Calibration of energy scale uncertainty based on understanding of Super-Kamiokande detector elements

Kotaro Shimizu (Kamioka Observatory) kshimizu@km.icrr.u-tokyo.ac.jp

Abstract

The Super-Kamiokande is the water Cherenkov detector. In this experiment, the energy scale uncertainty is about 2%. To improve this energy scale uncertainty to 0.5%, I compared Data and MC of PMT hit pattern by using Ni/Cf source. As a result, I eliminated the possibility that water absorption is the cause of Data/MC disagreement of hit pattern.

1. Motivation

Purposes of Super-Kamiokande (SK) experiment

- (1) Measure the oscillation parameters of neutrinos (solar ν , atmospheric ν and accelerator ν).
- (2) Observe neutrinos from supernova explosions.
- (3) Search for nucleon decay and verify the GUT etc.

Energy scale uncertainty

- 2% for O(GeV) energy scale in current SK
- In HK, 0.5% is required for CP phase angle measurement
 - \rightarrow want to demonstrate this goal in SK by ${\color{blue}bottom-up}$ method

2. Method

I used Ni/Cf source (emitting 9 MeV γ -rays, see Fig.1) Data to study the uncertainty of the detector response at hit level. To compare Data and MC, I analyzed Hit Rate* in each PMT.

<u>Ni/Cf source</u> (see Fig.2)

- (1) The position of the light source can be clearly identified.(2) Light can be isotropically emitted.
- <u>Data</u> ← Apply the individual QE of each PMT to MC and match it with Run#1 Data. I used two kinds of Ni/Cf source Data. (Period when water is convected.)
 - -Run#1 : source position Z=0 m (center of the tank)
 - -Run#2, #3 : source position Z=-12, +12 m (off-center)





Fig.1 Ni/Cf source

Run #3**0**,

Run #1

39.3 m

Run #2**O**(z= -12 m

Fig.2 source position

ДД

(Z=+12 m)

Hit Rate (i) = $\frac{1}{\sum [\text{Number of events with hit}(i) \times r(i)^2 / F(\theta(i))] / N}$ $\begin{bmatrix} i : \text{PMT cable number } (1 \sim 11146), r(i) : \text{ distatnce (source to each PMT)} \\ F(\theta(i)) : \text{ acceptance of a PMT as a function of incident angle, } N : \text{Number of PMTs (11146)} \end{bmatrix}$

Number of events with hit(i) $\times r(i)^2/F(\theta(i))$

3. Results

Fig.3 shows Hit Rate ratio (= Data/MC) as a function of distance (Ni/Cf source to each PMT). The Ni/Cf source is located at Z = -12 m and +12 m. This histgram is mean value of the ratio in every 100 cm. The <u>red</u>, <u>blue</u> and <u>black</u> lines represent the each PMTs located at the <u>Top</u>, <u>Bottom</u>, <u>Barrel</u> of the tank respectively.



Fig.3 Distance distribution of Hit Rate ratio (Data/MC)

For <u>Top</u> PMT in left figure and <u>Bottom</u> PMT in right figure, Hit Rate ratio decreases with increasing distance, whereas this tendency is not seen in <u>Barrel</u> PMT. So it can be said that Data/MC disagreement is <u>not</u> caused by water absorption parameter.

4. Conclusion and Future Step

So far, I compared the Data and MC of Hit Rate in off center using MC that was calibrated in center data. As a result, I eliminated the possibility that water absorption is the cause of Data/MC disagreement of hit pattern. I want to identify the cause of uncertainty from the tendency of Data/MC agreement and change the parameters (reflection, scattering, angle dependence of PMT etc..) and calculation methods of MC in low energy. After that, I want to apply that method to MC in high energy.

