

Measuring the birefringence of the sapphire mirrors installed in the KAGRA detector

Research Results Presentation Meeting
of the Inter-University Research Program
for Fiscal Year 2022

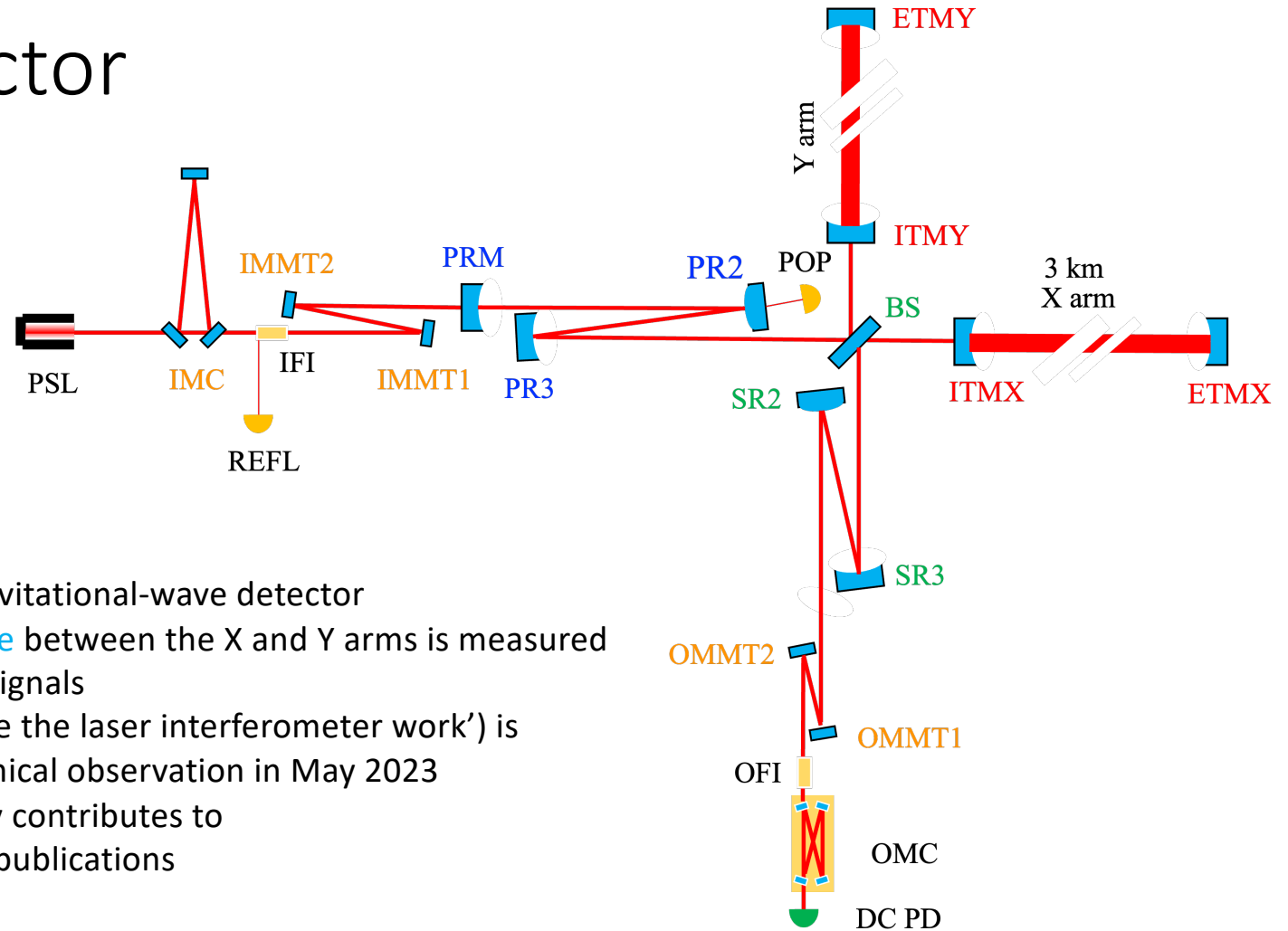
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Keiko Kokeyama, Cardiff University

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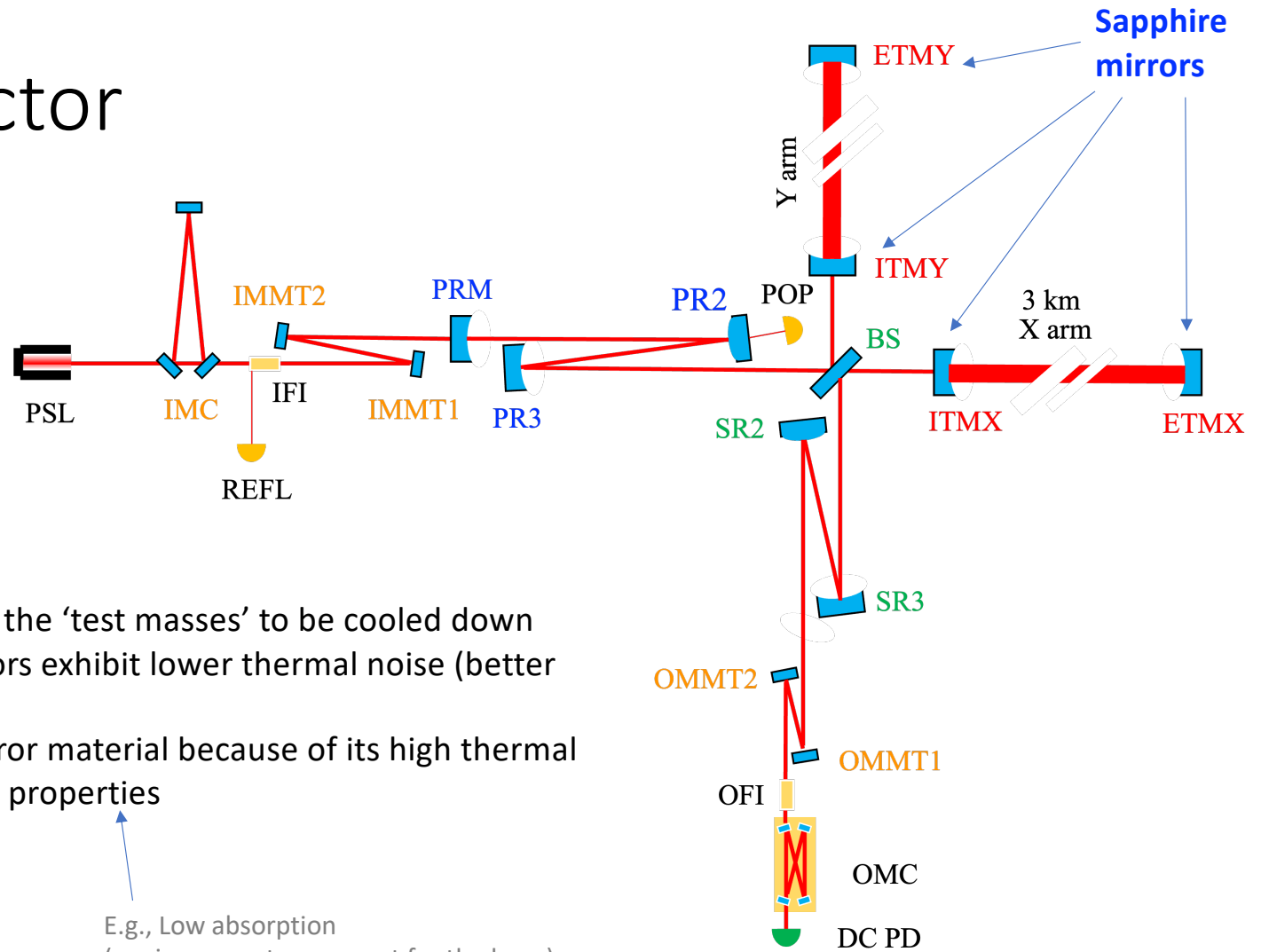
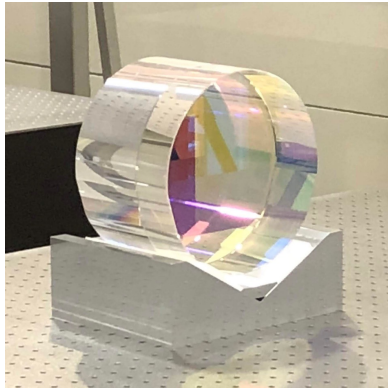
- KAGRA detector
- Sapphire mirrors
- Birefringence
- Polarization phase camera
- Status of the project

KAGRA Detector



- Large-scale and **cryogenic** gravitational-wave detector
- **Differential arm length change** between the X and Y arms is measured to detect gravitational wave signals
- Intense commissioning ('make the laser interferometer work') is ongoing to start the astronomical observation in May 2023
- Cardiff University significantly contributes to detector commissioning and publications

KAGRA Detector

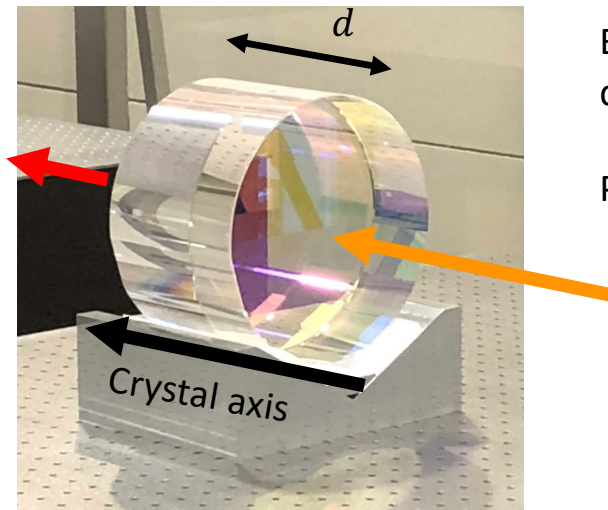


- **Sapphire mirrors** are used for the 'test masses' to be cooled down
- Lower the temperature, mirrors exhibit lower thermal noise (better sensitivity)
- Sapphire is chosen as the mirror material because of its high thermal conductivity and good optical properties

Heat induced by laser must be removed

E.g., Low absorption
(= mirrors are transparent for the laser)

Sapphire and Birefringence



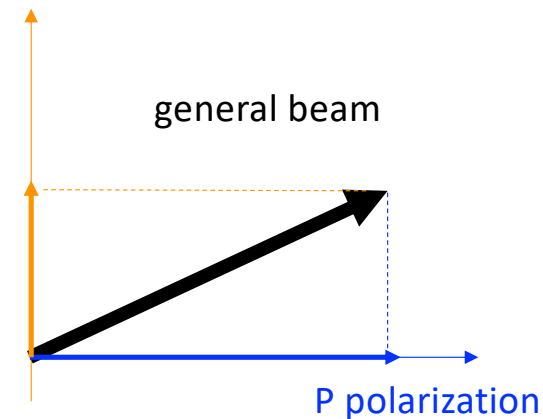
Birefringence materials have different refractive indexes for different light polarization.

Phase difference R between the blue- and orange-polarized beams:

$$R = \frac{2\pi d(n_o - n'_e)}{\lambda}$$

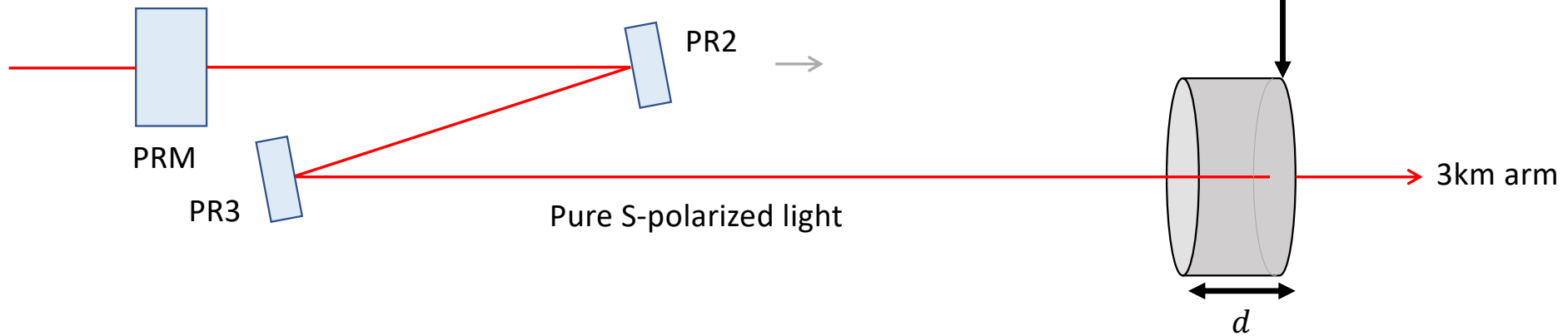
Because any general linear polarized beam is a combination of the orthogonal blue- and orange-polarized beam, **the transmitting beam's polarization state is changed if the crystal z-axis and the beam propagation axis don't match, or the crystal axis is inhomogeneous over the crystal**

S polarization



Birefringence Problems in KAGRA

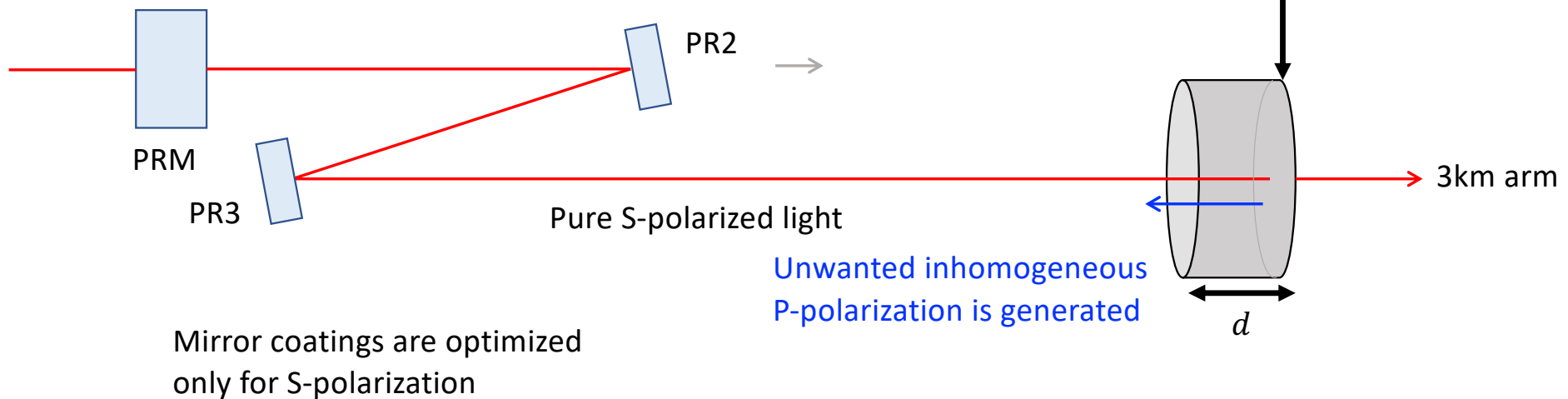
Birefringent input test masses (ITMs) generate unwanted polarization component in the power recycling cavity (PRC)



The crystal axis was found to be **inhomogeneous over the mirror**

Birefringence Problems in KAGRA

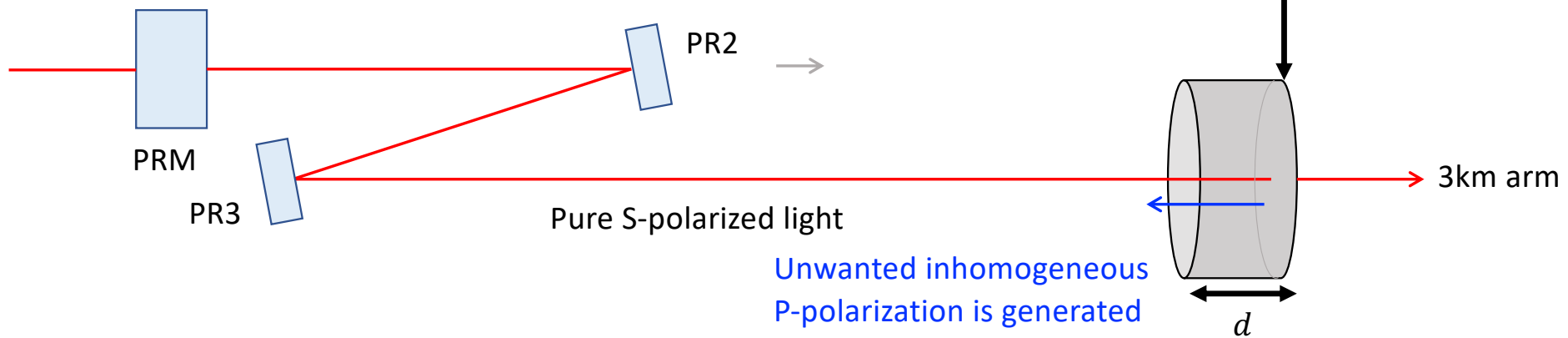
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Birefringence Problems in KAGRA

Birefringent input test masses (ITMs) generate unwanted polarization component in the power recycling cavity (PRC)

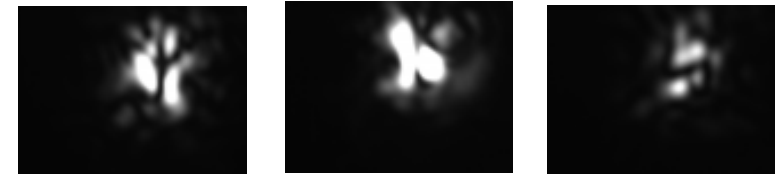


The crystal axis was found to be **inhomogeneous over the mirror**

Resulted in

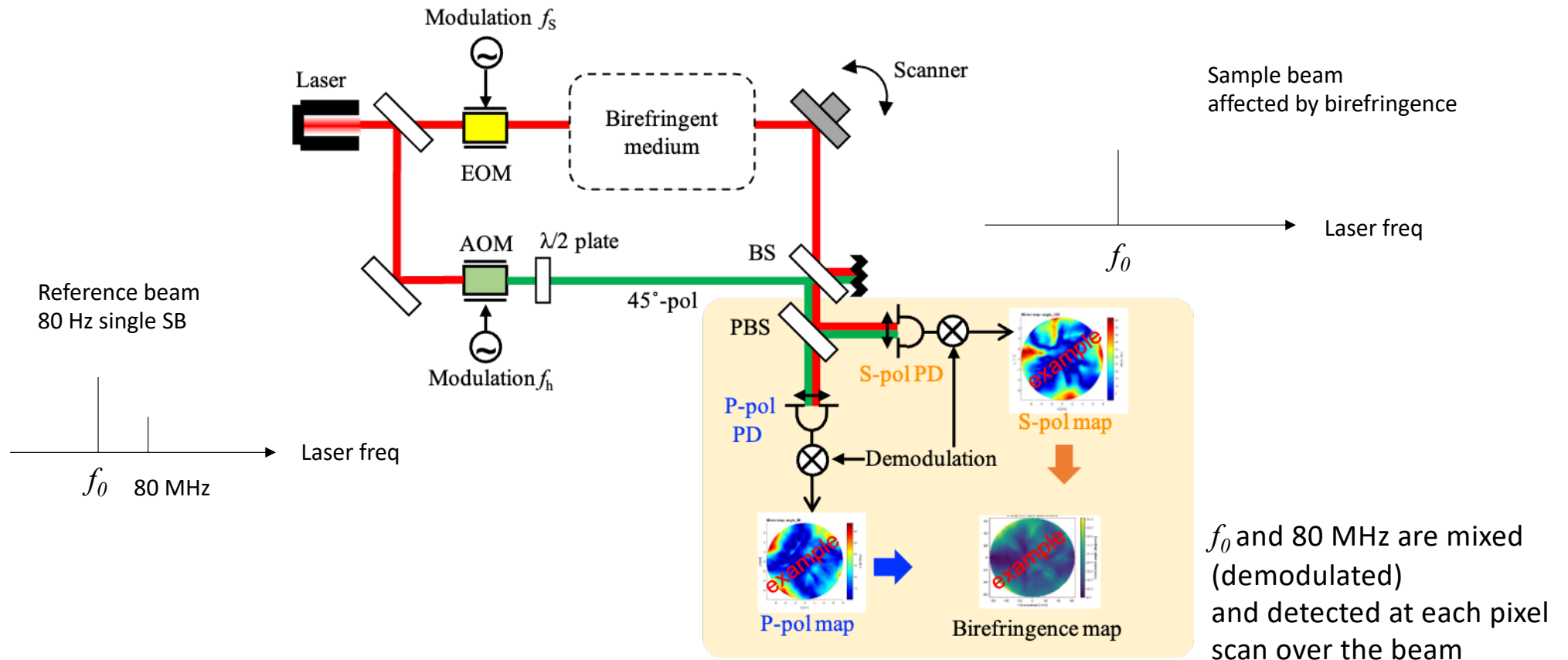
- Laser power loss
- Unstable error signals for the cavity length and angular controls
 → **Lower duty factor of the observation**

<https://klog.icrr.u-tokyo.ac.jp/osl/?print&r=9495>

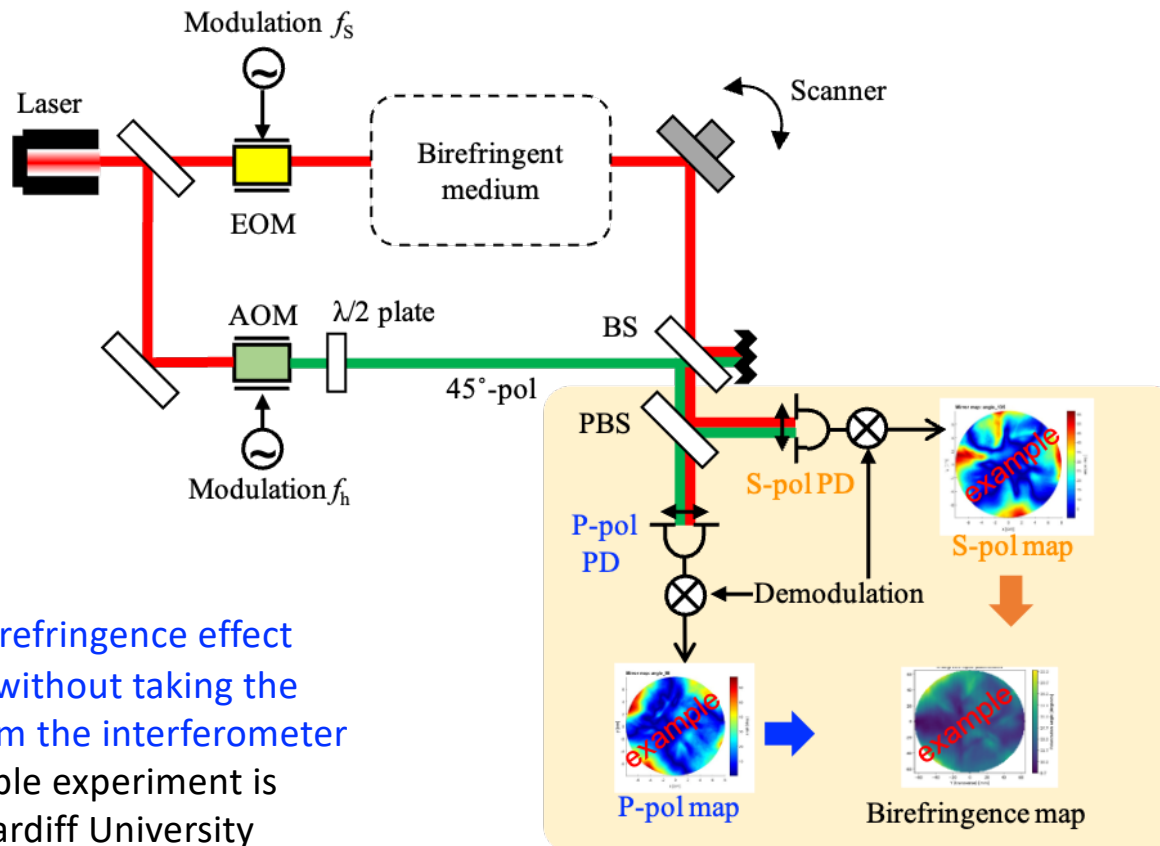


→ Transmitted beam of PR2 in the unwanted polarization

Principle of the polarized phase camera

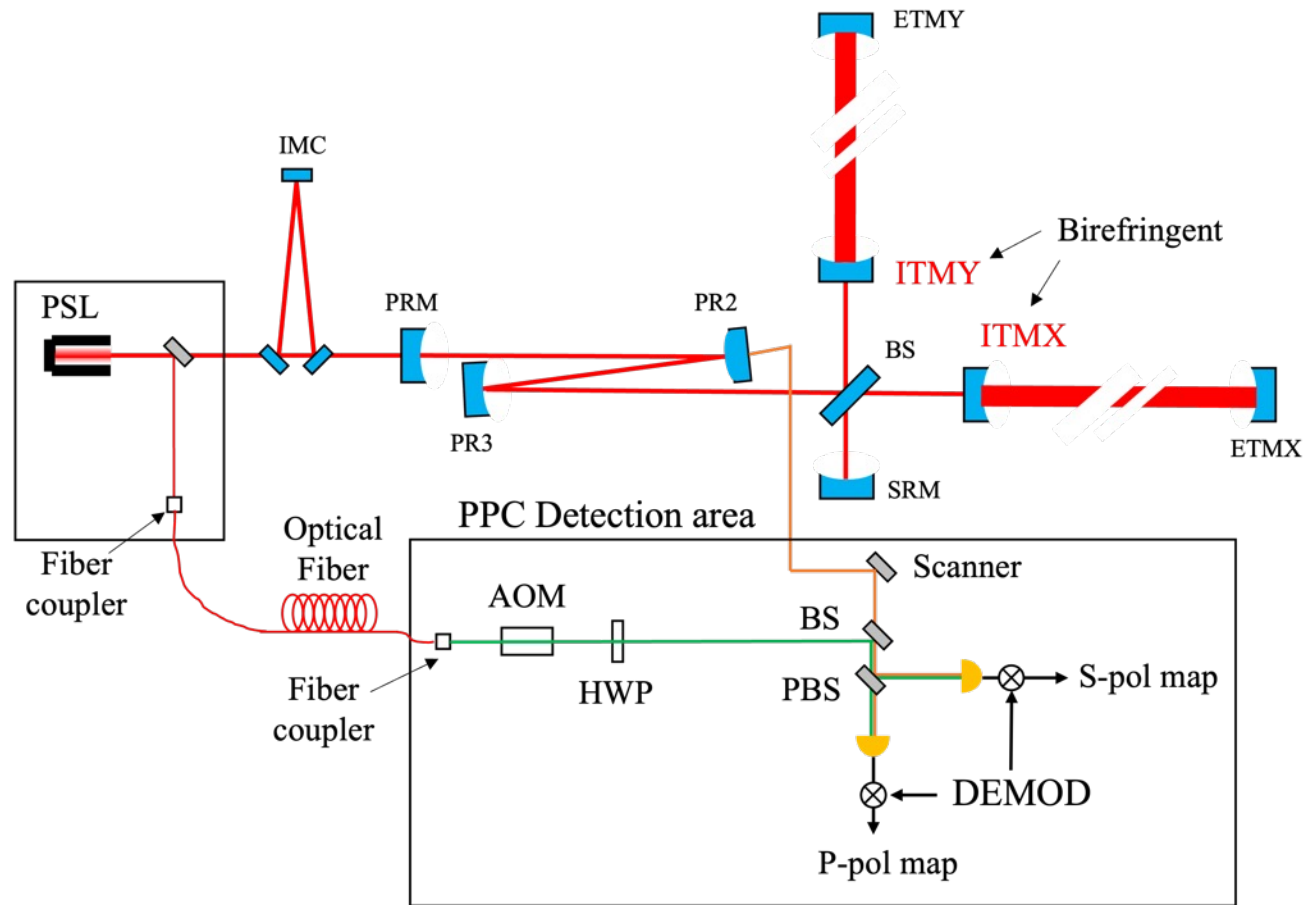


Principle of the polarized phase camera



- Enabling the birefringence effect measurement without taking the mirrors out from the interferometer
- Proof-of-principle experiment is underway at Cardiff University

FSY 2022: Application Design



Procurement of FSY2022

- Approved: JPY 300,000
- Used: JPY 144,188
 - Half wave plate
 - Fiber checker to align the fiber-coupled AOM

Plans for FSY2023

- Prepare the reference beam
 - Consider how to pick off from the pre-stabilized laser room (upstream of the interferometer)
 - Select the optical fiber and paths to send the beam to the PPC detection area (near PR2) by optical fiber
 - Purchase the AOM