

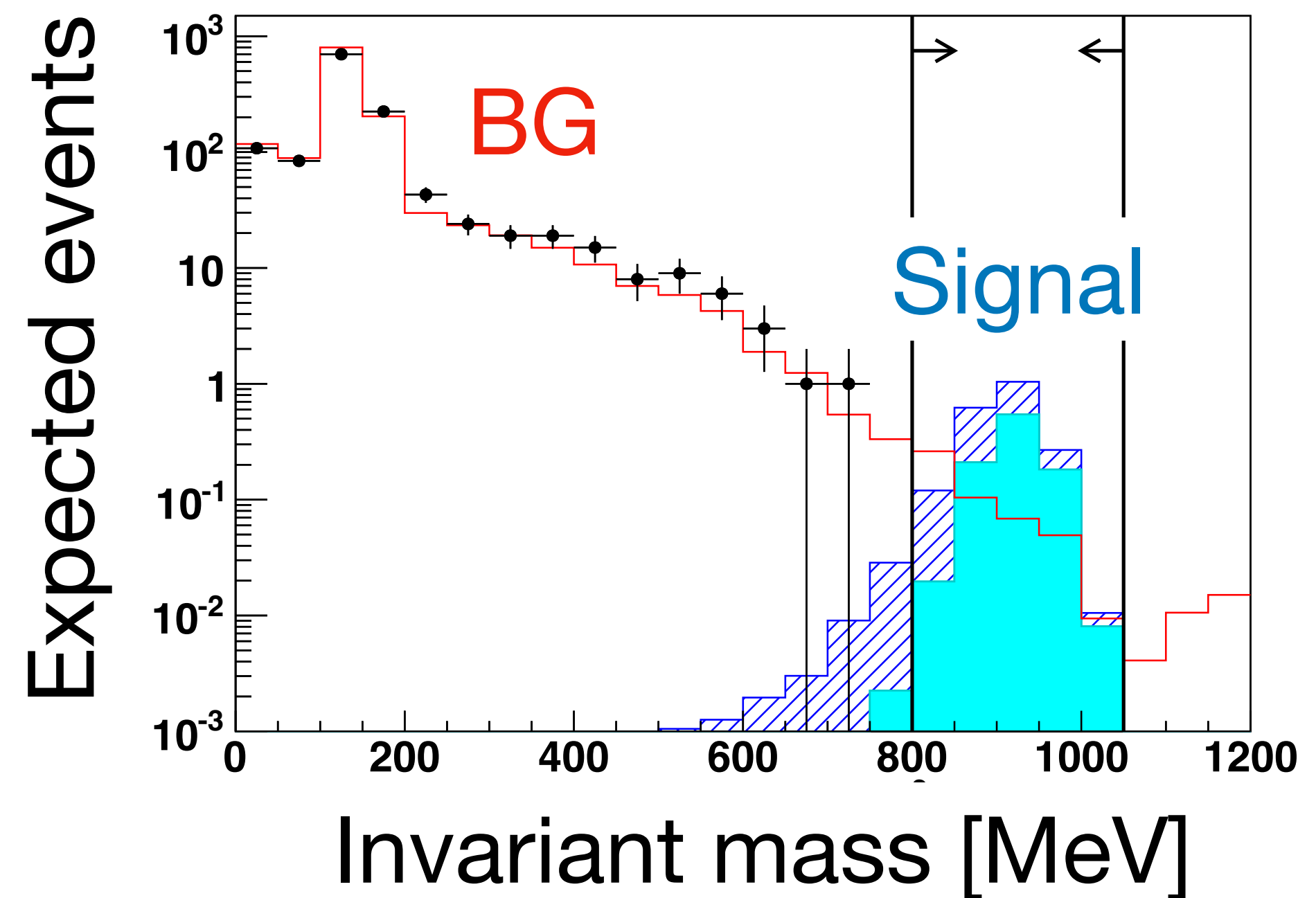
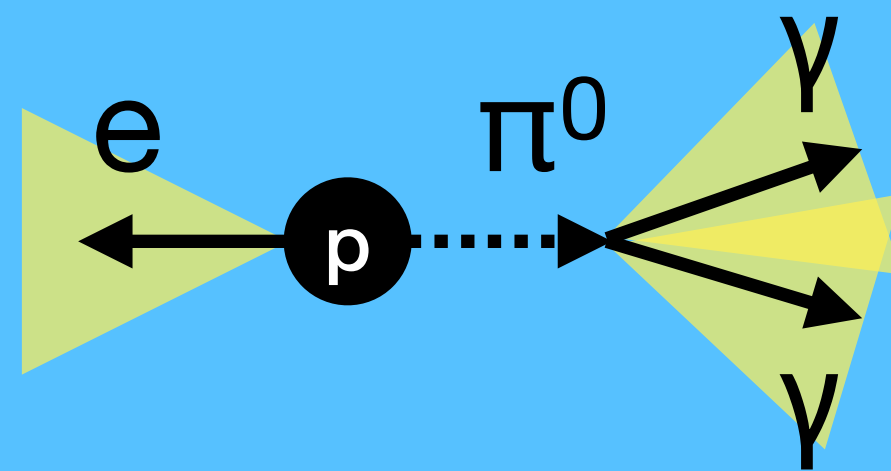
# Super-Kamiokande Results in FY2023: Proton Decays and Atmospheric Neutrinos

Seungho Han (ICRR)

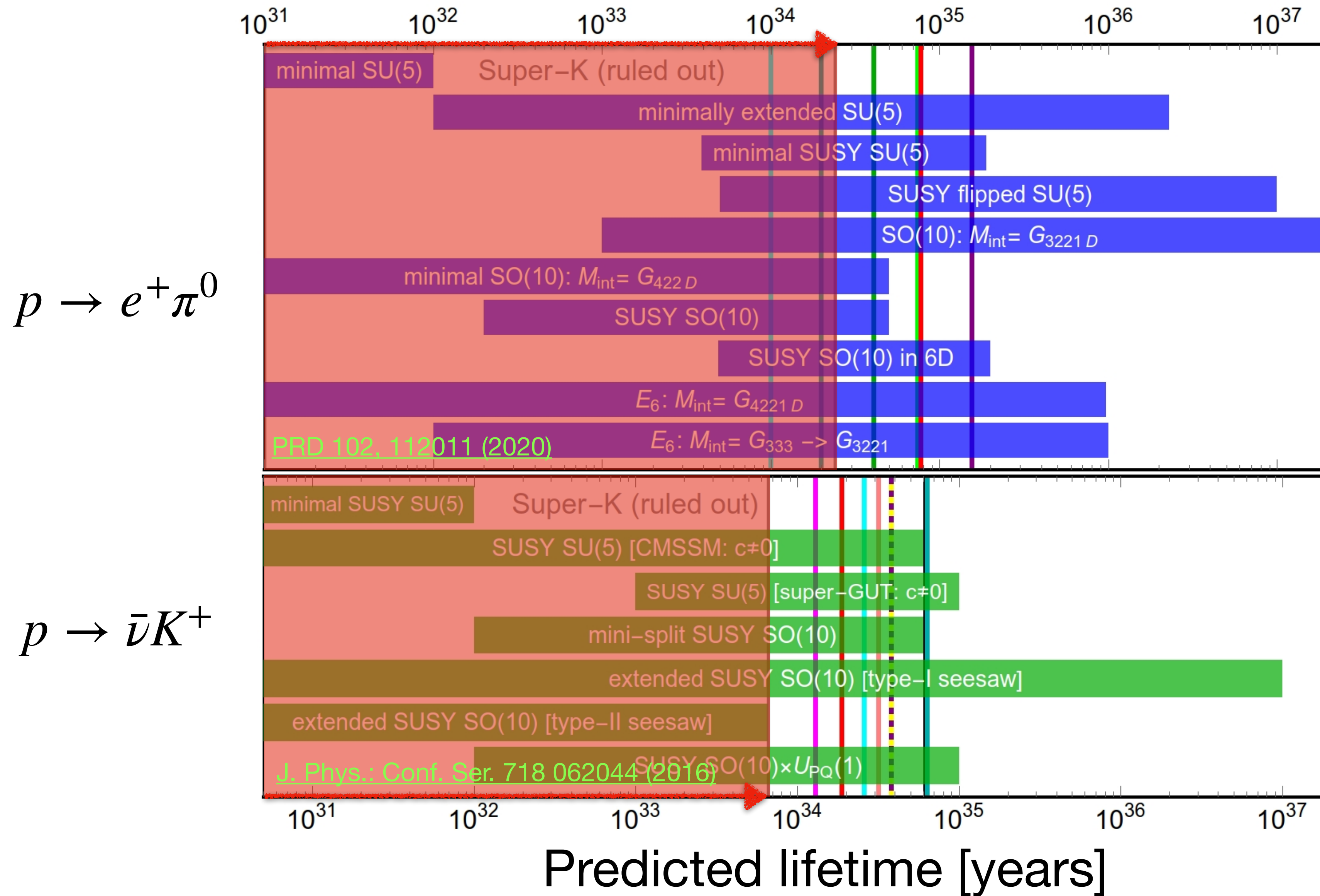
# Proton decay

- Quark  $\rightarrow$  lepton (e.g.  $p \rightarrow e^+ \pi^0$ ,  $p \rightarrow \bar{\nu} K^+$ )
- Lifetime:  $\tau_p \sim O(10^{34}-10^{35})$  years?
- Signal with invariant mass  $\sim m_p \sim 1$  GeV

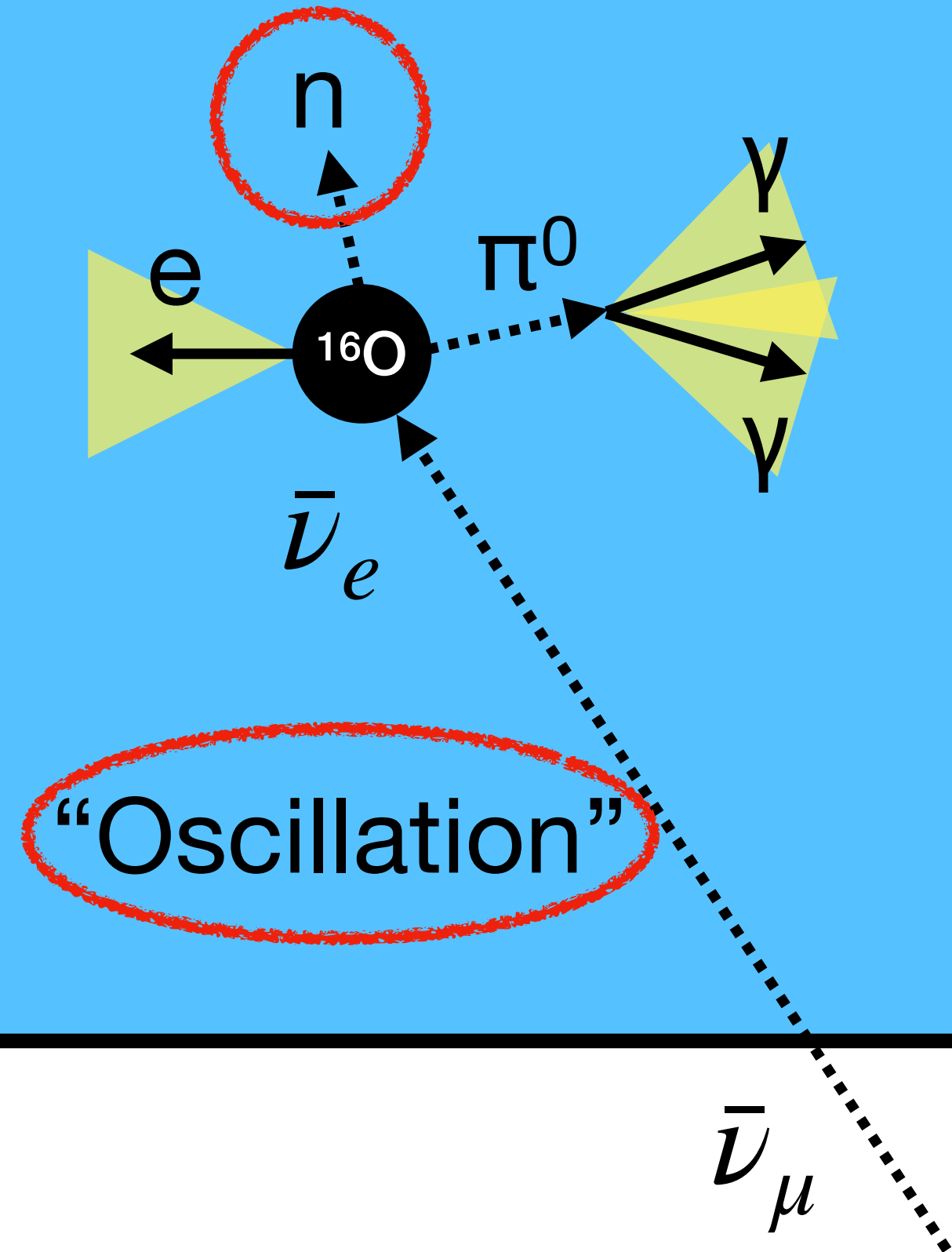
Example



# SK (90% CL)



Example



# Atmospheric neutrinos

- Byproduct of CR air showers
- All directions, easily pass through the Earth
- 😓 BKG for proton decays → **n-tag reduction**
- 😊 Abundant, diverse energy and path lengths

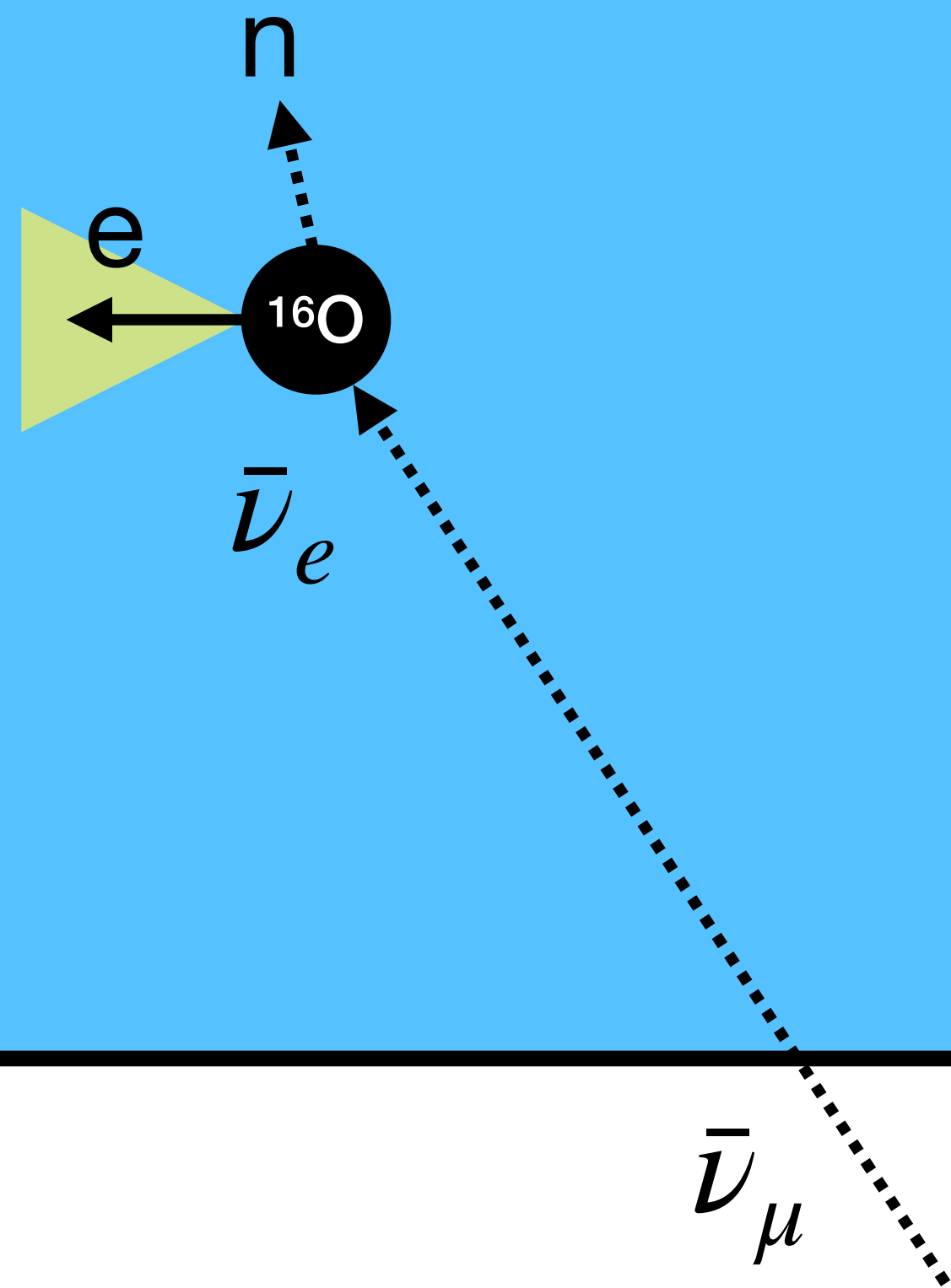
# Oscillation studies

$$P(\nu_\alpha \rightarrow \nu_\beta) \approx f(\theta_{ij}, \delta_{CP}, \Delta m_{ij}^2; L/E)$$

$\nu/\bar{\nu}$ , flavor      Parameters Kinematics

Mixing angles, CP phase, mass splittings

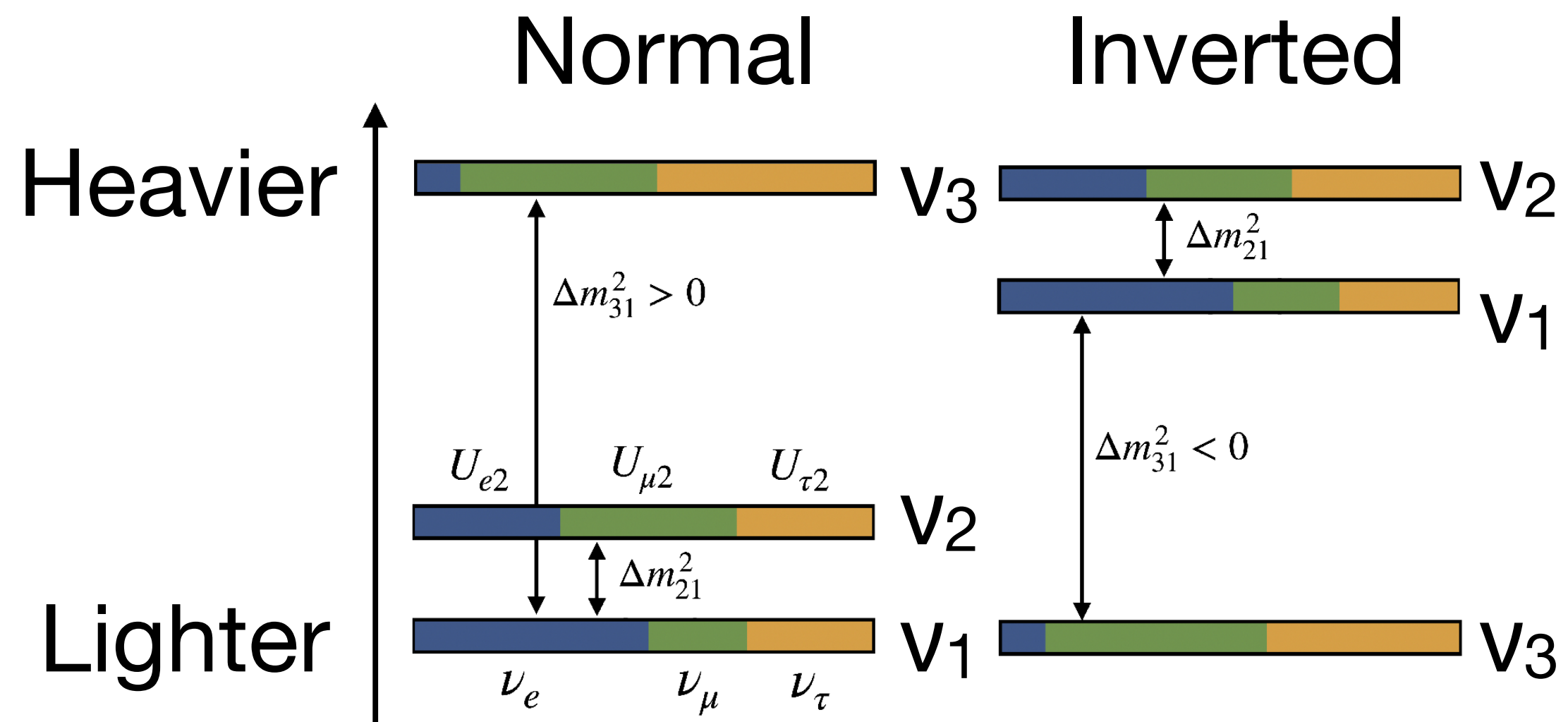
Example



# Important questions

## Mass Ordering

$$e < \mu < \tau \sim 1 < 2 < 3 ?$$



## CP symmetry

$$P(\nu_\alpha \rightarrow \nu_\beta) = P(\bar{\nu}_\alpha \rightarrow \bar{\nu}_\beta) ?$$



Second-order effects: require large statistics and accurate neutrino reconstruction

# Oscillation studies

$$P(\nu_\alpha \rightarrow \nu_\beta) \approx f(\theta_{ij}, \delta_{CP}, \Delta m_{ij}^2; L/E)$$

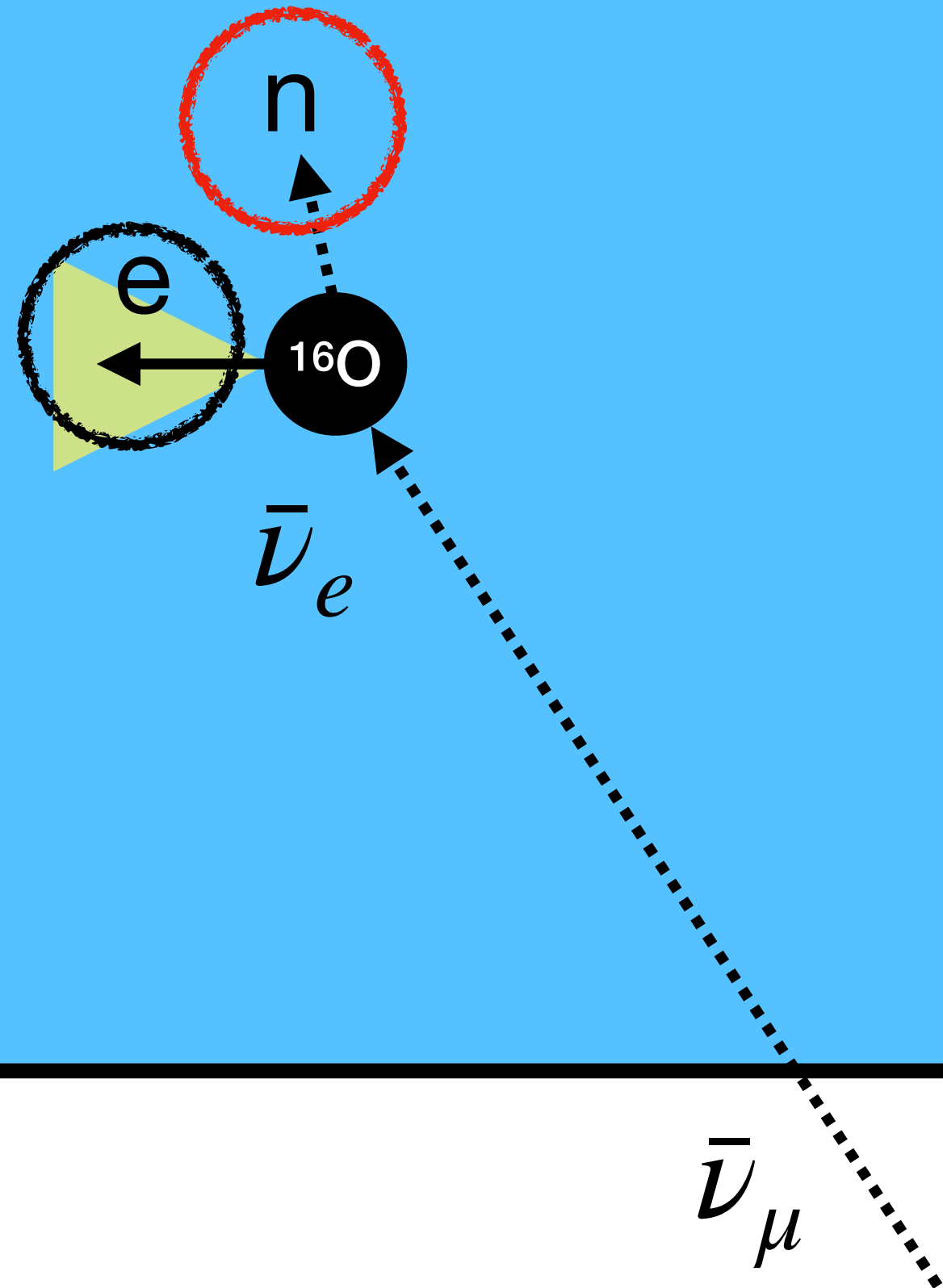
$\nu/\bar{\nu}$ , flavor      Kinematics

Rely solely on outgoing e/ $\mu$ , limits in accuracy

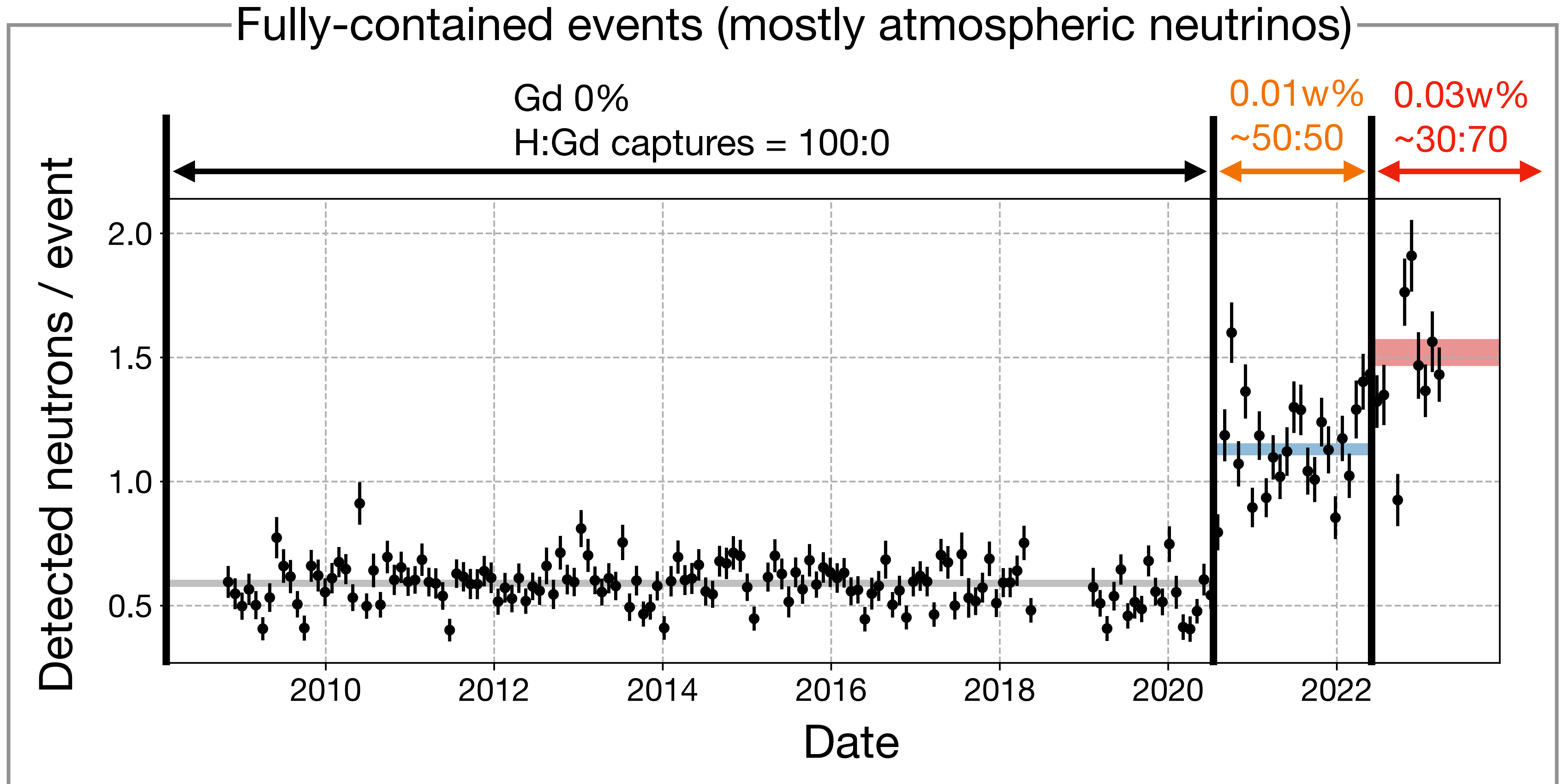
Difficult to access the “second-order effects”

Again, neutron-tagging can help improve  $\nu/\bar{\nu}$  ID and  $\nu$  kinematic reconstruction

Example



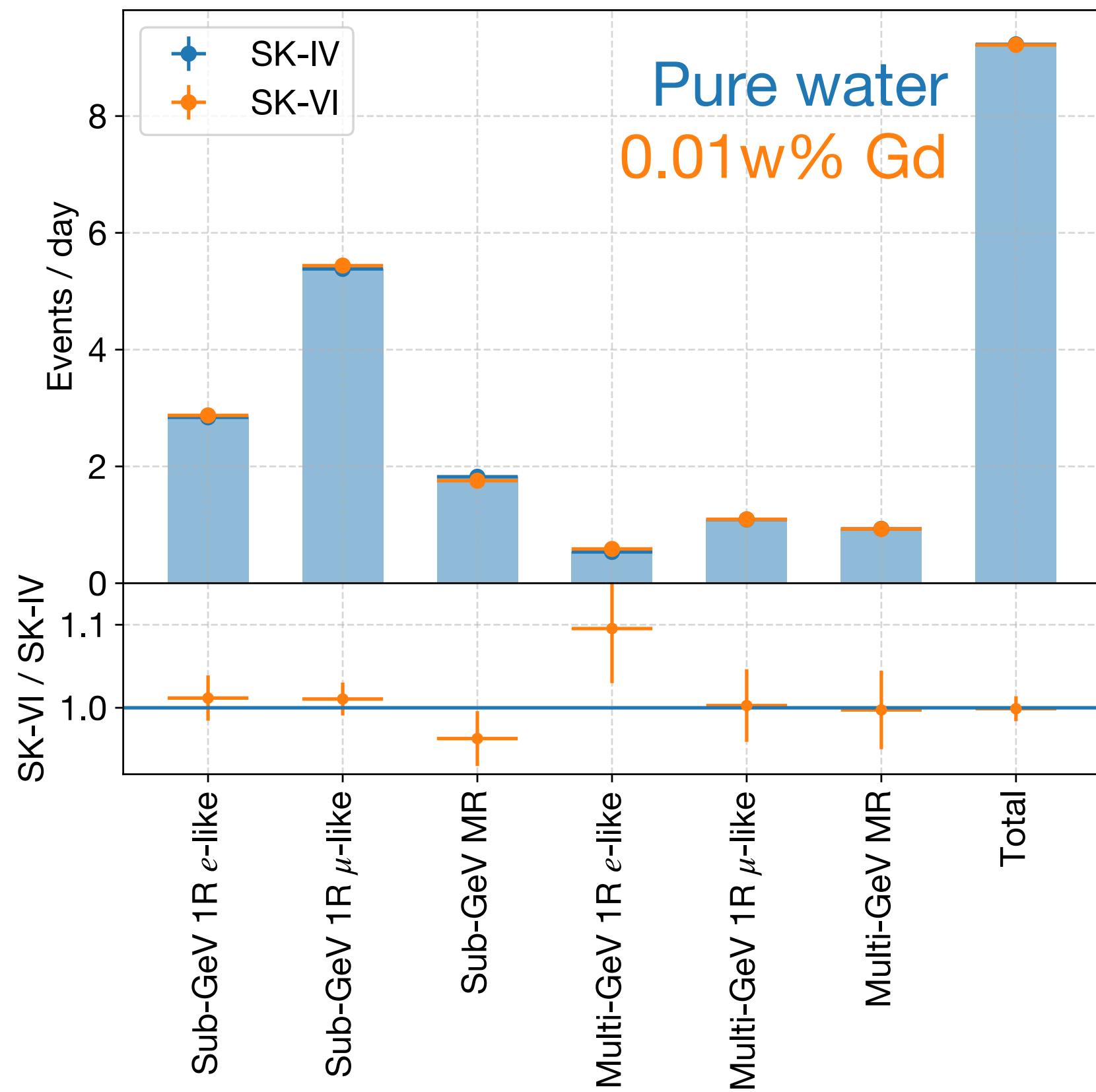
# Gd-loading effect on neutron detection rate



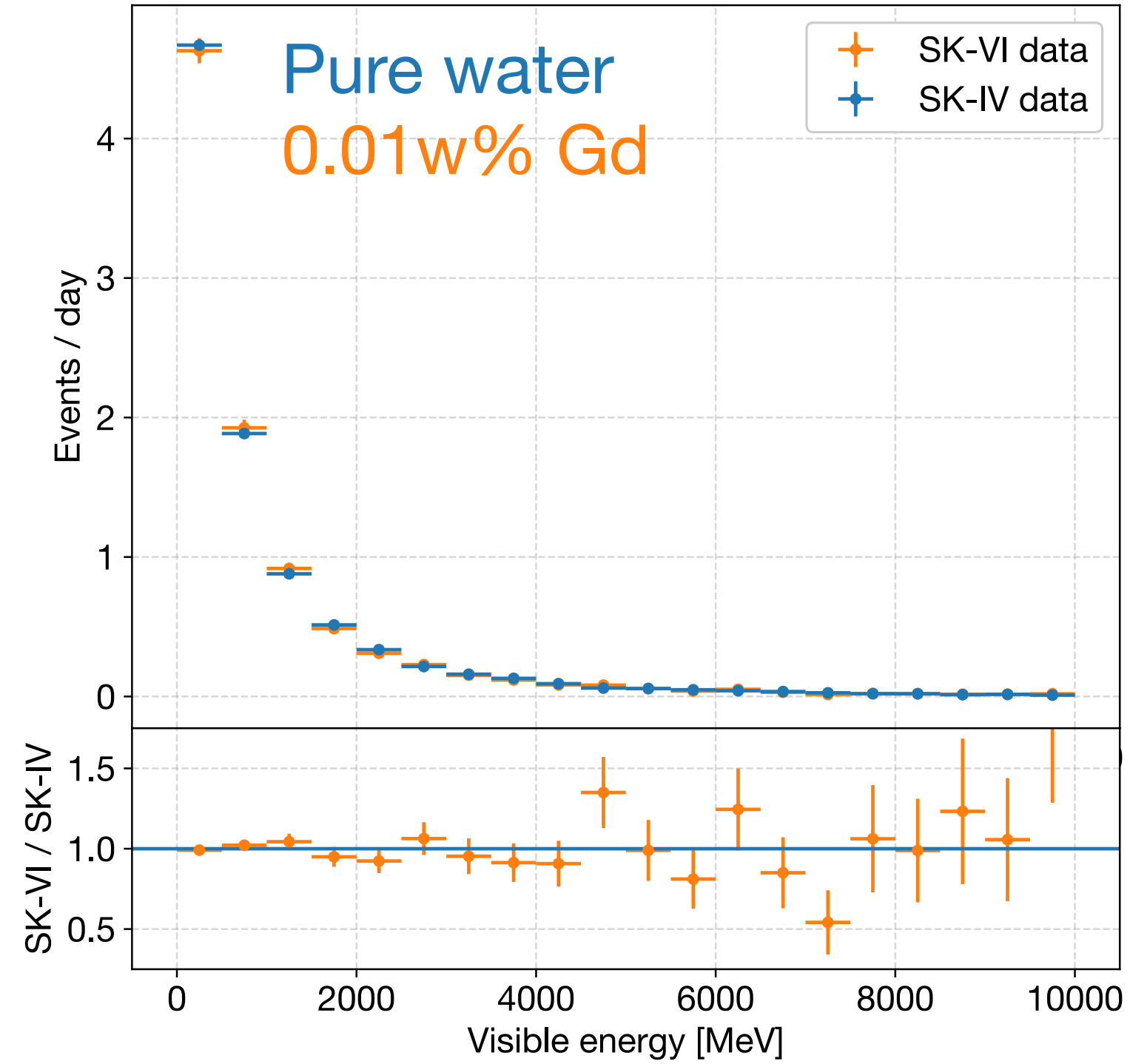


# Gd impact on data quality seems small

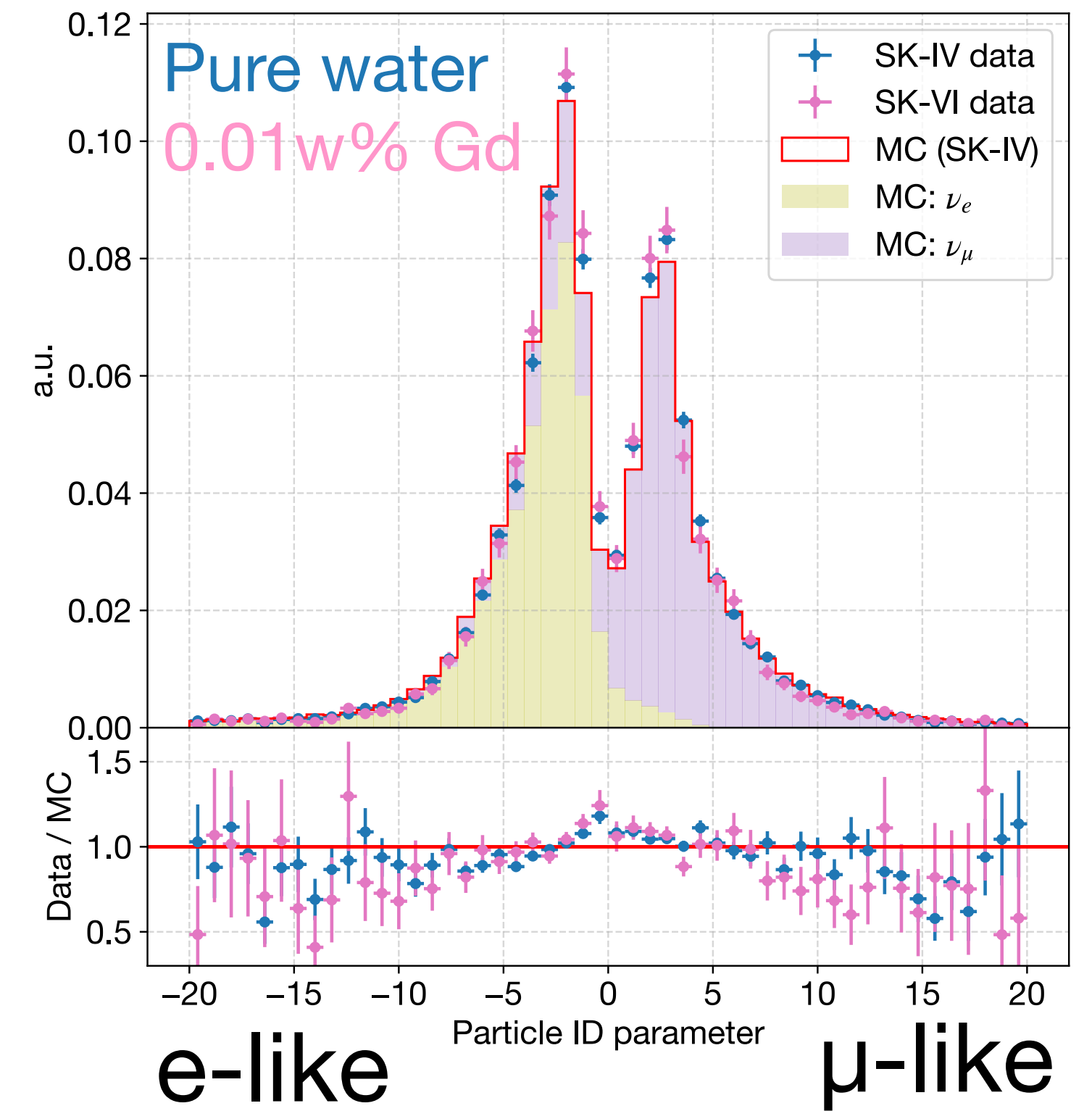
## Event rate



## Event visible energy



## Particle ID likelihood



# SK results in FY2023

## Proton decay

- Search for  $p \rightarrow e^+ \pi^0 \pi^0$ ,  $p \rightarrow \mu^+ \pi^0 \pi^0$  ([NNN23](#))

## Atmospheric neutrinos

- Neutrino oscillation parameter measurement ([arXiv:2311.05105](#))

- “NCQE” cross section measurement ([PRD 109, L011101](#))

- Neutron production measurement (paper in preparation)

↑ My PhD work

All results involve neutron-tagging data events

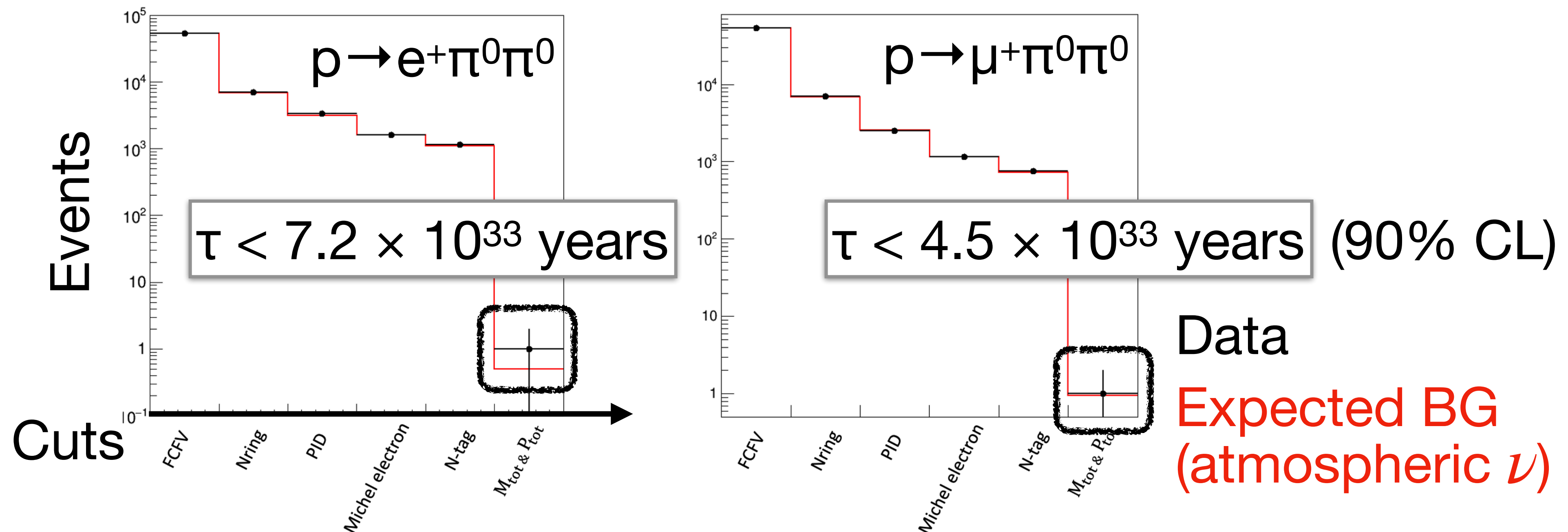
Pure  
water

0.01w%  
Gd

# Search for $p \rightarrow e^+ \pi^0 \pi^0$ , $p \rightarrow \mu^+ \pi^0 \pi^0$

Decay rates expected to be comparable to  $p \rightarrow e^+ \pi^0$ ,  $p \rightarrow \mu^+ \pi^0$

Full pure water data: SK I-V (6511.3 days, 0.40 Mt·yrs) + n-tag BG reduction



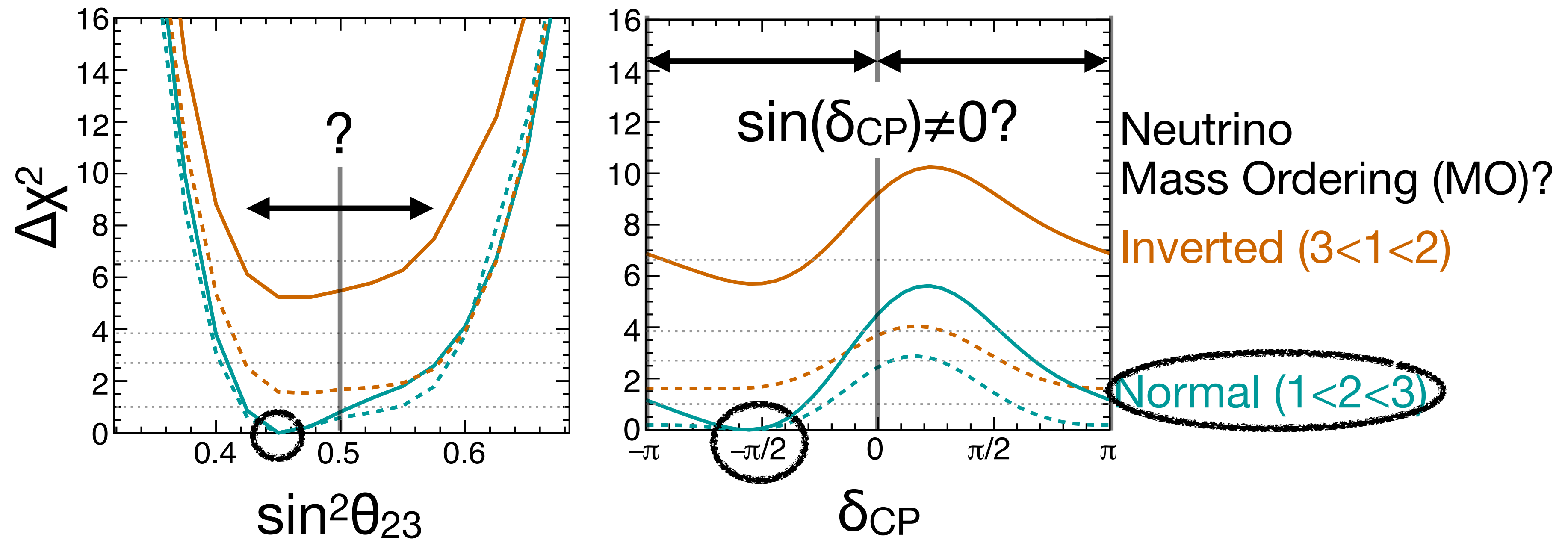
Improves IMB limits (1999) by factor of  $\sim 40$

# Updates in neutrino oscillation studies

Full pure water data: SK I-V (6511.3 days, 0.48 Mt·yrs) + external  $\theta_{13}$  constraint

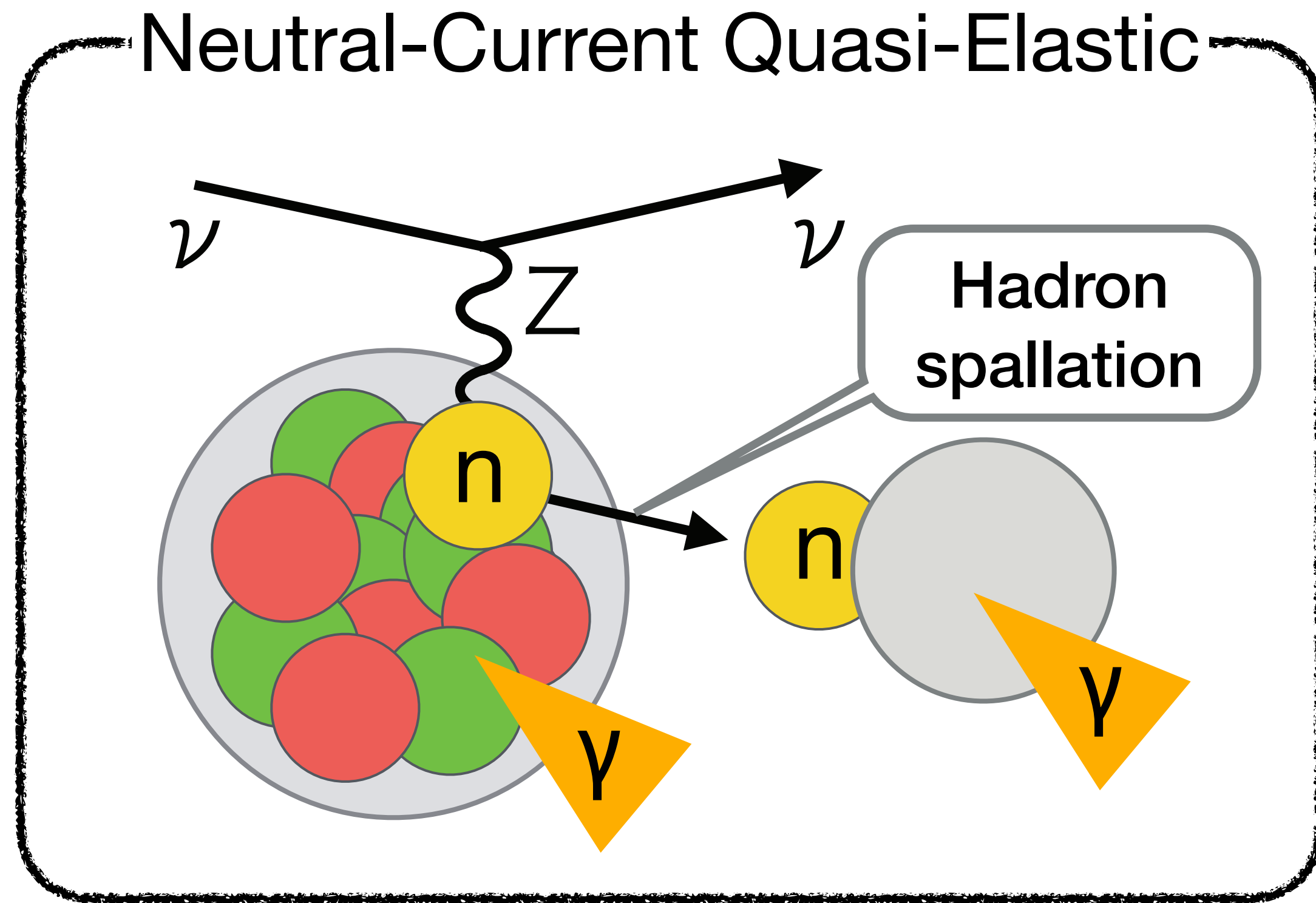
Compared to last publication (2018): 50% stat.  $\uparrow$  +  $\nu/\bar{\nu}$  separation w. n-tag

Expect MO/CP sensitivity  $\sim 10\% \uparrow$



SK (2023) favors: • Lower  $\theta_{23}$  octant • CP violation • Normal MO (92% CL)

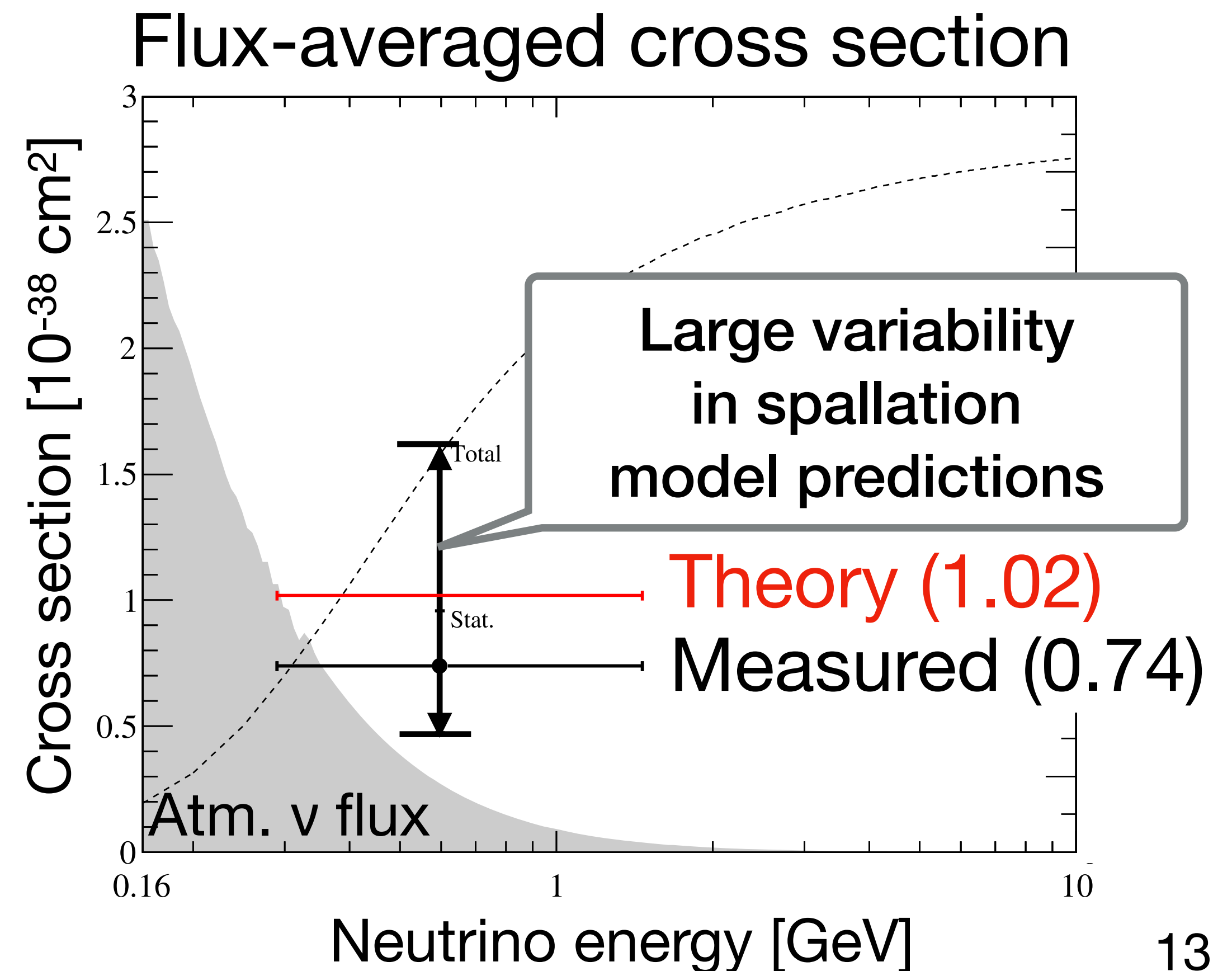
# “NCQE” cross section measurement



Dominant BG for **Supernova  $\nu$  signals**

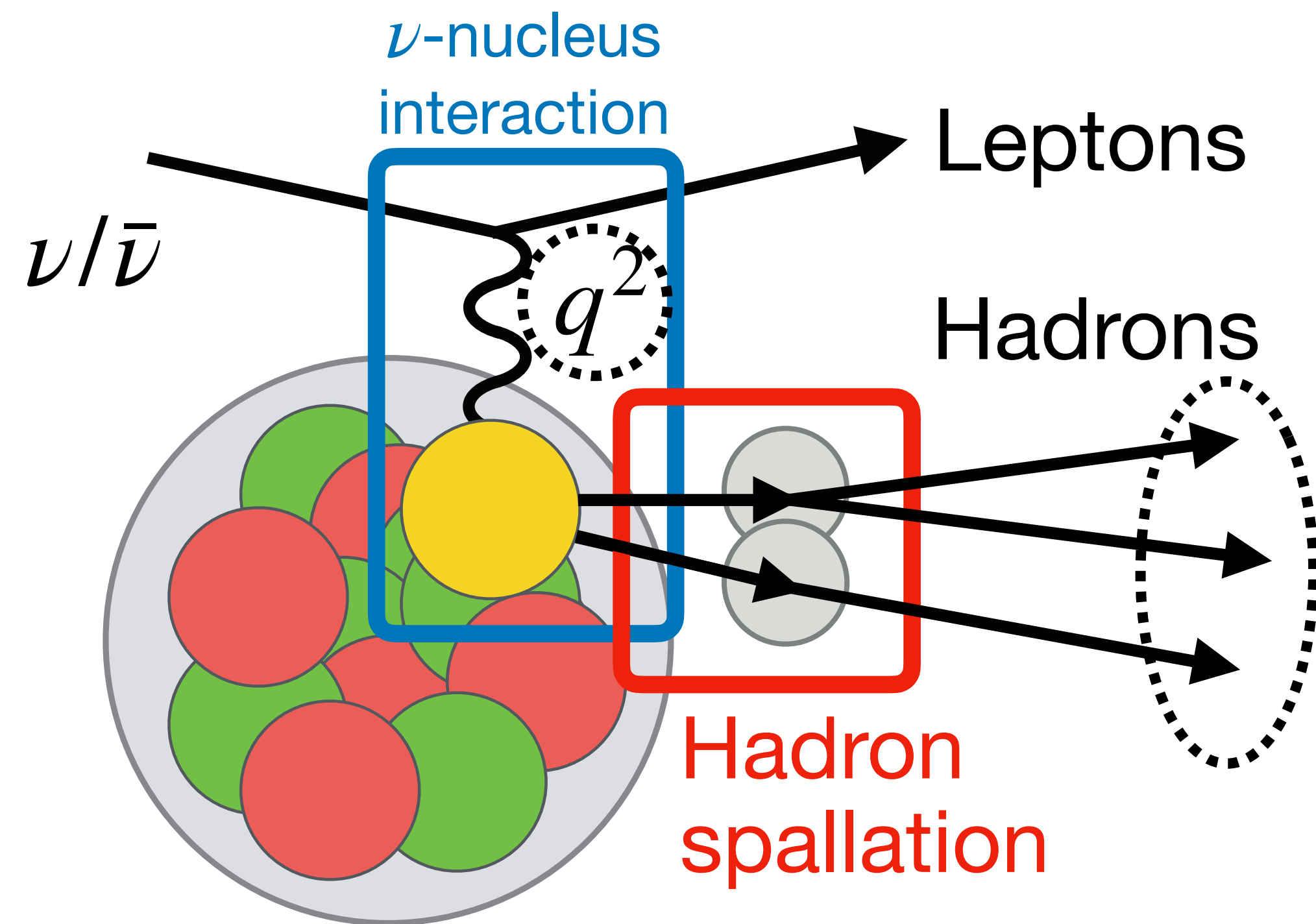
O(1-10) MeV Charged-Current  $\bar{\nu}_e p \rightarrow e^+ n$

38 NCQE-like events  
in 0.01w% Gd data (552.2 days)

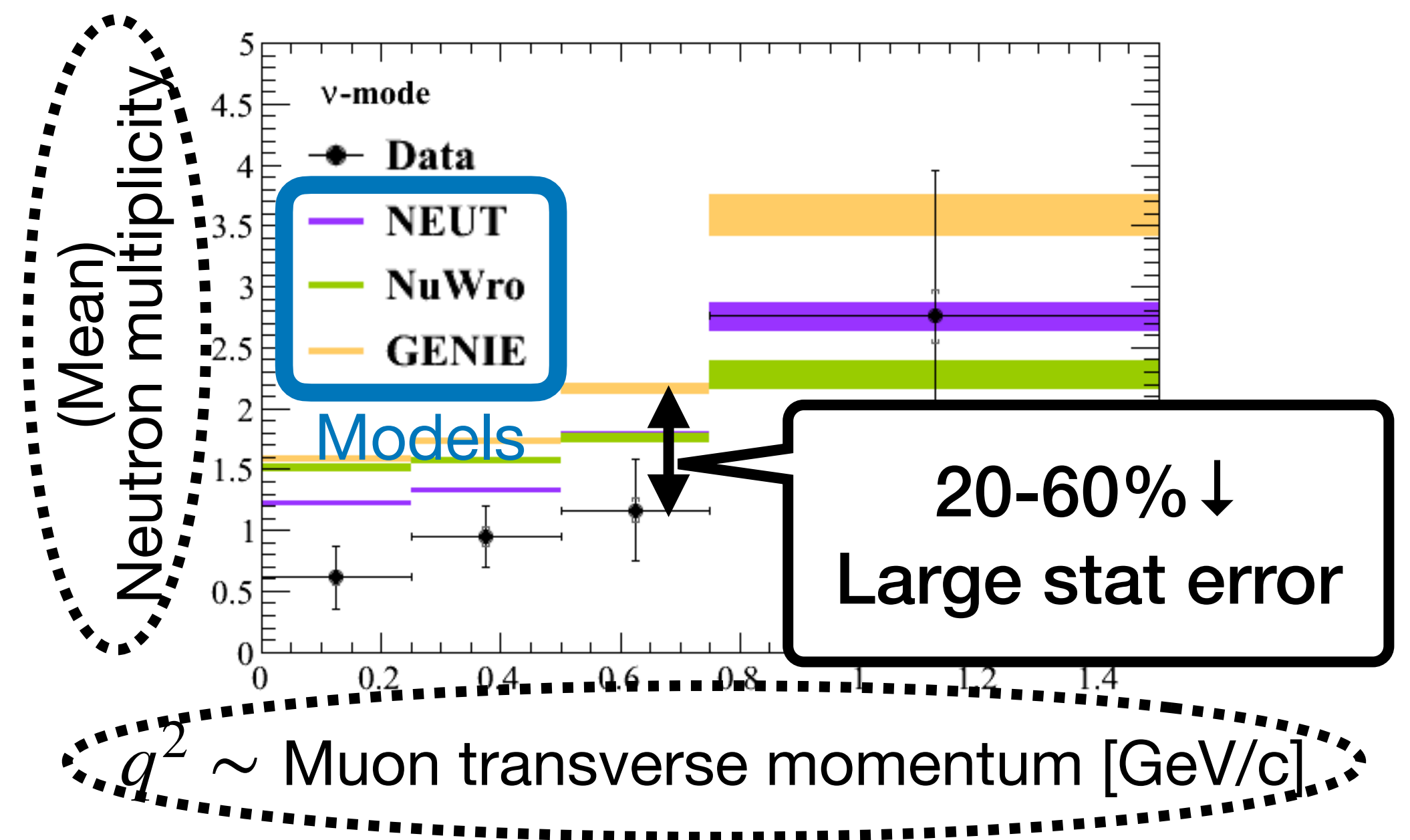


# **Measurement of neutron production in atmospheric neutrino events**

# Motivation: Validation of simulation models



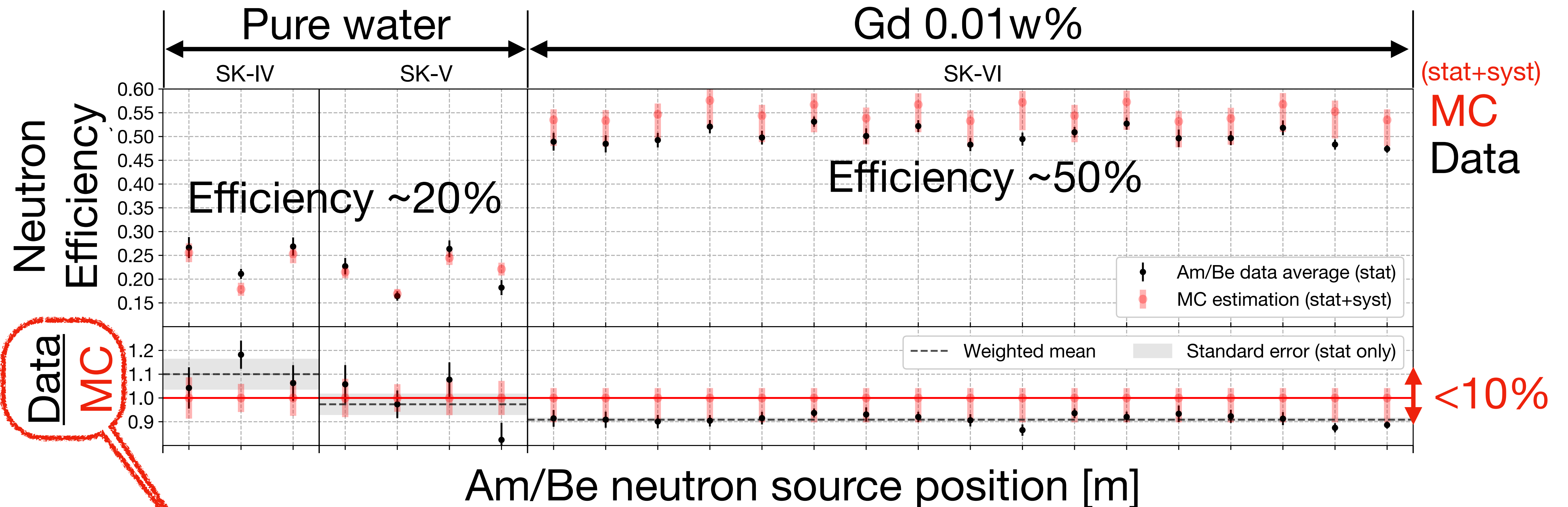
~200 artificial  $\nu_\mu$  events in water (T2K, 2019)



**This study** - Larger statistics (~40,000 atmospheric  $\nu$  events in water)

- Investigate variability in **hadron spallation models**

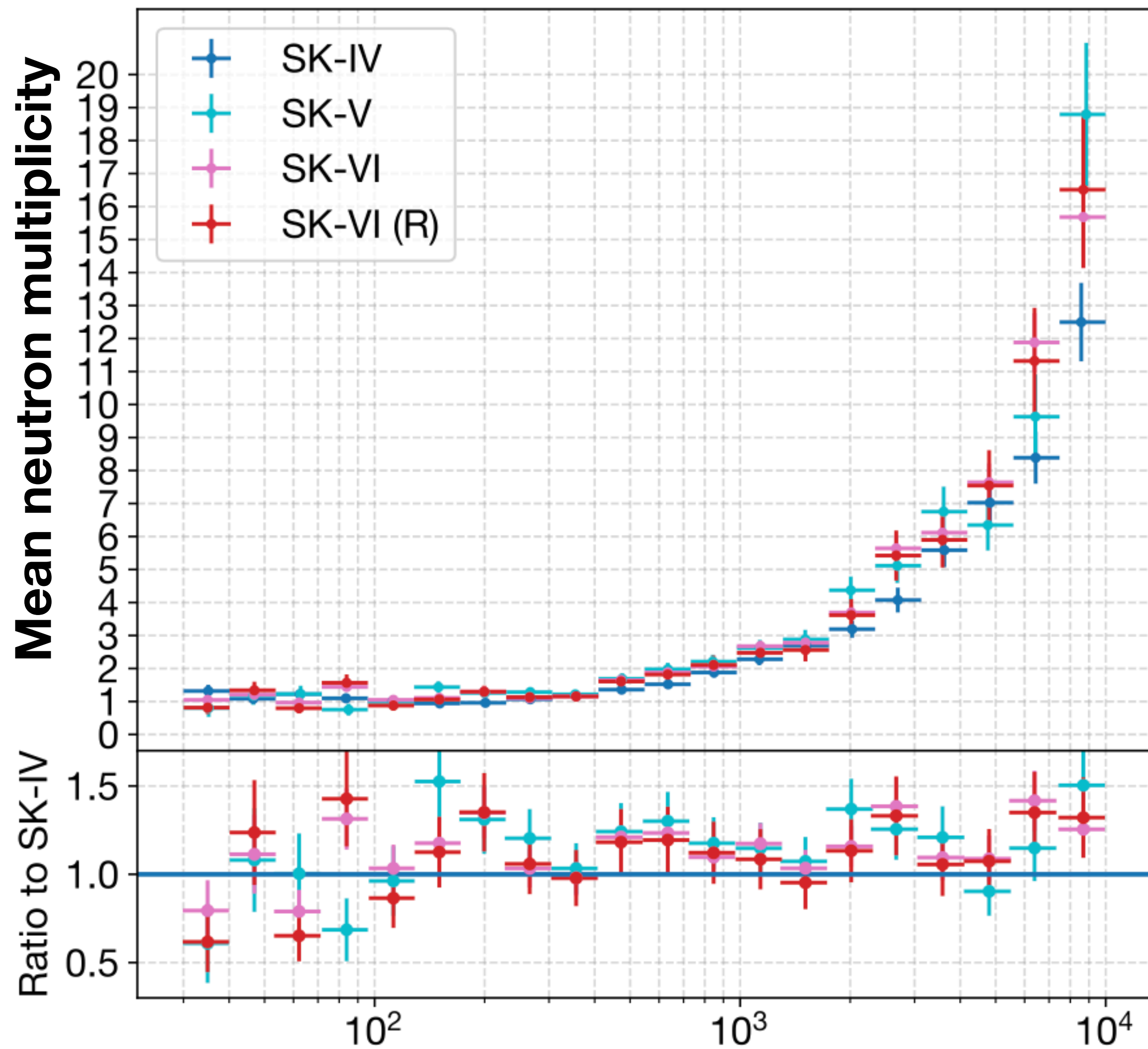
# Neutron efficiency: Calibration data vs. MC simulation



Used to adjust MC-expected efficiencies in neutrino events  
Assigned overall systematic uncertainty: ~15%



## Efficiency-corrected



# Neutron multiplicity: Phase consistency

Atmospheric neutrino data  
in different SK operational phases

++ Pure water: 2008-2019

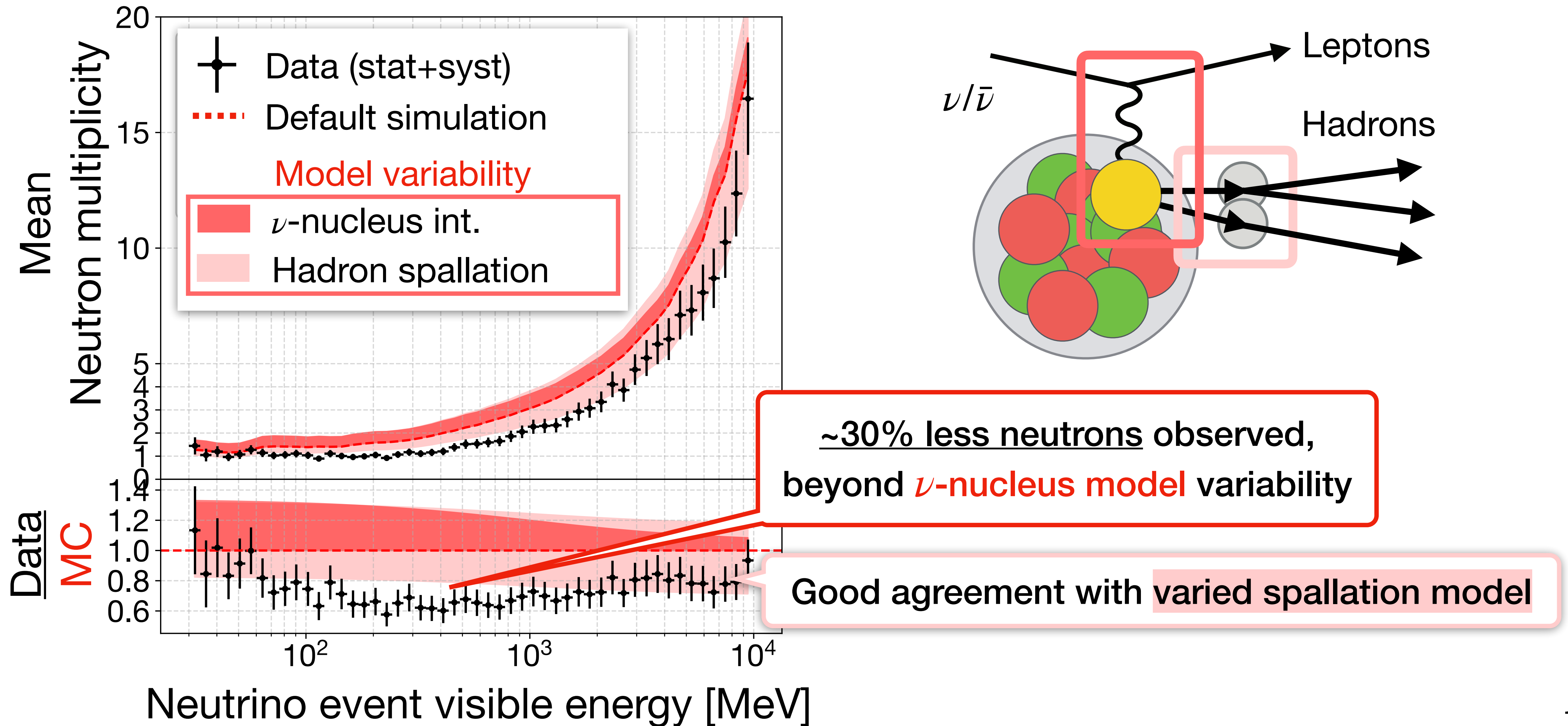
++ 0.01w% Gd-water: 2020-2021

Phases **with/without** Gd,  
having different neutron efficiencies,  
show consistent results

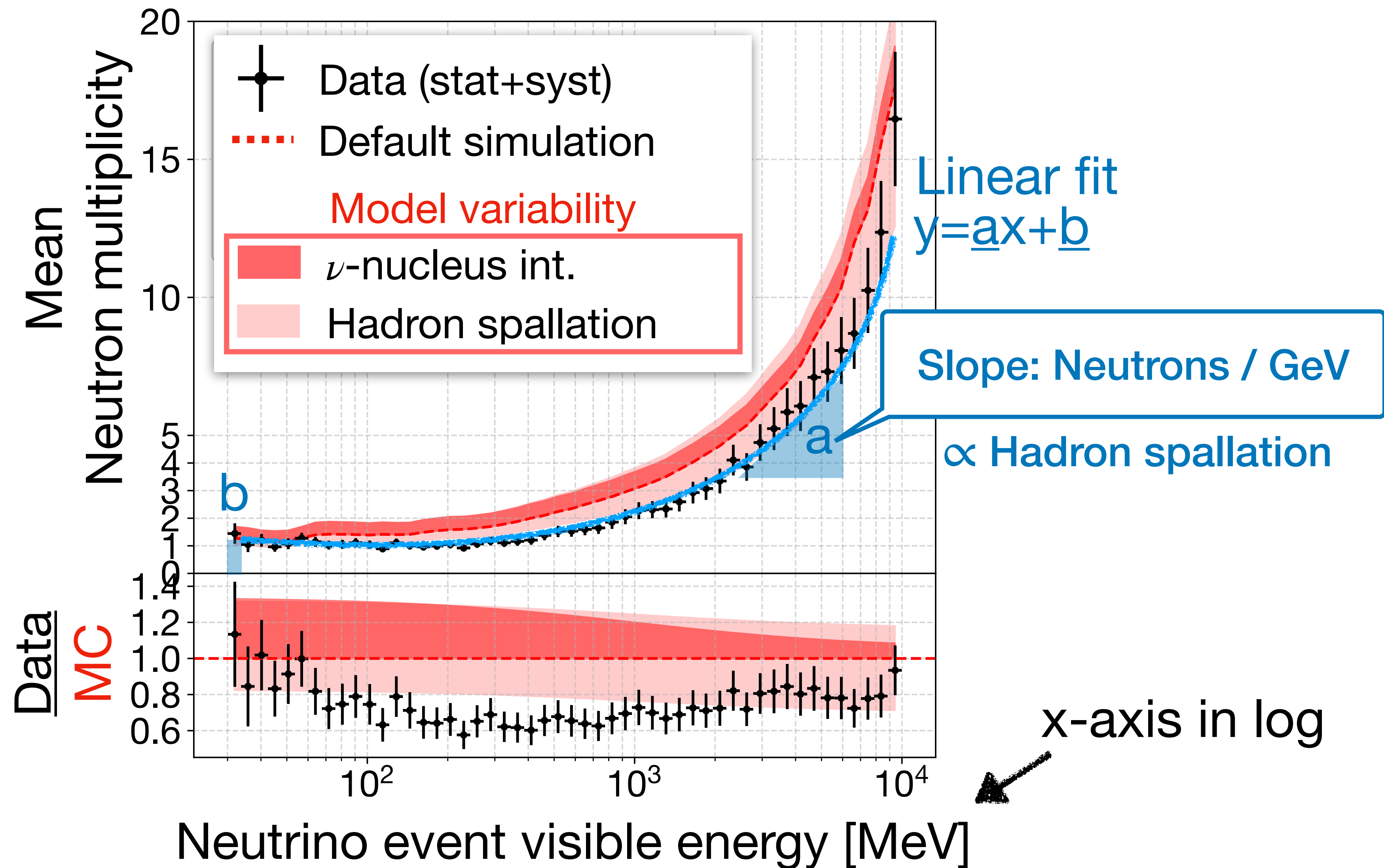
Also consistent with past study (SNO, 2019)

Visible energy [MeV]  $\sim$  Total PMT charge seen in interaction  $\sim q^2$

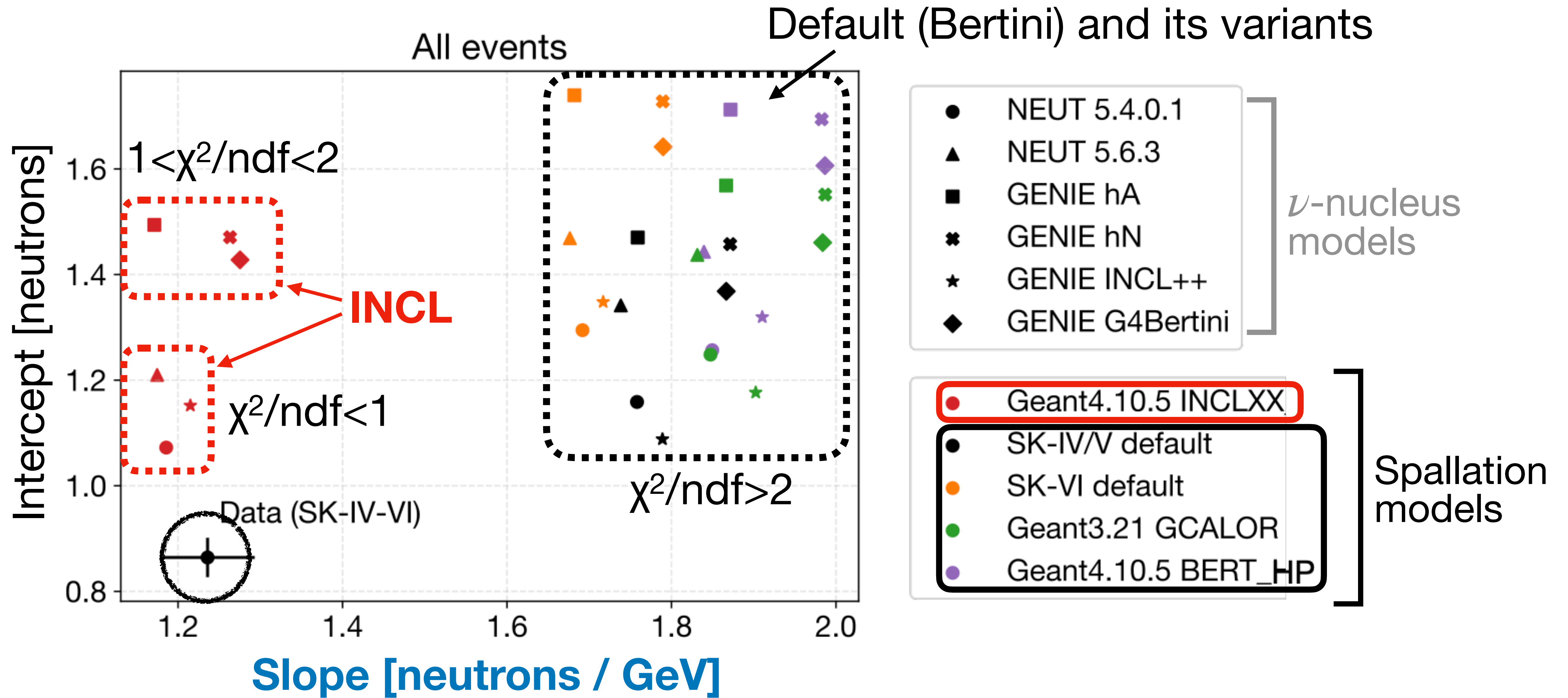
# Combined Data vs. Models

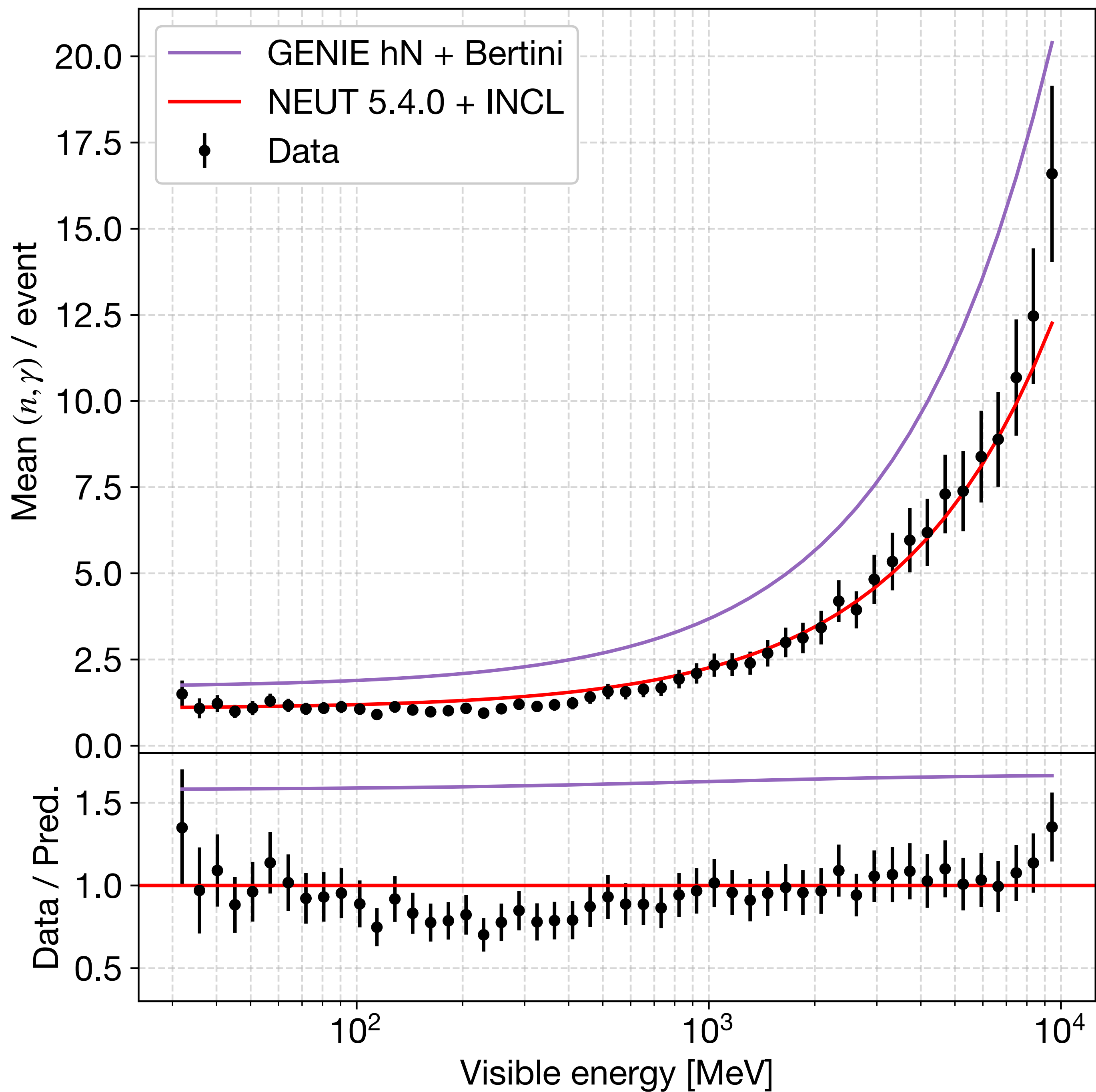


# Combined Data vs. Models



# Fitted Slope: Data vs. Models





## Default spallation model (Bertini)

**~O(1) GeV Intra-Nuclear Cascade model:**  
 Non-interfering 2-body collisions of hadrons

## Best spallation model (INCL)

+ “Nuclear effects” which tend to **reduce** the estimate of neutron production.

For example:

- ✓ Nucleon repulsion, interference
- ✓ “Pauli blocking” of low-energy recoil
- ✓ Cluster formation

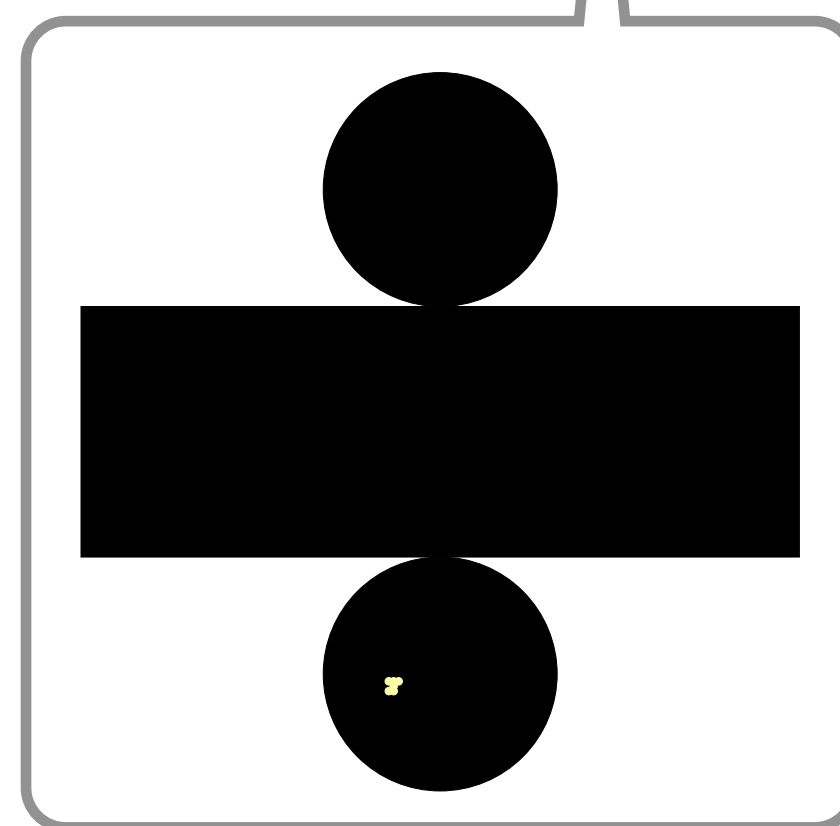
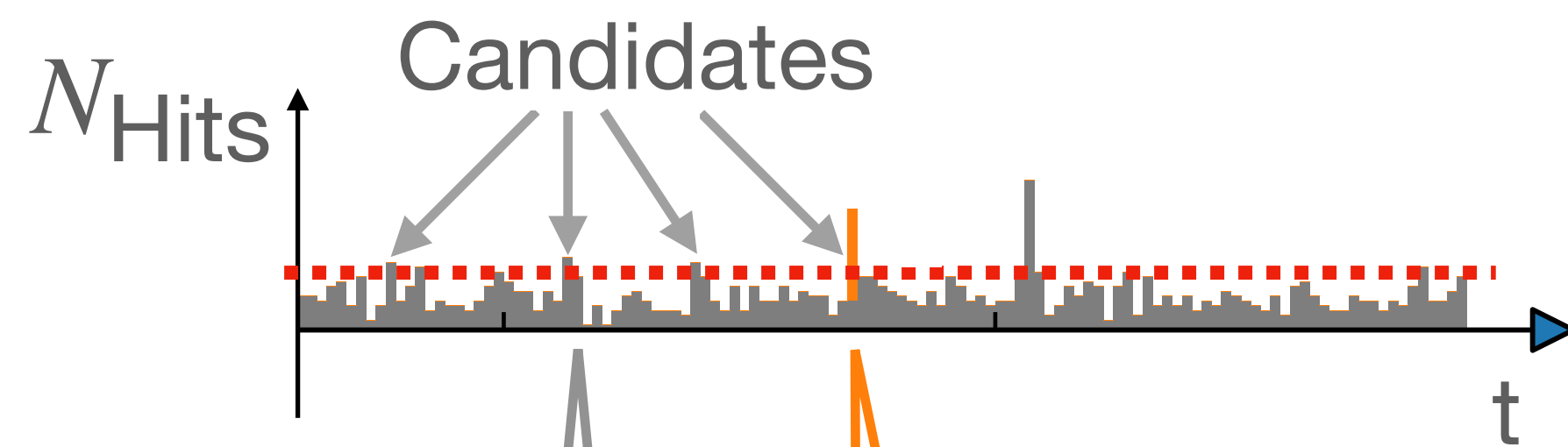
# Summary: SK Proton Decay, Atmospheric $\nu$

- **Neutron-tagging** has potential to **reduce BG for proton decays** and **add sensitivity to neutrino mass ordering and CP symmetry**; it is now applied to most analyses and Gd results are just coming up.
- Gd impact on analyses, including reconstructed variables and neutron efficiency, is constrained through data monitoring and calibration.
- **Measured neutron counts in neutrino data** to reduce large uncertainty; Accuracy in hadron spallation model was critical in explaining the previously reported neutron deficit.

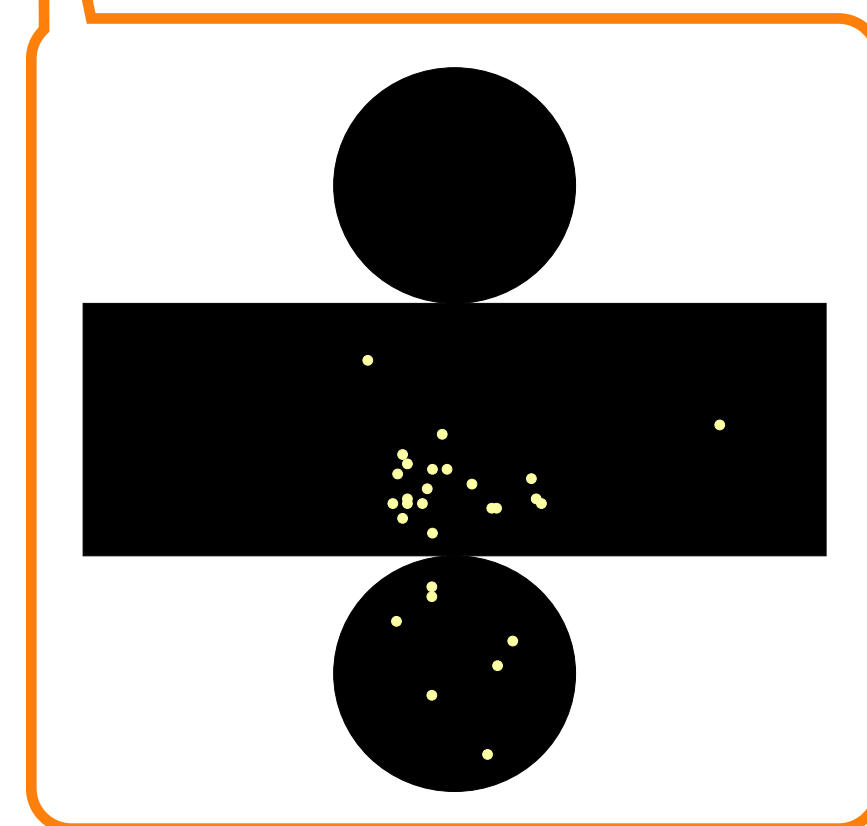
# Backup

# Neutron signal detection algorithm @ SK

## (Step 1) PMT hit trigger



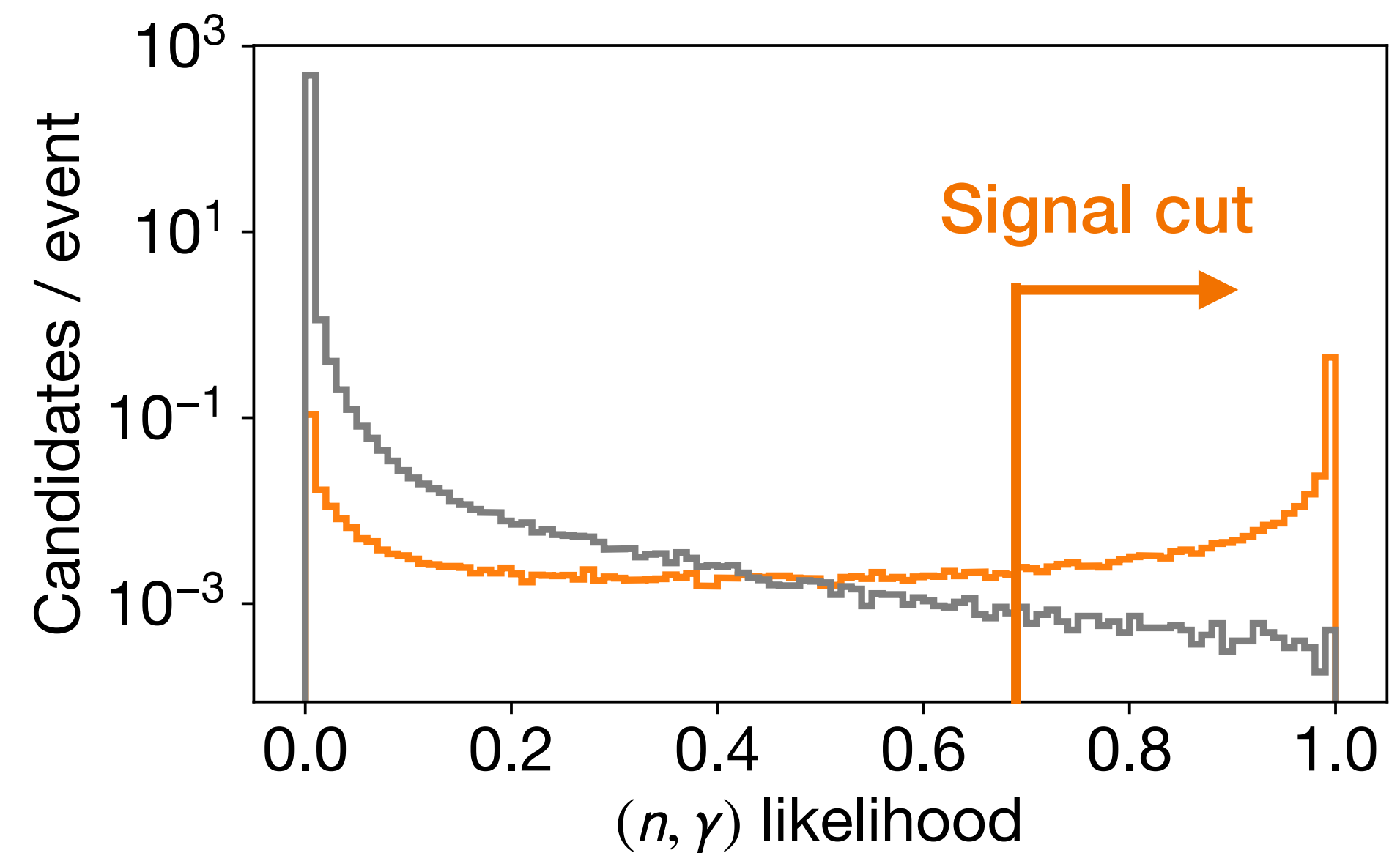
BG radioactivity



Neutron signal

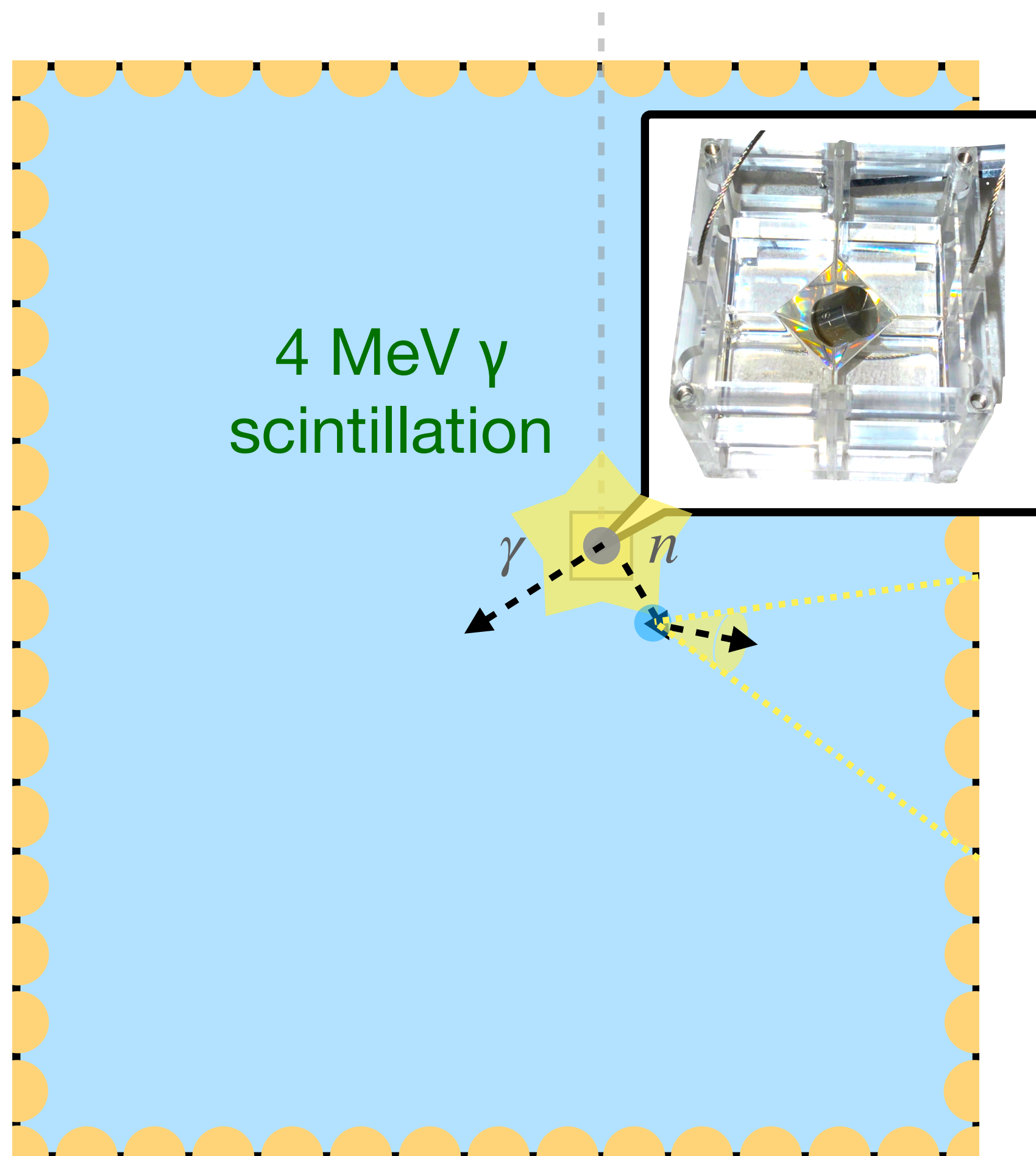
## (Step 2) Neural network

Noise vs. **Signal** classification using  $N_{\text{Hits}}$ , positional correlation, etc.





# Calibration using neutron point source



- Am/Be neutron source:  $1n + 1\gamma$  (4 MeV)
- Surround with scintillator for event trigger
- Take data for 0.5-1h at each source position

$$\text{Neutron efficiency} \approx \frac{\text{Detected neutron signals}}{4 \text{ MeV scintillation triggers}}$$

1n control

# Atmospheric neutrinos and neutron signals

