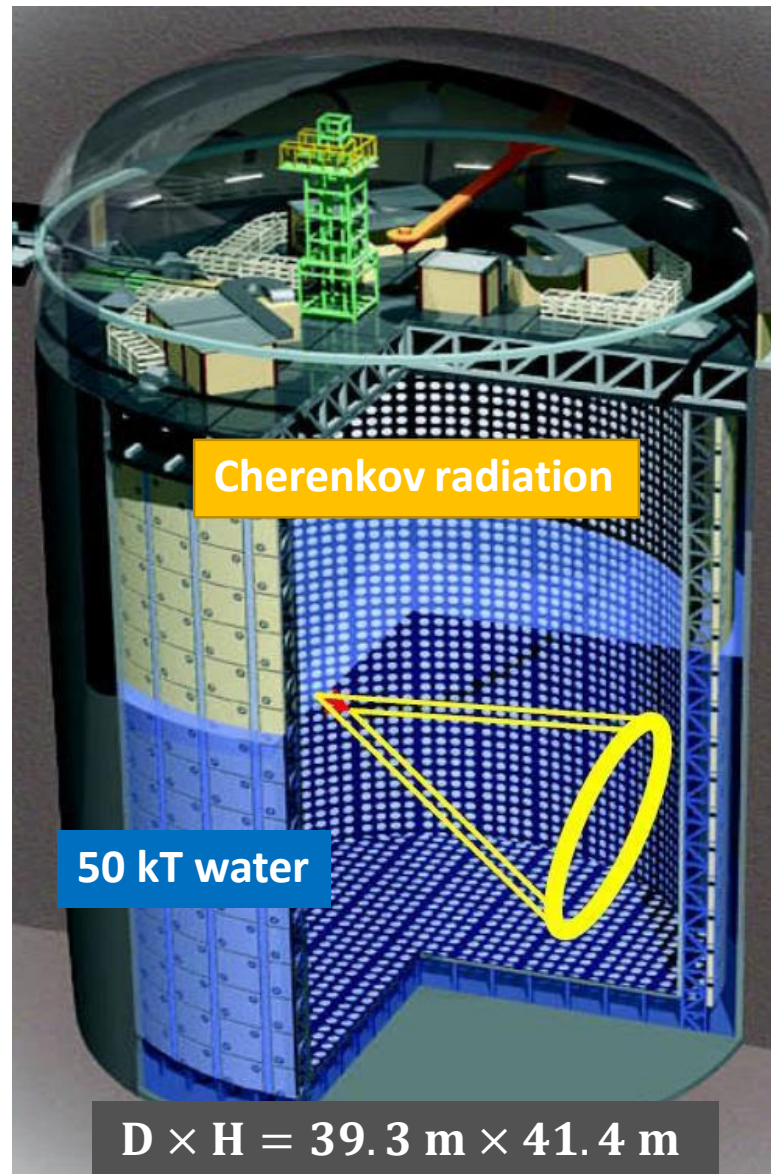
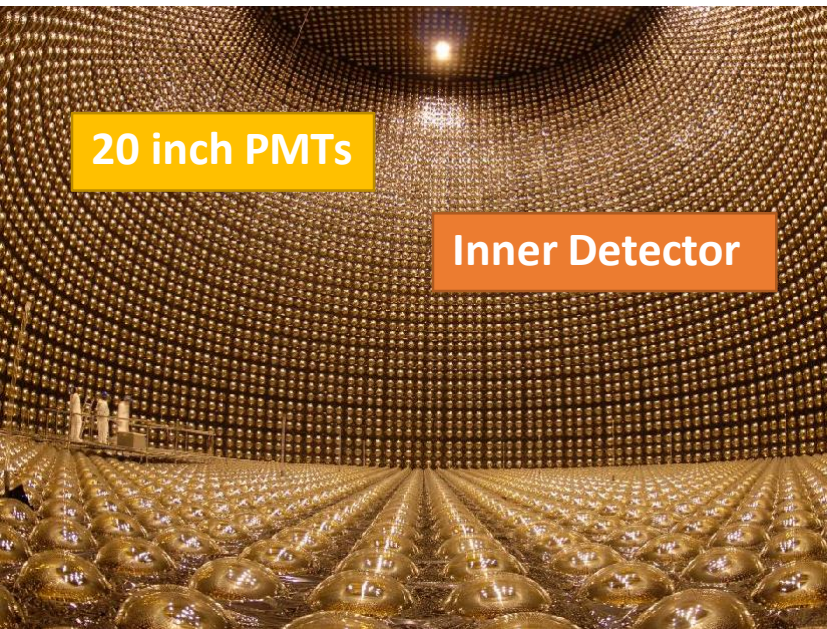




# スーパーカミオカンデ： 大気ニュートリノ、核子崩壊

三浦 真 (神岡宇宙素粒子研究施設, ICRR)

# The Super-Kamiokande Detector





# The Super-Kamiokande Collaboration

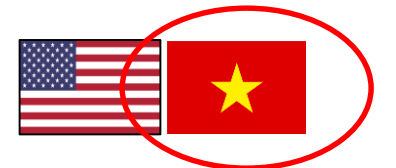


~230 collaborators  
from 51 institutes  
in 11 countries

Kamioka Observatory, ICRR, Univ. of Tokyo, Japan  
RCCN, ICRR, Univ. of Tokyo, Japan  
University Autonoma Madrid, Spain  
BC Institute of Technology, Canada  
Boston University, USA  
University of California, Irvine, USA  
California State University, USA  
Chonnam National University, Korea  
Duke University, USA  
Fukuoka Institute of Technology, Japan  
Gifu University, Japan  
GIST, Korea  
University of Hawaii, USA  
IBS, Korea  
IFIRSE, Vietnam  
Imperial College London, UK  
ILANCE, France

INFN Bari, Italy  
INFN Napoli, Italy  
INFN Padova, Italy  
INFN Roma, Italy  
Kavli IPMU, The Univ. of Tokyo, Japan  
Keio University, Japan  
KEK, Japan  
King's College London, UK  
Kobe University, Japan  
Kyoto University, Japan  
University of Liverpool, UK  
LLR, Ecole polytechnique, France  
Miyagi University of Education, Japan  
ISEE, Nagoya University, Japan  
NCBJ, Poland  
Okayama University, Japan  
University of Oxford, UK

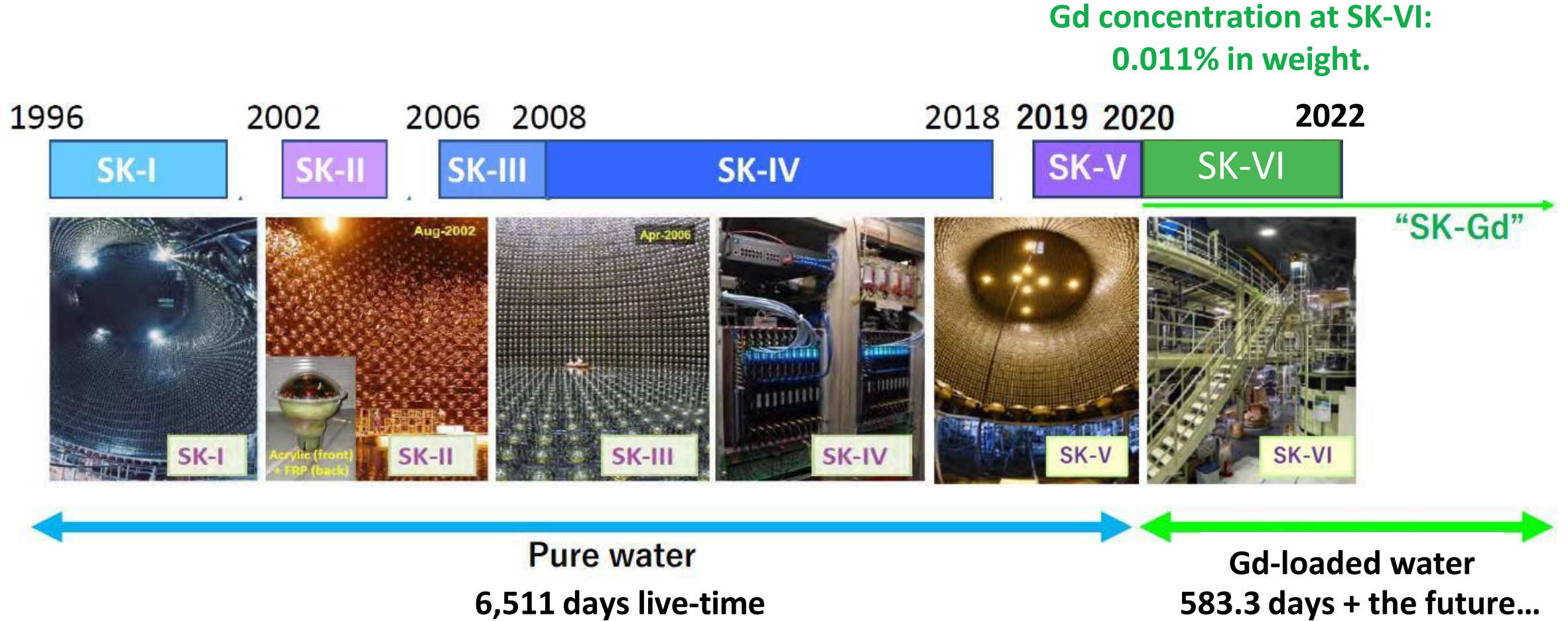
Rutherford Appleton Laboratory, UK  
Seoul National University, Korea  
University of Sheffield, UK  
Shizuoka University of Welfare, Japan  
Sungkyunkwan University, Korea  
Stony Brook University, USA  
Tohoku University, Japan  
Tokai University, Japan  
The University of Tokyo, Japan  
Tokyo Institute of Technology, Japan  
Tokyo University of Science, Japan  
TRIUMF, Canada  
Tsinghua University, China  
University of Warsaw, Poland  
Warwick University, UK  
The University of Winnipeg, Canada  
Yokohama National University, Japan



2023/02/22

Since 2021

# SK Data Taking Phases

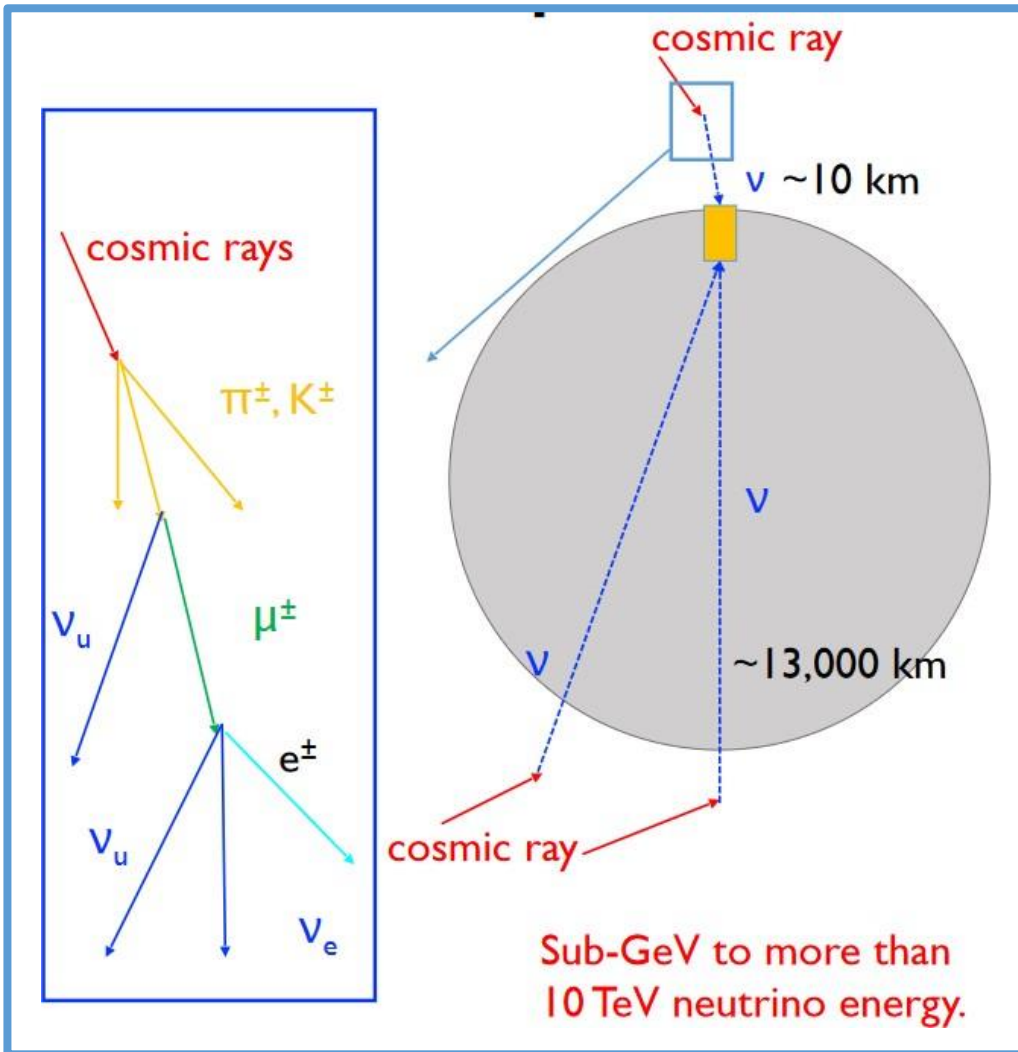




# New Results from SK

- **Atmospheric neutrino oscillation measurements**
  - SK-I through SK-V + Expanded FV
  - Three Flavor Oscillation with T2K Constraints
- **Recent publications:**
  - **Search for Cosmic ray Boosted Sub-GeV Dark Matter Using Recoil Protons (PRL 130, 031802 (2023))**
  - **$p \rightarrow \mu^+ K^0$  (PRD 106, 072003)**
- **Neutron capture on Gd in SK-VI/VII**

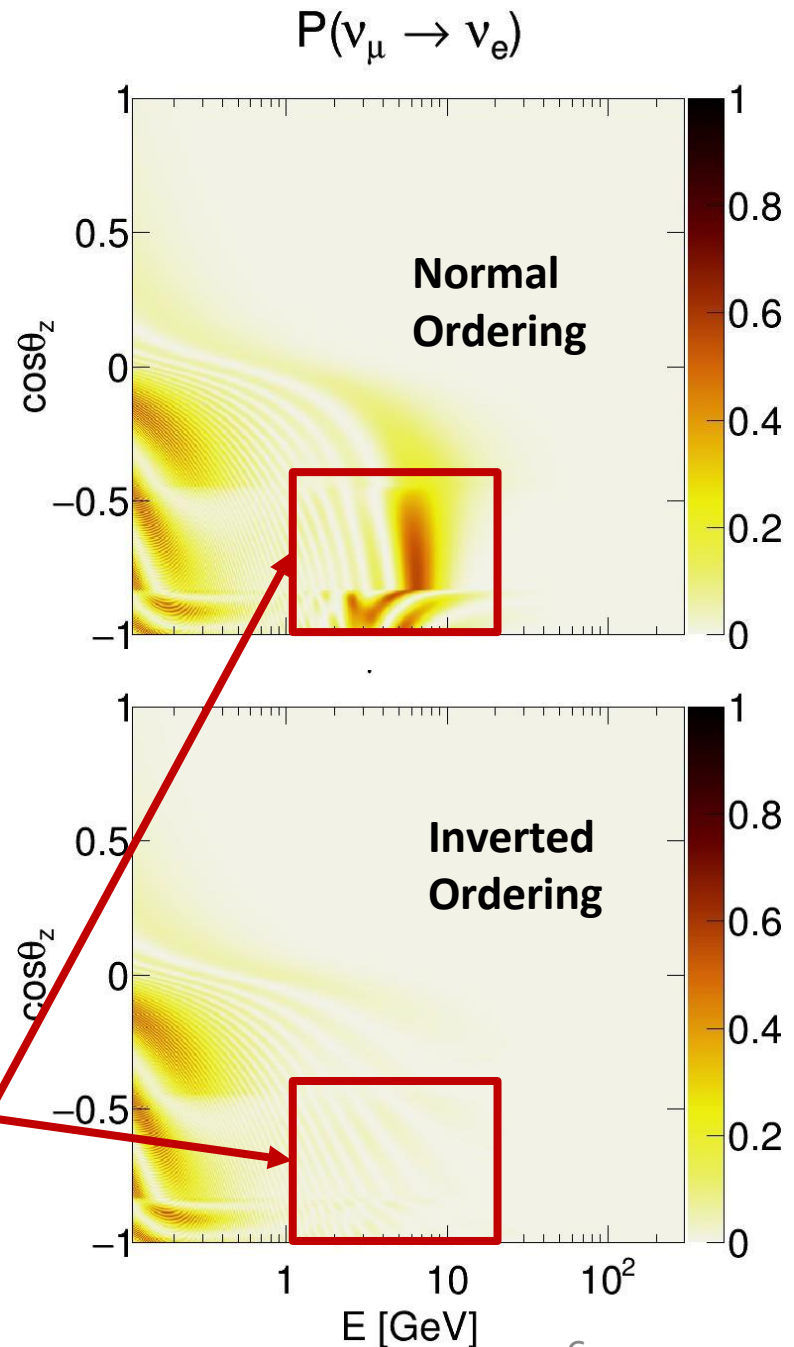
# Atmospheric Neutrino Oscillation



## Key measurements:

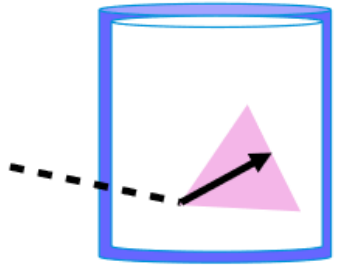
- $\nu_\mu$  disappearance
  - $\Delta m_{32}^2$
  - $\sin^2 \theta_{23}$
- $\nu_e$  appearance
  - CP violation  $\delta$
  - Mass-ordering

Matter Effect

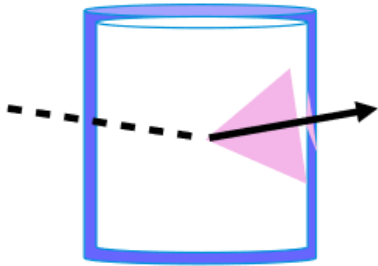


# Atmospheric Neutrino Analysis at SK

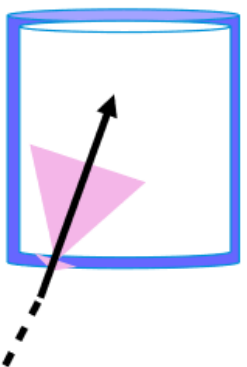
Fully Contained (FC)



Partially Contained (PC)



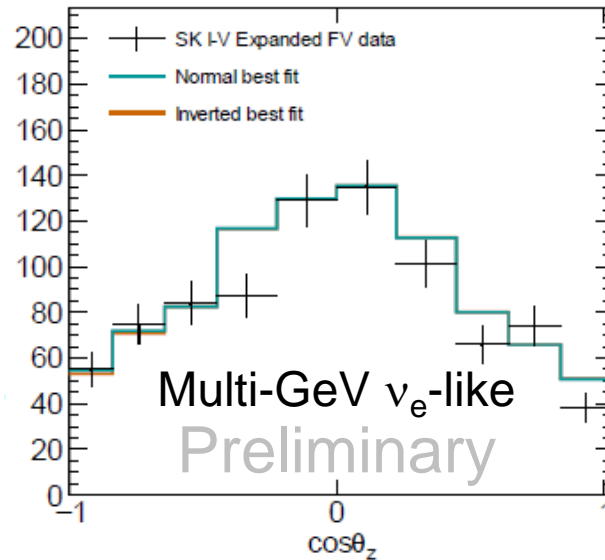
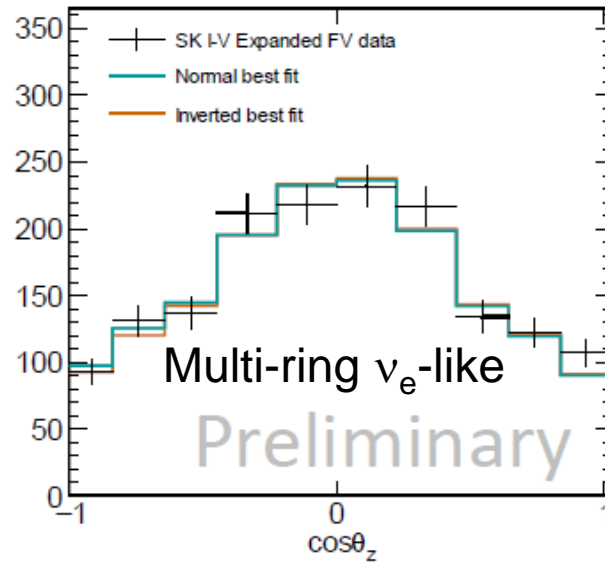
Upward-going Muons (Up- $\mu$ )



Reconstructed,  
Classified, and  
binned

2023/02/22

Multi-Ring e-like  $\nu_e$



**Total exposure:  
484.2 kiloton-years**

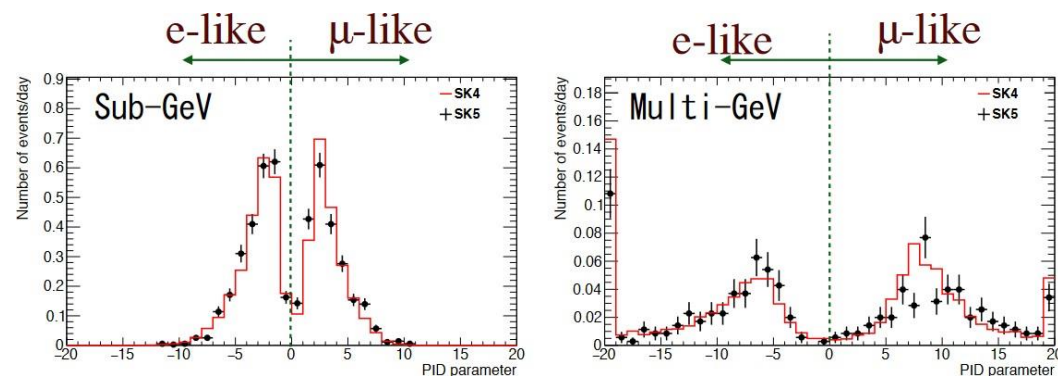
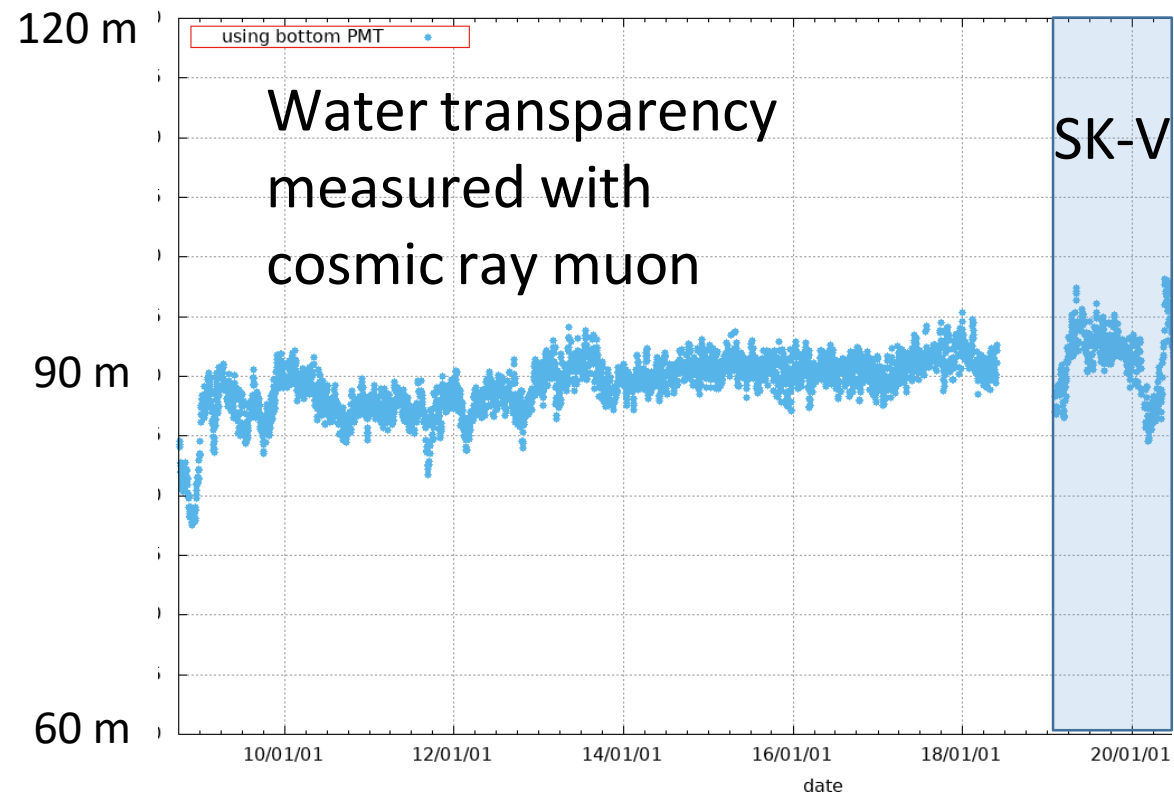
30% more data than 2020 analysis  
Using all of pure water data at SK

**New in this analysis:**

- SK-V data
- Expanded fiducial volume
- T2K model including  $\bar{\nu}$  mode
- New multi-ring selection
- Systematics improvements



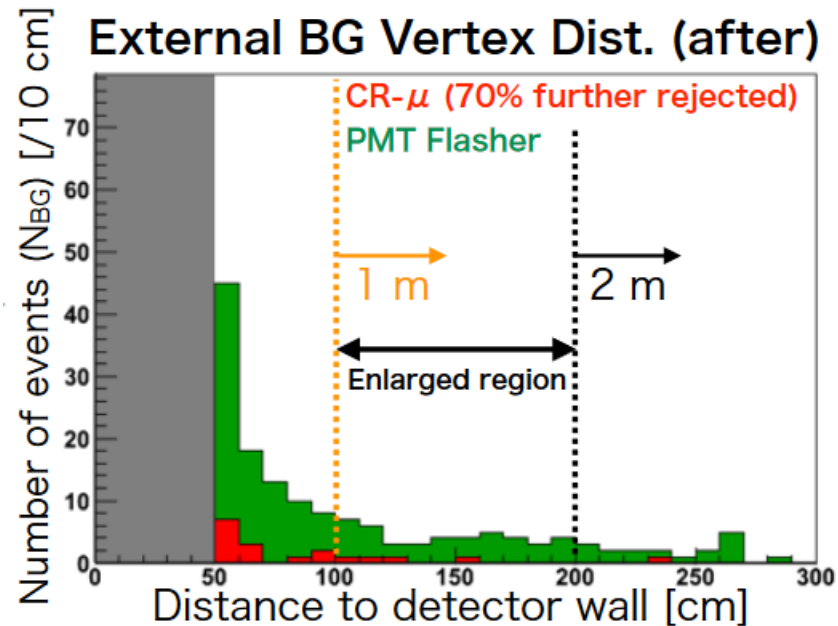
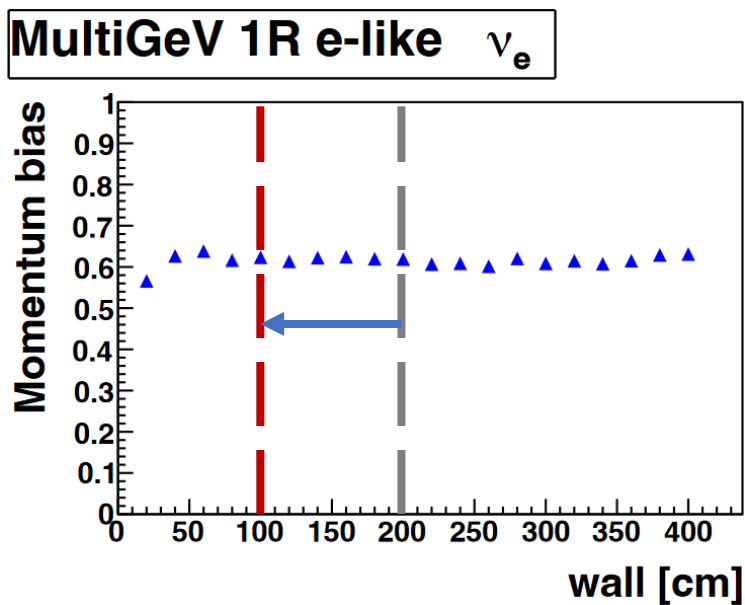
- The last SK phase with pure water
- Upgraded water system, replaced PMT, cleaned detector... Getting ready for Gd loading!
- Consistent data quality with SK4



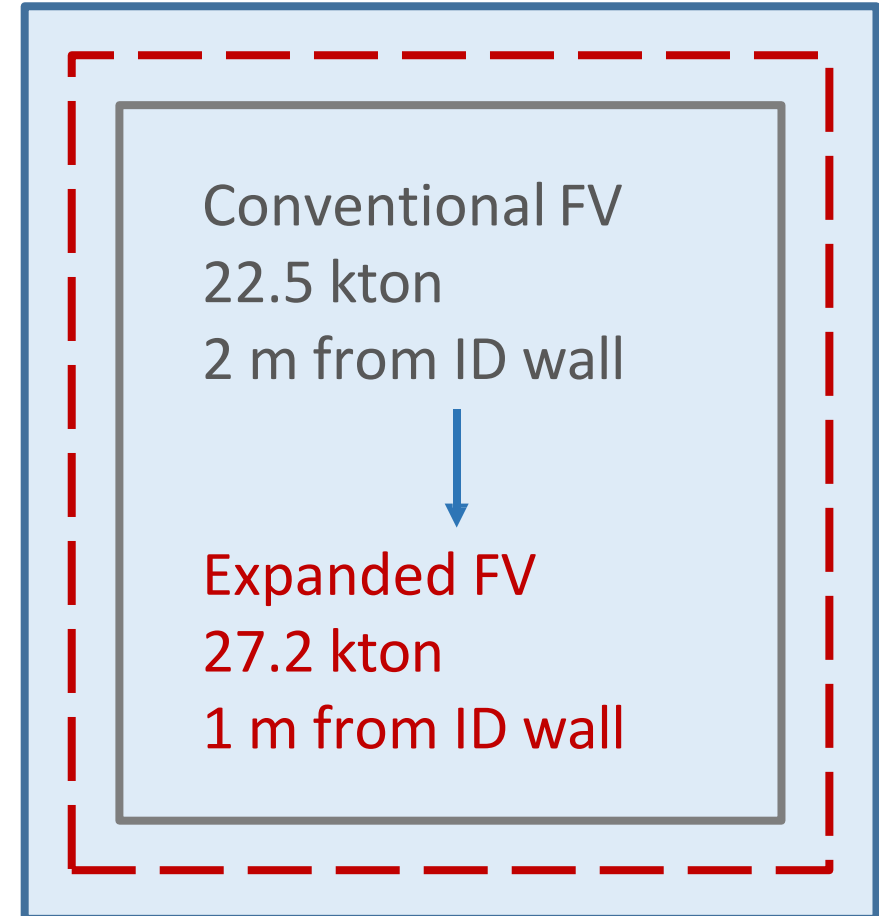


# Expanded Fiducial Volume

- Expanded fiducial volume
  - 22.5 kton  $\rightarrow$  27.2 kton, 20% increase
- No significant increase of external background
- No significant bias in reconstruction
- Systematics re-estimated for expanded FV



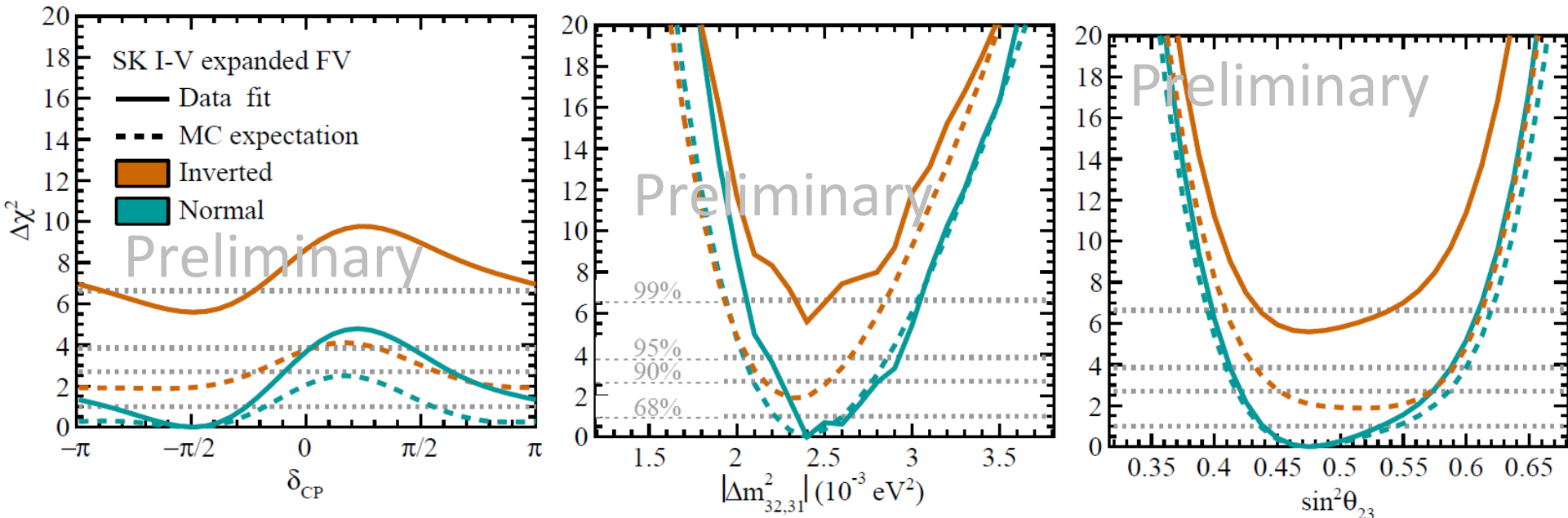
Inner detector (ID) wall



# Oscillation Measurements (SK only)

SK atmospheric neutrino data favors:

- maximal mixing
- NO ( $\Delta\chi^2 = 5.6$ )
- $\delta_{CP} \approx -\frac{\pi}{2}$

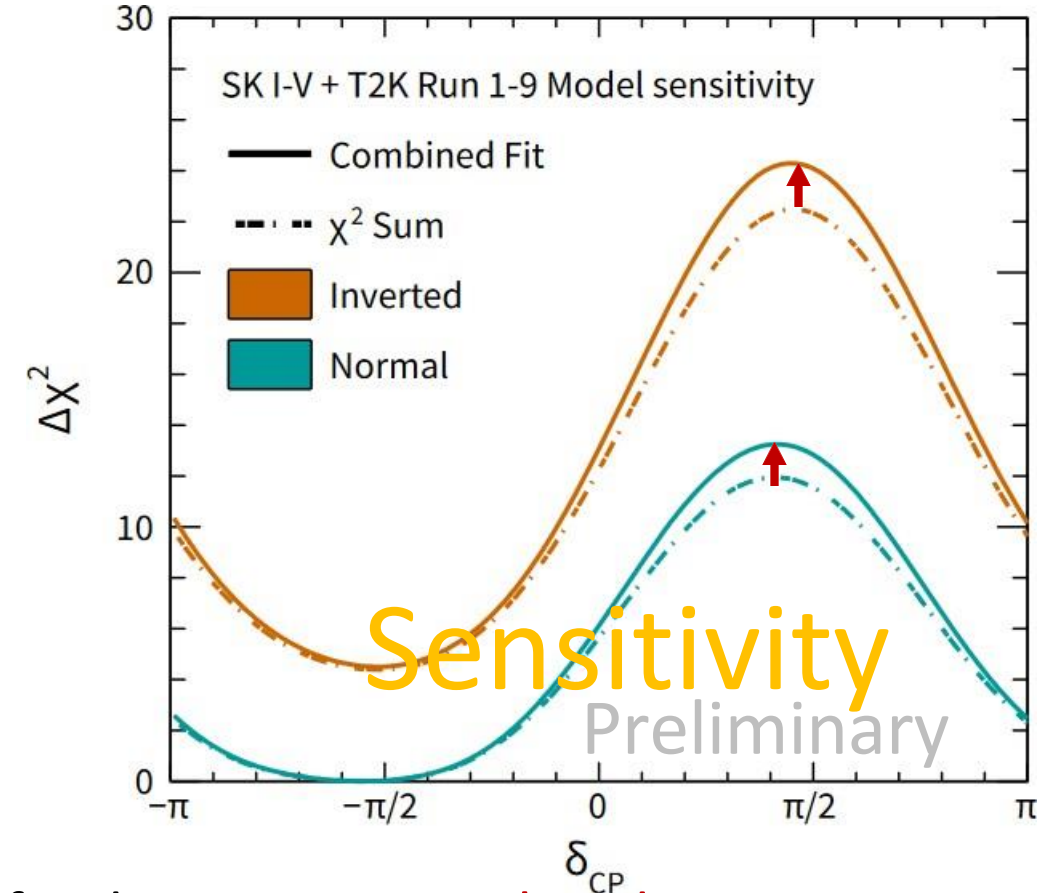
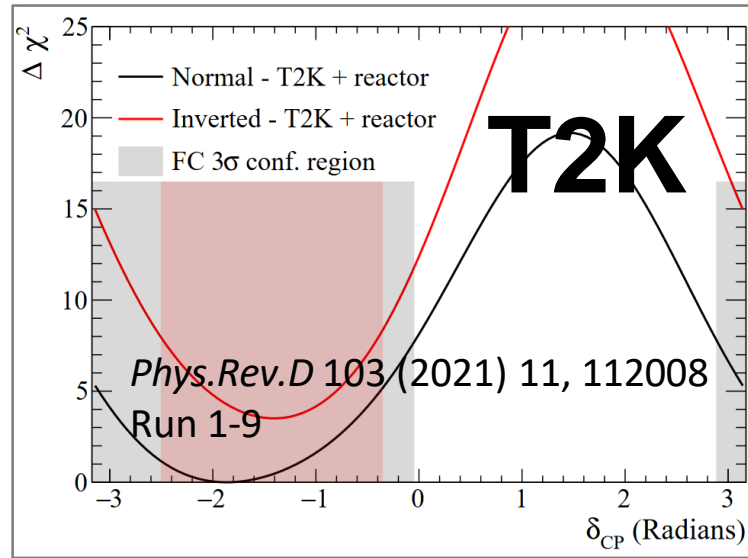
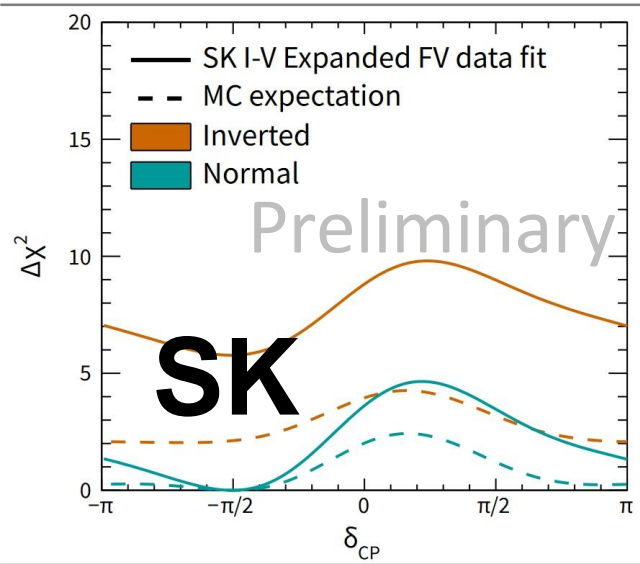


| 930 bins | $\chi^2$ | $\delta_{CP}$ | $\sin^2\theta_{23}$ | $\Delta m_{23}^2$                 |
|----------|----------|---------------|---------------------|-----------------------------------|
| SK NO    | 1004.56  | 4.71          | 0.48                | $2.4 \times 10^{-3} \text{ eV}^2$ |
| SK IO    | 1010.15  | 4.71          | 0.48                | $2.4 \times 10^{-3} \text{ eV}^2$ |

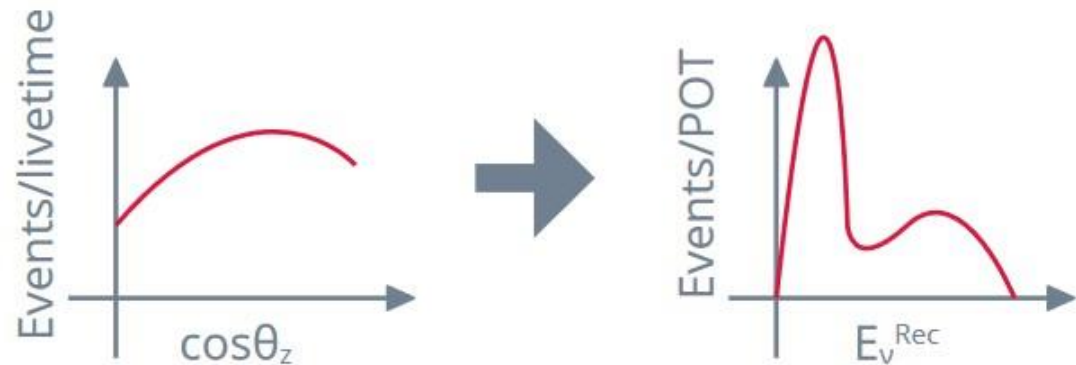
$\sin^2\theta_{13} = 0.0220 \pm 0.0007$

# Combining SK and External T2K Constraints

- SK sensitive on **mass ordering**, T2K sensitive on  $\delta_{CP}$



- To combine, reweight SK MC to T2K published data



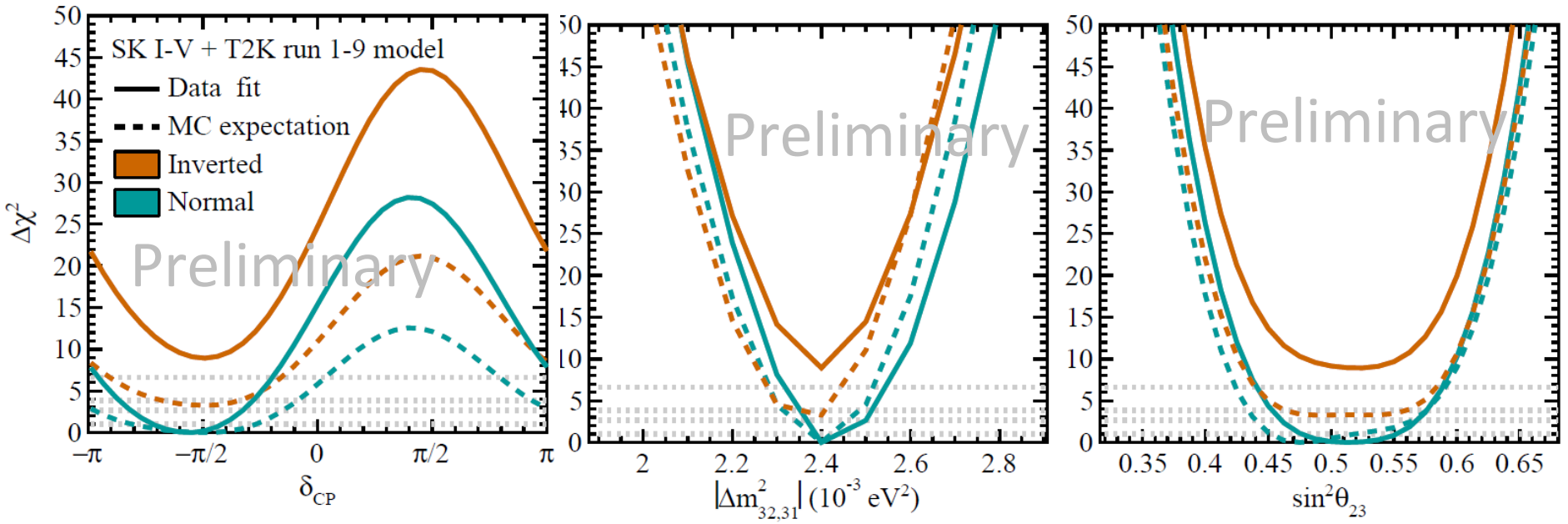
- SK = T2K far detector → **correlated cross-section**
- Additional sensitivity gained from combined fit with **correlated cross-section uncertainty**



# Oscillation Measurements (SK+T2K)

SK + external T2K constraints favor:

- maximal mixing
- NO ( $\Delta\chi^2 = 8.9$ )
- $\delta_{CP} \approx -\frac{\pi}{2}$

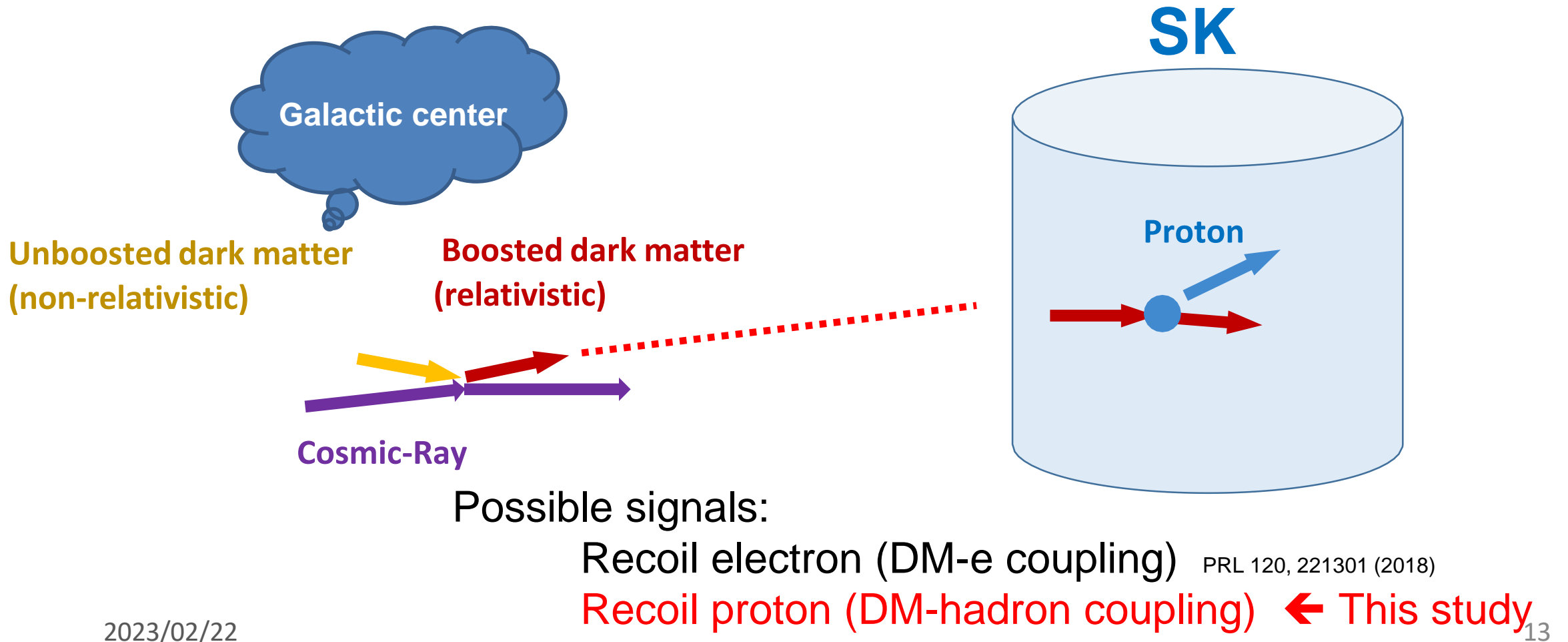


| 1020 bins | $\chi^2$ | $\delta_{CP}$ | $\sin^2\theta_{23}$ | $\Delta m^2_{23}$                 |
|-----------|----------|---------------|---------------------|-----------------------------------|
| SK+T2K NO | 1094.58  | 4.54          | 0.51                | $2.4 \times 10^{-3} \text{ eV}^2$ |
| SK+T2K IO | 1103.50  | 4.71          | 0.53                | $2.4 \times 10^{-3} \text{ eV}^2$ |

$\sin^2\theta_{13} = 0.0220 \pm 0.0007$

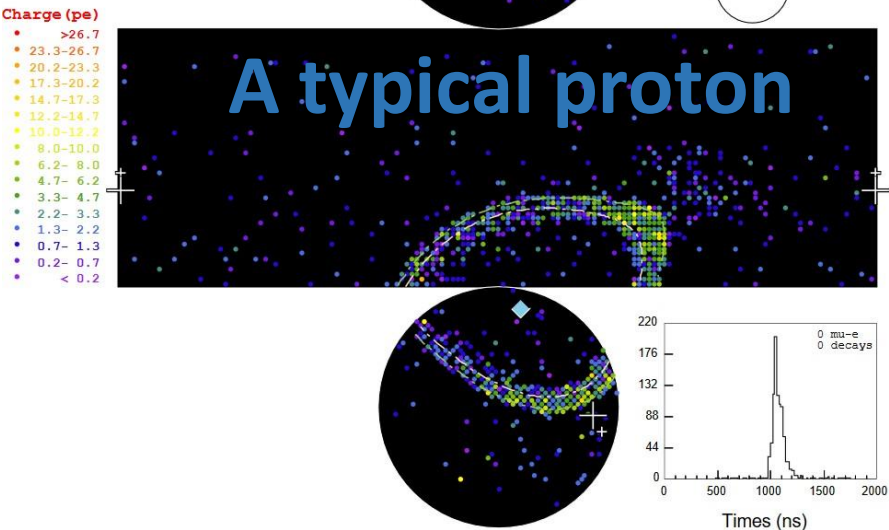
(Published paper, PRL 130,031802 (2023))

# Search for Cosmic-Ray Boosted Dark Matter (CRDM) Using Recoil Protons



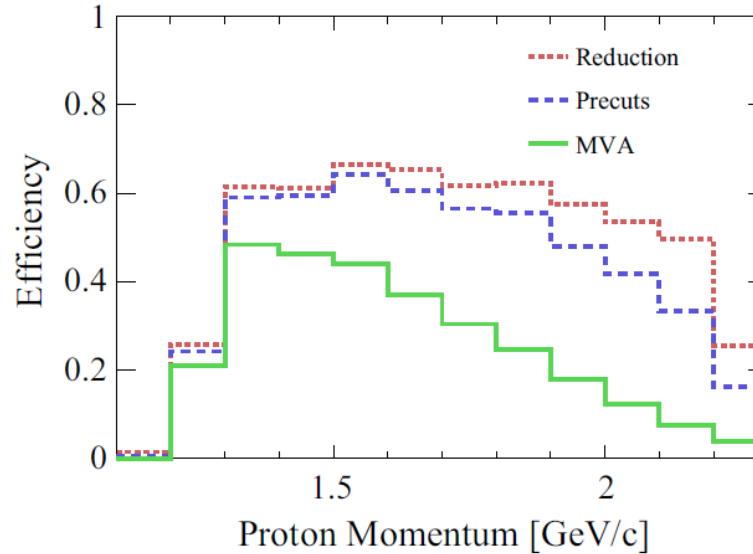
# CRDM search

Super-Kamiokande IV  
 Run 999999 Sub 12 Event 146  
 18-01-27:13:03:01  
 Inner: 928 hits, 2090 pe  
 Outer: 2 hits, 1 pe  
 Trigger: 0x07  
 D\_wall: 369.7 cm  
 Evis: 228.1 MeV  
 mu-like, p = 410.6 MeV/c

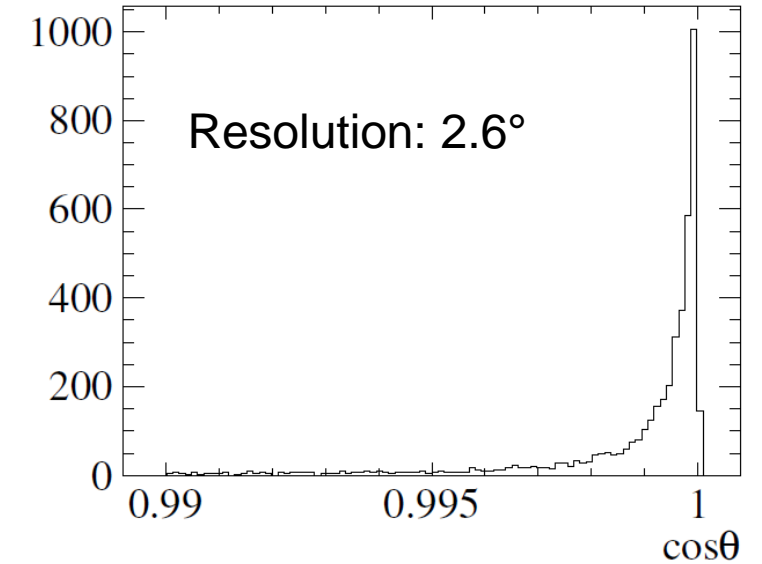


- Search for FCFV 1Ring proton-like events.
- Smaller opening angle, sharper ring edge than muon → make likelihood pattern fitting.
- Further  $\mu$ /proton separation by MVA

Proton detection efficiency



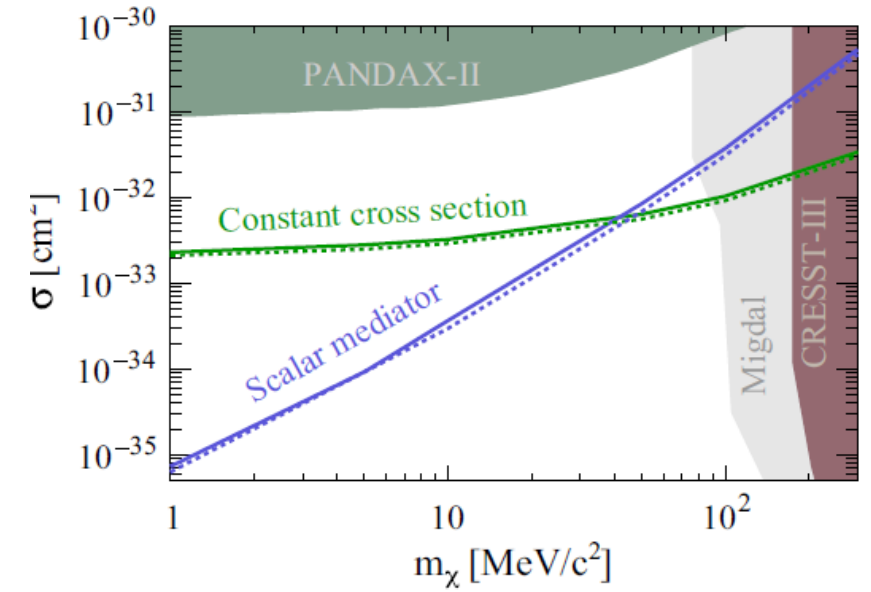
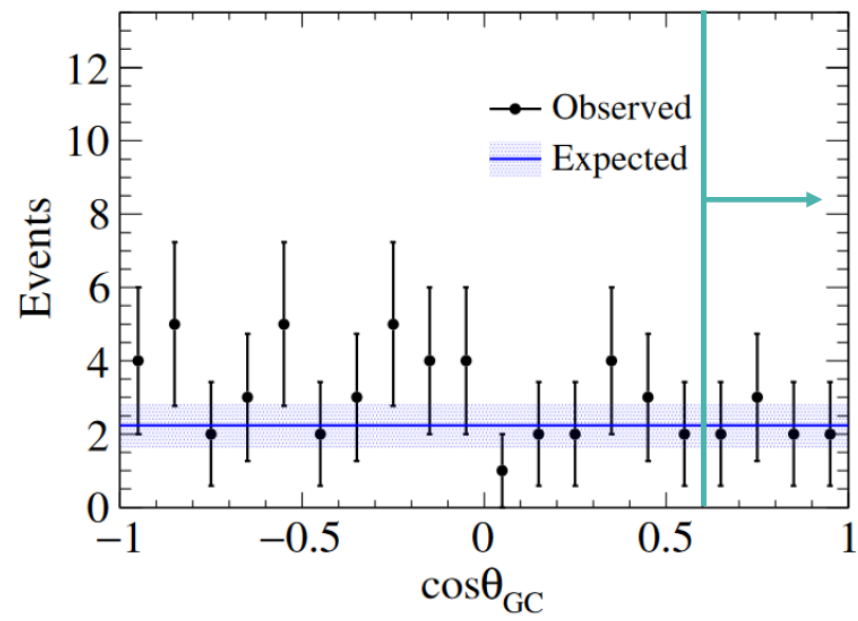
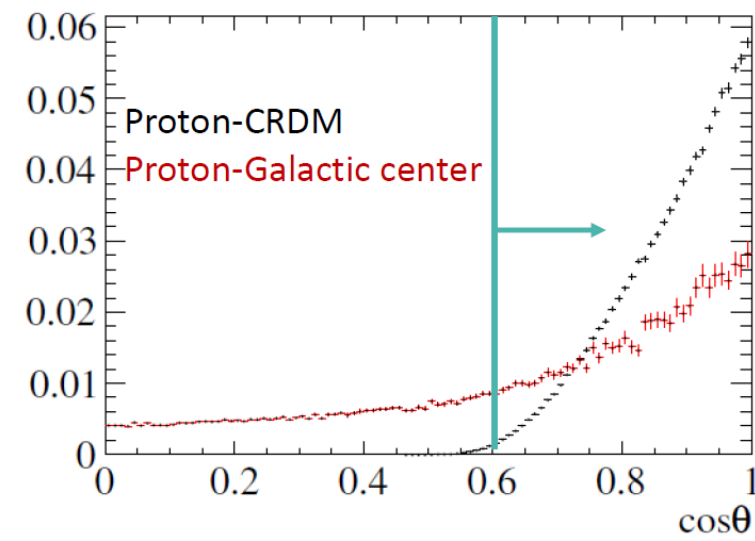
Angle between true and reconstructed proton direction



- CRDM is attenuated in the earth and only CRDM from above can reach to the detector.
- Apply zenith angle cut:  $\cos \theta_z > 0.2$



# CRDM search



Angular correlation between GC and proton direction. Signal region is defined as  $\cos\theta_{GC} > 0.6$

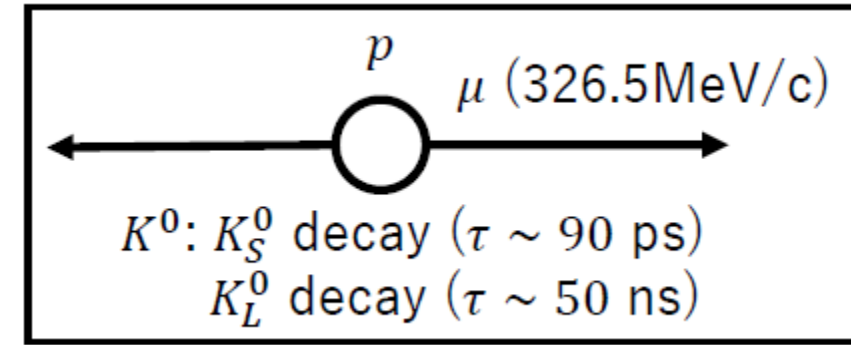
No excess found in GC direction.

Obtain constrain on dark matter-nucleon cross section. More than one order of magnitude better than existing results.

# (Published paper, PRD 106, 072003)

## Proton decay search: $p \rightarrow \mu^+ K^0$

- Favored by SUSY GUTs.
- $K^0$  is a mixing state of  $K^0_S$  and  $K^0_L$
- Results of SK-I~III (178 kt\*yr) have been published, (Phys.Rev.D85 (2012)112001)



- SK-IV data (200 kt\*yr) is newly analyzed with neutron tagging.

- Target  $K^0$  decays

➤  $K^0_S \rightarrow \pi^0 \pi^0, \pi^+ \pi^-$  ( $\sim 70 \text{ ps}$ )

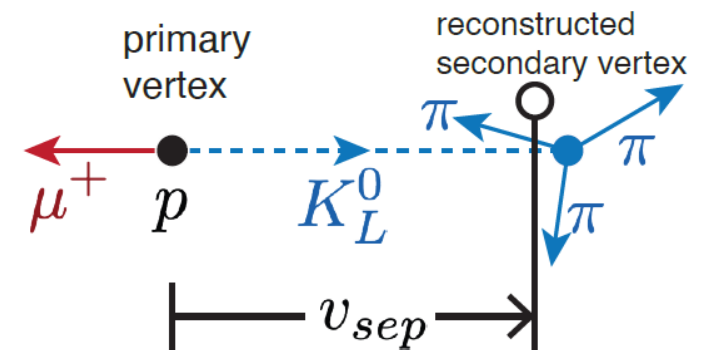
✓ All final particles can be detected in SK

➔ Reconstruct  $K^0$ , proton mass & momentum

➤  $K^0_L \rightarrow \pi^0 \pi^0, \pi^+ \pi^- \pi^0, \pi^\pm l^\mp \nu$  ( $\sim 50 \text{ ns}$ )

✓ (Challenging!) Have multiple vertices.

➔ Newly developed reconstruction tool for multi-vertex events.

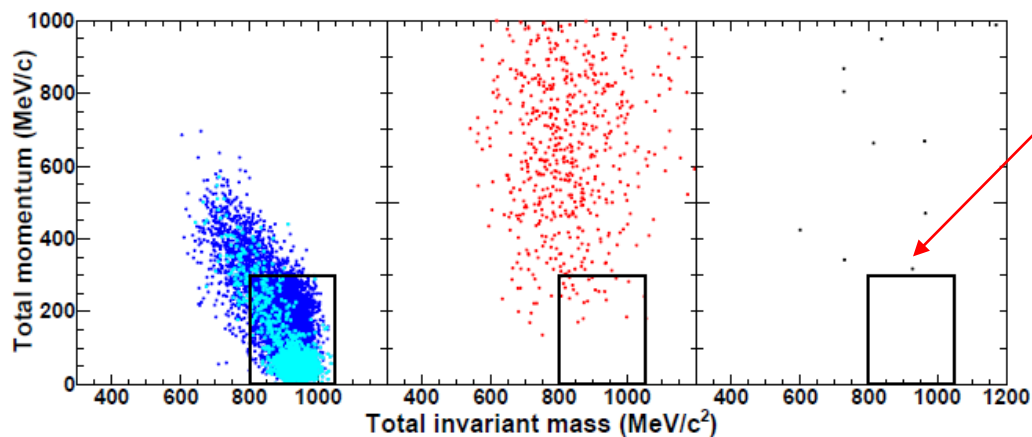
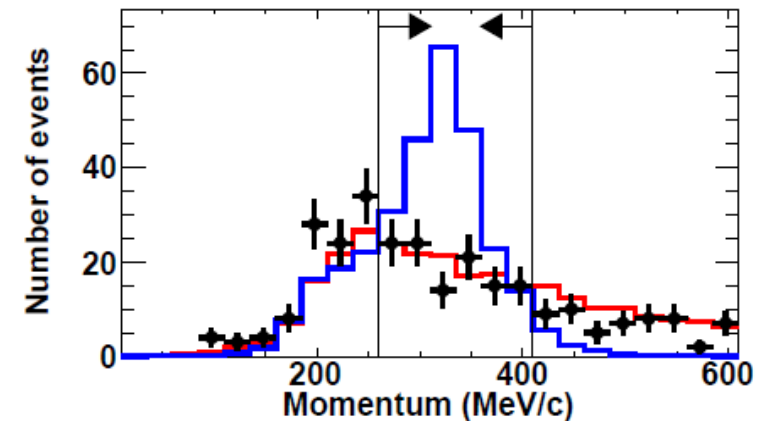
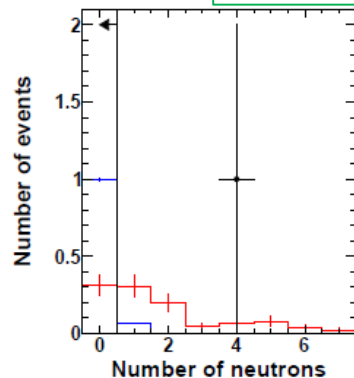
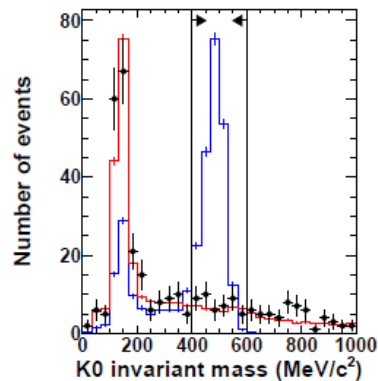
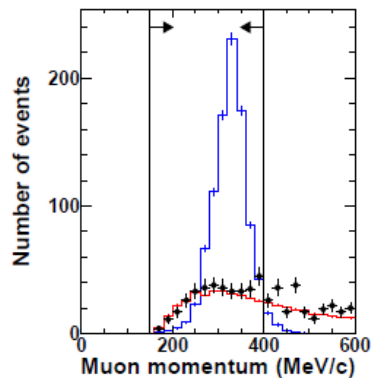


# Proton decay search: $p \rightarrow \mu^+ K^0$

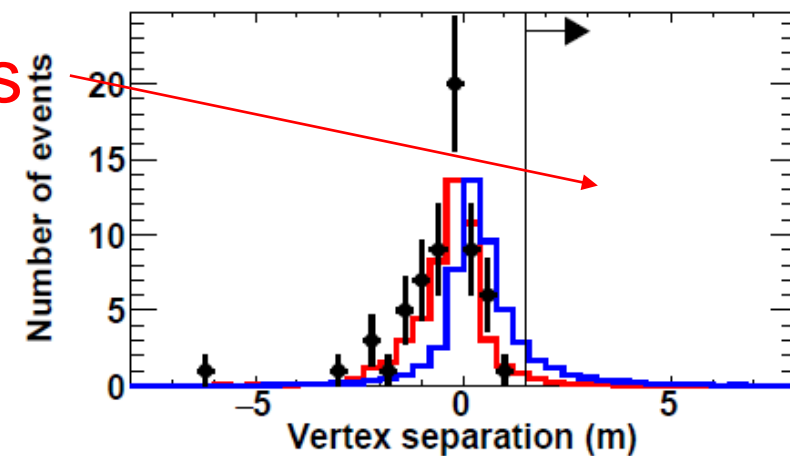
$K^0_S (\pi^0 \pi^0)$

Blue: Signal MC  
Red; Atm.v MC  
Black: Data

$K^0_L$



No candidates found.





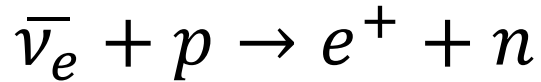
# Proton decay search: $p \rightarrow \mu^+ K^0$

## Results (SK-IV)

| Search mode                           | Efficiency (%)  | Background (events) | Candidates (events) | Lower limit ( $10^{33}$ years) |
|---------------------------------------|-----------------|---------------------|---------------------|--------------------------------|
| $K_S^0 \rightarrow 2\pi^0$            | $9.7 \pm 1.0$   | $0.31 \pm 0.14$     | 0                   | 2.7                            |
| $K_S^0 \rightarrow \pi^+ \pi^-$       | $4.98 \pm 0.54$ | $0.8 \pm 0.2$       | 0                   | 1.4                            |
| $K_L^0 \rightarrow \pi^\pm l^\mp \nu$ | $0.91 \pm 0.17$ | $1.0 \pm 0.3$       | 0                   | 0.2                            |
| $K_L^0 \rightarrow 3\pi^0$            | $0.36 \pm 0.06$ | $0.12 \pm 0.06$     | 0                   | 0.1                            |
| $K_L^0 \rightarrow \pi^+ \pi^- \pi^0$ | $0.18 \pm 0.04$ | $0.16 \pm 0.07$     | 0                   | 0.05                           |

Proton lifetime (90 % CL, SK-I~IV combined):  $> 3.6 \times 10^{33}$  years  
(Previous paper:  $> 1.6 \times 10^{33}$  years)

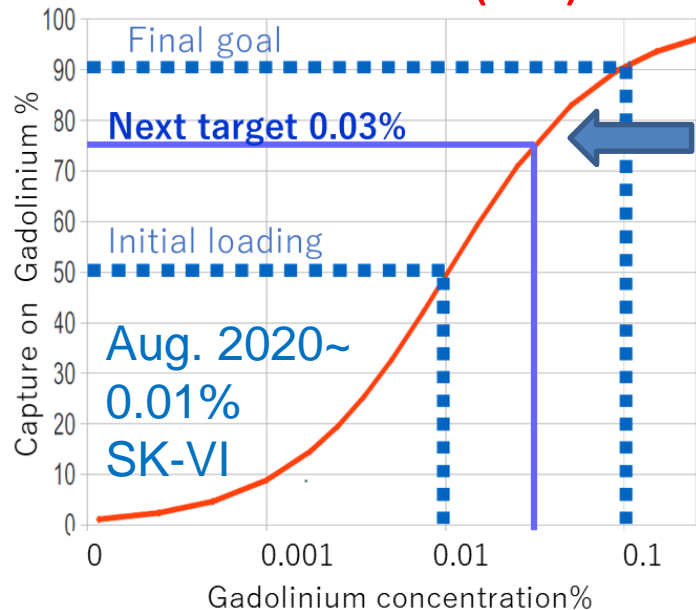
# Evolution of Super-K: SK-Gd



Gadolinium captures neutron and emit ~ 8 MeV  $\gamma$

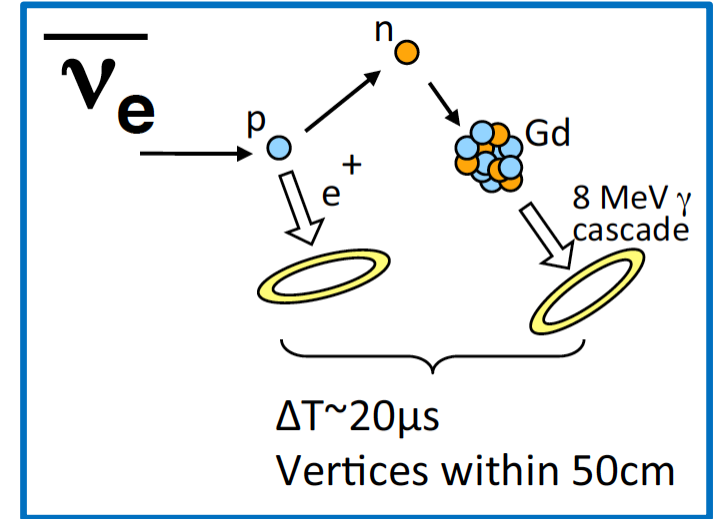
Detection efficiency of 8MeV  $\gamma$  ~ 100%

Add Gadolinium (Gd) to the SK water.



The latest news:  
We finished dissolving  
on July 5<sup>th</sup> !

**SK-VII started !**

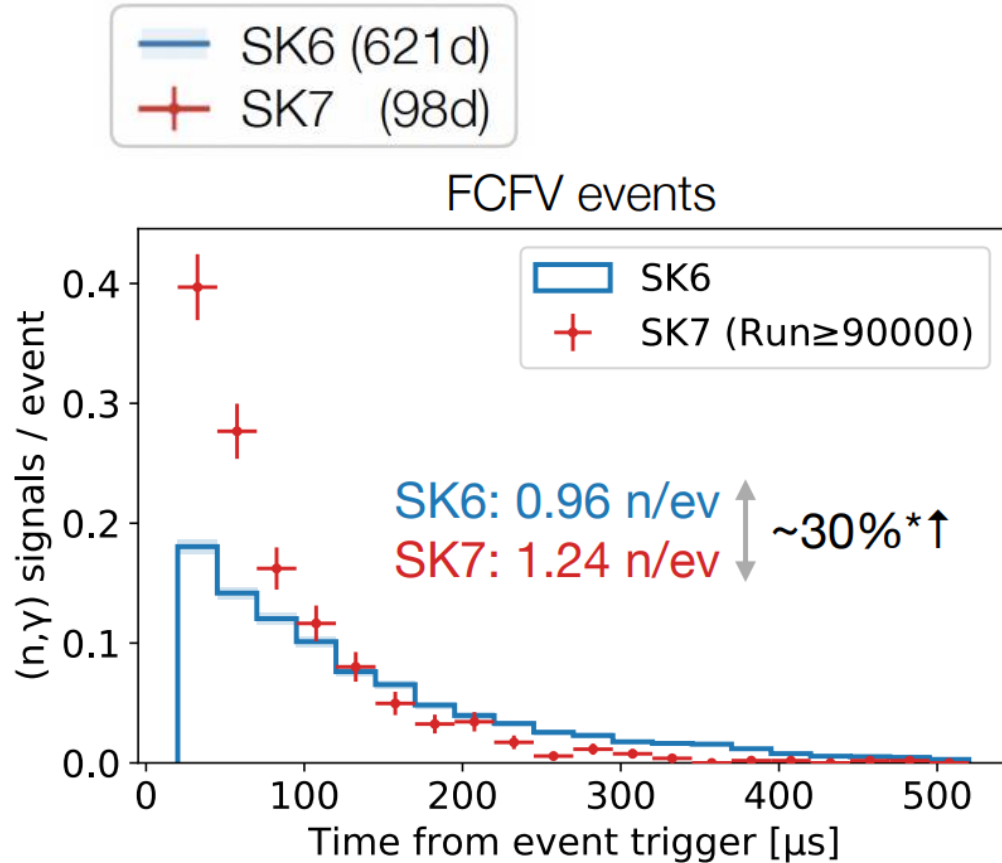


Neutrino / anti-neutrino discrimination

- Discovery of supernova(SN) diffuse  $\nu$  search and pointing accuracy improvement for SN burst
- Improve Discrimination power of  $\nu$  and  $\bar{\nu}$  in T2K and atmospheric neutrino analyses
- Nucleon decay background rejection

# Neutron capture time in SK6/7

# Atmospheric $\nu$ FCFV data



Am-Be neutron source

SK6  $\tau$ :  $116.5 \pm 0.2 \mu$ s

SK7  $\tau$ :  $61.6 \pm 0.1 \mu$ s

Fit  $Ae^{-t/\tau} + B$

SK6  $\tau$ :  $119.0 \pm 3.3 \mu$ s

SK7  $\tau$ :  $60.4 \pm 1.6 \mu$ s



*Faster captures with more Gd in SK7,  
consistent with calibration data*

Preparing MC for analysis.

# Summary

- Atmospheric neutrino oscillation:
  - Analyze all pure water phase (SK-I ~ V).
  - Expand fiducial volume.
  - favoring NO,  $\delta_{CP} \approx -\frac{\pi}{2}$ , and  $\sin^2\theta_{23} \approx 0.5$
- Published paper
  - Search for Cosmic ray Boosted Dark matter using recoil protons
  - Proton decay search  $p \rightarrow \mu^+ K^0$
- SK-Gd
  - SK-VI (Gd ~0.01 %) analysis is on going.
  - SK-VII (Gd ~0.03%) just started from July !