# SEARCH FOR NEUTRINOS EVENTS COINCIDENT WITH GRB 221009A IN SK

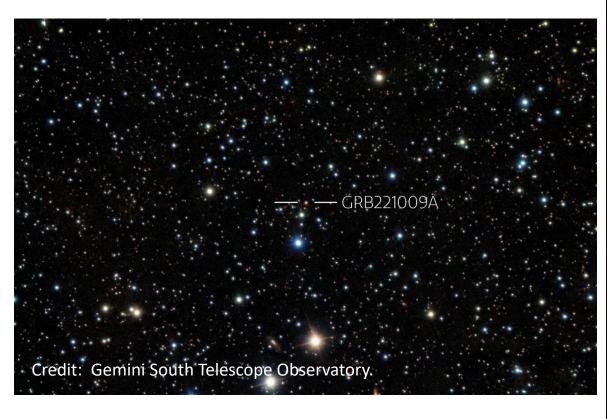
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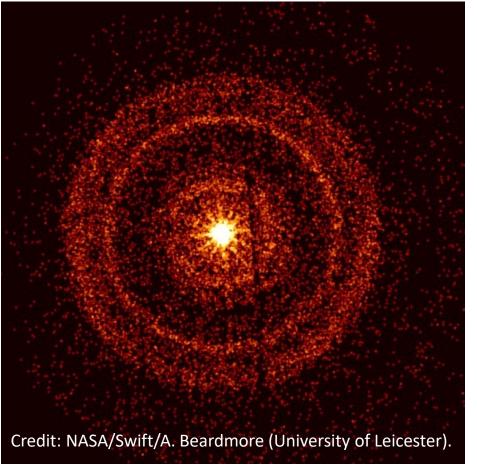


#### INTRODUCTION

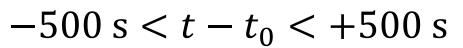


On October 9, 2022, a gamma ray burst GRB 221009A was observed by several observatories. This is one of the closest and most luminous gamma ray bursts, and is also the first time to detect photons above 10 TeV from a gamma ray burst.

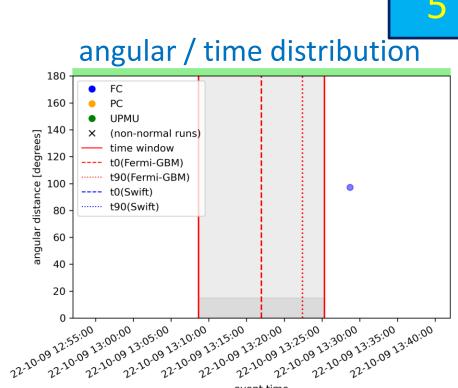




#### **RESULTS**

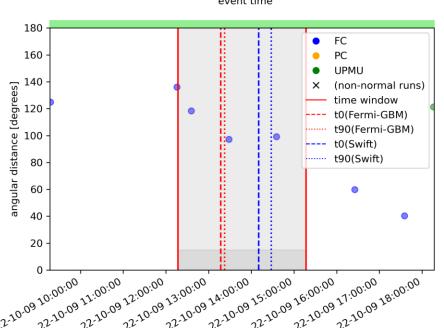


	FC	PC
Observed	0	0
Expected	0.10	$6.0 \times 10^{-3}$
p value	100%	100%
N90	2.3	2.3



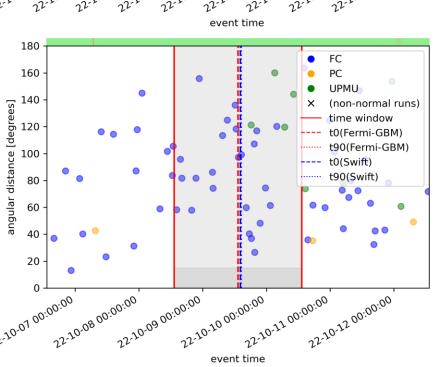
<b>-</b> 1	h	<	t	 $t_{0}$	<	+2	h
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<b>O</b>				
	FC	PC		
Observed	3	0		
Expected	1.1	$6.5 \times 10^{-2}$		
p value	10.4%	100%		
N90	5.6	2.3		



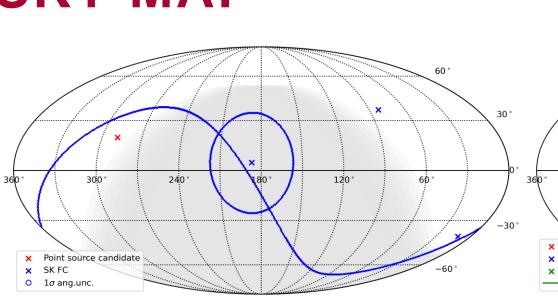
$$-1 d < t - t_0 < +1 d$$

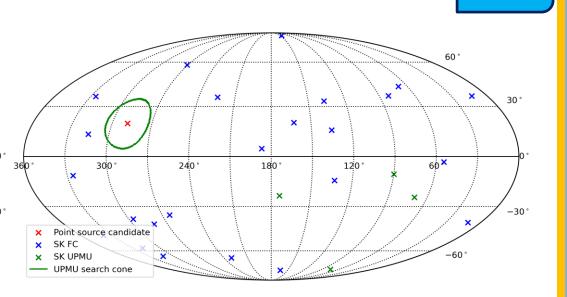
	FC	PC	UPMU
Observed	24	0	0
Expected	17.9	1.0	0.01
p value	9.8%	100%	100%
N90	13.7	2.3	2.3



- ◆ GRB 221009A is out of UMPU sensitive region during time window[-500s, +500s] and [-1h, +2h].
- ◆ Number of expected background events was estimated using off-time data.
- According to p value no significant neutrino emission excess against background was found, thus number of signal event upper limit was set.

# SKY MAP





- $-1 \text{ h} < t t_0 < +2 \text{ h}$
- $-1 d < t t_0 < +1 d$
- Gray shadow area is UPMU sensitive region.
- ◆ Angular uncertainty is estimated using SK4 MC.
- ♦ 15° search cone is set for UPMU.

#### **OBSERVATION**

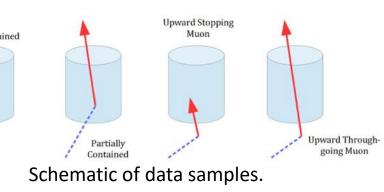


- **□** Swift-BAT:
  - > Trigger time = 2022/10/09 14:10:17 UTC
  - $\rightarrow$  T90 = 1068.4s
- □ Fermi-GBM:
  - > Trigger time = 2022/10/09 13:16:59 UTC
  - $\rightarrow$  T90 = 325.8s
- □ LHASSO:
  - > Photos above 10 TeV observed.
- □ IceCube:
  - > No significant neutrino emission was observed.

### **DATA SAMPLE**



- ☐ Using high energy neutrino data above 30 MeV in SK:
  - ➤ Fully contained: > 0.03 GeV
  - Partially contained: > 0.35 GeV
  - ➤ Upward-going muon: > 1.6 GeV



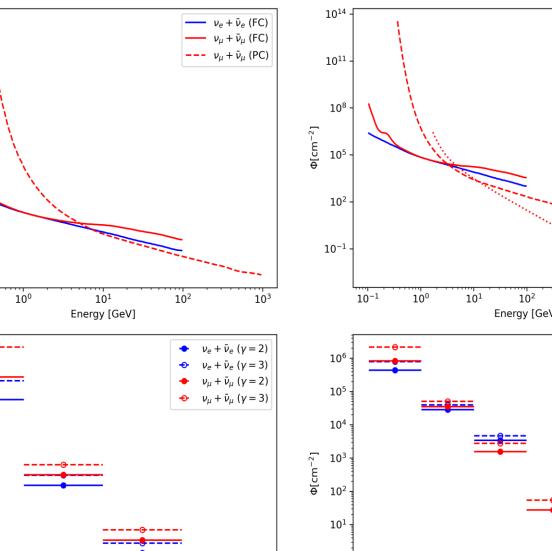
---  $\nu_e + \bar{\nu}_e$  (FC)

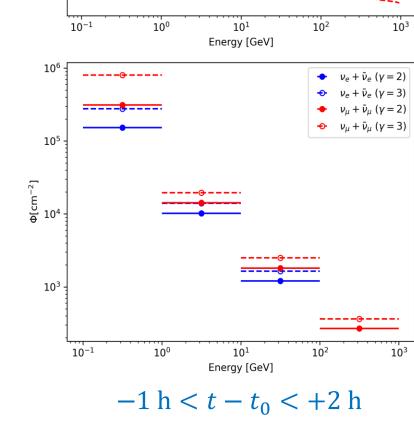
## **SEARCH TIME WINDOW**

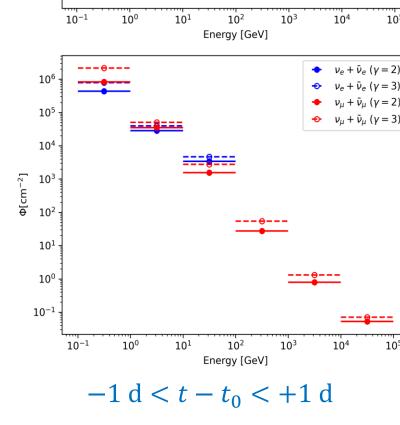


- $\Box$  Center of time  $(t_0)$  chosen as trigger time of Fermi-GBM.
  - > +/- 500s (Normal GRB search time window)
  - > +2h/-1h (Consistent with IceCube search)
  - > +/-1d (Extended search time window)

# **FLUENCE UPPER LIMIT**







- ◆ Model independent fluence limit (top).
- Model dependent fluence limit (assume  $E^{-\gamma}$  spectrum) (bottom).

# **SUMMARY**



Searched for neutrino events in coincidence with GRB 221009A with ATMPD:

Three time-window ([-1h, +2h], [-500s, +500s], [-1d, +1d]) used.

No significant excess of neutrino emission found.

Two types of flux upper limit calculated:

- 1. Model independent fluence
- 2. Fluence assuming power law spectrum emission model