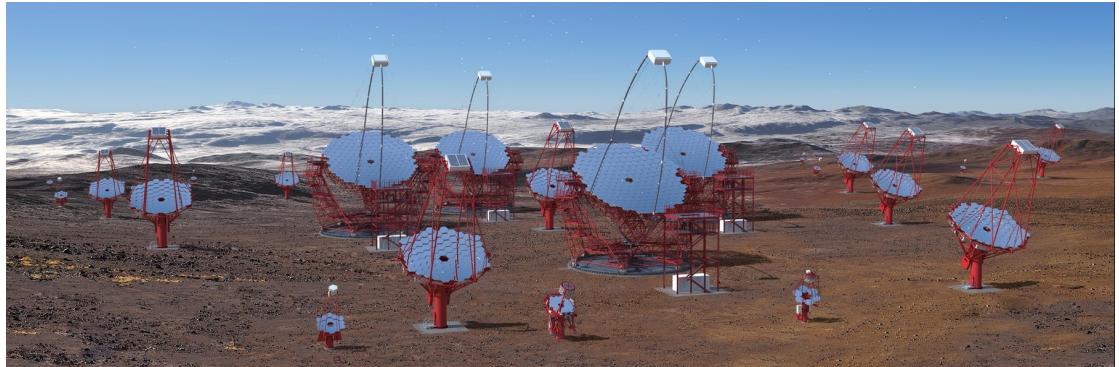
プロジェクトの熟成度

Alpha Configuration (CTAO Council, CTAO BGR) is decided with the financial constraints

Roque de los Muchachos Observatory
La Palma, Spain



Paranal, Chile



プロジェクトの成熟度

2006 HESS-MAGIC CTA Meeting, Project 議論を開始

2008 ESFRI Projectとして認定を受ける

2008-2012 Baseline Design, Telescope R&D 開始

2012 CTAC設立、CTA Resource board 設置

CTAO設置、STAC設置、LST/MST/SST Projects設置

2014 CTA South, CTA North sites 決定 (in CTA Resource Board)

2014 CTAO Council 設置

2014 LST/MST/SST Critical Design Review

2015 Business Plan 策定(柏ミーティング) , CTAO Council

2017 CTAO ERIC を目指して、CTAO BGR が作られる

2018 LST-1 が最初のCTA望遠鏡として建設される

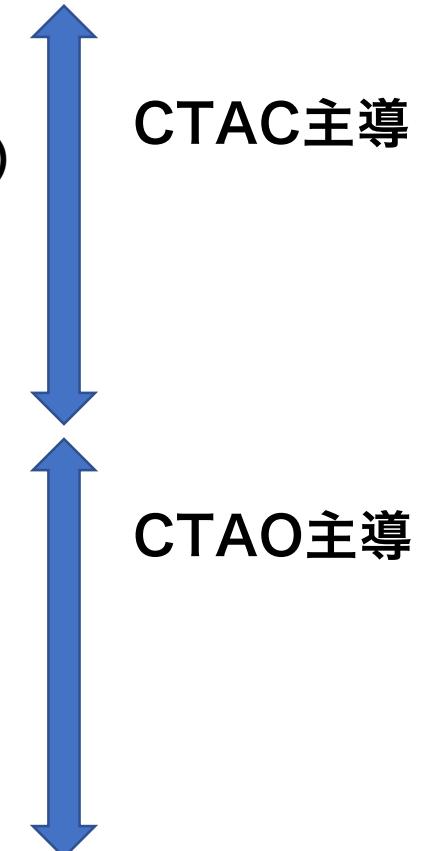
2018 ESFRI Landmark Project として認定される

2019 CTAO

2020 LST CDR を通過

2021 Alpha Config, CTAO Statutes 定款をまとめ、EC に10月提出

2022 CTAO ERIC 発足予定、CTAO建設開始



CTA Timeline

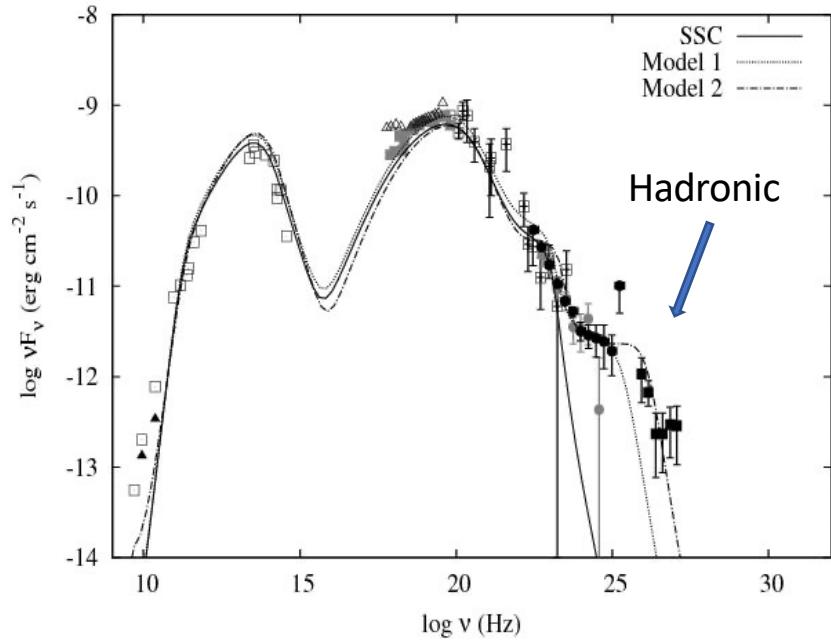
- 2016-2018 大口径望遠鏡 LST1 が最初のCTAサイトに建設される
- 2021-2022 Array配置、各国からの予算が決定され、CTAO ERIC が設立される。
- 2021-2024 CTA North にLST 4基のアレイが建設・完成され、運用が開始される。
- 2024-2028 14MSTs、37MSTs が建設される。

	2020	2021	2022	2023	2024	2025	2026	2027	2028	2029	2030					
Organization	CTAO gGmbH (Heidelberg)															
				CTAO ERIC (European Research Infrastructure Consortium)												
Alpha Config	2020	2021	2022	2023	2024	2025	2026	2027	2028	2029	2030					
LST North	Comissioning and Operation of LST1					Operation as 4 LST Array				Observatory Operation						
	CDR	Deployment of LST2-4														
MST North	Design and Finance		Construction of 9MSTs													
CTA South	Array config, Finance and CDR		INFRA	Construction and Deplyment of 14 MSTs												
				Construction and Deployment of 37 SSTs												
Extension	2020	2021	2022	2023	2024	2025	2026	2027	2028	2029	2030					
LST South		Advanced Design and Proto / Finance / CDR					Construction of 4LSTs				Operation					

LST South 緊急性

- 南半球でのサイエンスが不十分
 - AGNフレアー、GRB 観測の重要観測が半数
 - 100GeV-1TeV 領域でDM探索感度が不十分
 - G.C. (Soft → Hard at 100GeV)の観測ができない
 - Cen A (Fermi-HESS tension at 100GeV) → Two components
- 予想外の速さで CTAO ERIC の設立が進みつつある。
- First Phaseの後、Second Phase があるのか?
 - Second Phase がない可能性が高い。
 - 5–6年以内に、LST South の建設を行わないといけない。
 - First Phase 後では、マンパワーも拡散してしまう。
 - LSTを建設する技術が喪失される。
- 日本の貢献は10%レベルで、ドイツ、フランス、イタリア、スペインと比べ劣り、バランスが悪い。
 - LST South への貢献で20%レベルとなる。
 - MST、SSTへの貢献は他国が主要な貢献をしている。

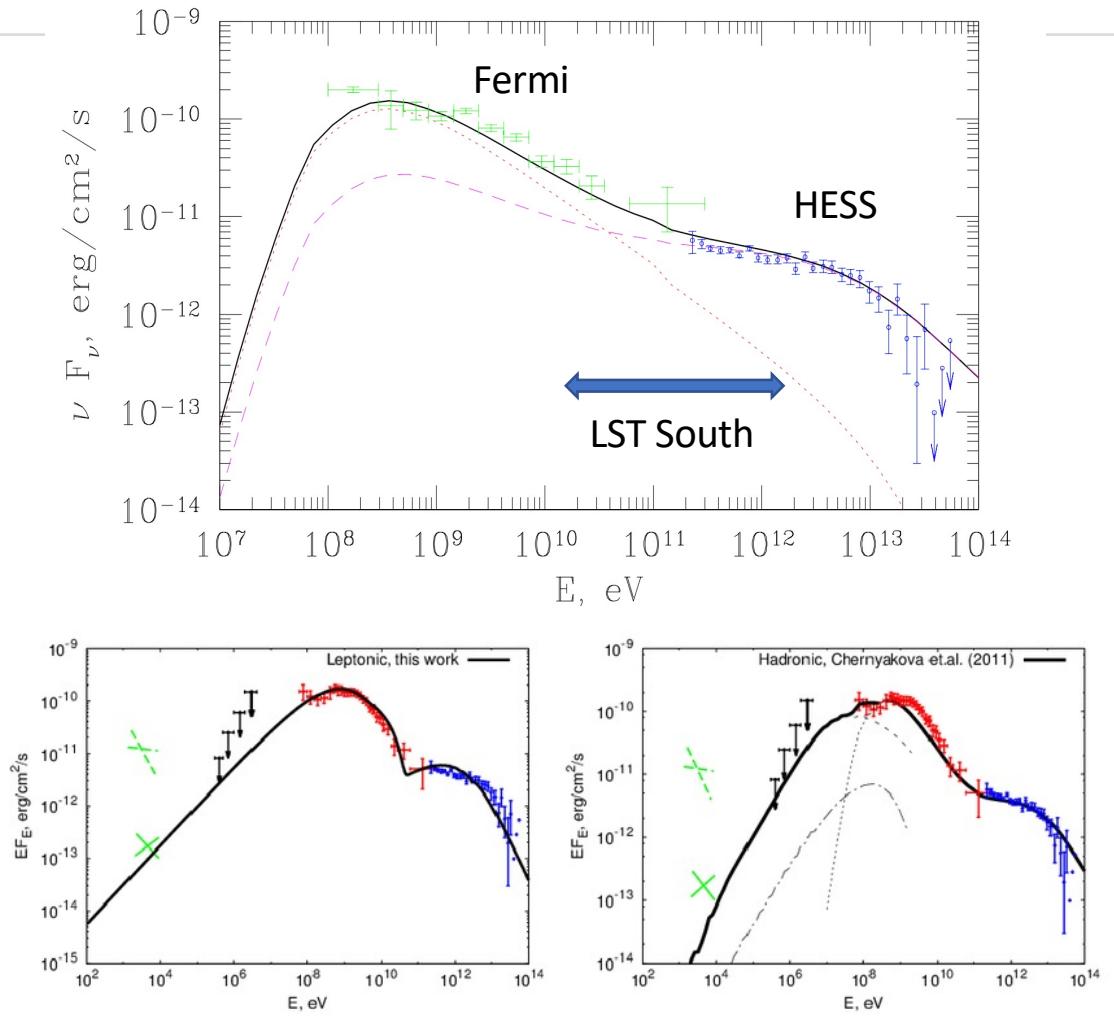
Cen A and Galactic Center



Cen A, Lepto hadronic Model

$N(P)/N(e) \sim 50$

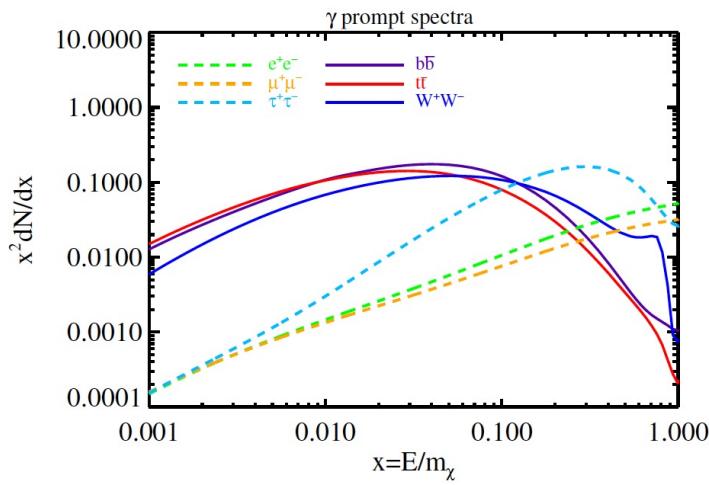
By M. Petropoulou et al. 2014 A&A



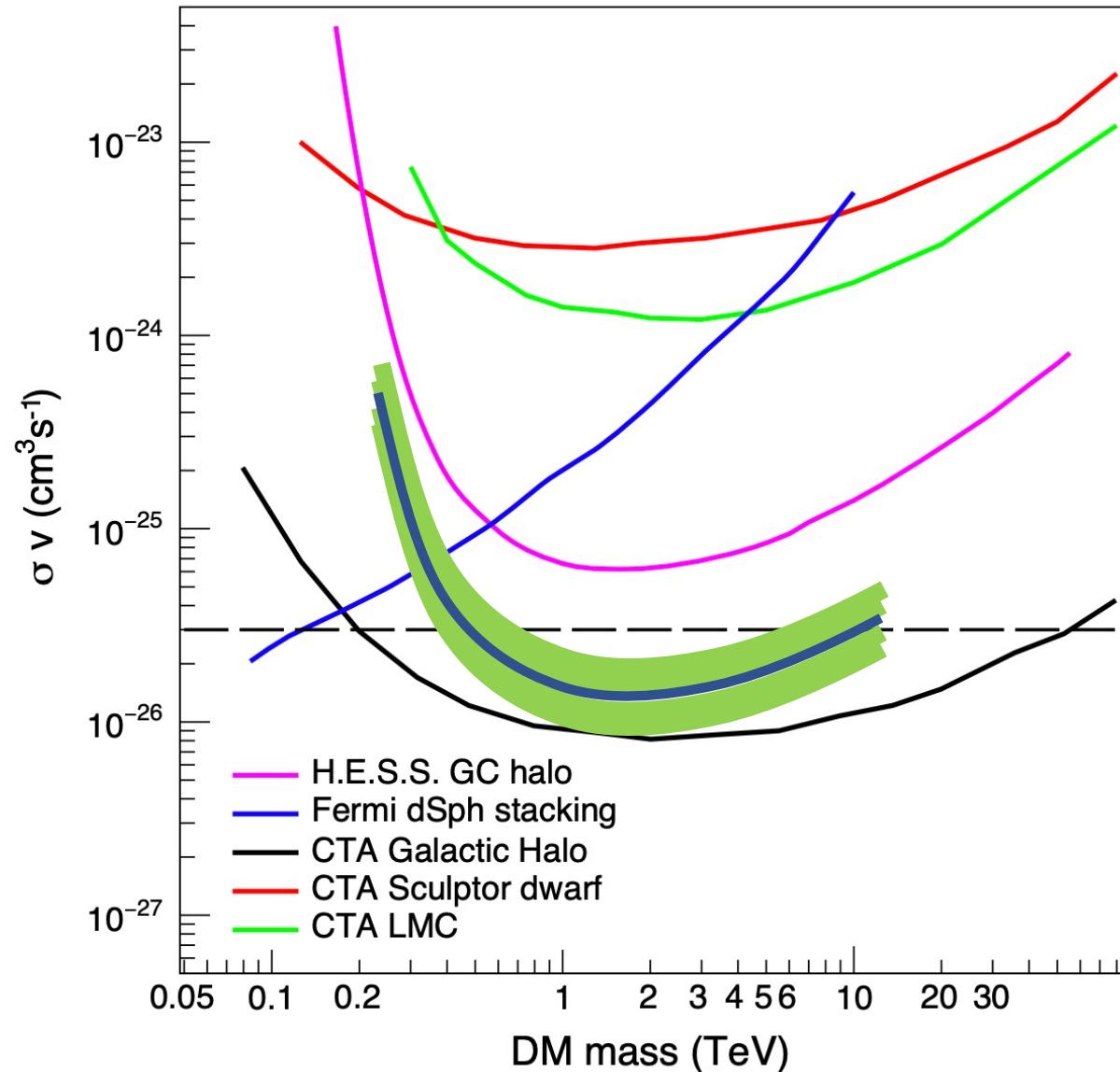
Spectrum of Galactic Center
By Chernyakova 2015

Dark Matter Sensitivity w/o LSTs and SSTs,

Teshima's Guess 2021, Surely we need MC results.



Gamma rays from Annihilation produce the bump around $1/10 - 1/20 M_\chi \rightarrow 20\text{GeV}-1\text{TeV}$ gamma





cherenkov
telescope
array

Multi-messenger and Multi-wavelength Astrophysics

Wave
AstroPhysics

ASTRO-PARTICLE PHYSICS
Cosmic Ray Physics
High Energy Astrophysics

Particle Physics

ASTRO-PHYSICS
Gamma Ray Bursts, Black holes,
Neutron Stars, Space and Time

PARTICLE PHYSICS
Dark Matter, Neutrino
Energy Frontier

