

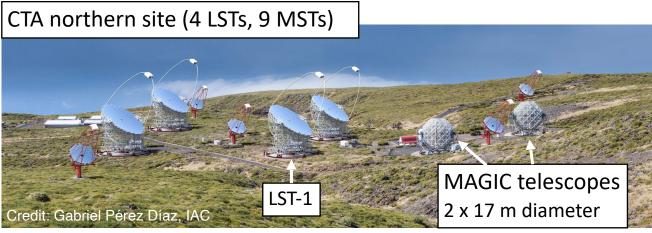


# CTA大口径望遠鏡初号機と MAGIC望遠鏡による研究成果

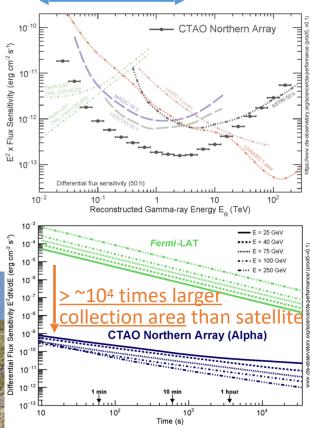
#### Ryuji Takeishi for the CTA-LST and MAGIC Collaborations 2024.2.21 令和5年度共同利用研究成果発表会

## Cherenkov Telescope Array (CTA)

- <u>New gamma-ray observatory</u> under construction
- 4 LSTs will be set at northern site in La Palma, Spain, alongside 9 MSTs.
- Compared to current telescopes,
  - 10 times better sensitivity
  - 10 times wider energy range: 20 GeV 300 TeV
- We started LST-1 operation from 2018.

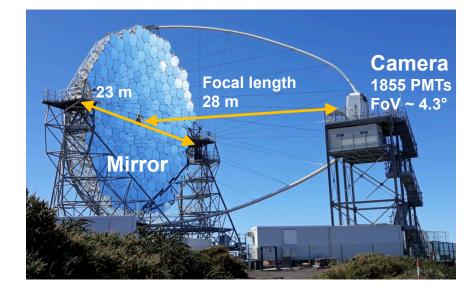


#### LST energy range

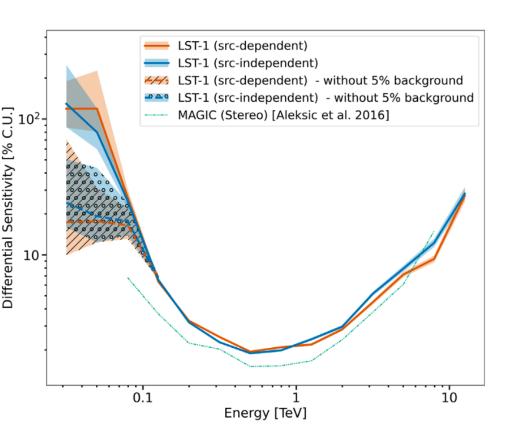


# CTA large-sized telescope (LST)

- 23 m diameter: over 400 m<sup>2</sup> mirror area
- Targeting an energy threshold ~20 GeV
- Stereo observations at lowest energy ever observed from ground
- Ability to reposition to any point in the sky within 20 seconds
- Ideal for fast transients and soft sources



## LST-1 performance



#### Flux Sensitivity

(zenith angle < 35 deg, 50 hours)

- Energy range widened to lower energy compared to MAGIC
- ~1.5 times less sensitive than MAGIC (stereo system)

   consistent with single telescope performance vs stereoscopic system
- Astrophysical Journal **956**, 80 (2023)

## Crab Nebula and Pulsar

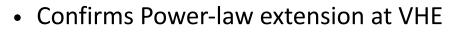
Crab Nebula spectrum

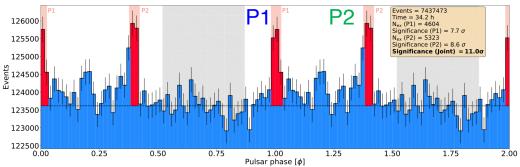
- 34.2 hours of data with zenith < 35°</li>
- Systematic errors: gray points • correspond to the effect of +1% background
- Compatible with MAGIC and *Fermi*-LAT •

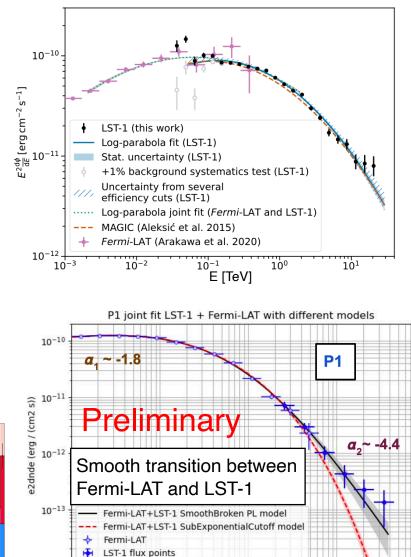


#### Crab pulsar

Clear detection down to a few tens of GeV







 $10^{-14}$ 

 $10^{-1}$ 

100

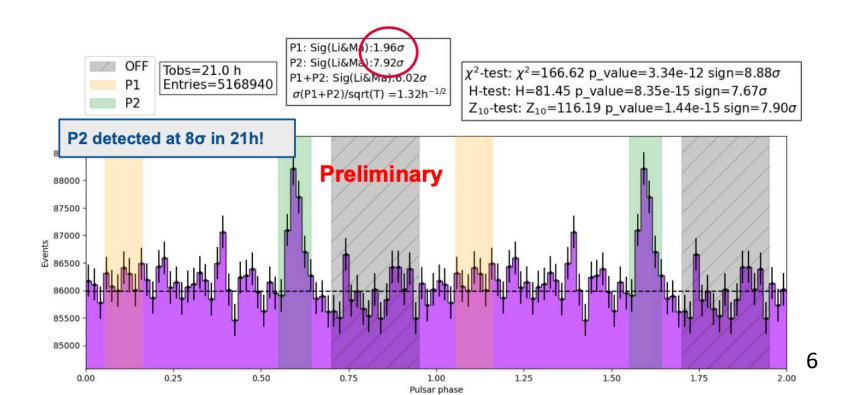
10<sup>1</sup>

Energy [GeV]

10<sup>2</sup>

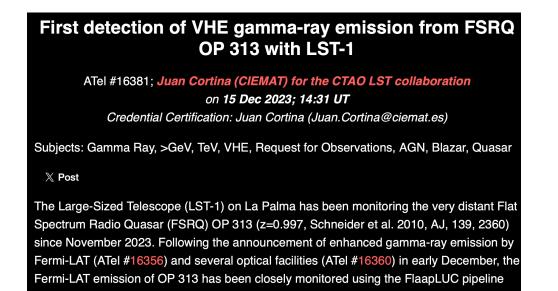
### Geminga Pulsar

- One of the brightest pulsars detected by Fermi-LAT at few GeV
- 8σ detection for P2 in 21 hours of data
   MAGIC: 6.2σ sigma in ~80 hours
- This shows excellent performance of LST-1 at ~10 GeV.



#### Discovery of AGN OP 313 at VHE gamma-rays

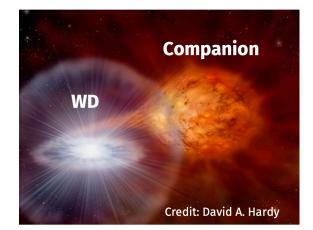
- Flat Spectrum Radio Quasar (FSRQ), Redshift: z = 0.997
- Detected Dec. 2023 flare (ATel #16381)
- Most distant source with LST-1.
- This shows excellent performance of LST-1 for distant sources.
- Study of EBL is ongoing.

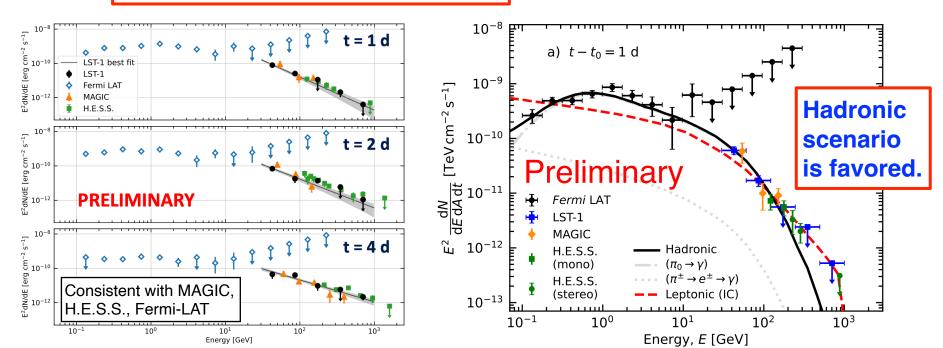


### Detection of Nova RS Ophiuchi

- recurrent symbiotic nova, d~2.69 kpc
- First detected recurrent nova in VHE gamma rays by 2021 outburst (H.E.S.S. and MAGIC)
- LST-1 took part in the first VHE gamma-ray detection with >~6σ in each night

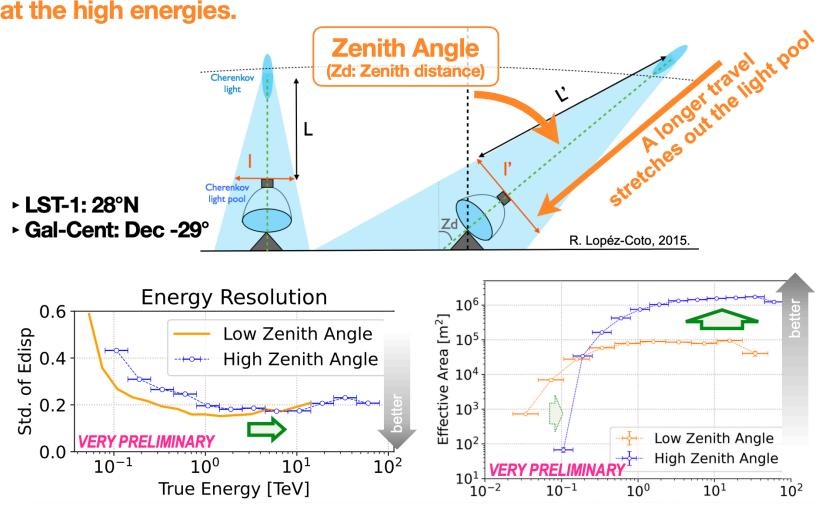
Novae established as new source class at VHE gamma rays





### Large-zenith angle observation for Galactic Center

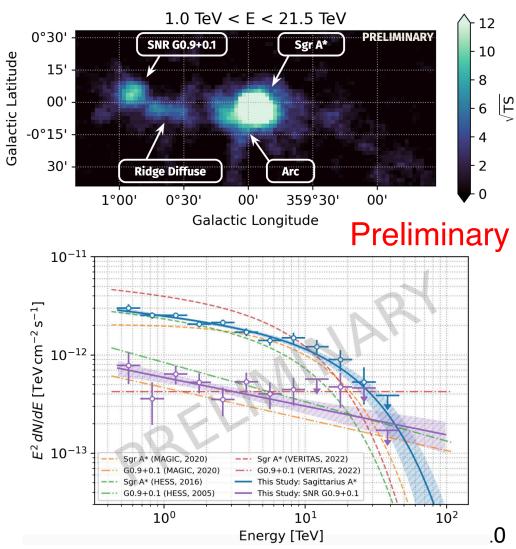
Large-zenith-angle observations (55-70 deg) enlarge the effective area at the high energies.



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### **Galactic Center**

- **39-hour data** in 2021/2022 after selection
- Analysis is carried out through a dedicated special background modeling.
- Sgr A\* & SNR G0.9+0.1 SEDs are in line with results from other telescopes.
- Successful extended-source observations
- Observation is ongoing.



### **BL** Lacertae

2.00

1.75

1.50

1.25

1.00

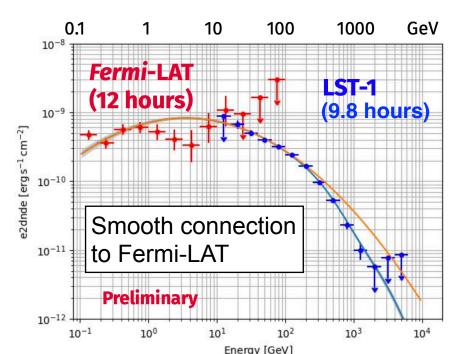
0.75

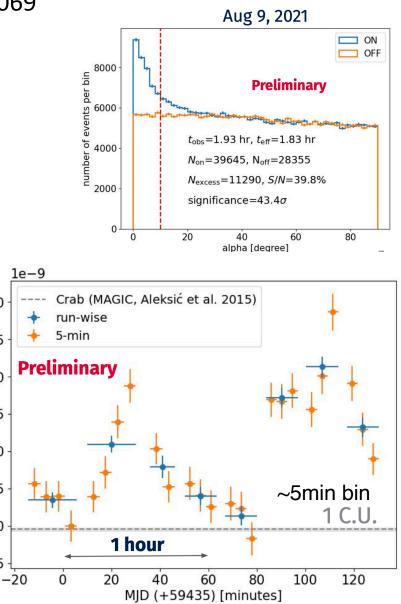
0.50

0.25

 $IuX_{E > 100 \text{ GeV}} (cm^{-2} \text{ s}^{-1})$ 

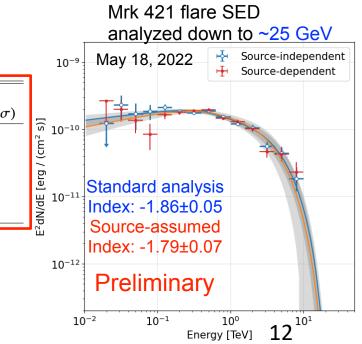
- Intermediate frequency peaked BL Lac at z = 0.069
- Flare in 2021 July and August was detected.
- Energy spectrum down to ~30 GeV
- Intra-night variability with sub-hour-scale was detected up to 3-4 Crab unit.
- MWL SED study is ongoing.
- QG and Relativity tests are ongoing.





#### AGN zoo

- LST array aims to observe AGN up to z~2.
- More than one thousand hours of data taken with LST-1 from 2020: Near sources: Mrk 421, Mrk 501, 1ES 1959+650, ...
   Distant sources: 1ES 0647+250, PG 1553+113, ...
- All known TeV blazers, and detected (>5 $\sigma$ ) up to a z~0.45

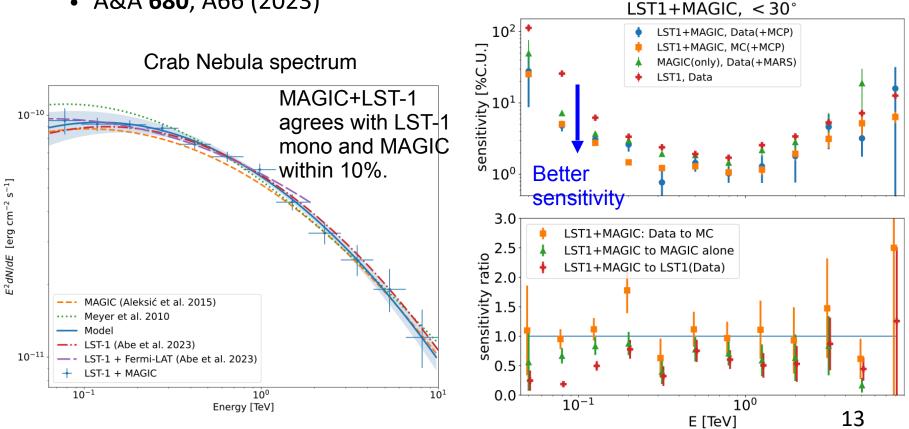


Observation summary

Source	Observation date	Redshift	Observation time	Detection
			before/after cut (h)	significance ( $\sigma$ )
Mrk 421	2020 Dec. 12 - 2022 May 23	0.031	68.5 / 31.9	53
Mrk 501	2020 July 10 - 2022 May 22	0.034	67.2 / 39.7	21
1ES 1959+650	2020 July 11 - 2022 May 5	0.048	21.3 / 11.8	13
1ES 0647+250	2020 Dec. 16 - 2020 Dec. 21	$0.45 \pm 0.05$	8.8 / 8.2	7
PG 1553+113	2021 Apr. 8 - 2022 May 23	0.433	12.2 / 9.9	16

#### LST+MAGIC

- LST and MAGIC joint observation and analysis method
- Allow detection of 30% (40%) lower fluxes than MAGIC alone (LST-1) alone) (Current best sensitivity at tens of GeV in northern IACT)
- A&A 680, A66 (2023)



#### MAGIC publication in refereed journals (2023)

#### 8 papers:

First characterization of the emission behavior of Mrk421 from radio to VHE gamma rays with simultaneous X-ray polarization measurements MAGIC collaboration, Abe *et al.* Accepted in 2023 for publication in Astronomy & Astrophysics

MAGIC detection of GRB 201216C at z=1.1 MAGIC collaboration, Abe *et al.* Monthly Notices of the Royal Astronomical Society **527** 5856–5867 (2024), accepted in 2023

A lower bound on intergalactic magnetic fields from time variability of 1ES 0229+200 from MAGIC and Fermi/LAT observations MAGIC collaboration, Acciari *et al.* Astron. & Astroph. **670** (2023) A145

Long-term multi-wavelength study of 1ES 0647+250 MAGIC collaboration, Acciari *et al.* Astron. Astrophys. **670** (2023) A49

MAGIC observations provide compelling evidence of the hadronic multi-TeV emission from the putative PeVatron SNR G106.3+2.7 MAGIC collaboration, Abe *et al.* Astron. & Astroph. **671** (2023) A12

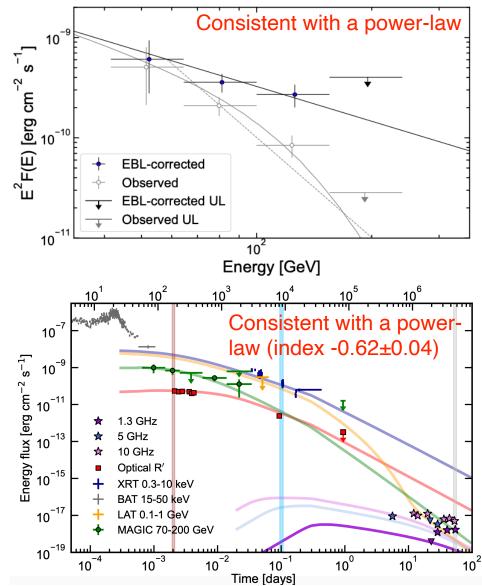
Multimessenger Characterization of Markarian 501 during Historically Low X-Ray and gamma-Ray Activity MAGIC collaboration, Abe *et al.* Astroph.J. Suppl. **266** (2023)

Search for Gamma-ray Spectral Lines from Dark Matter Annihilation up to 100 TeV towards the Galactic Center with MAGIC MAGIC collaboration, Abe *et al.* Phys. Rev. Lett. **130** (2023) 061002

Study of the GeV to TeV morphology of the gamma-Cygni SNR (G78.2+2.1) with MAGIC and Fermi-LAT MAGIC collaboration, Acciari *et al.* Astron. & Astroph. **670** (2023) A8

#### MAGIC highlights Detection of GRB 201216C

- Most distant GRB detected in VHE range, z = 1.1
- MAGIC observed from T<sub>0</sub> + 56 sec and detected between 70 and 200 GeV.
- From the light curves of the optical/sub-TeV fluxes, wind-like ambient medium density profile is suggested.
- MNRAS 527, 5856–5867 (2024)



# Summary

- LST-1 is continuing observation and performing scientific observation.
- Crab Nebula, pulsar: Significant detection down to few tens of GeV
- Geminga pulsar: significant detection of soft spectrum source
- Nova RS Ophiuchi: LST-1 took part in the first VHE gamma-ray detection with >~6σ in each night
- Blazar BL Lacertae: Flux variability with sub-hour-scale is observed.
- AGN zoo: detected (>5 $\sigma$ ) up to a z~0.45
- Galactic center: successful extended-source observation
- MAGIC+LST method: 30-40% better sensitivity was achieved
- Observing FSRQ OP 313
- Observing GRB following burst alerts